

Optics Thermal Noise Work at American

**Gregg Harry, Sam Hickey,
Jonathan Newport, Hannah Fair**

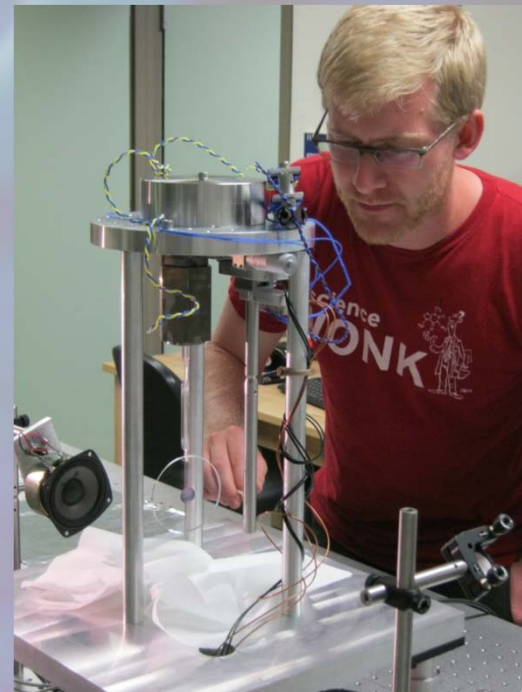
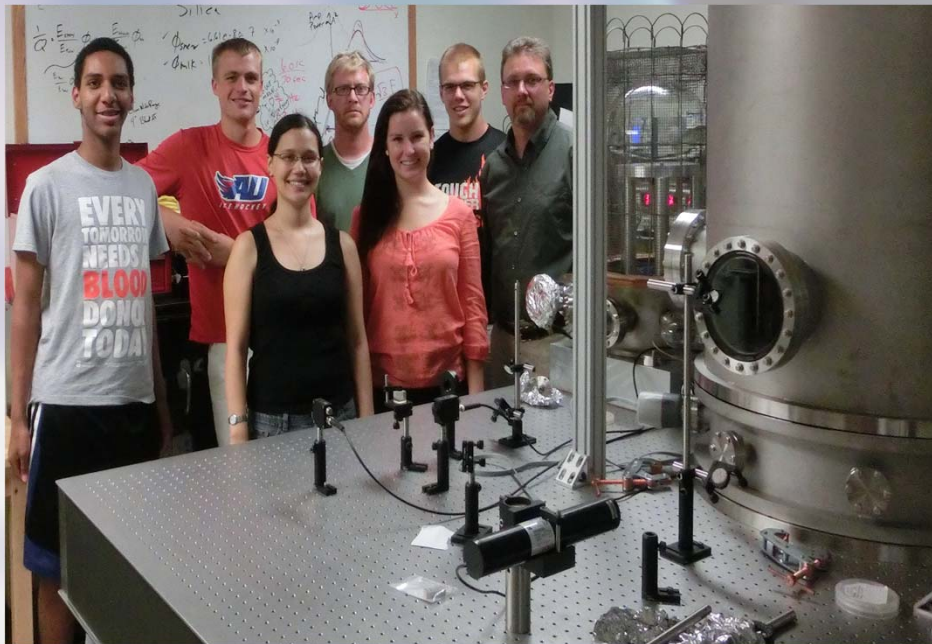
American University

OWG Session, August 2014 LVC Meeting



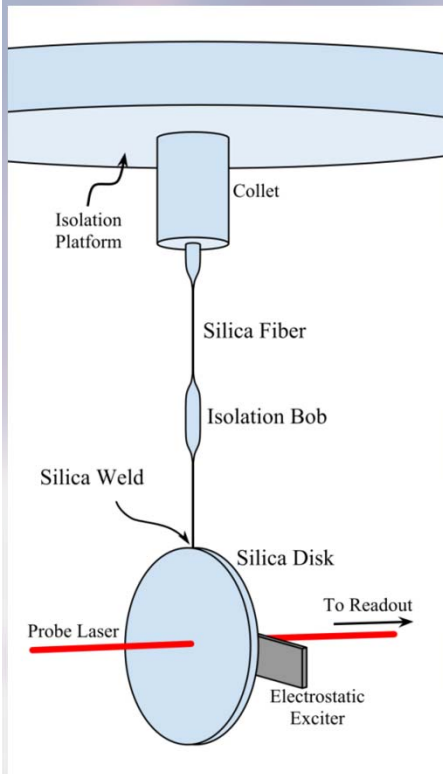
Overview of AU Lab

- The AU Thermal Noise lab has two bell jars dedicated to studying thermal noise in gravitational wave detector optics
 - Dedicated finite element modeling computer
- PI Gregory Harry, Lab Specialist Jonathan Newport
 - Typically two students, currently Hannah Fair and Sam Hickey



Q Measuring

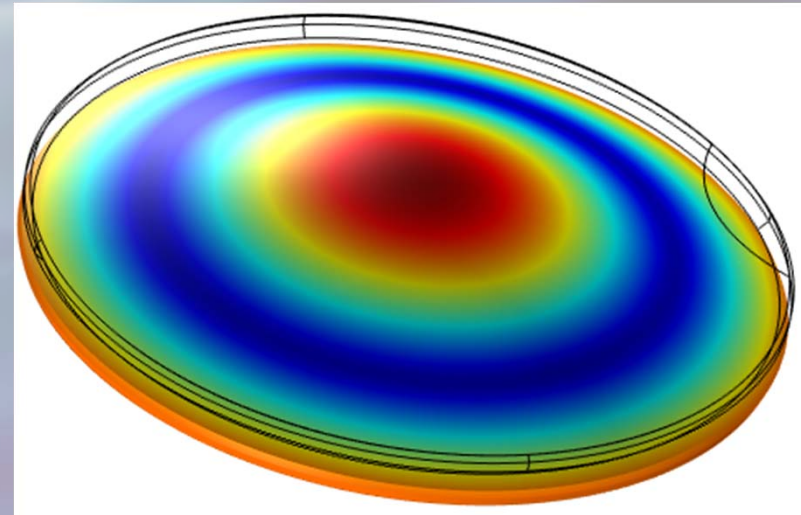
- Measure modal Q's of sample optics



- Determine mechanical loss from Q's
- In vacuum, silica fiber suspension, electrostatic excitation, birefringence laser readout

Finite Element Model

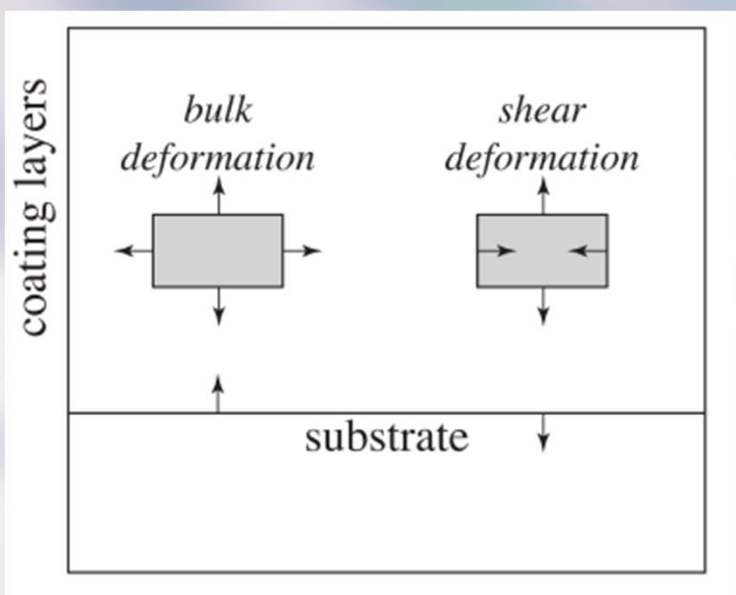
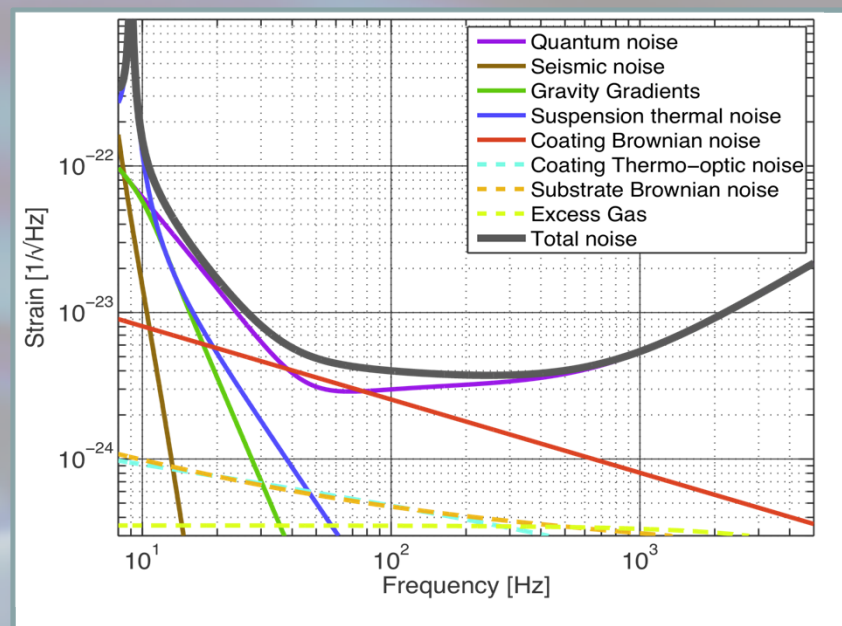
- Energy distribution in sample by FEA
- Determine modally (frequency), spatially, tensor elements





Coating Thermal Noise

- Coating thermal noise caused by mechanical loss
- Sensitivity limiting in aLIGO
 - Titania doping, et al.
- Continuing in 3rd generation
 - Many approaches



- Hong et al, Phys. Rev. D **87**, 082001
- Coating thermal noise depends on shear & bulk mechanical loss
 - Hong et al estimate 37% uncertainty in thermal noise from ignorance of these values in aLIGO



Ti-Ta Sample

- Nominal thickness 0.5 μm
 - Being measured at ERAU
- Using standard values for
 - Young's Modulus 140 GPa
 - Poisson ratio 0.23
 - Density 2200 kg/m^3



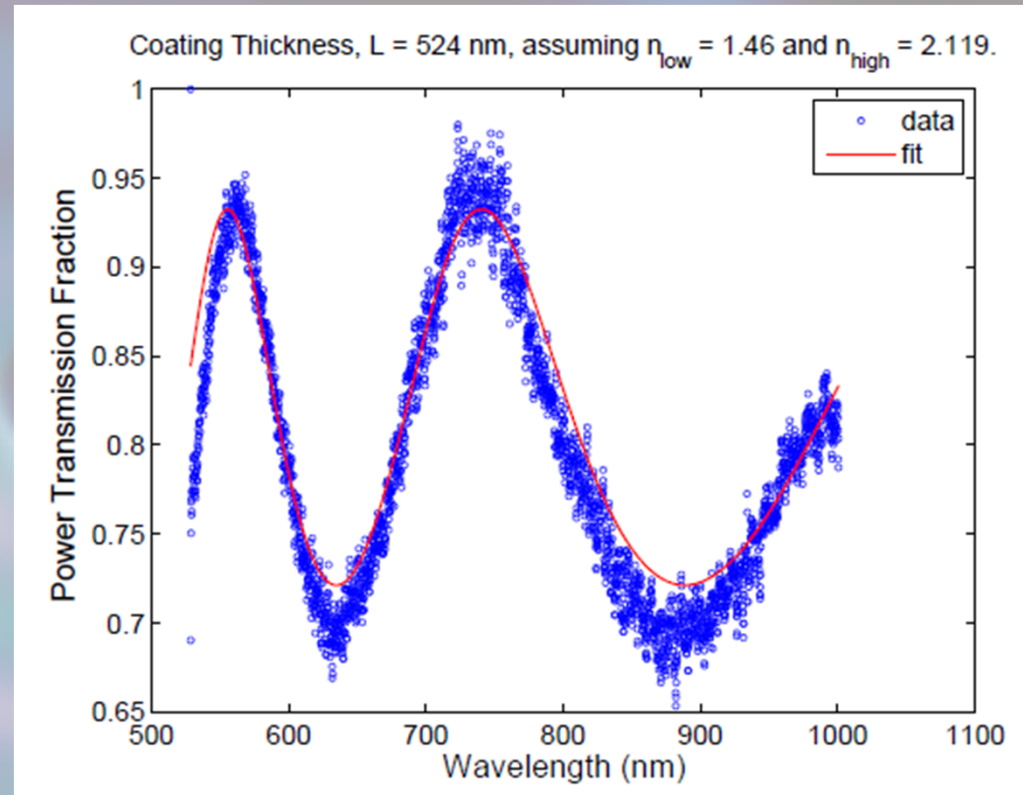
- Jonathan Newport has been doing FEA work

| Mode | Frequency | Q | δQ | $E_{\text{shear}}/E_{\text{tot}}$ | $E_{\text{bulk}}/E_{\text{Tot}}$ |
|------|-----------|--------------------|-------------------|-----------------------------------|----------------------------------|
| BF | 2773 Hz | 1.14×10^6 | 2.5×10^3 | 9.93×10^{-3} | 6.58×10^{-4} |
| DH | 4178 Hz | 8.70×10^5 | 2.6×10^4 | 6.85×10^{-3} | 4.79×10^{-3} |
| Hex | 6307 Hz | 9.32×10^5 | 3.2×10^4 | 9.37×10^{-3} | 9.65×10^{-4} |
| DDH | 9707 Hz | 9.16×10^5 | 3.8×10^3 | 7.68×10^{-3} | 3.62×10^{-3} |
| Oct | 10943 Hz | 8.87×10^5 | 1.6×10^4 | 9.05×10^{-3} | 1.18×10^{-3} |



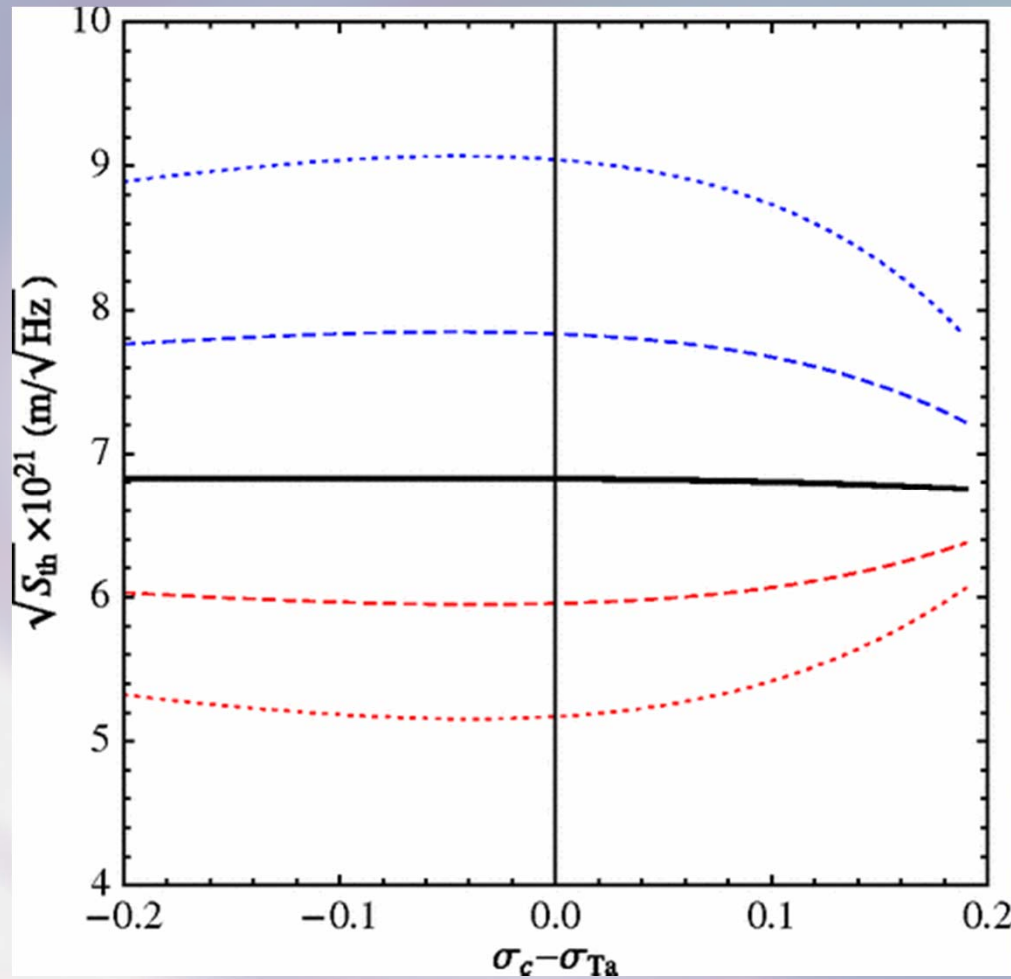
Coating Thickness

- Transmission vs angle at ERAU by Andri Gretarsson



- $n = 2.119$ for 25% ti in ta from CGQ 24, 405 2007.
- Single layer fit gives $d = 0.524 \mu\text{m}$

Poisson Ratio

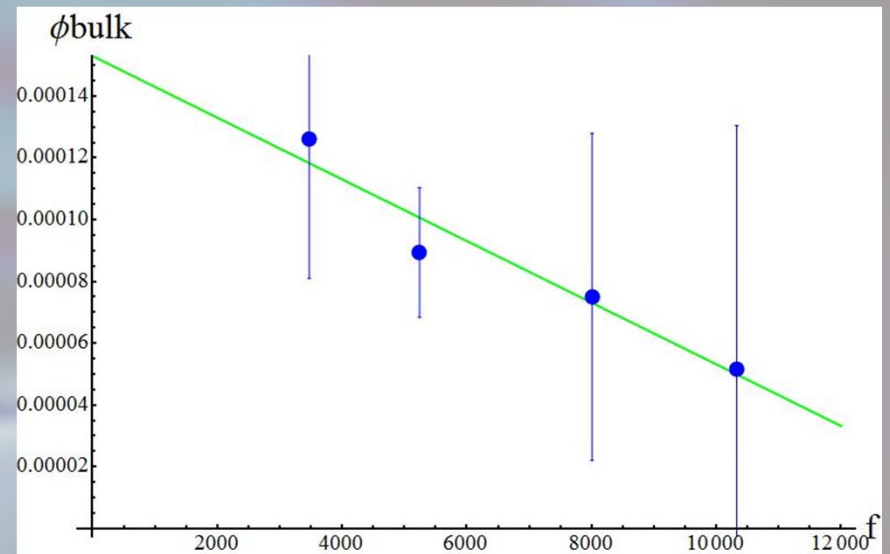
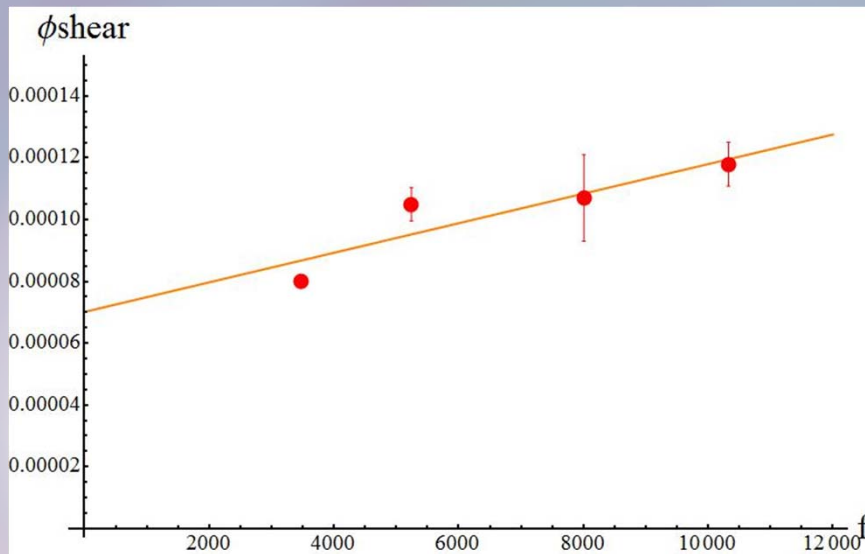


- Uncertainty in $\nu_i - \nu_a$ Poisson ratio significant contributor to uncertainty in energy ratios
- Hong et al estimate it accounts for $\sim 10\%$ uncertainty in thermal noise



Results

Bulk and shear ϕ as functions of frequency



$$\varphi_{shear} = 7.0 \times 10^{-5} + 4.8 \times 10^{-9} f$$

$$\varphi_{bulk} = 1.5 \times 10^{-4} - 1.0 \times 10^{-8} f$$

Random uncertainties from energy ratios,
systematic uncertainty from Poisson ratio



Thermal Noise

Noise prediction from
Hong paper

- With $\varphi_B(f)$

100 Hz $\varphi_B/\varphi_S = 2.2$

1000 Hz $\varphi_B/\varphi_S = 1.9$

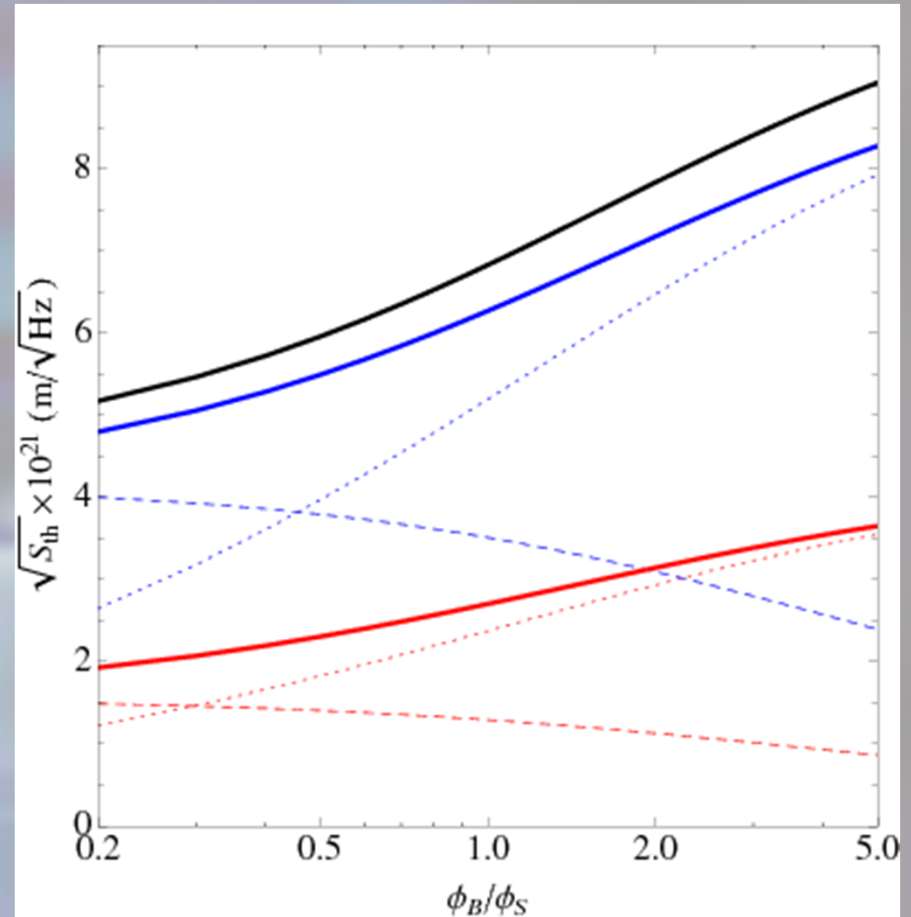
→20% worse aLIGO thermal noise

- With φ_B constant

100 Hz $\varphi_B/\varphi_S = 1.3$

1000 Hz $\varphi_B/\varphi_S = 1.3$

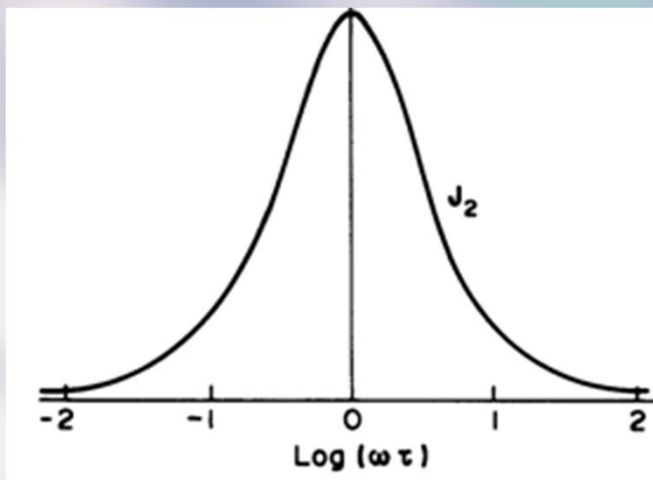
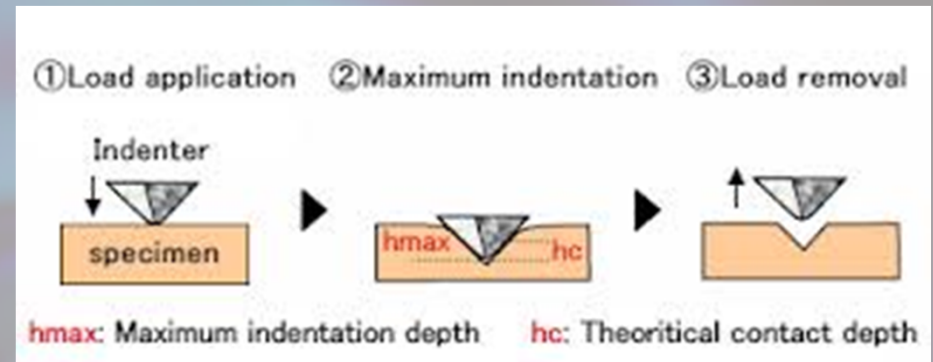
→2% worse aLIGO thermal noise



- Matt A found $\varphi_B/\varphi_S \approx 2$
in G1300063

Outstanding Issues

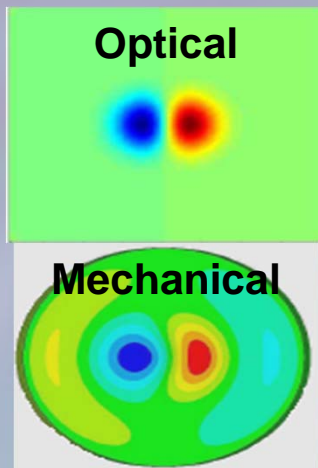
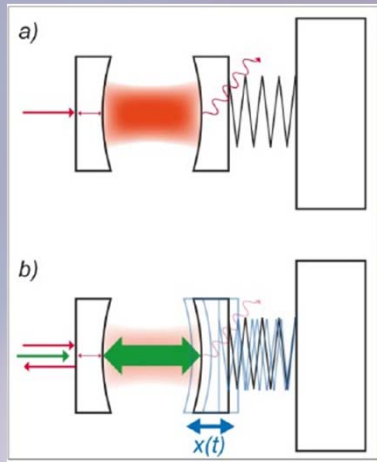
- Young's modulus and Poisson ratio
 - Nanoindenter
 - Acoustic reflection
 - Biggest source of uncertainty in aLIGO thermal noise



- More precise FEA model
 - Errors from energy ratios
- Linear loss model
 - Improve approximation
- Thermal noise predictions
 - Compare to prototypes



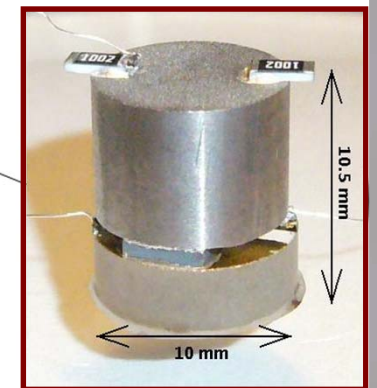
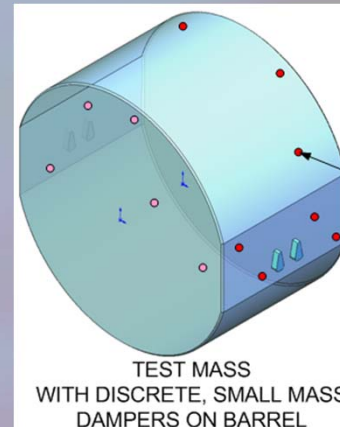
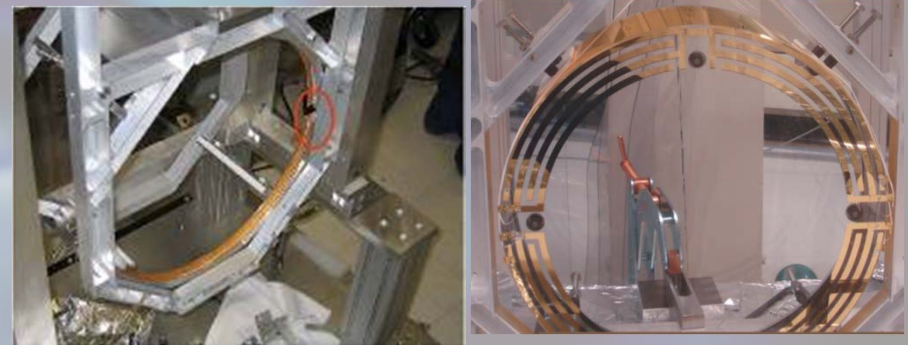
Parametric Instability



- Possibly problem in aLIGO from exchange of energy between optical cavities and mirror acoustic modes

Possible solutions

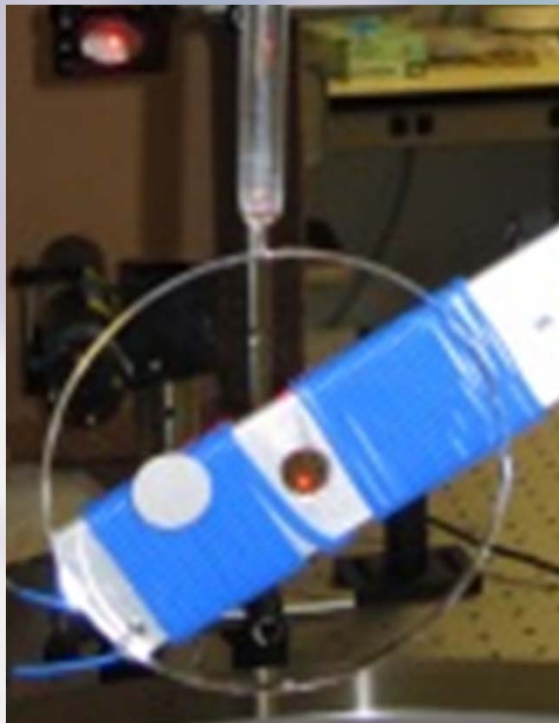
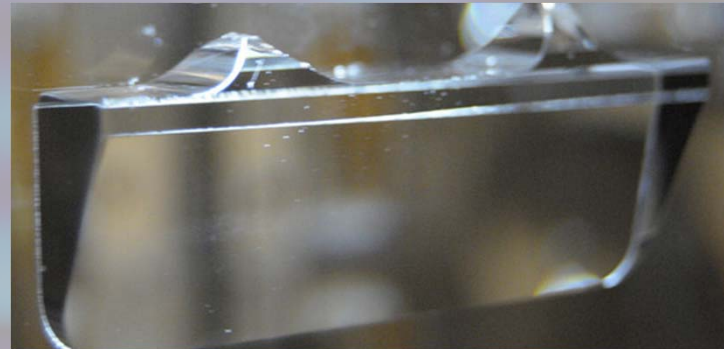
- Ring heaters to adjust mirror mode frequencies
- Electrostatic drive to actuate on and damp acoustic modes
- **Acoustic mass dampers to passively damp acoustic modes**
- Optical mode active damping





Epoxy Thermal Noise

- Retrofit rules out silicate bonding, best solution
- Use epoxy to connect dampers to mirrors



- Predict thermal noise for dampers
- Need mechanical loss of epoxies

– All bad, some more than others

- Measure Q's, do FE analysis
- See H. Fair poster for details

Mechanical Loss of Vacuum Compatible Epoxies for Tuned Mass Dampers
 Hannah Fan, Gregory Harris, Jonathan Newport, Sam Healey from American University, Sławek Grzes, Peter Fritzsche from MIT, Bill Keith from Caltech, LIGO-G540038

Parametric Instability
 The higher optical power in the Fabry Perot cavities of Advanced LIGO inevitably introduces the issue of parametric instability, the exchange of energy between the optical modes of the cavity and the acoustic modes of the mirrors. Energy transfer from the optical cavities to the mirrors could ring up the mirror's elastic modes, increase the ring lifetime, and potentially lead to control problems and lock loss.

Interpretation into Mechanical Loss
 After obtaining the Qs of the silica disk with epoxy on them, we use the FEA program COMSOL to obtain our mechanical loss plots. The program takes into account the properties of the silica disk and wafer, the Young's modulus of the epoxy, the Qs of the different frequencies, and the bulk and shear properties to determine the plot. This is an ongoing program, as we still need to find the thickness, Young's modulus and related parameters to some of the epoxies.

Technique
 We find the mechanical Q of different silica disks with different epoxies between the disks and a thin top layer of silica. The position of the epoxies on the silica disk is determined via Finite Element Modeling (FEA).

Epoxy Solution
 One possible solution is to damp the test mass modes with tuned mass dampers. There is a concern that damping the system would increase mechanical loss and result in higher thermal noise. However, the test mass modes are above 50kHz, while thermal noise is a concern around 100Hz. Parametric instability can be mitigated if a damping device that damps higher frequencies more than lower frequencies can be created. Because these dampers must be retrofitted onto the AIGO mirrors, silicate bonding is not possible. At American University, we are measuring the mechanical loss of different epoxies in order to find a suitable candidate for attaching these dampers to the test masses.

Epoxies in Progress

| Name | Frequency (Hz) | Quality Factor (Q) |
|--------------------------|---------------------|---------------------|
| Hysol EA 9313 (centered) | 2892 | 4.5x10 ⁵ |
| 6114 Hz | 1.2x10 ⁶ | |
| 8115 Hz | 1.1x10 ⁶ | |
| 9370 Hz | 1.2x10 ⁶ | |
| PP230 (off center) | 2083 | 35,000 |
| 4077 Hz | 300 | |
| 9987 Hz | 14,000 | |

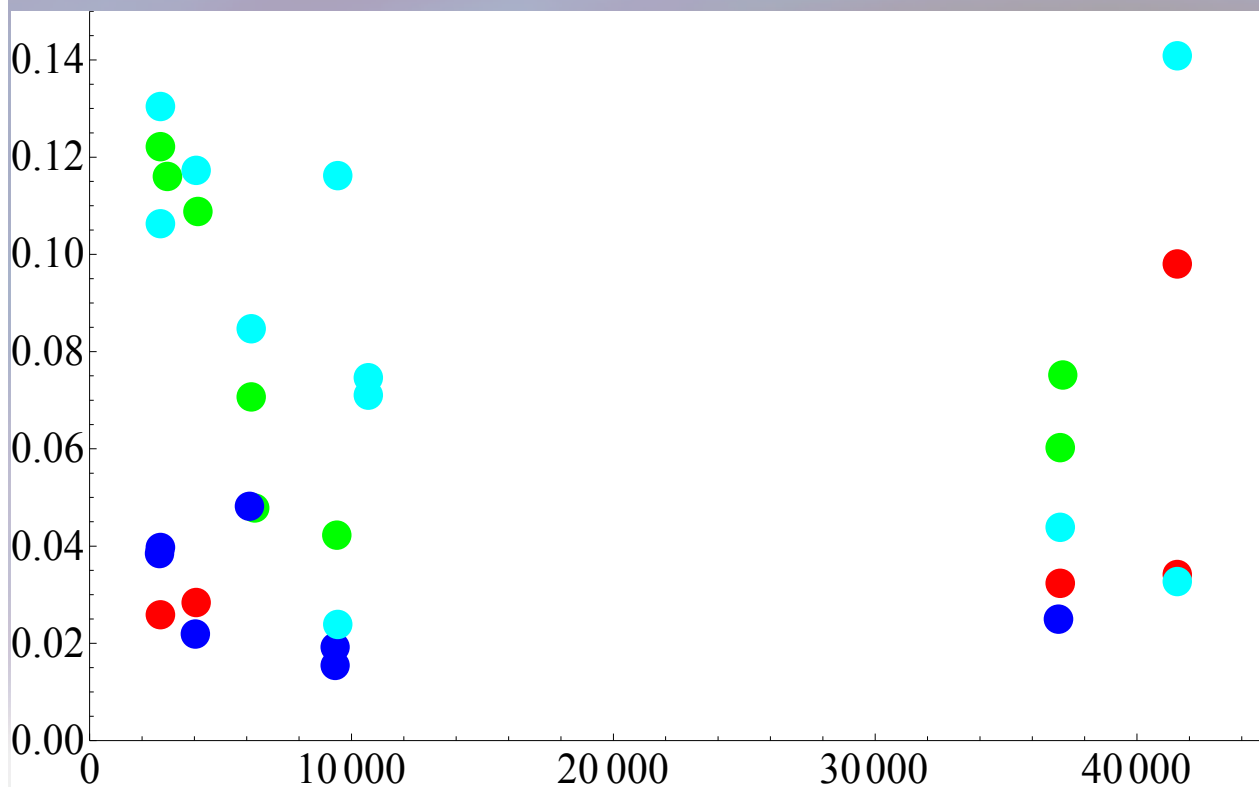
Epoxies analyzed with FEA

| Name | Thickness (µm) | Young's Modulus (GPa) | Viscosity (Pa·s) | φ |
|---------------------|----------------|-----------------------|------------------|---------------|
| Hysol Tri-Dent 2902 | 47 | 4.8 | 20Pa | 0.05 φ < 0.1 |
| EM Optocast 3553V | 10 | 3.4 | 5 Pa | 0.01 φ < 0.5 |
| Masterbond EP30 | 10 | 2.9 | 02 Pa | 0.02 φ < 0.54 |
| Epotek 353ND | 10 | 3.7 | 12.5 Pa | φ < 0.3 |

Future plans with Epotek 353ND
 We are adding carbon to Epotek in an attempt to make it conductive, and then re-test the new mixture for mechanical loss.



Epoxy Results



Green – Tra-Duct 2902, conducting

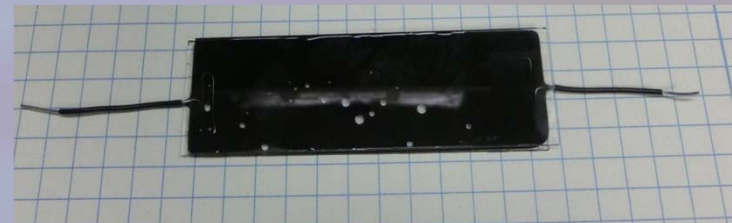
Red – Epotek 353ND, Nergis' lab

Blue – Optocast 3553LV, aLIGO OMC

Cyan – EP 30, aLIGO standoffs

Also Hysol EA9313, EP1730, superglue

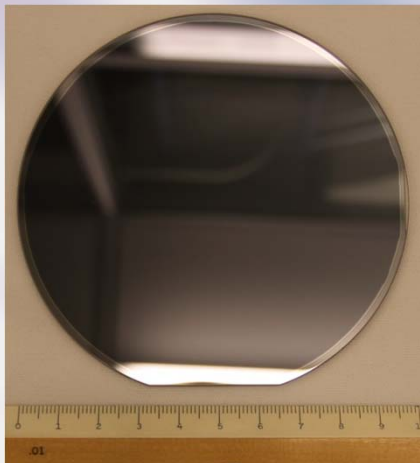
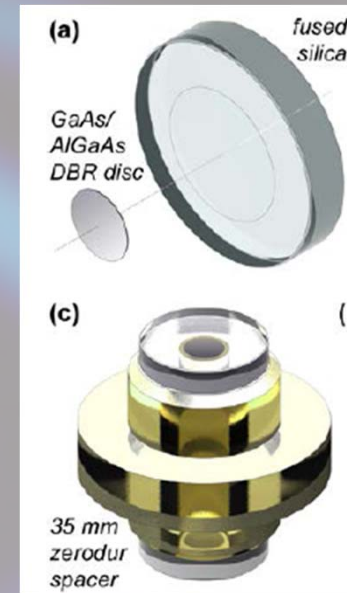
- Epotek 353ND now aLIGO vacuum approved
- Adding carbon to Epotek 353ND
 - Make conductive
 - First attempt $> 1 \text{ M}\Omega \cdot \text{m}$





Aluminum Gallium Arsenide (AlGaAs)

- Mechanical loss/thermal noise results
 - Low TN in quantum experiments
 $\phi \approx 2.5 - 4 \times 10^{-5}$
 - Two Q results on silica substrates
 $\phi \approx 1 - 2 \times 10^{-4}$



- Crystalline Mirror Solutions recently able to make larger diameters, up to 10 cm
- Improvement in bond strength
- Flaw in AU/HWS sample due to scratch during transport/handling