

Pseudo-quasinormal modes used in SEOBNRv3 (LIGO DCC: LIGO-T1600554-v2)

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Similarly to what is done in the underlying nonprecessing EOBNR model [1], pseudo-quasinormal modes (QNMs) are introduced in the precessing EOBNR model to bridge the gap between the end-of-inspiral and the least-damped-QNM frequency, especially for large mass ratios and large spin magnitudes, and replace some of the highest physical overtones. The pseudo-QNMs are functions that depend of the GW frequency at $t = t_{\text{match}}$ (the time where the ringdown is attached to the inspiral-plunge signal), the least-damped QNM frequency ω_{220} and decay time τ_{220} , and the symmetric mass ratio. Let M be the total mass of the binary, M_f be the mass of the remnant BH, ω_{22} be the instantaneous GW frequency. Let us define

$$\omega_{1,0} = \frac{1}{3} \left[2\omega_{22}(t_{\text{match}}) \frac{M}{M_f} + \omega_{220} \right], \quad (1)$$

$$\tau_{1,0} = 0.257 \tau_{220}, \quad (2)$$

$$\omega_{2,0} = \frac{1}{4} \left[3\omega_{22}(t_{\text{match}}) \frac{M}{M_f} + \omega_{220} \right], \quad (3)$$

$$\tau_{2,0} = 0.286 \tau_{220}. \quad (4)$$

Let us introduce the spin combination

$$\chi \equiv \frac{1}{1-2\nu} \frac{(\mathbf{S}_1 + \mathbf{S}_2) \cdot \hat{\mathbf{L}}_{\text{match}}}{M^2}, \quad (5)$$

where $\mathbf{S}_{1,2}$ are the component BH spins, ν is the symmetric mass ratio, $\hat{\mathbf{L}}_{\text{match}}$ is the direction of the orbital angular mo-

mentum at $t = t_{\text{match}}$, and the following functions:

$$f_\omega = 0.7 + 0.3 e^{100(\nu-1/4)}, \quad (6)$$

$$f_{\tau_1} = [0.5 (1 + 267 \nu^2)^{1/2} - 0.125] e^{-(\nu-0.005)/0.03}, \quad (7)$$

$$f_{\tau_2} = [0.5 (1 + 22.2 \nu^{3/2})^{2/3} - 0.2] e^{-(\nu-0.005)/0.03}. \quad (8)$$

By default, when building the QNM spectrum for the (2, 2) mode, we introduce two pseudo-QNMs whose frequencies and decay times are $\omega_i^{\text{pQNM}} = \omega_{i,0}$ and $\tau_i^{\text{pQNM}} = \tau_{i,0}$ ($i = 1, 2$), respectively. We have a special treatment for other parts of parameter space. When $q < 10$ and $\chi \geq 0.8$, we prescribe that $\omega_i^{\text{pQNM}} = f_\omega \omega_{i,0}$ and $\tau_1^{\text{pQNM}} = f_{\tau_1} \tau_{1,0}$, and $\tau_2^{\text{pQNM}} = 0.95 f_{\tau_2} \tau_{2,0}$. When $10 \leq q < 30$ and $\chi \geq 0.8$ or when $q \geq 30$ and $0.8 \leq \chi < 0.9$, we prescribe that $\omega_i^{\text{pQNM}} = f_\omega \omega_{i,0}$ and $\tau_i^{\text{pQNM}} = f_{\tau_i} \tau_{i,0}$ ($i = 1, 2$), and we also introduce two additional pseudo-QNMs whose frequencies and decay times read $\omega_{i+2}^{\text{pQNM}} = 0.4 (1 + f_\omega) \omega_{i,0}$ ($i = 1, 2$), $\tau_3^{\text{pQNM}} = 1.5 f_{\tau_1} \tau_1^{\text{pQNM}}$, and $\tau_4^{\text{pQNM}} = 2.5 f_{\tau_2} \tau_2^{\text{pQNM}}$, respectively. Finally, when $q \geq 30$ and $\chi \geq 0.9$, we slightly modify the third and fourth pseudo-QNMs just discussed, namely we put $\omega_3^{\text{pQNM}} = 0.4 (1 + f_\omega) \omega_{1,0}$, $\omega_4^{\text{pQNM}} = 0.44 (1 + f_\omega) \omega_{2,0}$, $\tau_3^{\text{pQNM}} = 1.575 f_{\tau_1} \tau_1^{\text{pQNM}}$, and $\tau_4^{\text{pQNM}} = 2.625 f_{\tau_2} \tau_2^{\text{pQNM}}$. When building the QNM spectrum for the (2, 1) mode, we introduce one pseudo-QNM whose frequency and decay time read $\omega_1^{\text{pQNM}} = \omega_{22}(t_{\text{match}})$ and $\tau_1^{\text{pQNM}} = 0.5 \tau_{220}$, respectively.

[1] A. Taracchini et al., Phys. Rev. **D89**, 061502 (2014), 1311.2544.