

Time Delay Of Gravitational Waves Propagating Through a Galaxy Toward a Non-Stationary Observer and Other Problems

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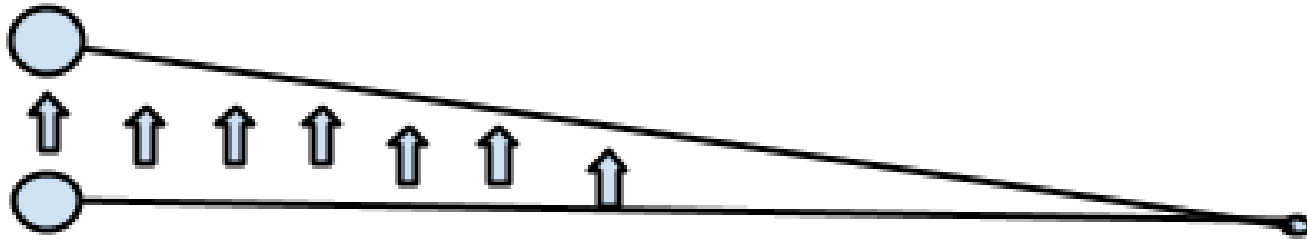
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Original Plans and Motivation

- “Amplitude and phase fluctuations for gravitational waves propagating through inhomogeneous mass distribution in the universe” – Takahashi [1]
- “Hunting for topological dark matter with atomic clocks” – Derevianko and Pospelov [2]
- Simulation papers on domain walls as dark matter candidates – effectively excluded [3,4,5]
- “Domain Wall Dominated Universes” – Battyre, Bucher, and Spergel [6]

Time Delay Calculations

• $\Phi = \int \frac{M(r)}{r^2} dr$ $g = \Delta\Phi$ $t_{delay} = \int v \cdot g dl$



Results

- Simple model of our galaxy: $1/r$ mass distribution out to $5 * 10^{20} m$, normalized to $6 * 10^{42} kg$ total, Earth positioned $2.6 * 10^{20} m$ from origin
- Maximum $\dot{t}_{delay} \sim 9.2 * 10^{-9} s/s \sim 0.29 s/yr$ for the case of a source at infinity
- Maximum $\dot{t}_{delay} \sim 9.0 * 10^{-9} s/s \sim 0.28 s/yr$ for the case of a source at a distance of 5 Mpc
- Maximum $\dot{t}_{delay} \sim 9.1 * 10^{-9} s/s \sim 0.28 s/yr$ for the case of a source at a distance of 100 Mpc

Future work

- Find and use a more accurate mass distribution function for the galaxy
- Check on how significantly the slight discontinuity at the cutoff distance is effecting the results
- Other related problems

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References

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