

Physical-Statistical Analysis of Scatter in Febry-Pérot Arm Cavity of aLIGO

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LIGO

Talk Outline

- Introduction
 - » Advanced LIGO
 - » Scatter characterization in Fabry-Perot arm cavity
 - » Scatter measurements
- Analysis of Steady-State Scatter
 - » Scatter during locks
 - » Scatter change through Observation Run 2 (11/30/16 8/25/17)
 - » Comparison with simulation
- Analysis of Transient Scatter
 - » Coupling with beam position
- Regression Analysis
 - » Train model to predict future beam positions

LIGO

LIGO

Laser Interferometer Gravitational-Wave Observatory:

- Measuring the space-time ripple (strain $\sim 10^{-21}$)
- Twin detectors

Intro

LIGO Hanford, WA

LIGO Livingston, LA



(Credit: LIGO Lab)

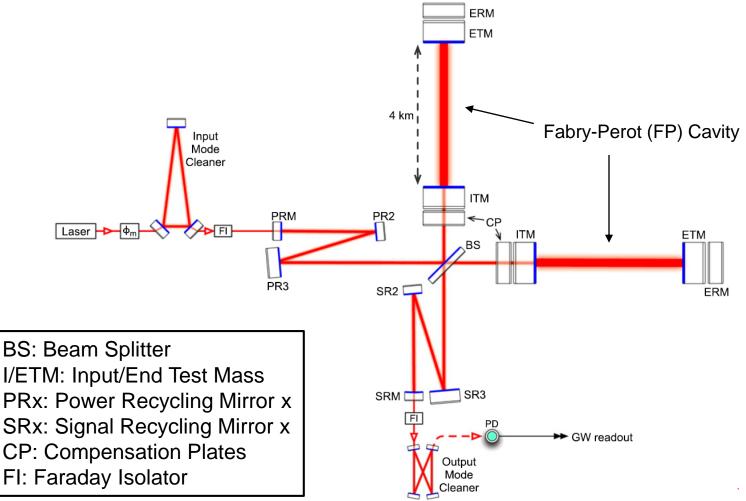
Regression Co

Q&A

LIGO

Advanced LIGO (aLIGO)

A dual-recycled interferometer:



(Credit: LIGO Scientific Collaboration)

LIGO-T1800224 Intro

LIGO

Steady-state

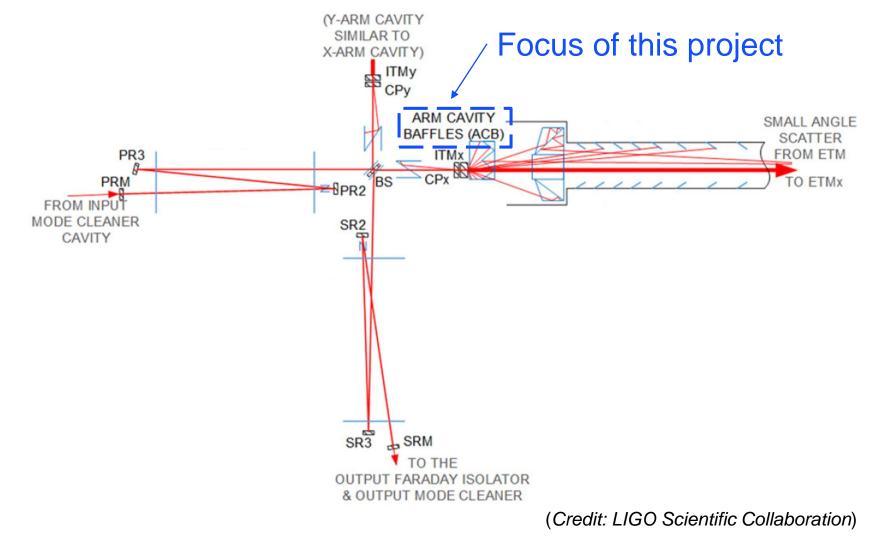
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sion Conclus

&A 5

Scattered Light Baffles in aLIGO

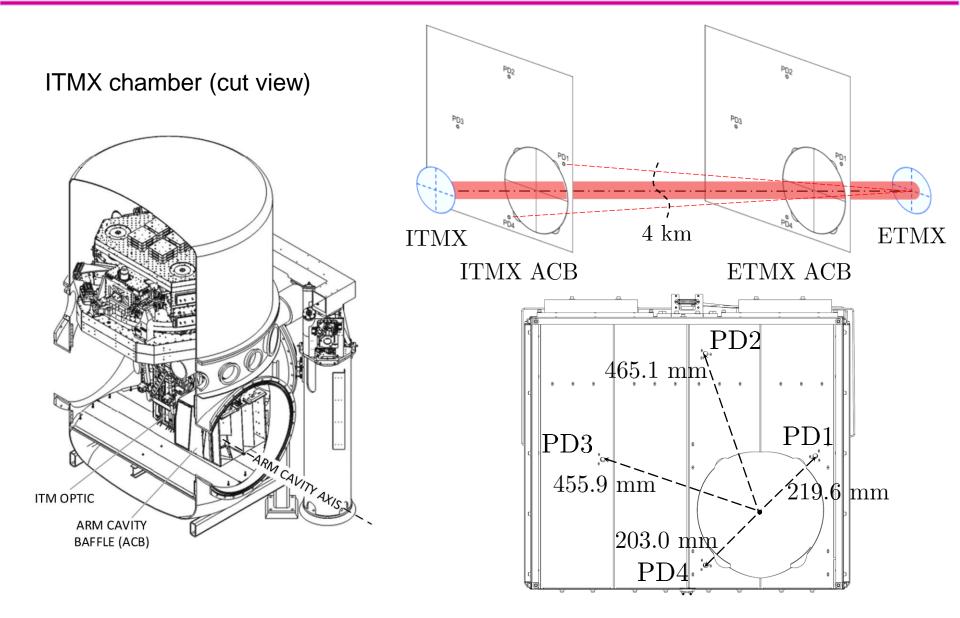
Various baffles are designed to block and measure scatter.



LIGO

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Scatter Measurement



LIGO

Transient

Regression (

Q&A

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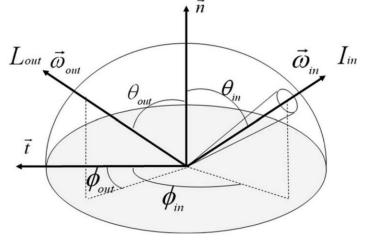
Imperfection of ETM

Spiral patterns on the end test mass (ETM): ETM09 Coated, 160mm dia. S1 -T,P,A LIGO +1.39 10² -ITM04 polished ITM04 coated 10^{1} ETM07 polished nm ETM07 coated 10⁰ PSD (nm² mm) 10^{-1} -2.52 10^{-2} ETM07 S1 Coated 160mm dia LIGO +1.67 10^{-3} 10^{-4} nm 10⁻² 10^{-1} 10⁰ spatial frequency (1/mm) (Credit: GariLynn, Hiro) -2.37

LIGO

Scatter Characterization

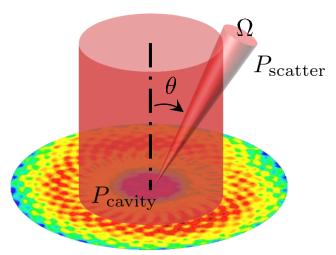
Bidirectional Reflectance Distribution Function (BRDF): a method to characterize scatter on the surface



 $BRDF(\theta_{in}, \theta_{out}, \phi_{in}, \phi_{out}) = \frac{dL_{out}(\theta_{out}, \phi_{out})}{dI_{in}(\theta_{in}, \phi_{in})}$

 L_{out} : Outgoing radiance $[W/(sr \cdot m^2)]$ I_{in} : Incoming irradiance $[W/m^2]$

(Credit: Duck Bong Kim)



In LIGO FP arm, it can be simplified:

$$\mathrm{BRDF}(\theta) = \frac{P_{\mathrm{scatter}}}{P_{\mathrm{cavity}}\Omega}$$

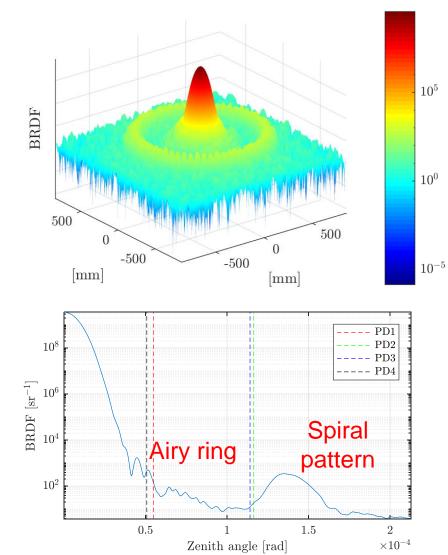
Regression Conclusion

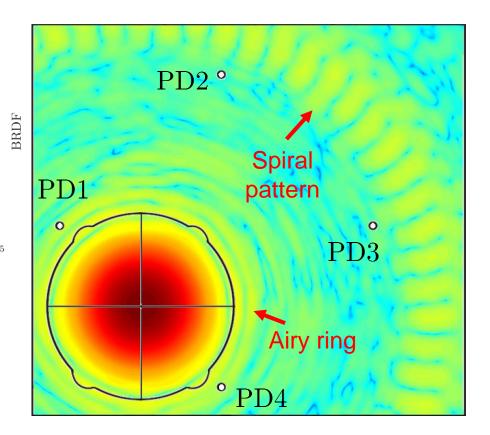
&A 9



BRDF Simulation

Estimation of ETMX BRDF from SIS (scatter on ITMX)





(Credit: Hiro Yamamoto)

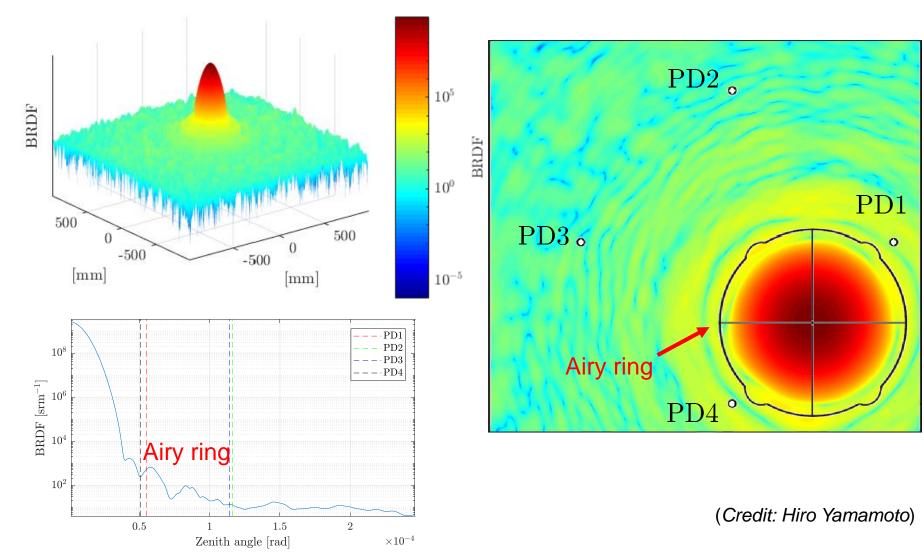
Regression Conclusion

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BRDF Simulation

Estimation of ITMX BRDF from SIS (scatter on ETMX)



Intro Steady-state

Transient

nt Regressio

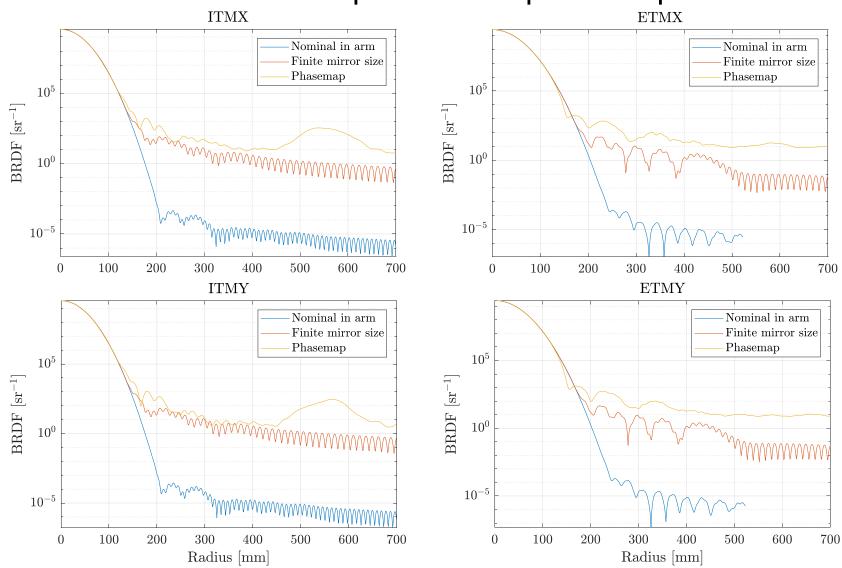
Conclusion Q&A

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LIGO Static Inter

Static Interferometer Simulation

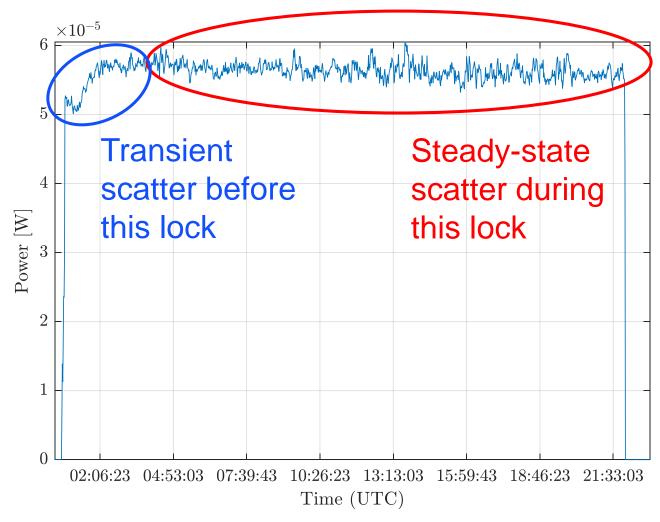
Effect of finite mirror aperture and phasemap:



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Scatter Measurement

Sample minute-trend data of ITMX PD1 on 7/1/17:



LIGO-T1800224 Intro

Steady-state

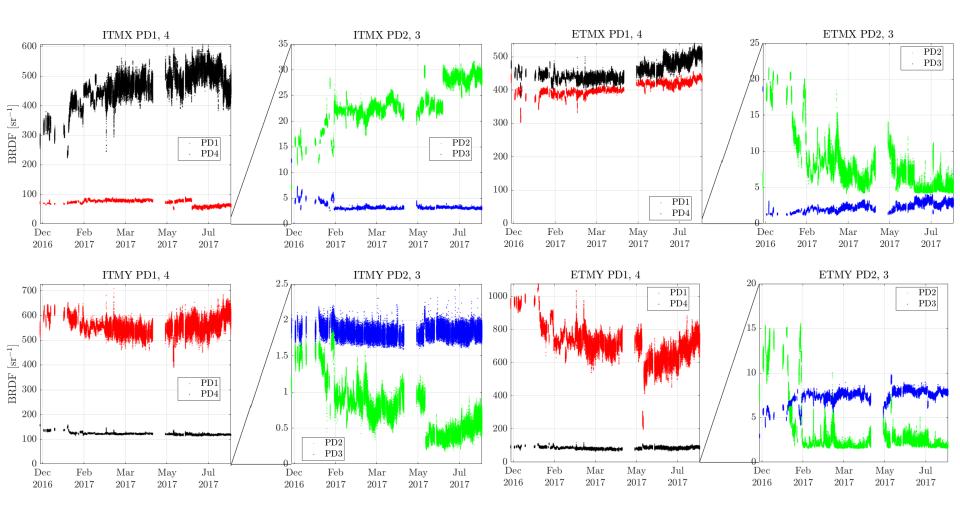
ransient

Regression (

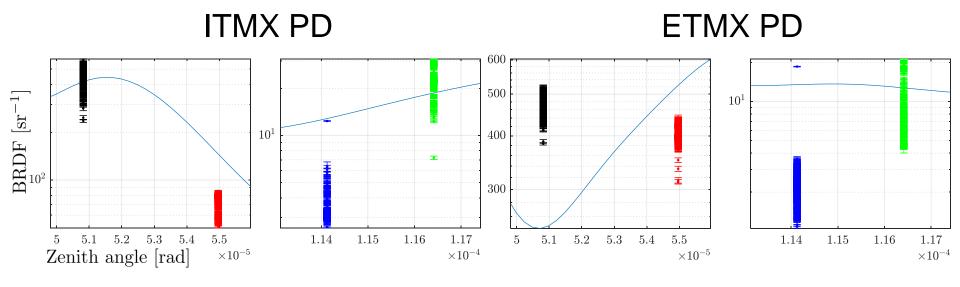
Q&A 13

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Overall Scatter Data in O2



[R G B K] corresponds to PD 1,2,3,4



ITMY PD

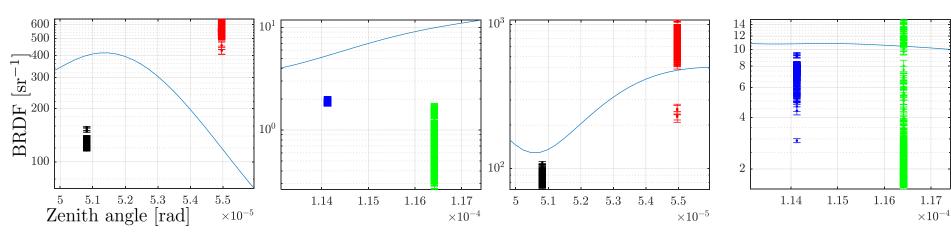
Steady-state

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ETMY PD

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Steady-state

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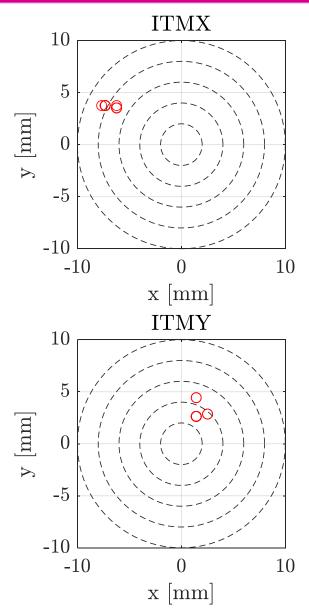
Regression

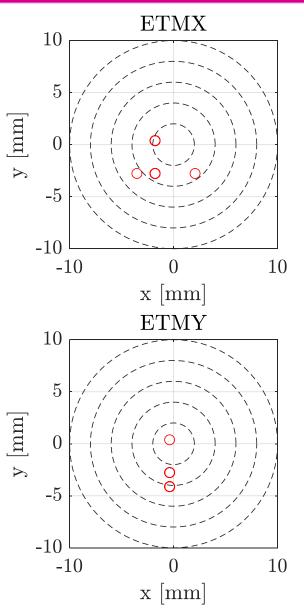
lusion Q&A

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Beam Position in O2 Locks





Steady-state

ransient

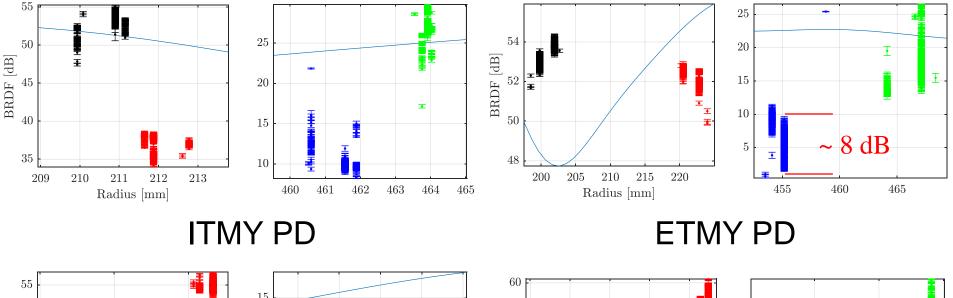
Regression

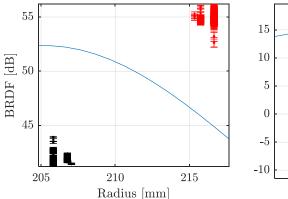
sion Q&A

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Scatter with Beam Positions in O2

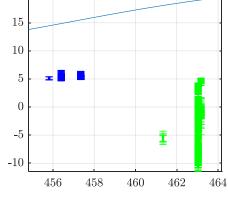
Scatter simulation (all beams at center) versus scatter data. ITMX PD ETMX PD

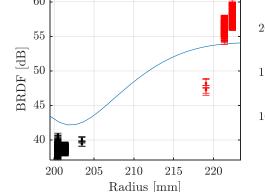


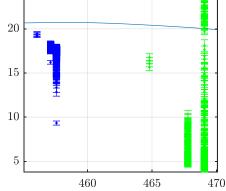


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ro Steady-state

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sion Conclusion Q&A

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Scatter Growth with Time

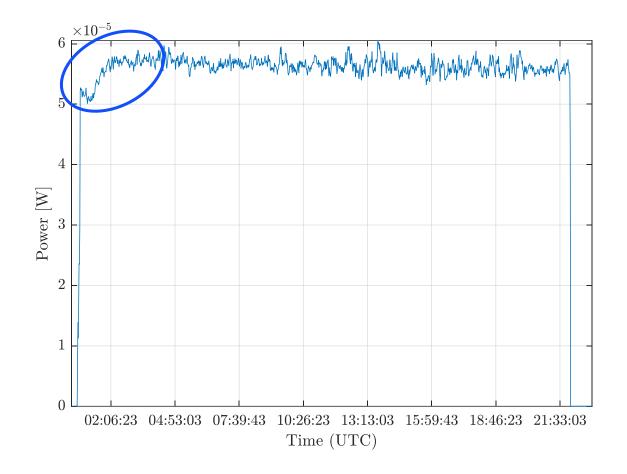
 Choose a longest period (2/25-6/6, 3 month) when beam positions are controlled.

	ITMX	ETMX	ITMY	ETMY
PD1	-16±1%	21.7±0.5%	-8±2%	-13±2%
PD4	33.6±0.2%	13.7±0.5%	-4±7%	-20±20%

Difficult to conclude overall scatter growth.
Something else is changing to cause scatter change.



 More data on beam positions and scatter coupling during transience of lock. (ADS beam alignment)

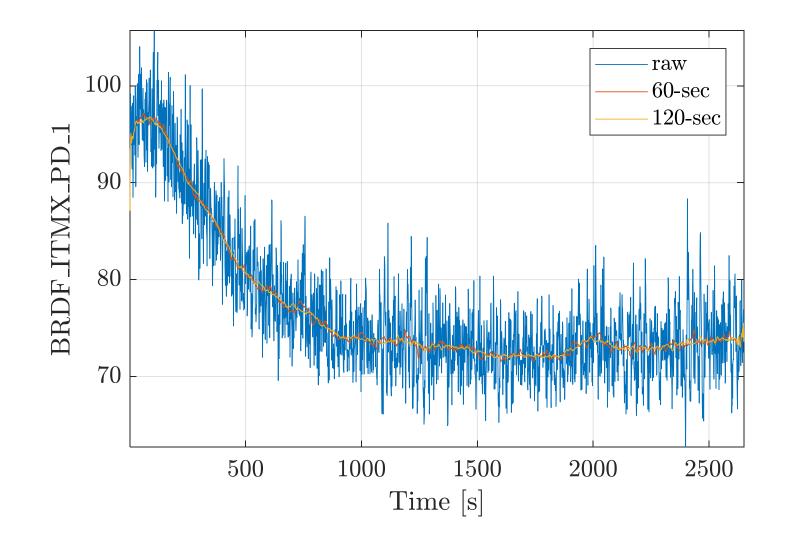


LIGO

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Scatter Data Processing

Moving-average filter (sample transience on 3/29/17)



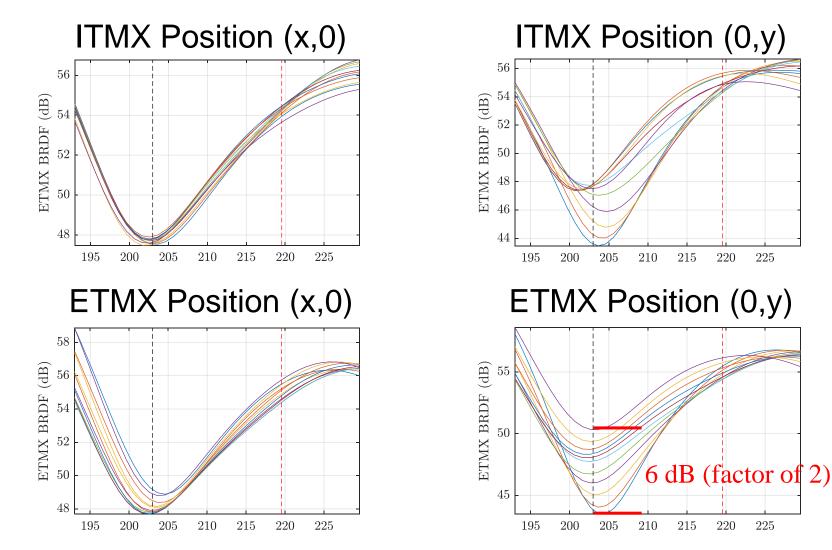
LIGO

Intro Steady-state

Transient

Scatter versus Beam on Both Sides

Range of change: [-10:2:10] mm



Intro

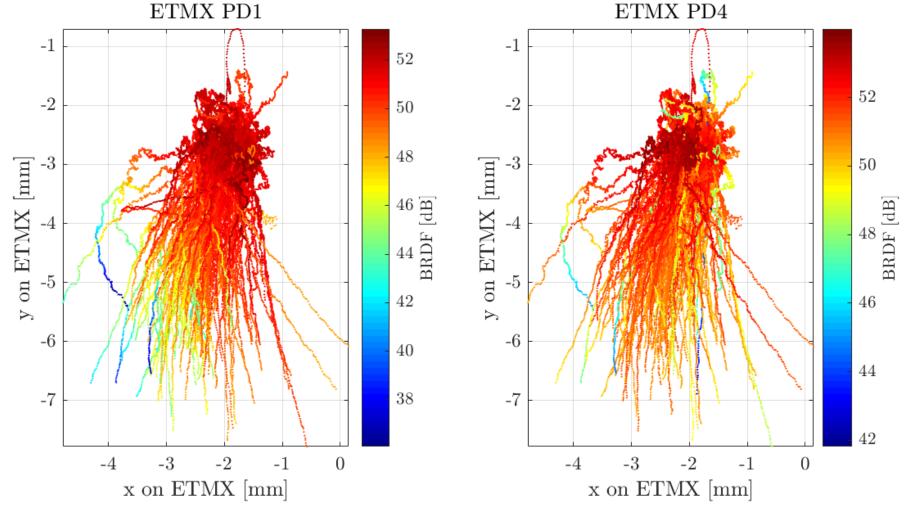
Transient

Regression Conclusion Q&A 21

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Transient Scatter in LSD

ETMX PD measurements (120-s MA filtered) through each alignment process, in LSD (2/25/17 - 6/6/17)



LIGO

Steady-state

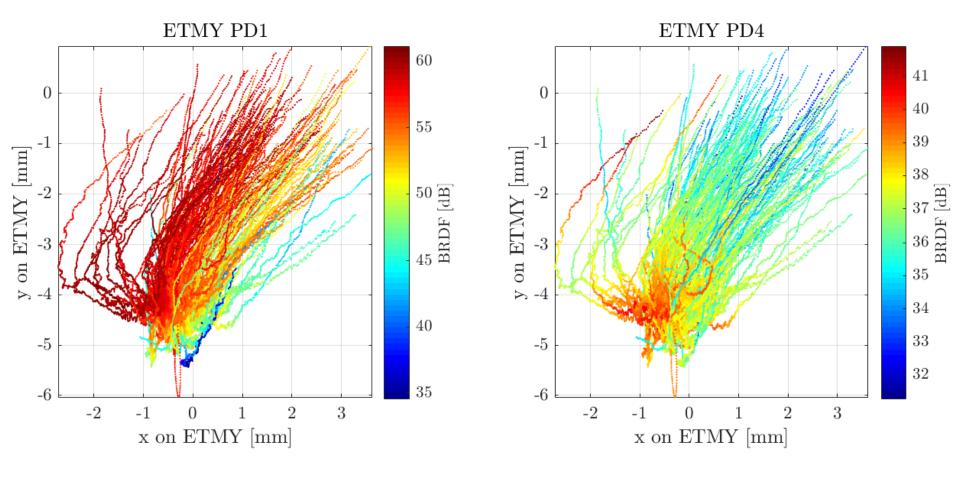
Transient

t Regression Conclusion

h Q&A 22

Transient Scatter in LSD

ETMY PD measurements (120-s MA filtered) through each alignment process, in LSD (2/25/17 - 6/6/17)



Transient LIGO-T1800224 Regression Conclusion 23 LIGO Scatter versus Beam on Both Sides 01-Mar-2017 Experimental data showing different scatter: 11-Mar-2017 25-Mar-2017 ITMY PD1 ITMY PD4 12-Apr-2017 475446 52[[[]] 50 48 48 BRDF [dB] 45 $4 \, \mathrm{dB}$ 447 dB 434642443 -1 2.5x on ETMY [mm] y on ETMY [mm] -2 21.5-3 1 0.5-4 0 -5 15001000 15002000 5001000 2000 25005002500Time [s]

Time [s]

Steady-state

Scatter versus Beam on Both Sides

- SIS doesn't agree with the measured scatter at the precision of ~3 dB and beam positions of ~5 mm.
- The SIS result depends on the beam position on both test mass.
- Use regression model to predict beam positions from scatter data.

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Ensemble Methods

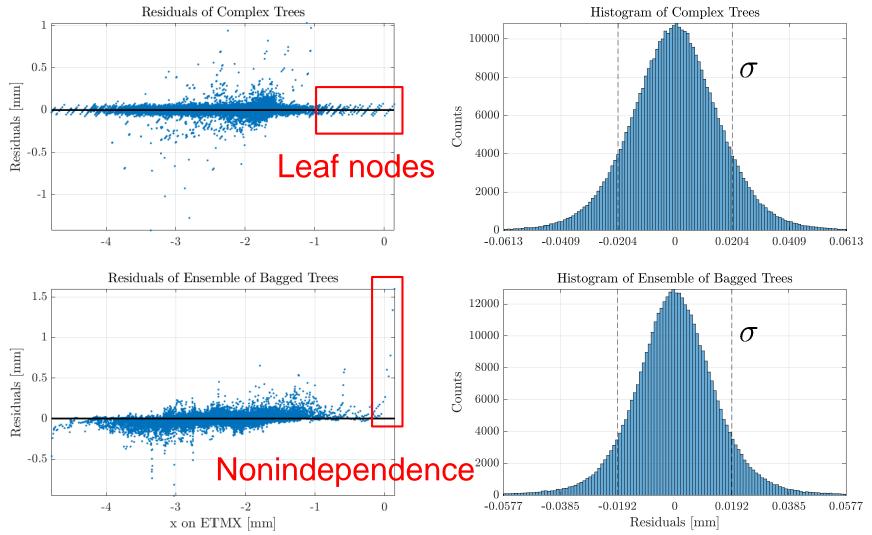
- Used measured scatter of 4 ETMX PD to predict beam positions measured by ADS error signal.
- Train the model with all 3-month data

Model	R ²	RMSE	Prediction Speed
Stepwise Linear (order of 5)	0.64	0.17	~1.6e6 obs/s
Stepwise Interaction Linear (order of 5)	0.73	0.15	~1.3e6 obs/s
Simple Tree (> 32 obs on node)	0.94	0.07	~1.7e6 obs/s
Complex Tree (> 4 obs on node)	0.98	0.04	~1.6e6 obs/s
Ensemble of Boosted Trees	0.62	0.17	~2.2e5 obs/s
Ensemble of Bagged Trees	0.99	0.03	~6.7e4 obs/s



Trees versus Ensemble

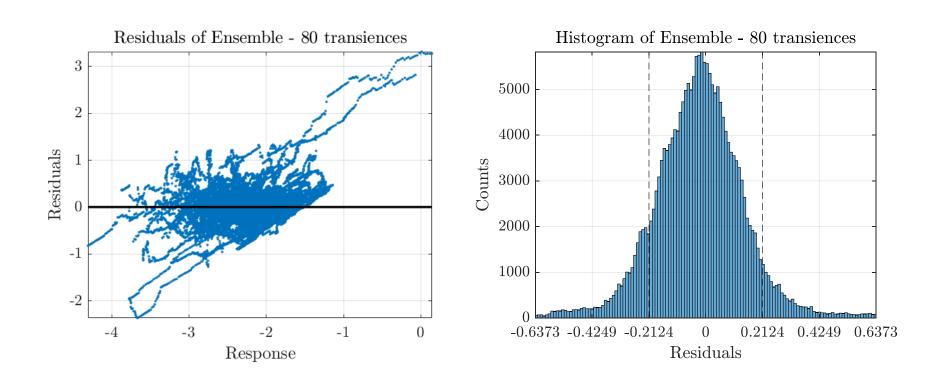
Residuals plot of the two best models with highest R^2

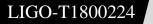


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Model Training

- In practical, we won't have $\sim 10^5$ data and 159 transience data before training model in O3 scatter.
- Test ensemble method with first half of data, and test it with another half of data.



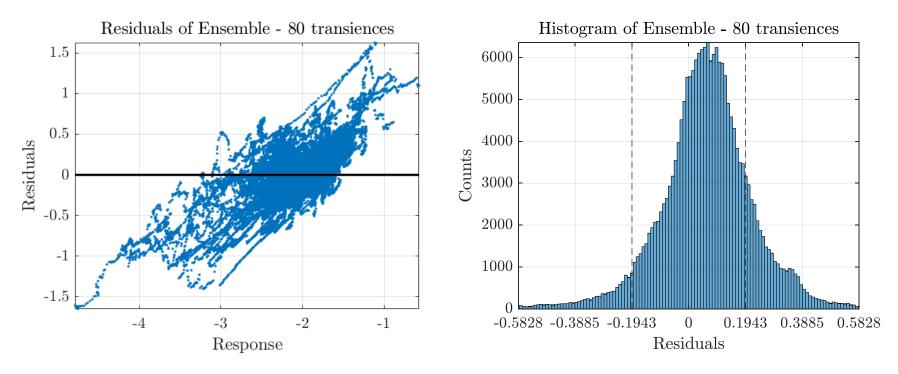


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Model Training

 Training the other way around: use last-half to train and predict first-half data

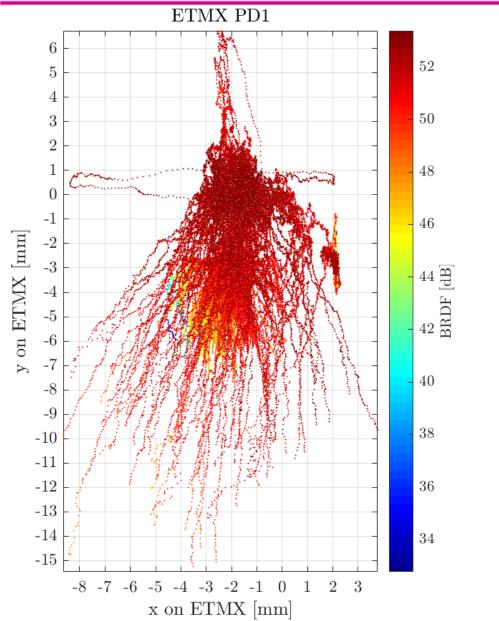


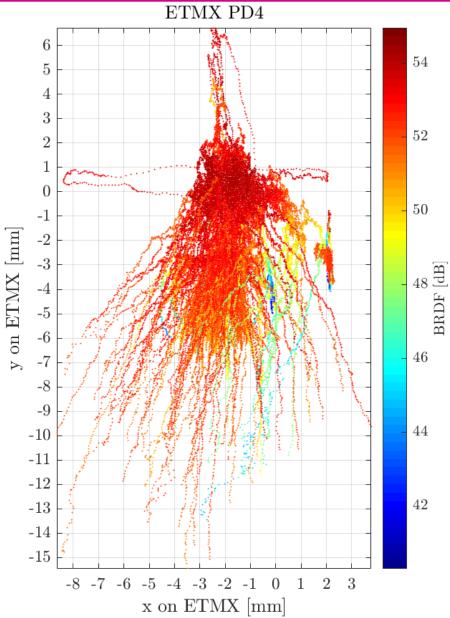
Use diverse data to have better model (avoid selecting same path)

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Transient Scatter in O2

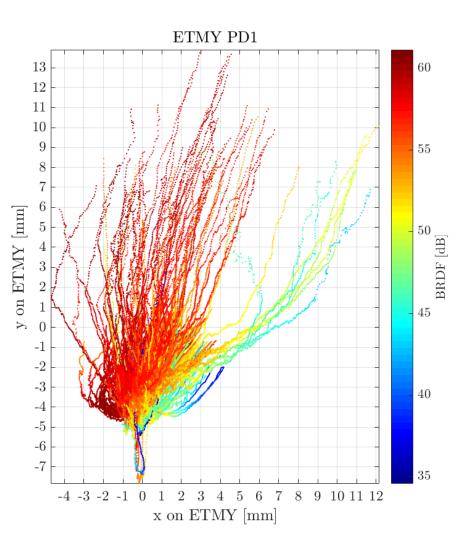


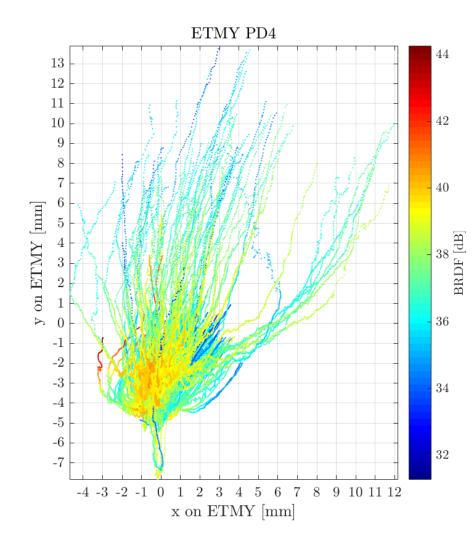


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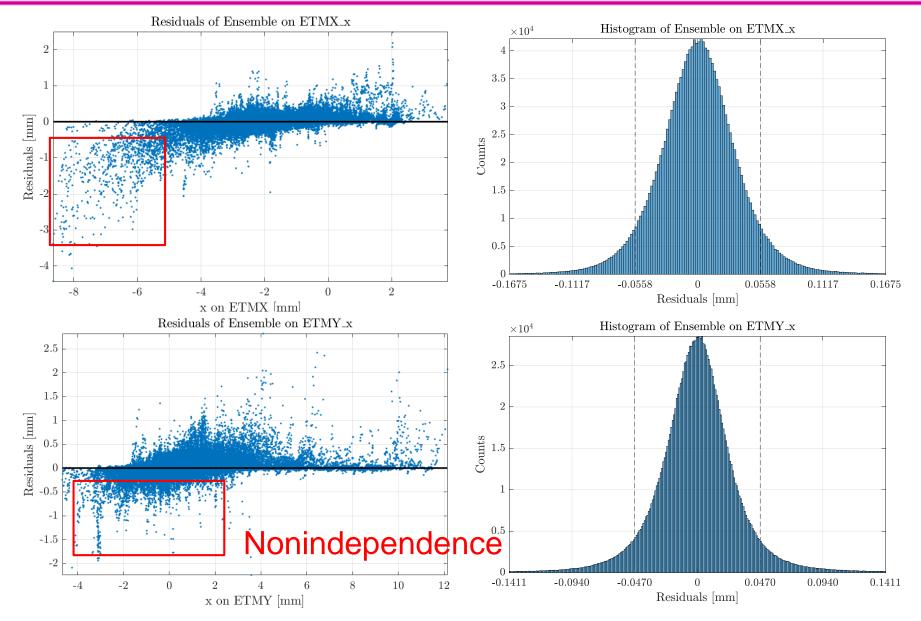
Transient Scatter in O2







Ensemble on O2 Data

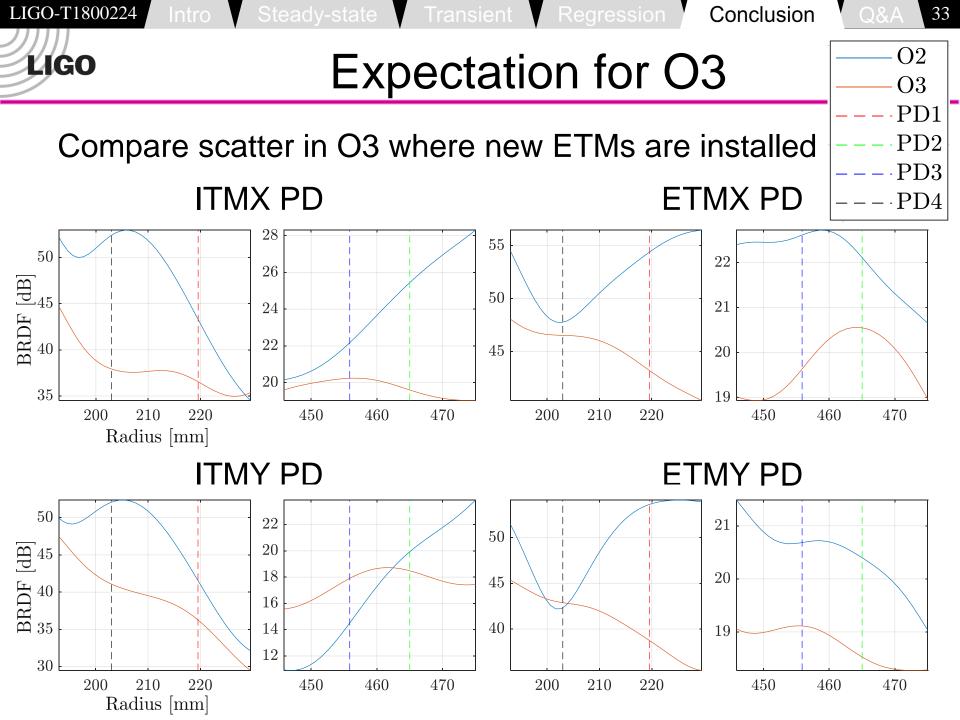






Conclusion

- Compare measured scatter to the simulation. The difference is at most 10 dB.
- Compare scatter with respect to beam position using transient data in the scale of millimeter and 5 dB. It is shown that the scatter depends upon beam positions on test masses on both sides.
- Proposed regression model to predict beam positions using scatter. The standard deviation of error is within 0.5 mm. We are still working on solving nonindependence problem.



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Acknowledgement

I am very grateful to Anamaria Effler and Valera Frolov for patient mentoring, Hiro Yamamoto for consistent help on SIS, and all LLO fellows for answering questions! Special thanks to California Institute of Technology and NSF to make SURF happen!

Thank you for listening!

Q&A



Intro

n Q&A

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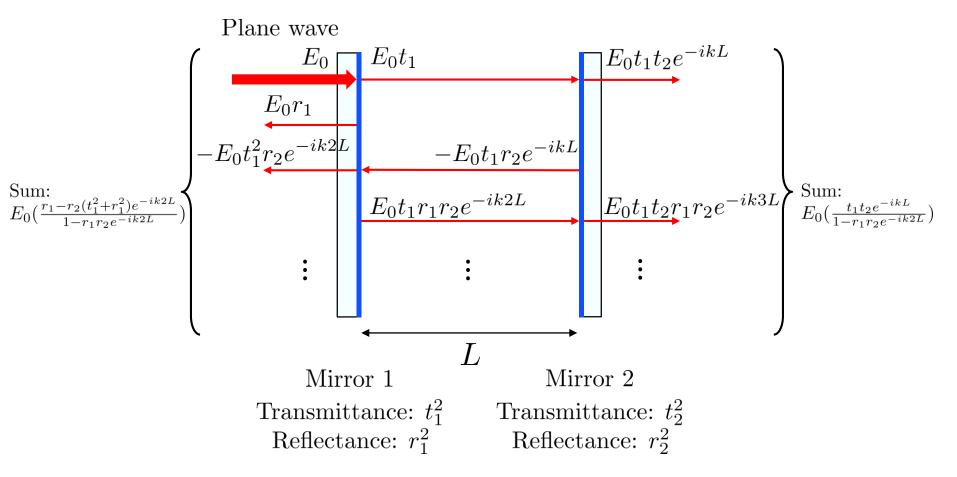


Backup Slides

LIGO

Simple Fabry-Perot Cavity

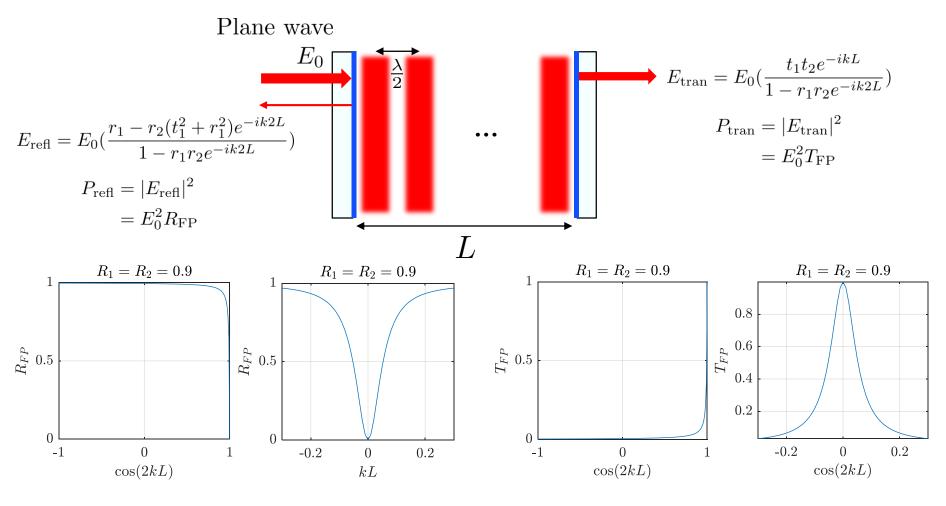
Fabry-Perot (FP) is composed of two highly reflective mirrors. A standing-wave is maintained in resonance.



LIGO

Simple Fabry-Perot Cavity

Fabry-Perot (FP) is composed of two highly reflective mirrors. A standing-wave is maintained in resonance.



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LIGO

Transient R

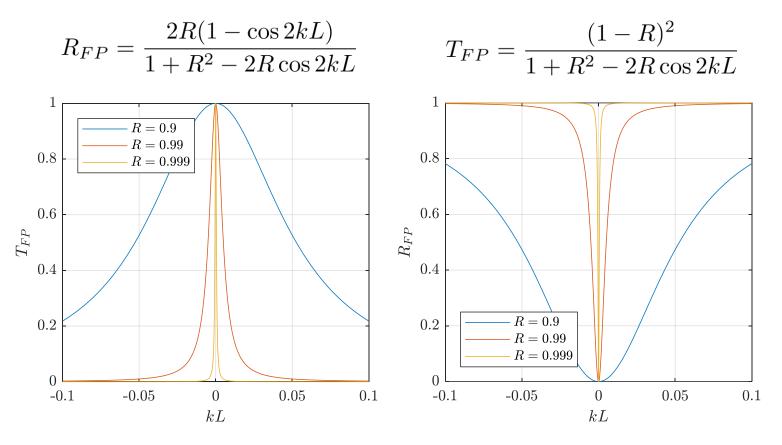
Regression

Q&A 38

Simple Fabry-Perot Cavity

The role of transmittance and reflectance in FP cavity:

For simplicity, let $r_1 = r_2 = \sqrt{R}$:



aLIGO FP Cavity

To achieve the nominal FP cavity, incredible engineering has been made:

- Stabilized laser at $\lambda = 1064 \text{ nm}$
 - » Relative intensity noise $< 2 \times 10^{-9} \ 1/\sqrt{\mathrm{Hz}}$ at 10 Hz.
 - » Frequency noise $< 1 \times 10^{-3}~{\rm Hz}/\sqrt{\rm Hz}$ at 100 Hz.
- Ultra-high vacuum in cavity: $< 4 \times 10^{-7}$ Pa
- Perfect optics (test mass):
 - » Smoothness: peak-to-peak $\sim 1~\mathrm{nm}$ (ITM), $\sim 4~\mathrm{nm}$ (ETM)
 - » Low transmission coating: ~ 0.014 (ITM), $\sim 5 \times 10^{-6}$ (ETM)
- Vibration isolation:
 - » Longitudinal noise $\sim 1 \times 10^{-19} \ m/\sqrt{Hz}$ at 10 Hz.

LIGO

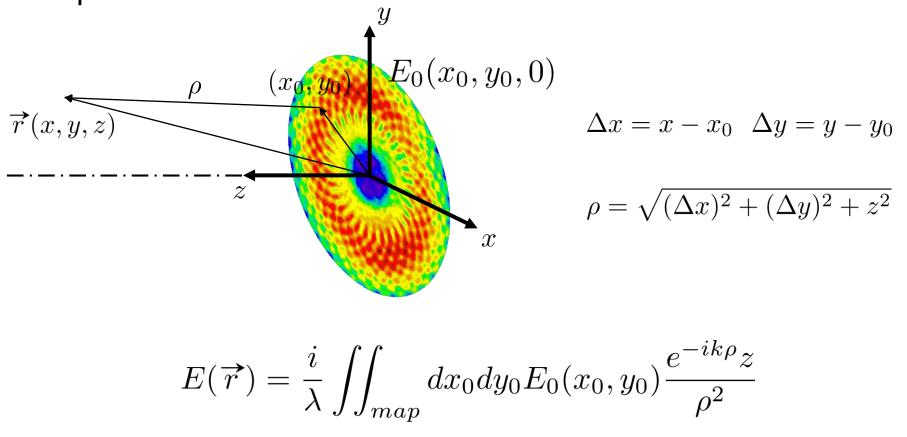
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Field Propagation in Free Space

Huygen's Principle:

Intro

Every point on the wavefront is the source of spherical wavelets.



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Field Propagation in Free Space

Fresnel approximation:

Intro

$$\rho = \sqrt{(\Delta x)^2 + (\Delta y)^2 + z^2} \approx z \left(1 + \frac{(\Delta x)^2 + (\Delta y)^2}{2z^2} \right)$$

So

$$E(\vec{r}) = \frac{i}{\lambda} \iint_{map} dx_0 dy_0 E_0(x_0, y_0) \frac{e^{-ikz}}{z} e^{-ik((\Delta x)^2 + (\Delta y)^2))/(2L)}$$

$$=\frac{ie^{-ikz}}{\lambda z}\iint_{map}dx_0dy_0E_0(x_0,y_0)e^{-ik\left((x-x_0)^2+(y-y_0)^2\right)/(2L)}$$

Convolution!

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Intro

Field Propagation in Free Space

- Optics can be understood as a space-invariant linear system.
- The diffraction is the convolution of starting field with paraxial diffraction kernel K

$$\tilde{E}_0(f_x, f_y) = \iint E_0(x, y) e^{-i2\pi(f_x x + f_y y)} dx dy$$

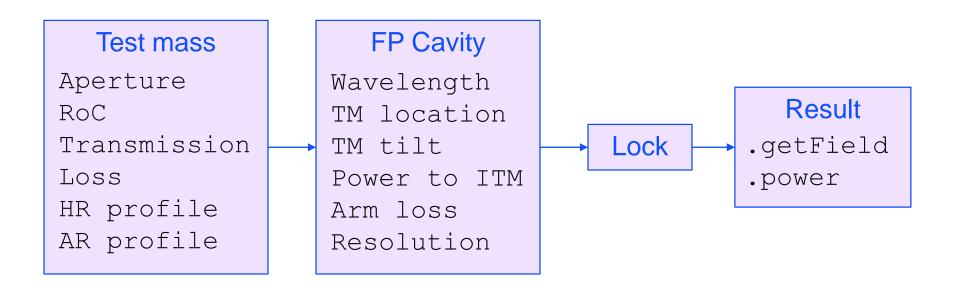
$$\tilde{K}(f_x, f_y) = \frac{ie^{-ikz}}{\lambda z} \iint e^{-ik(x^2 + y^2)/(2L)} e^{-i2\pi(f_x x + f_y y)} \, dx \, dy$$

 $\tilde{E}(\vec{r}) = \tilde{E}_0 \cdot \tilde{K}$

LIGO

Static Interferometer Simulation

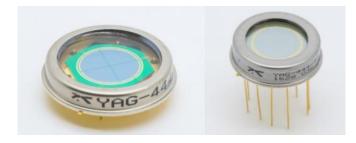
Given the dimensional and optical properties as well as the FP cavity configuration, the field amplitude distribution can be computed with a FFT-based tool: SIS.

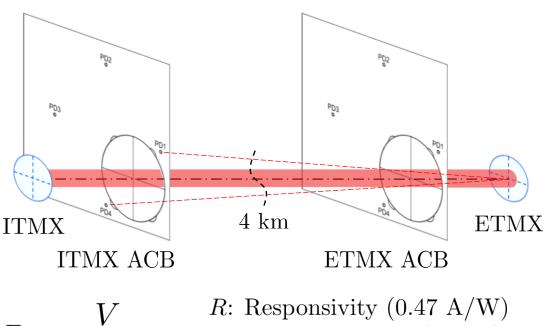


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Scatter Measurement

Quadrant photo diode: Excelitas YAG-444-AH

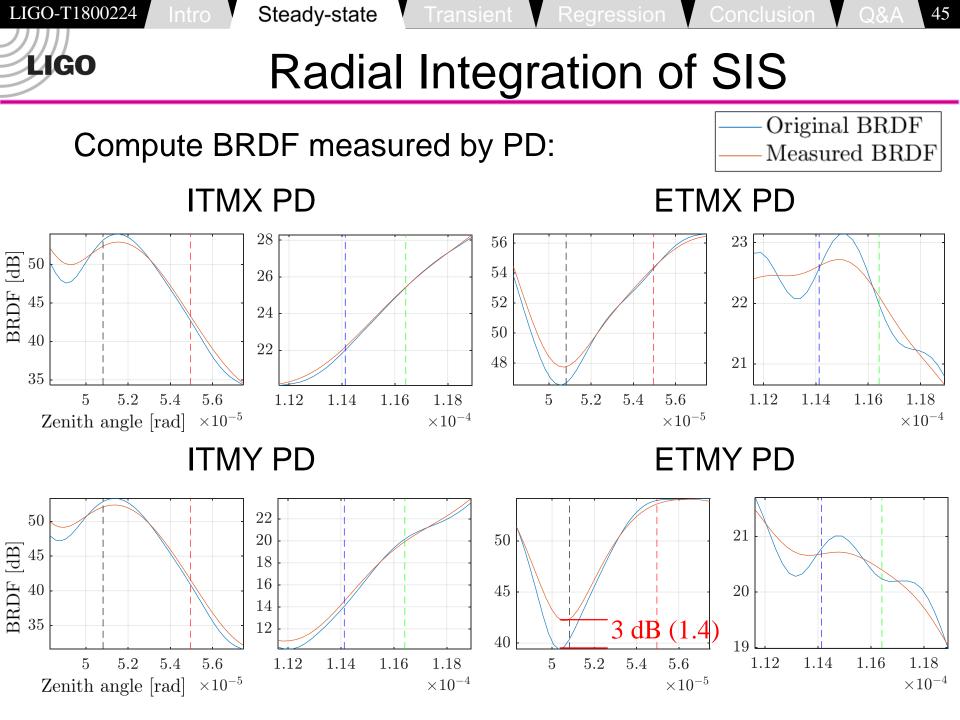




$$P = \frac{V}{RTG}$$

- *R*: Responsivity (0.47 A/W) *T*: Transimpedance (20 k Ω)
- G: Analog gain

2 arms \implies 4 baffles \implies 16 PD

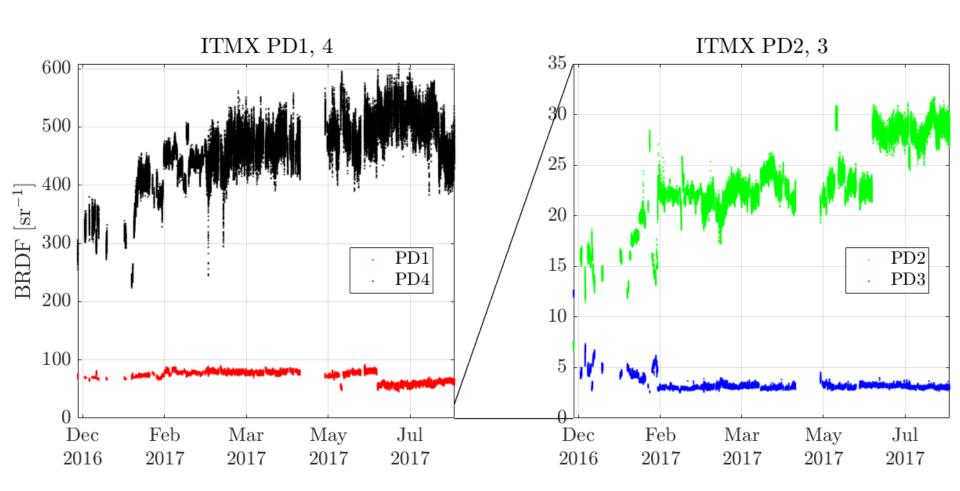


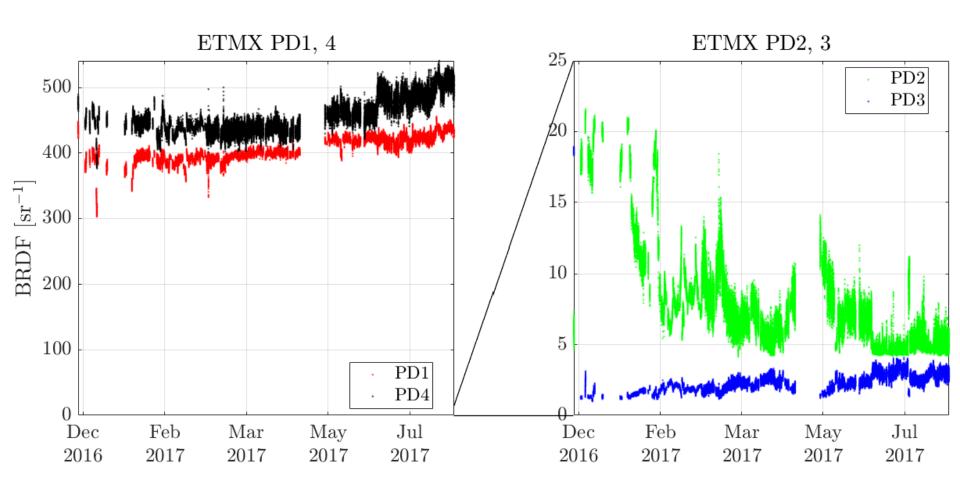


Scatter Data Processing

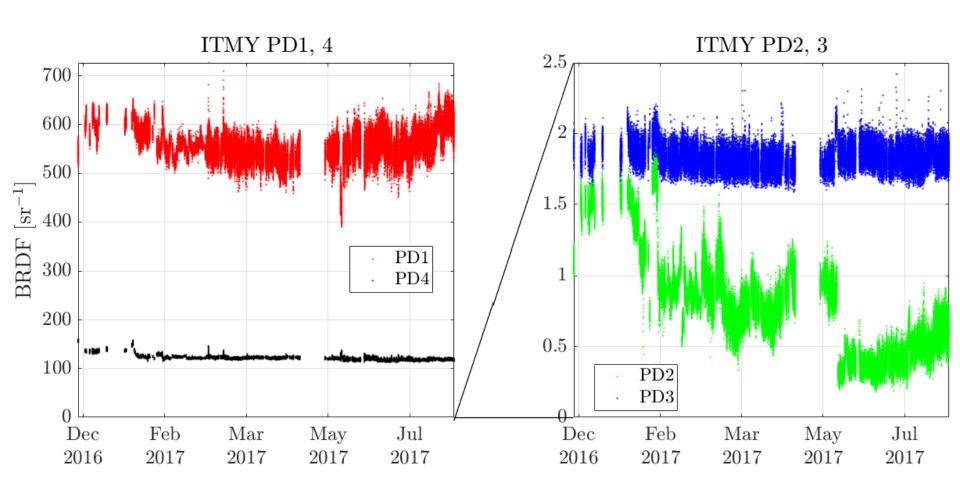
Factors	Channel Names
Lock State	L1:GRD-ISC_LOCK_STATE_N
Science Run Mode	L1:ODC-OPERATOR_OBSERVATION_READY
Power in PRC	L1:LSC-POP_A_LF_OUTPUT
Test Mass Attitude	L1:SUS-*TM*_L2_DRIVEALIGN_P2L_GAIN L1:SUS-*TM*_L2_DRIVEALIGN_Y2L_GAIN
PR2 Attitude	L1:SUS-PR2_M3_DRIVEALIGN_P2L_GAIN L1:SUS-PR2_M3_DRIVEALIGN_Y2L_GAIN

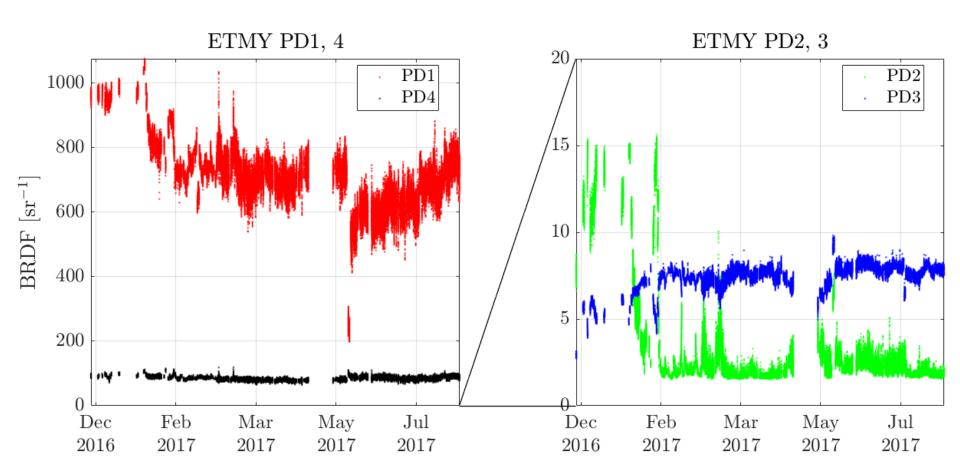
- Remove 3600 second transience on lock state, science run and all drivealign periods.
- Select locks with remaining time longer than 1 hour.

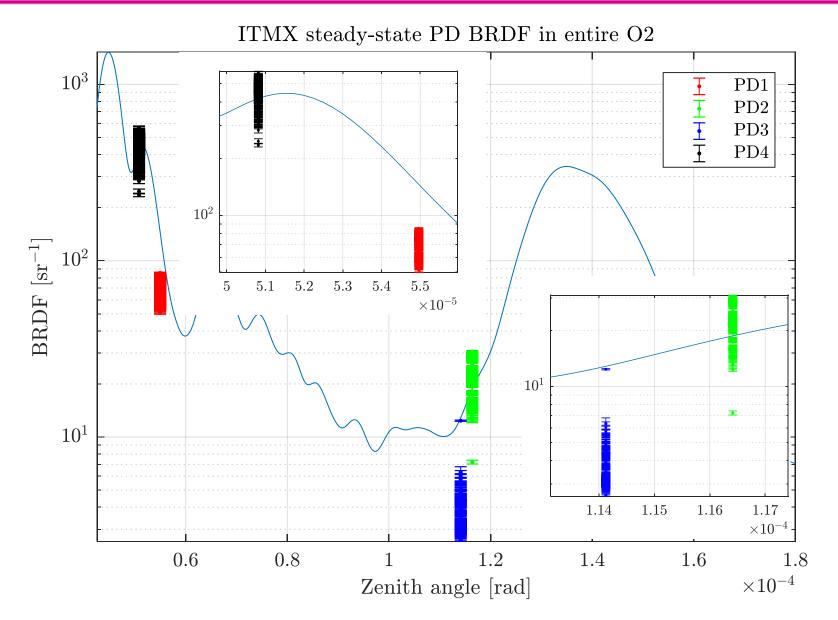




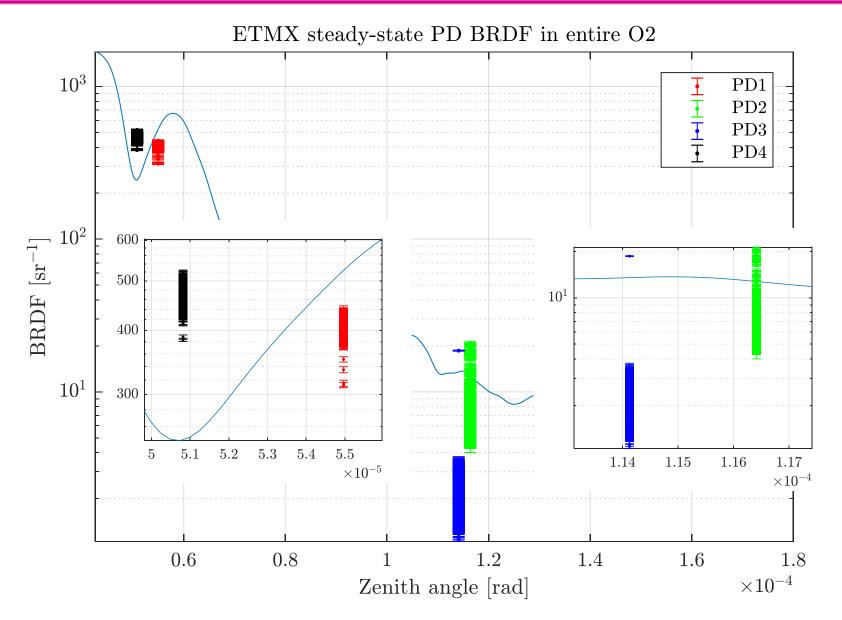
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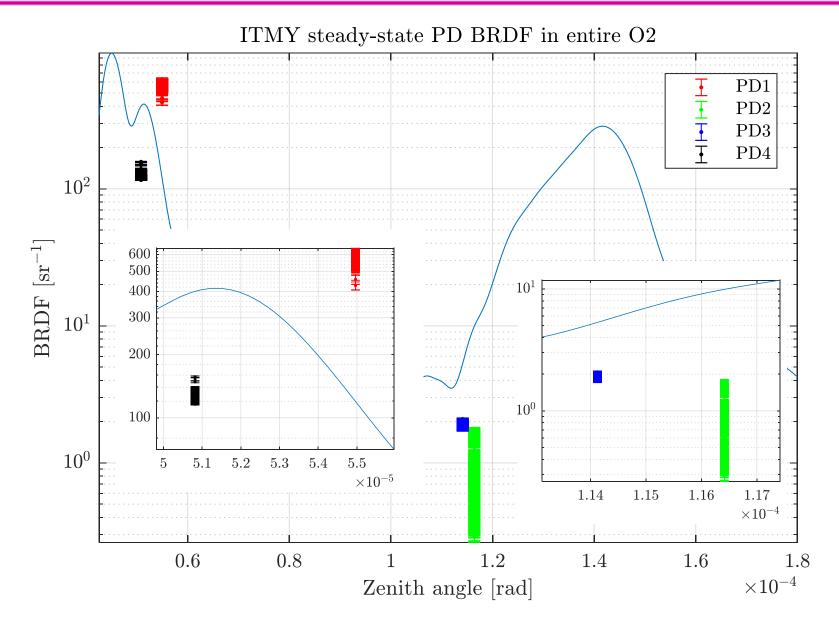




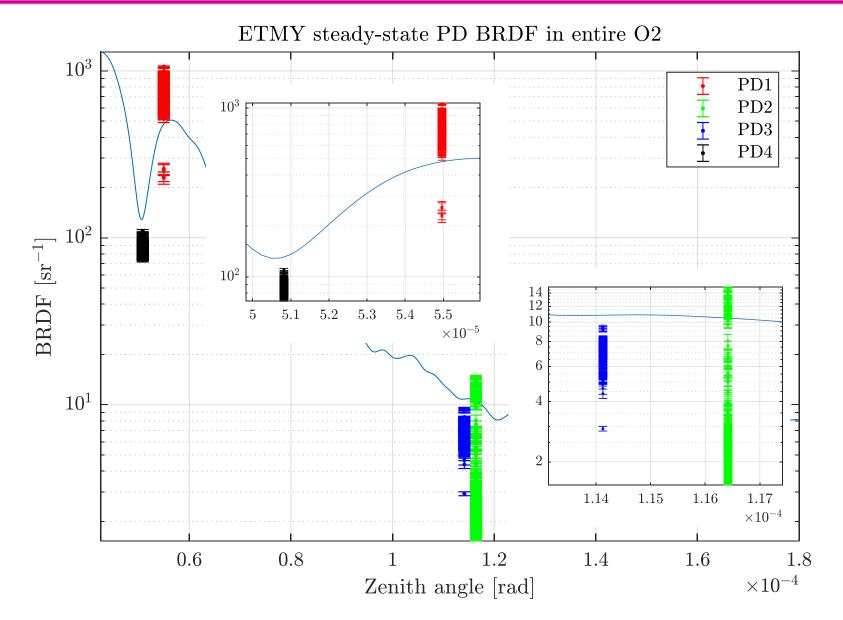


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Steady-state

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Regression

on Q&A

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Scatter with Beam Positions in O2

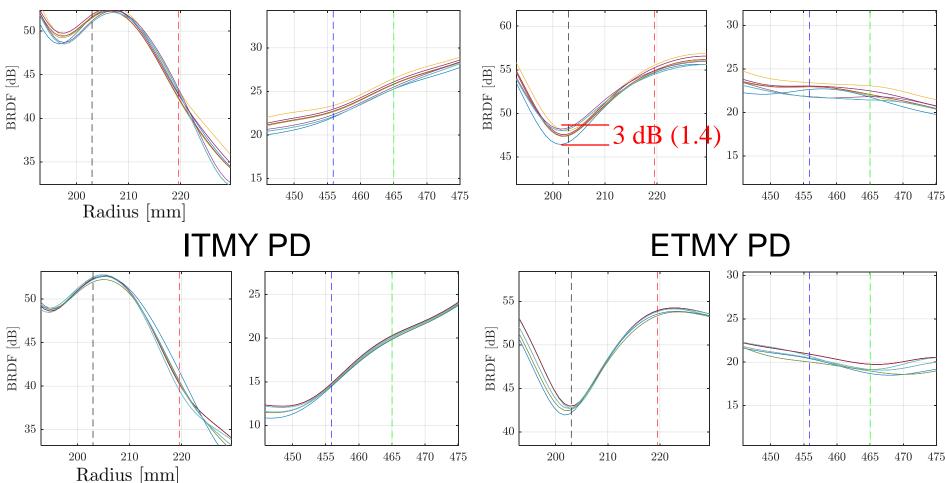
Scatter simulation with all lock beam positions (mirror tilting)

ITMX PD

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ETMX PD

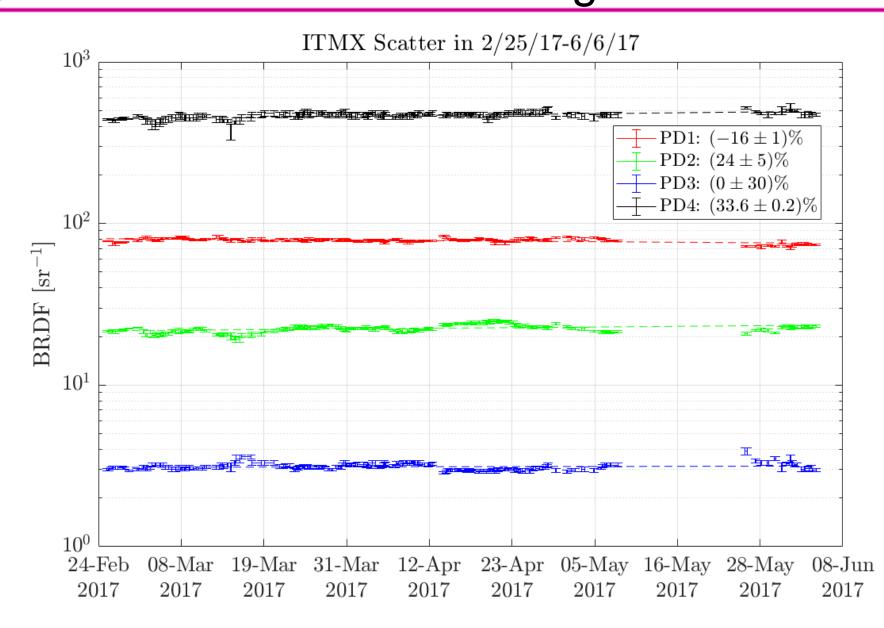


Longest Stable Duration in O2

02-Dec-2016: -6.21 3.50 2.10 -2.80 1.46 2.60 -0.35 -4.15 07-Dec-2016: -6.21 3.50 -1.75 -2.80 1.46 2.60 -0.35 -4.15 15-Dec-2016: -6.21 3.50 -3.50 -2.80 1.46 2.60 -0.35 -4.15 15-Dec-2016: -6.21 3.50 -3.50 -2.80 1.46 2.60 -0.35 -4.15 15-Dec-2016: -6.21 3.50 -1.75 -2.80 1.46 2.60 -0.35 -4.15 24-Jan-2017: -7.64 3.73 -1.75 -2.80 2.54 2.82 -0.35 -4.15 17-Feb-2017: -7.30 3.73 -1.75 -2.80 2.55 2.82 -0.35 -4.15 25-Feb-2017: -7.30 3.73 -1.75 -2.80 1.46 2.60 -0.35 -4.15 06-Jun-2017: -7.30 3.73 -1.75 0.35 1.46 2.60 -0.35 -2.80 02-Jul-2017: -7.30 3.73 -1.75 0.35 <td< th=""><th></th><th>ITMX_x</th><th>ITMX_Y</th><th>ETMX_x</th><th>ETMX_Y</th><th>ITMY_x</th><th>ITMY_Y</th><th>ETMY_x</th><th>ETMY_Y</th></td<>		ITMX_x	ITMX_Y	ETMX_x	ETMX_Y	ITMY_x	ITMY_Y	ETMY_x	ETMY_Y
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	02-Dec-2016:	-6.21	3.50						
15-Dec-2016:-6.213.50-1.75-2.801.462.60-0.35-4.1524-Jan-2017:-7.643.73-1.75-2.802.542.82-0.35-4.1517-Feb-2017:-7.303.73-1.75-2.802.552.82-0.35-4.1525-Feb-2017:-7.303.73-1.75-2.801.462.60-0.35-4.1506-Jun-2017:-7.303.73-1.750.351.462.60-0.350.3508-Jun-2017:-7.303.73-1.75-2.801.462.60-0.35-2.8002-Jul-2017:-7.303.73-1.750.351.462.60-0.35-2.8002-Jul-2017:-7.303.73-1.750.351.462.60-0.35-2.8002-Aug-2017:-6.213.73-1.750.351.464.40-0.35-2.80	07-Dec-2016:	-6.21	3.50	-1.75	-2.80	1.46	2.60	-0.35	-4.15
24-Jan-2017:-7.643.73-1.75-2.802.542.82-0.35-4.1517-Feb-2017:-7.303.73-1.75-2.802.552.82-0.35-4.1525-Feb-2017:-7.303.73-1.75-2.801.462.60-0.35-4.1506-Jun-2017:-7.303.73-1.750.351.462.60-0.350.3508-Jun-2017:-7.303.73-1.75-2.801.462.60-0.35-2.8002-Jul-2017:-7.303.73-1.750.351.462.60-0.35-2.8002-Jul-2017:-7.303.73-1.750.351.462.60-0.35-2.8002-Aug-2017:-6.213.73-1.750.351.464.40-0.35-2.80	15-Dec-2016:	-6.21	3.50	-3.50	-2.80	1.46	2.60	-0.35	-4.15
17-Feb-2017:-7.303.73-1.75-2.802.552.82-0.35-4.1525-Feb-2017:-7.303.73-1.75-2.801.462.60-0.35-4.1506-Jun-2017:-7.303.73-1.750.351.462.60-0.350.3508-Jun-2017:-7.303.73-1.75-2.801.462.60-0.35-2.8002-Jul-2017:-7.303.73-1.750.351.462.60-0.35-2.8002-Jul-2017:-7.303.73-1.750.351.462.60-0.35-2.8002-Aug-2017:-6.213.73-1.750.351.464.40-0.35-2.80	15-Dec-2016:	-6.21	3.50	-1.75	-2.80	1.46	2.60	-0.35	-4.15
25-Feb-2017:-7.303.73-1.75-2.801.462.60-0.35-4.1506-Jun-2017:-7.303.73-1.750.351.462.60-0.350.3508-Jun-2017:-7.303.73-1.75-2.801.462.60-0.35-2.8002-Jul-2017:-7.303.73-1.750.351.462.60-0.35-2.8002-Jul-2017:-7.303.73-1.750.351.462.60-0.35-2.8002-Aug-2017:-6.213.73-1.750.351.464.40-0.35-2.80	24-Jan-2017:	-7.64	3.73	-1.75	-2.80	2.54	2.82	-0.35	-4.15
06-Jun-2017:-7.303.73-1.750.351.462.60-0.350.3508-Jun-2017:-7.303.73-1.75-2.801.462.60-0.35-2.8002-Jul-2017:-7.303.73-1.750.351.462.60-0.35-2.8002-Aug-2017:-6.213.73-1.750.351.464.40-0.35-2.80	17-Feb-2017:	-7.30	3.73	-1.75	-2.80	2.55	2.82	-0.35	-4.15
08-Jun-2017: -7.30 3.73 -1.75 -2.80 1.46 2.60 -0.35 -2.80 02-Jul-2017: -7.30 3.73 -1.75 0.35 1.46 2.60 -0.35 -2.80 02-Jul-2017: -6.21 3.73 -1.75 0.35 1.46 4.40 -0.35 -2.80	25-Feb-2017:	-7.30	3.73	-1.75	-2.80	1.46	2.60	-0.35	-4.15
02-Jul-2017: -7.30 3.73 -1.75 0.35 1.46 2.60 -0.35 -2.80 02-Aug-2017: -6.21 3.73 -1.75 0.35 1.46 4.40 -0.35 -2.80	06-Jun-2017:	-7.30	3.73	-1.75	0.35	1.46	2.60	-0.35	0.35
02-Aug-2017: -6.21 3.73 -1.75 0.35 1.46 4.40 -0.35 -2.80	08-Jun-2017:	-7.30	3.73	-1.75	-2.80	1.46	2.60	-0.35	-2.80
	02-Jul-2017:	-7.30	3.73	-1.75	0.35	1.46	2.60	-0.35	-2.80
03-Aug-2017: -7.30 3.73 -1.75 0.35 1.46 4.40 -0.35 -2.80	02-Aug-2017:	-6.21	3.73	-1.75	0.35	1.46	4.40	-0.35	-2.80
	03-Aug-2017:	-7.30	3.73	-1.75	0.35	1.46	4.40	-0.35	-2.80
04-Aug-2017: -7.30 3.73 -1.75 0.35 1.46 2.60 -0.35 -2.80	04-Aug-2017:	-7.30	3.73	-1.75	0.35	1.46	2.60	-0.35	-2.80

- In the longest stable duration (LSD), the beam positions are constant in all locks.
- It would be interesting to see how scatter varied during LSD

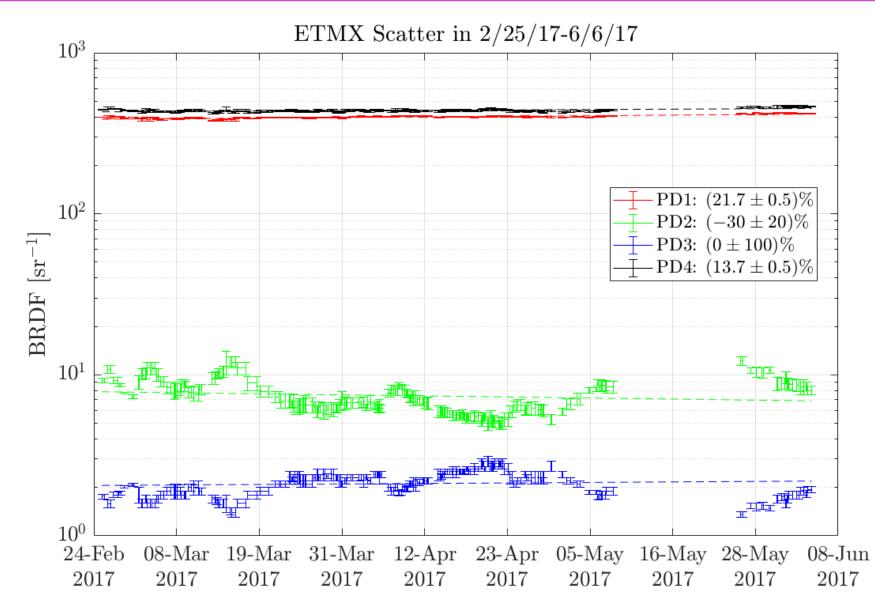
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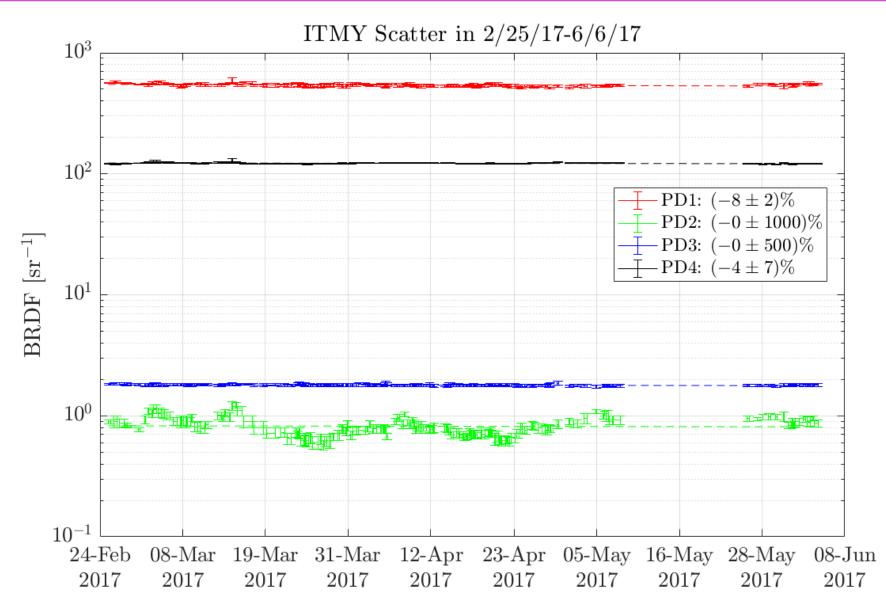


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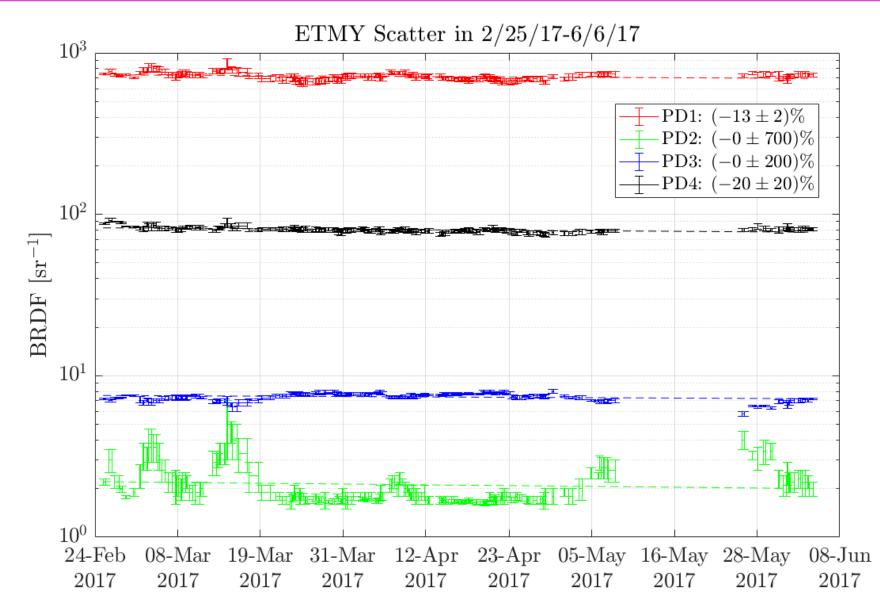
LIGO

LIGO-T1800224





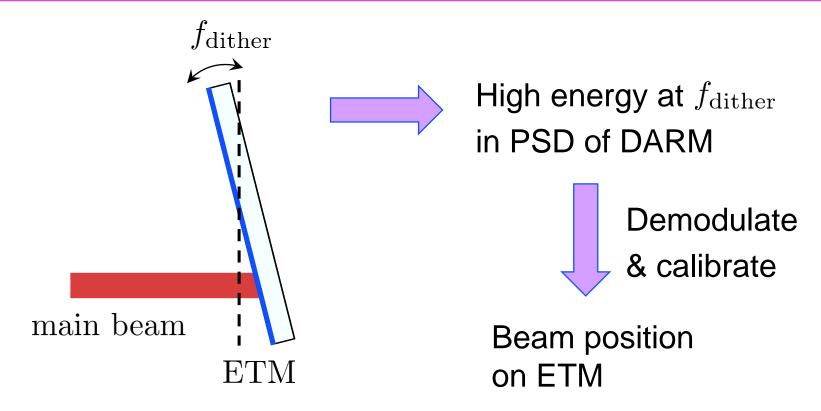
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Alignment Dither System



Demodulated amplitude ∝ dither amplitude
∝ beam offset

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Steady-state

Transient

Regression Conclusion Q&A

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Alignment Dither System





Scatter Data Processing

Factors	Channel Names				
General feedback gain	L1:ASC-ADS GAIN				
ETMX(Y/PR2) Pitch dither magnitude	L1:ASC-ADS PIT3(4/5) OSC CLKGAIN				
ETMX(Y/PR2) Pitch error	L1:ASC-ADS PIT3(4/5) DOF INMON				
ETMX(Y/PR2) Yaw dither magnitude	L1:ASC-ADS YAW3(4/5) OSC CLKGAIN				
ETMX(Y/PR2) Yaw error	L1:ASC-ADS YAW3(4/5) DOF INMON				

- 25 W-laser output alignment starts when LOCK_STATE_N = 1731
- Select locks with duration longer than 1 hour.

X/Y offset [m] =
$$\frac{INMON}{(CLKGAIN)(ADSGAIN)}k, \quad k = \begin{cases} 0.98, & \text{ETMX pitch} \\ -0.72, & \text{ETMX yaw} \\ 0.95, & \text{ETMY pitch} \\ -0.76, & \text{ETMY yaw} \end{cases}$$

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LIGO

Regression Model

- Nonparametric regression model
 - » High-dimensional predictors
 - » Complex
- Parametric regression model
 - » Physically interpretable
 - » Computationally efficient
 - » Easy diagnostics
 - » Most common one: linear model

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General Linear Regression

- General linear regression model
 - » Response: y
 - » *p*-dimensional predictors: $x_i, i = 1, 2, ..., p$
 - » With *n*-observation: $\mathbf{y} = \mathbf{X}\boldsymbol{\beta} + \boldsymbol{\varepsilon}$, $n \sim 10^5$ in LSD the design matrix (without interaction):

$$\mathbf{X} = \begin{bmatrix} 1 & x_{11} & x_{11}^2 & \dots & x_{11}^k & x_{21} & \dots & x_{21}^k & \dots & x_{p1}^k \\ \vdots & \vdots & \vdots & & \vdots & \vdots & & \vdots & & \vdots \\ 1 & x_{1j} & x_{1j}^2 & \dots & x_{1j}^k & x_{2j} & \dots & x_{2j}^k & \dots & x_{pj}^k \\ \vdots & \vdots & \vdots & & \vdots & & \vdots & & \vdots & & \vdots \\ 1 & x_{1n} & x_{1n}^2 & \dots & x_{1n}^k & x_{2n} & \dots & x_{2n}^k & \dots & x_{pn}^k \end{bmatrix}$$

with interaction:

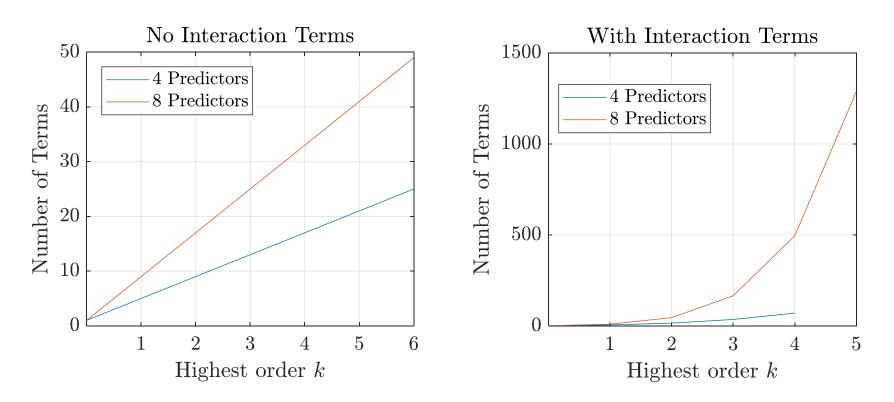
$$\mathbf{X}_j = \begin{bmatrix} 1 & x_{1j} & \dots & x_{1j}^{k_1} x_{2j}^{k_2} \dots x_{pj}^{k_p} & \dots & x_{pj}^k \end{bmatrix}$$

where $k_1, k_2...k_p \in \{0, 1...k\}$ and $\sum_{i=1}^p k_i \leqslant k$

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Stepwise Linear Regression

- Backward stepwise regression
 - » Dimensional reduction
 - » Trim model parameters to avoid overfitting and complexity
 - » 1 response: x/y beam position
 - » Use 4 predictors (4 PD) to avoid blowing up terms



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Conclusion Q&A 67

Stepwise Linear Regression

Use measured beam position (*x*) on ETMX from 4 ETMX PD:

