



Improved Targeted Sub-Threshold Search for Strongly Lensed Gravitational Waves with Sky Location Constraint

Progress Update 25/4

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Mentors during LIGO SURF 2022

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Likelihood Ratio

- The likelihood ratio of a trigger that is produced by a real gravitational wave is given by:

$$\mathcal{L} = \frac{P(\vec{D}_H, \vec{O}, \vec{\rho}, \vec{\xi}^2, [\Delta\vec{t}, \Delta\vec{\phi}] | \vec{\theta}, \text{signal})}{P(\vec{D}_H, \vec{O}, \vec{\rho}, \vec{\xi}^2, [\Delta\vec{t}, \Delta\vec{\phi}] | \vec{\theta}, \text{noise})} \cdot \frac{P(\vec{\theta} | \text{signal})}{P(\vec{\theta} | \text{noise})}$$

- $\vec{\Delta t}$: arrival time difference between detectors
- $\vec{\Delta\phi}$: arrival phase difference between detectors
- By constraining the sky location, we are constraining these 2 terms
- Originally calculated in `inspiral_extrinsics.py`



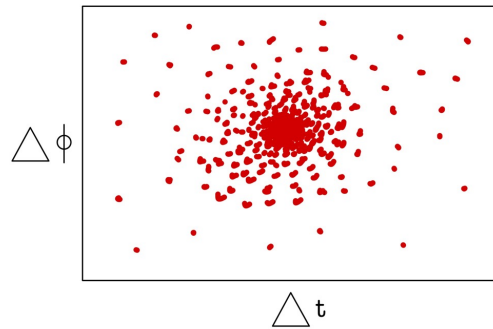
Replacing Original code with a Sampler

- Inspiral_extrinsics.py is the original code in GstLAL to calculate $P(\Delta t, \Delta\phi)$
- Limitations
 - **Low** and **fixed** pixel density: 3072 pixels for the skymap, 33 angles for the inclination and polarisations...
 - One mass fits all searches
 - Complicated and lengthy calculations
 - Computationally costly
 - Generate a map takes $O(\text{hours})$.
 - **Hard to modify** it for other applications that also requires $P(\Delta t, \Delta\phi)$, e. g. targeted searches

Advantage of a GW Sampler

- Easy to modify the domain of parameters
- Easy to change pixel density (just sample more in a smaller area)
- Easily modified to suit different needs
 - Adding sky location constraints (EM counterpart)
 - Changing physical models (cosmology/lensing)
- **Simple!**
 - Fast computation (?) (we still do not know how many samples we need)
 - Generating a sample takes ~ 1 s

Getting a PDF Map

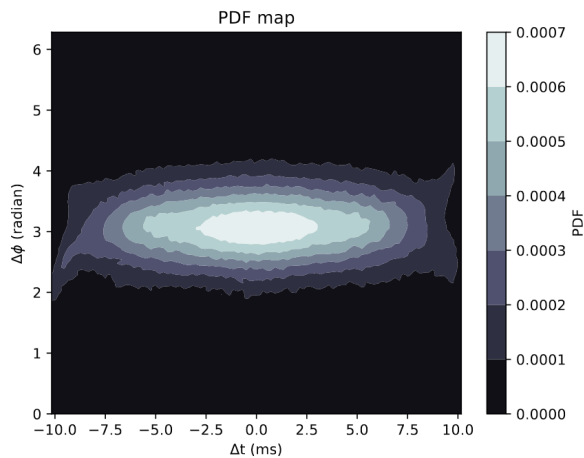
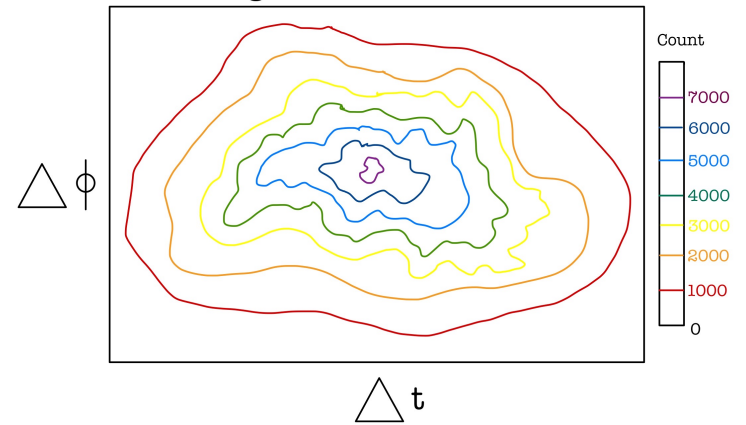


Sample a lot of data

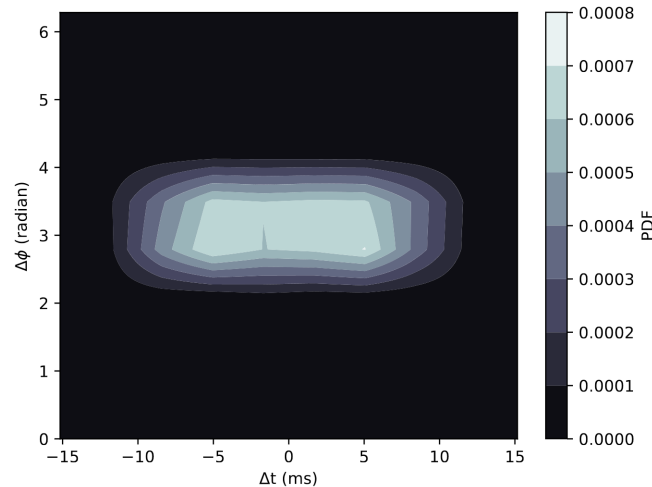


Get a histogram

Histogram -> Contour Plot



(inspiral_extrinsics)



(my generator)

Normalise by total count and perform interpolation to get a PDF map





Scripts Overview

- **GW_generator.py**

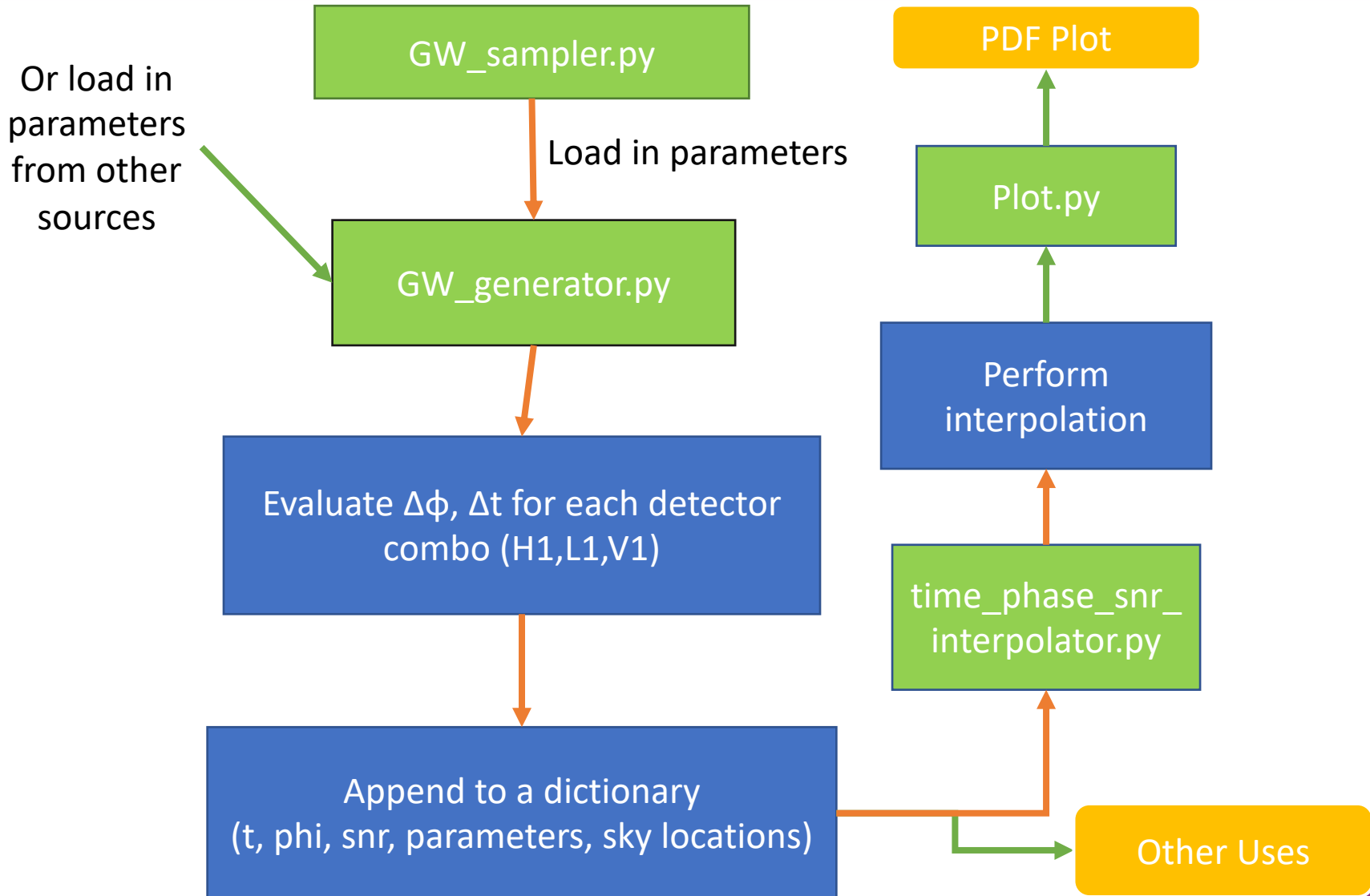
- Storing lots of functions for sampling different parameters and simulating GWs using LALSim -> calculate t, phi, snr
- Mass
- Spin
- Distance
- Inclination
- Polarization
- Waveform approximants
- Sky location (either by loading in a healpy skymap or choose a specific coordinate)
- Target trigger time (in progress)
- Lensing image types (in progress)



Scripts Overview (cont.)

- **GW_sampler.py**
 - Call `GW_generator.py` and store the samples in a dictionary
- **time_phase_snr_interpolator.py**
 - Find detector pairs
 - Calculate dt , $d\phi$, $dsnr$ for detector pairs
 - `rate.py` for binning
 - Interpolate using `scipy`
- **Plot.py**
 - Output a plot of the PDF
 - Store the data for plotting in a `nd.array`

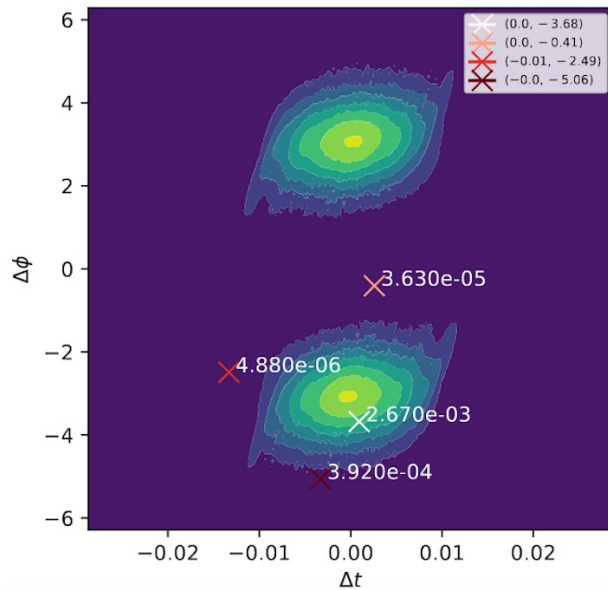
Programme Flow



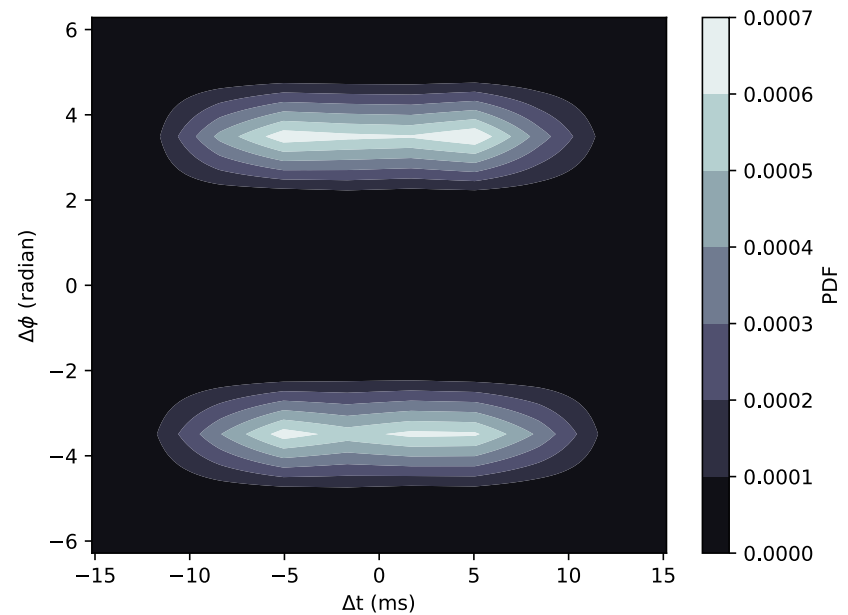
Current Progress

H1 and L1

(From original code)



(From my generator)



Samples: 10k

Interpolation: 10 steps

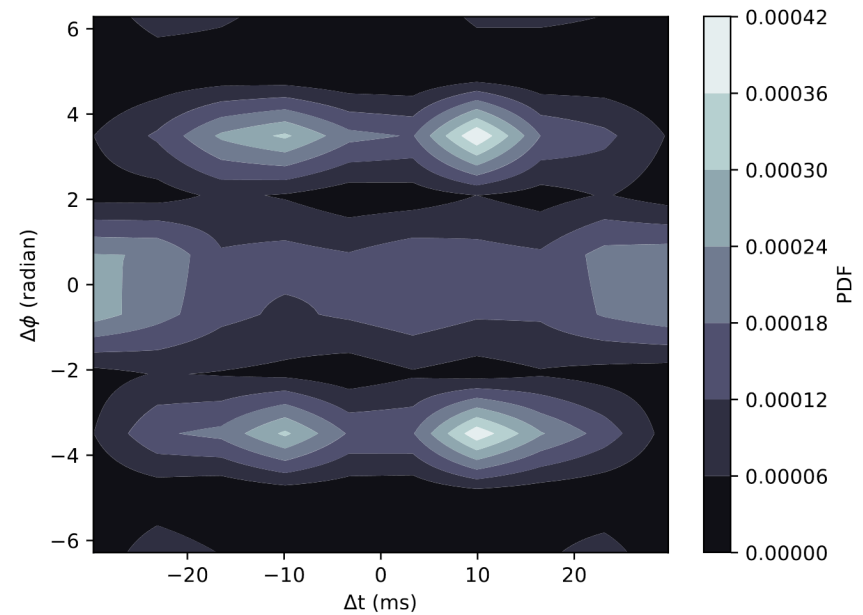
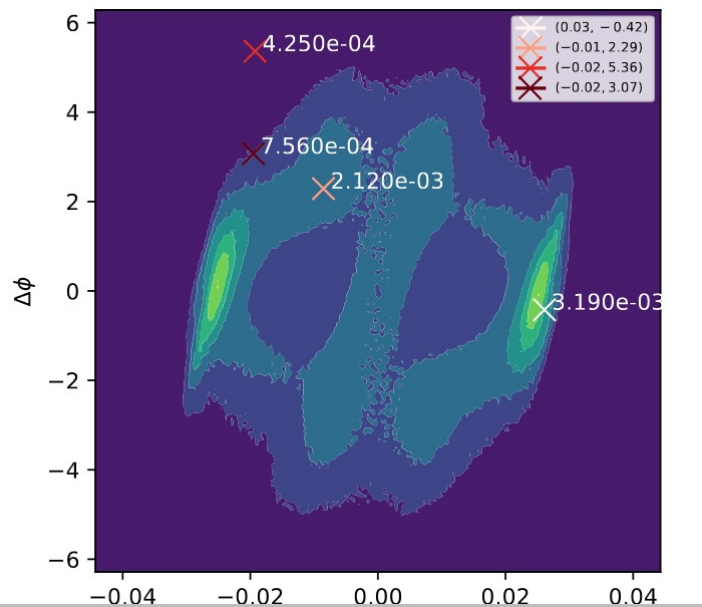
Snr ratio between detectors: 1

Current Progress

H1 and V1

(From original code)

(From my generator)

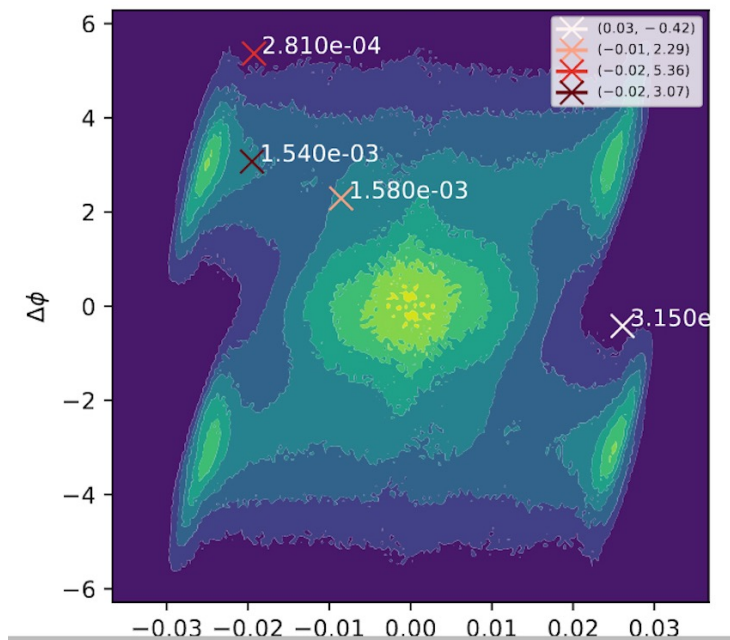


- We need more samples

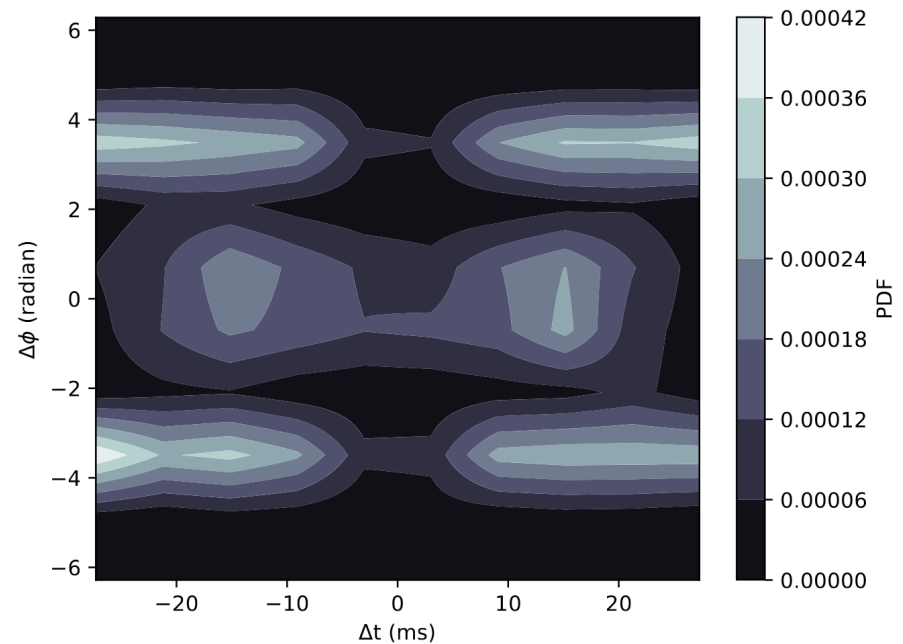
Current Progress

L1 and V1

(From original code)



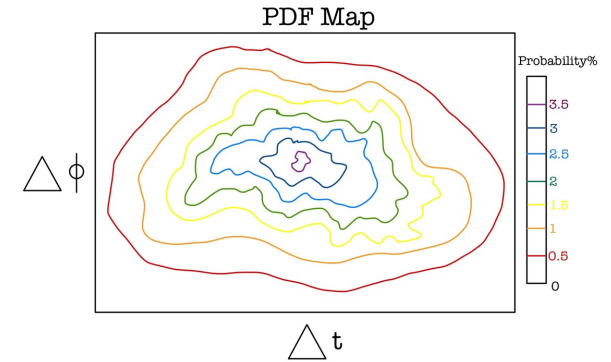
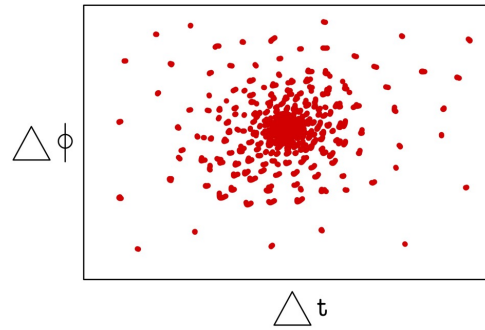
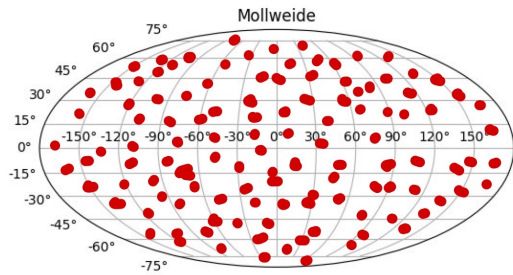
(From my generator)



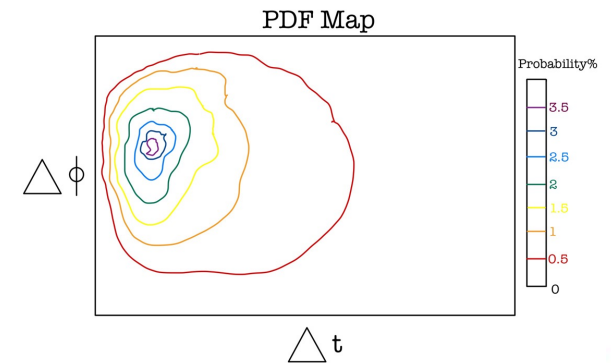
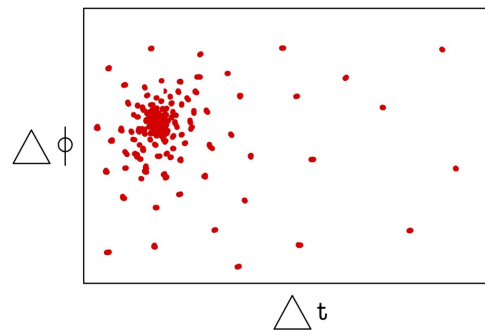
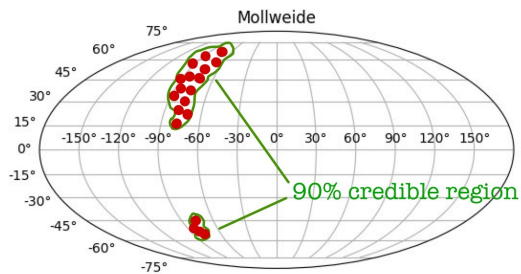
- We need more samples

General Search vs Targeted Search

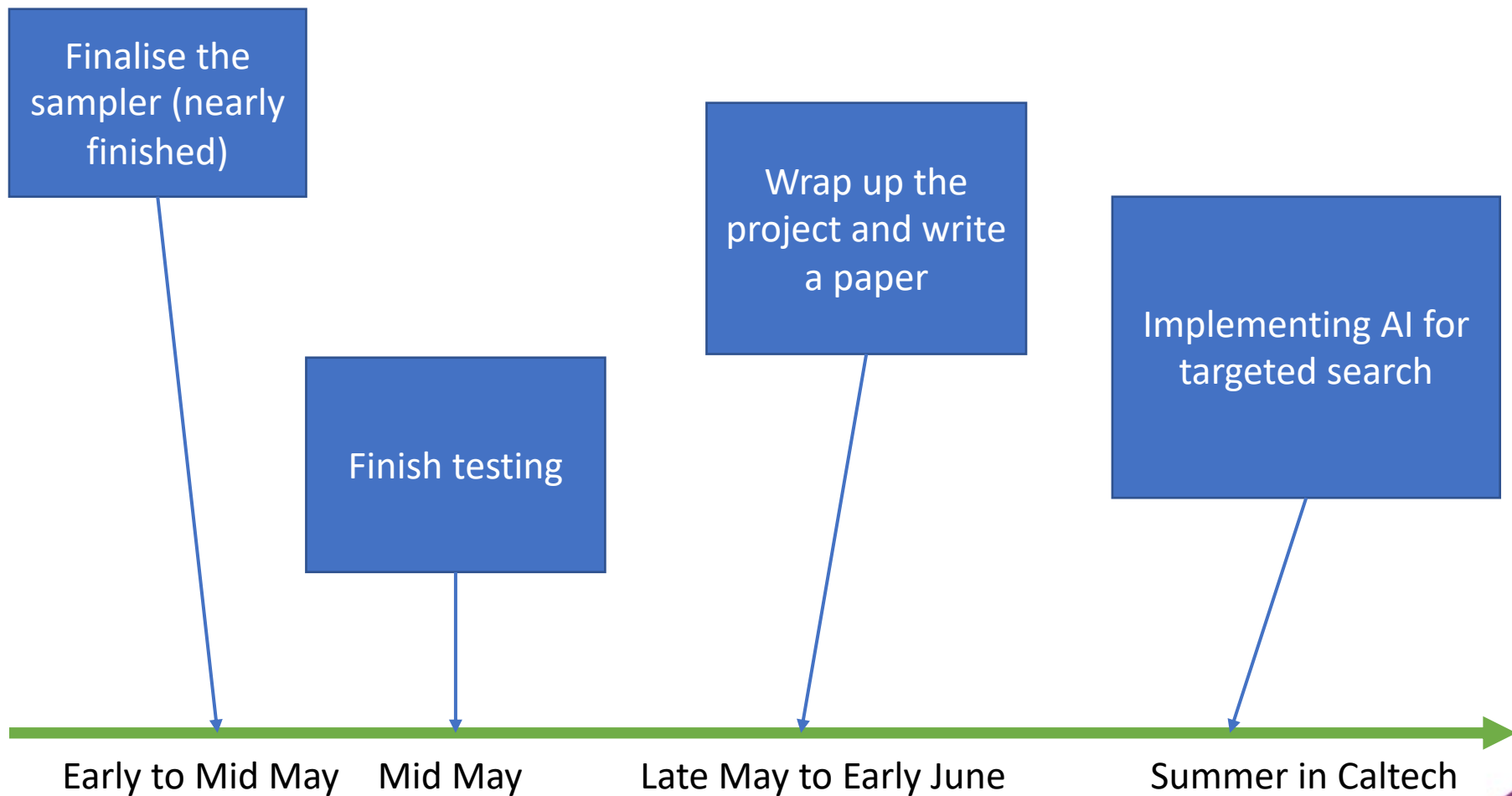
General Search



Targeted Search



Timeline





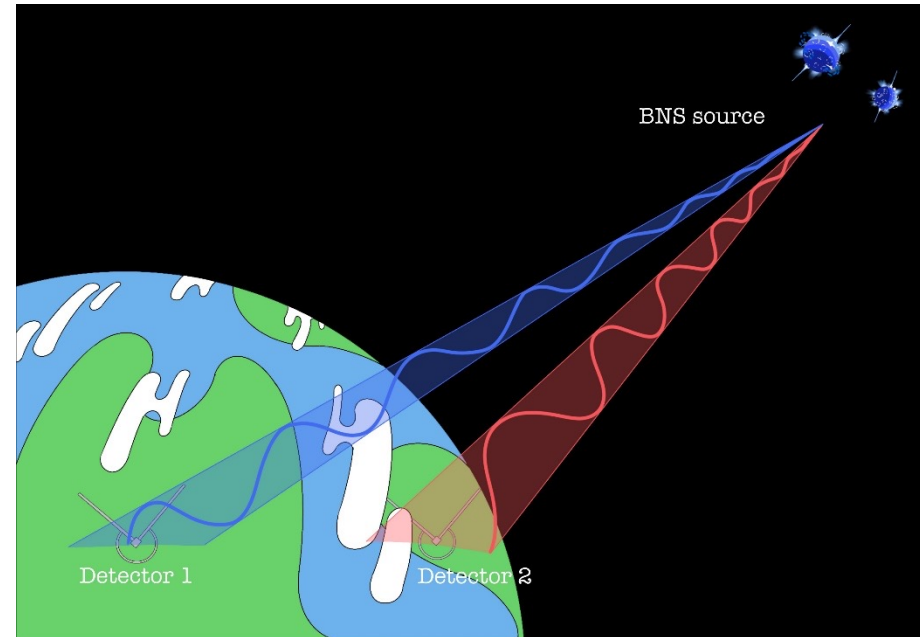
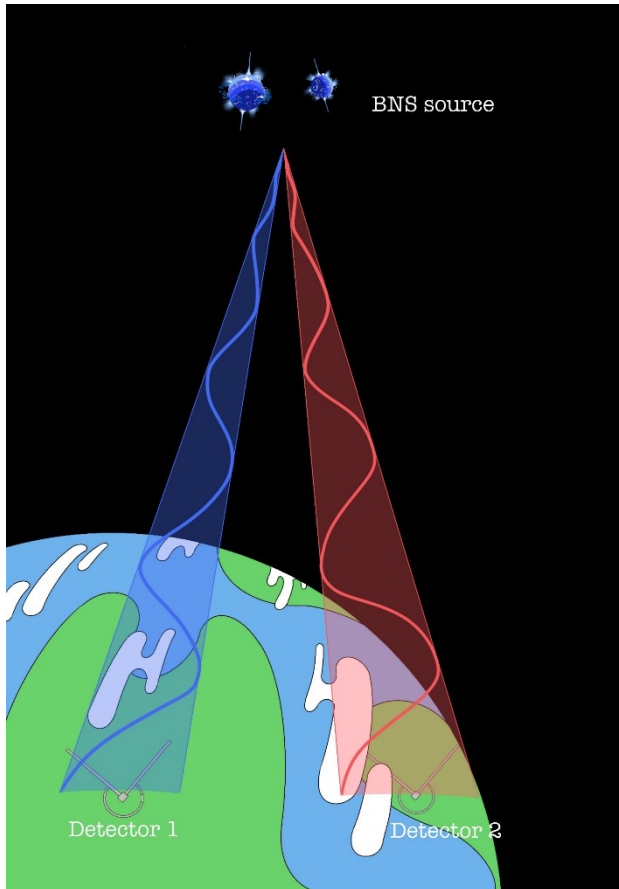
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The End



Different Sky Location

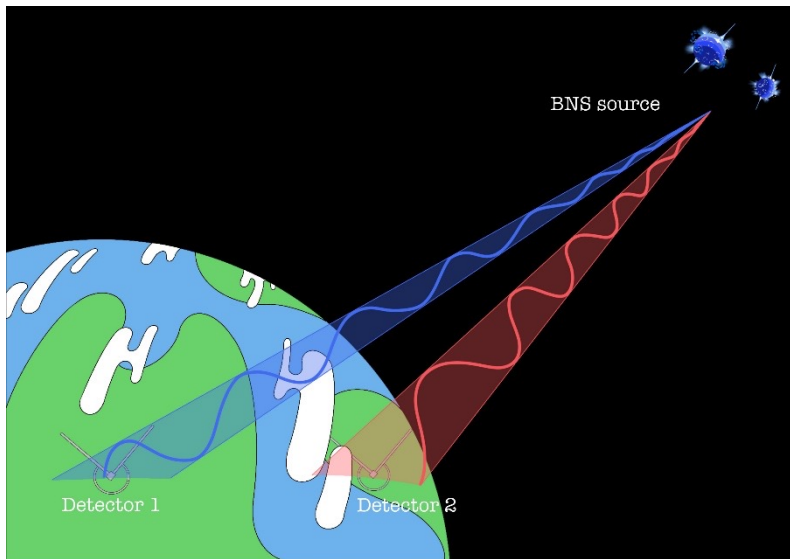
$$\overline{\Delta t} = 0, \overline{\Delta \phi} = 0$$



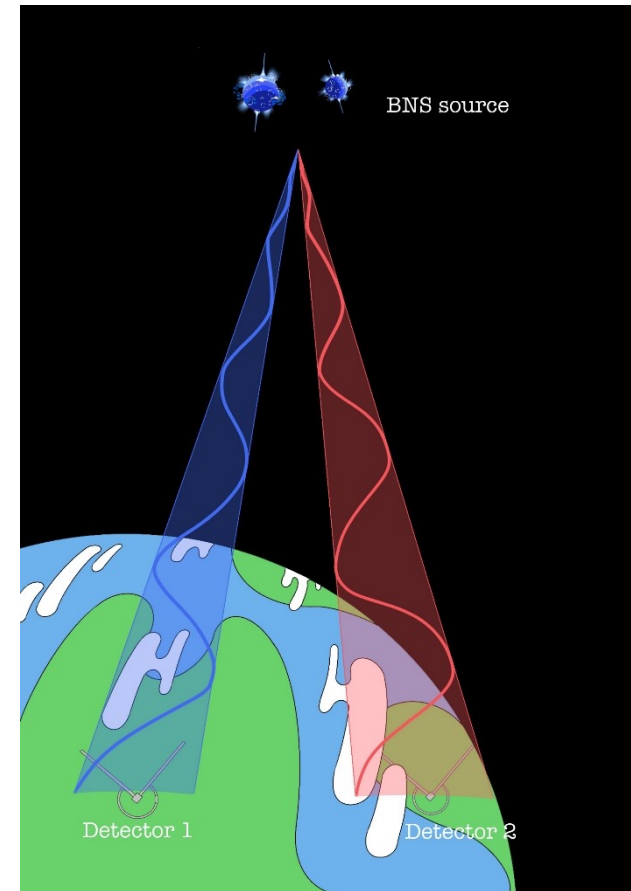
$$\overline{\Delta t} > 0, \overline{\Delta \phi} > 0$$

Earth Rotation

- $\Delta t, \Delta\phi$ would be different at different times
- need PDFs at different times
- Lensing might cause delays in $O(\text{months})$
- E.g. the time delay between 2 of the lensed images of SN Refsdal (a supernova) is 42 days.

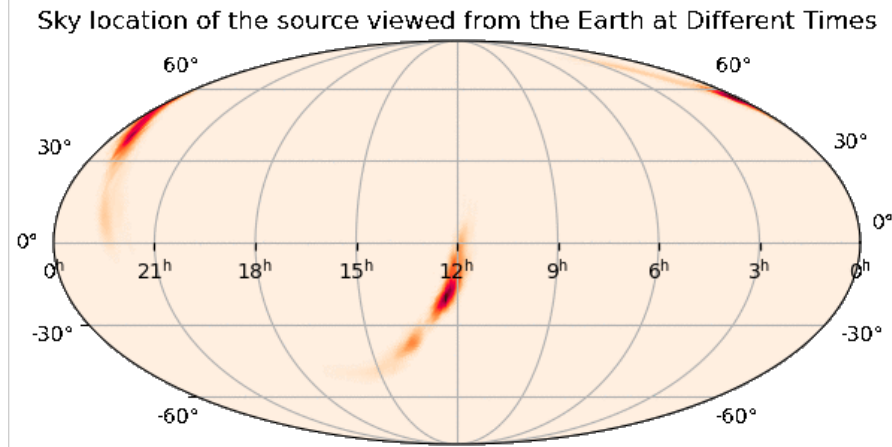
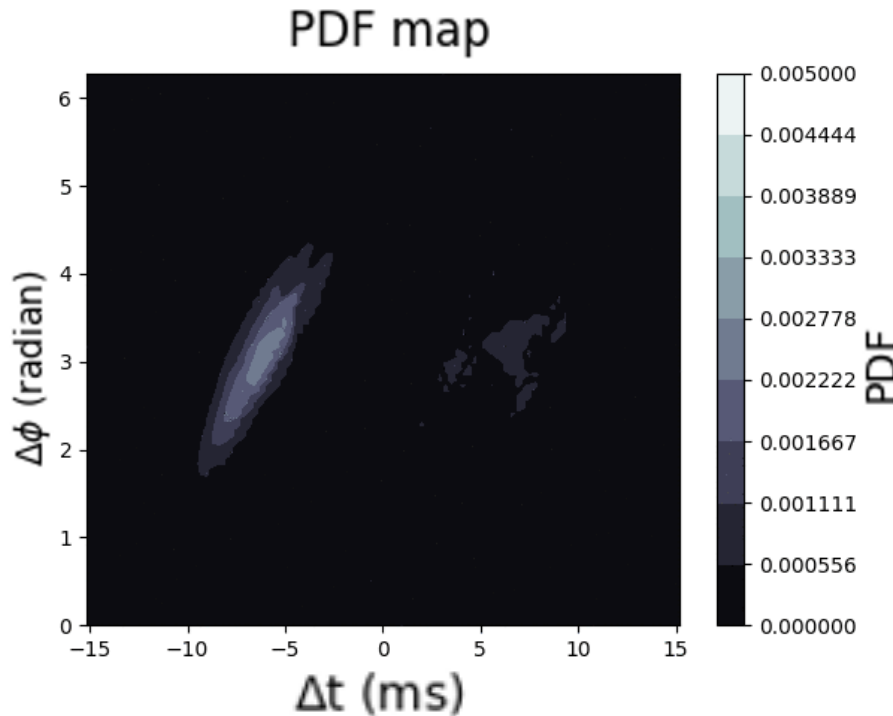


9:00am



12:00pm

Rotating for a whole cycle



Event: GW190519
Frequency: 240 steps per sidereal day rotation
Detectors: H1 and L1
SNR = 10

Programme Flow

