

Predicting Remnant Parameters in Black Hole Binaries

Using Generic Machine Learning Approaches

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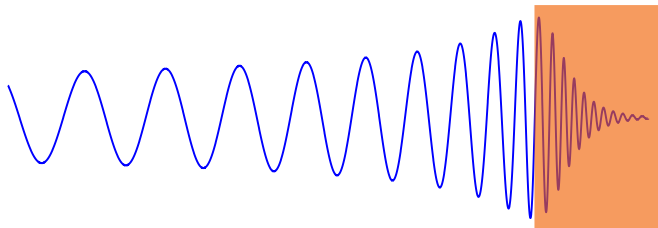
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Overview

- Predictions of mass and spin of remnant after binary black hole (BBH) merger
 - Numerical relativity (NR) can do this, but slow
 - Approximate fits: quicker, less computation
- **Goal:** improve fit accuracy
 - more data
 - new techniques
- Compare new fits with previous work
 - Healy and Lousto (2016)
 - Current formulas in LSC Algorithm Library (LAL)

Application



- Effective One Body (EOB) formalism
 - Compute waveform during merger and ringdown
 - LAL uses mass and spin fits for this

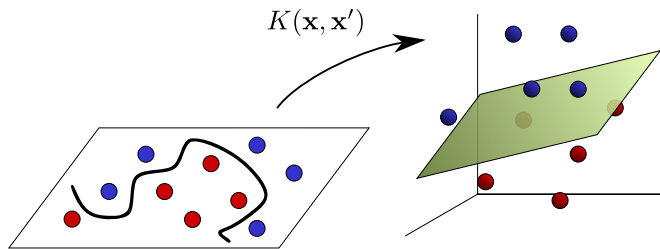
Data and fitting methods

- Data: SXS catalog of BBH simulations
 - Public: ~ 300
 - Not yet public: ~ 1100
- Generic regression techniques

Parametric vs Non-parametric Fit

- **Parametric** fits optimize function parameters
 - (e.g. $y = A \cos kx + B$)
 - Polynomial expansions (**previous fits**)
- **Non-parametric** fits use training data itself to make predictions
 - Gaussian Process Regression (GPR)
 - Decision Tree Regression

Gaussian Process Regression (GPR)



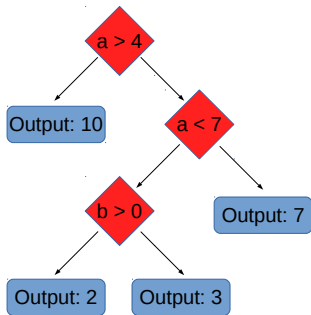
- **Kernel:** nonlinear data transformation
 - RBF kernel

$$K(\mathbf{x}, \mathbf{x}') = \exp(-\gamma \|\mathbf{x} - \mathbf{x}'\|^2) \quad (1)$$

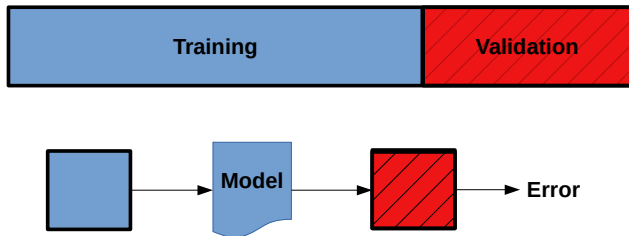
- Training = set kernel hyperparameters (γ)

Decision Tree Regression

- Predictions based on successive **decisions** about input parameters
- Training = construction of **decision tree** from data
- **Ensemble** = combine multiple decision trees

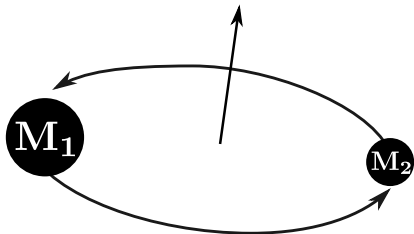


Measuring error



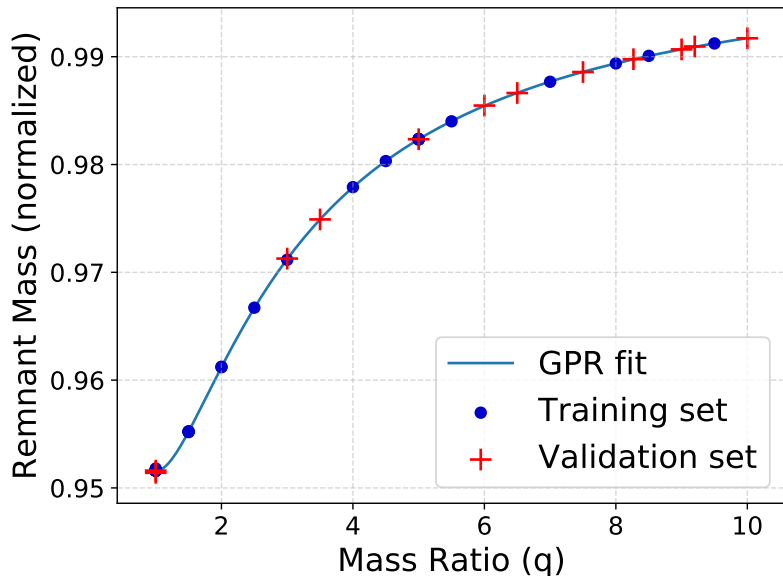
- Training and Validation set partition
 - Designate subset of data not used in training

Spinless (~ 50 simulations)

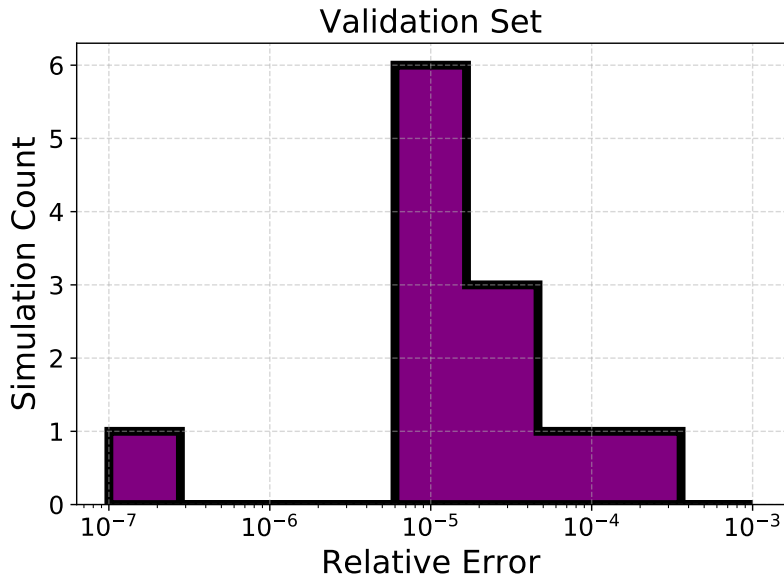


- Only one input parameter (mass ratio $q = \frac{M_1}{M_2}$)
- Fit with GPR

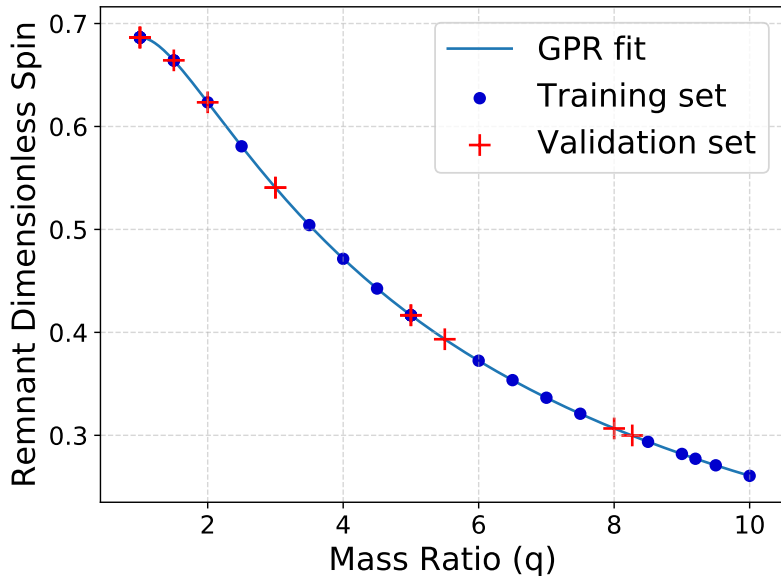
Spinless: remnant mass plot



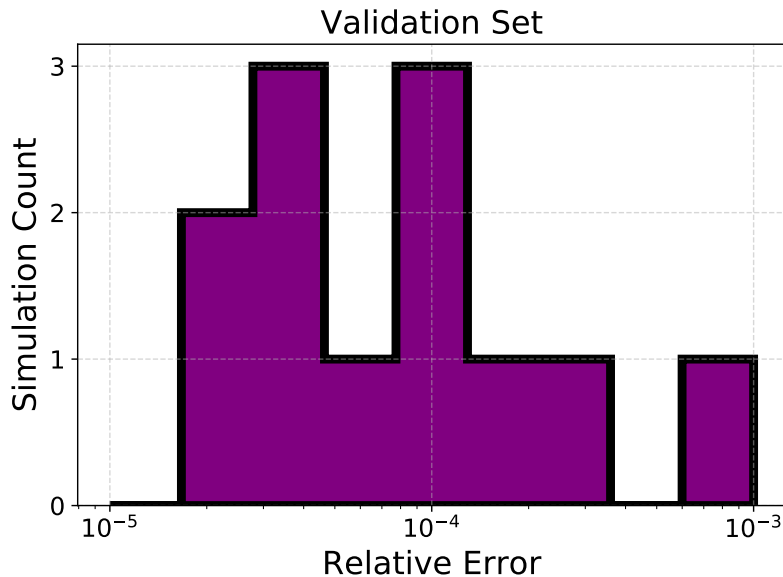
Spinless: remnant mass residuals



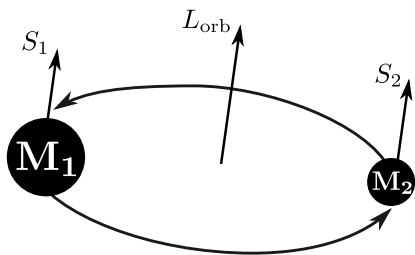
Spinless: remnant spin plot



Spinless: remnant spin residuals

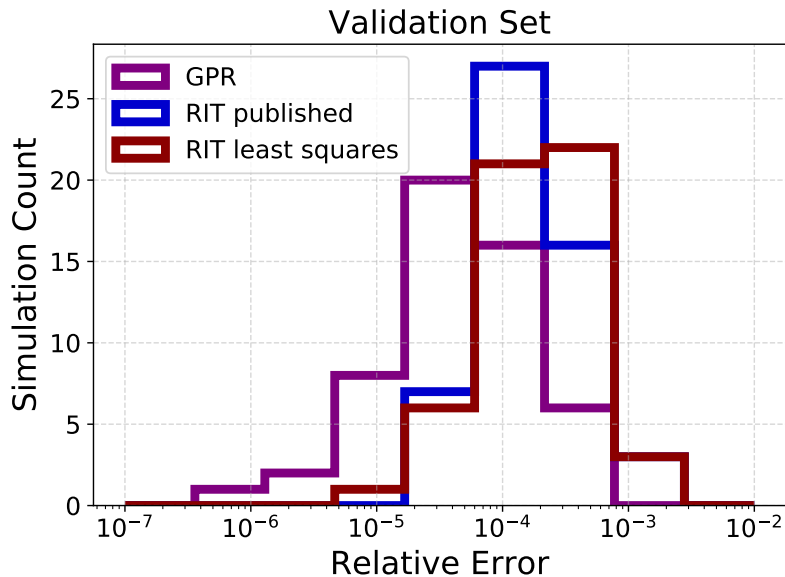


Aligned (~ 200 simulations)

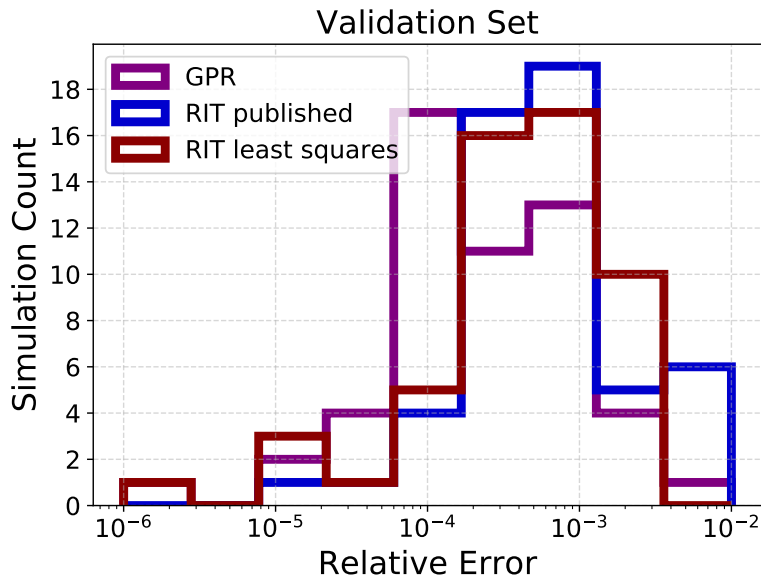


- **Three** input parameters (q, S_1, S_2)
- GPR vs Healy and Lousto (2016) - RIT

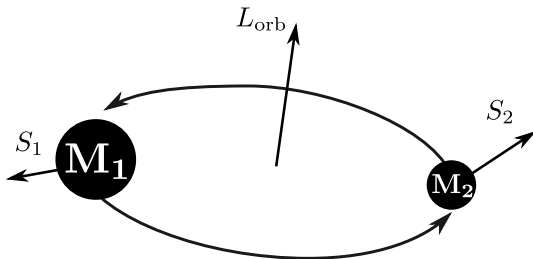
Aligned mass



Aligned spin



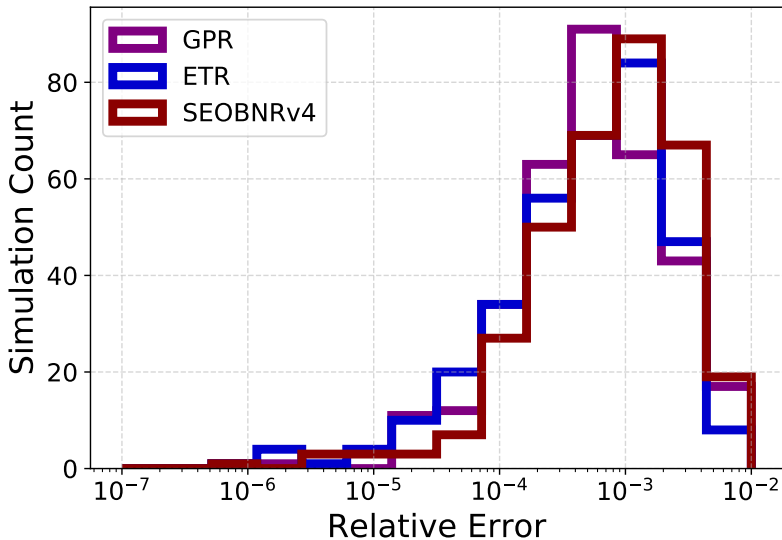
Generic spins (~ 1400 simulations)



- **7 dimensional** input space
- Comparing:
 - GPR
 - Extremely Randomized Trees (ensemble)
 - EOBFits implemented in LAL

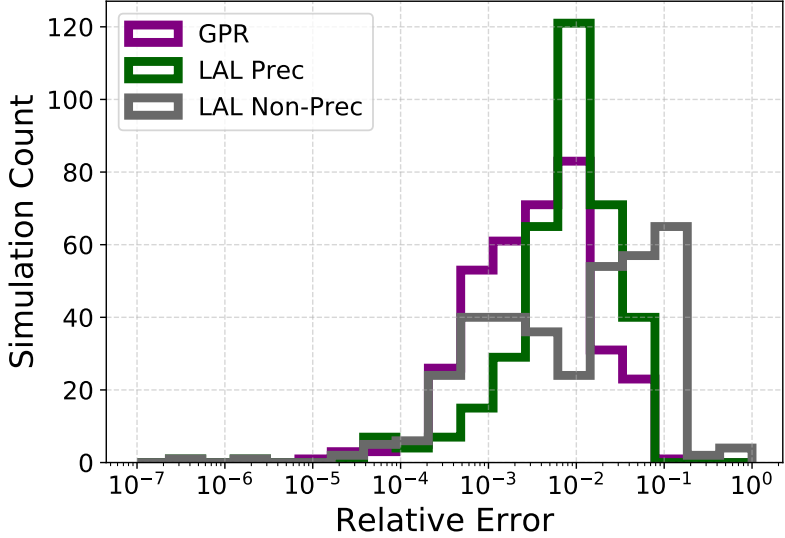
Remnant mass

Validation Set



Remnant spin magnitude

Validation Set



Summary

- Remnant parameter predictions needed
- Can predict with NR simulations
 - Accurate, but slow and computationally intensive
 - LIGO data analysis requires approximate fits
- Non-parametric methods can improve accuracy

Acknowledgments

