



**Institute of
Applied Physics**

Friedrich-Schiller-Universität Jena

Coating-reduced interferometer optics

Resonant waveguide gratings

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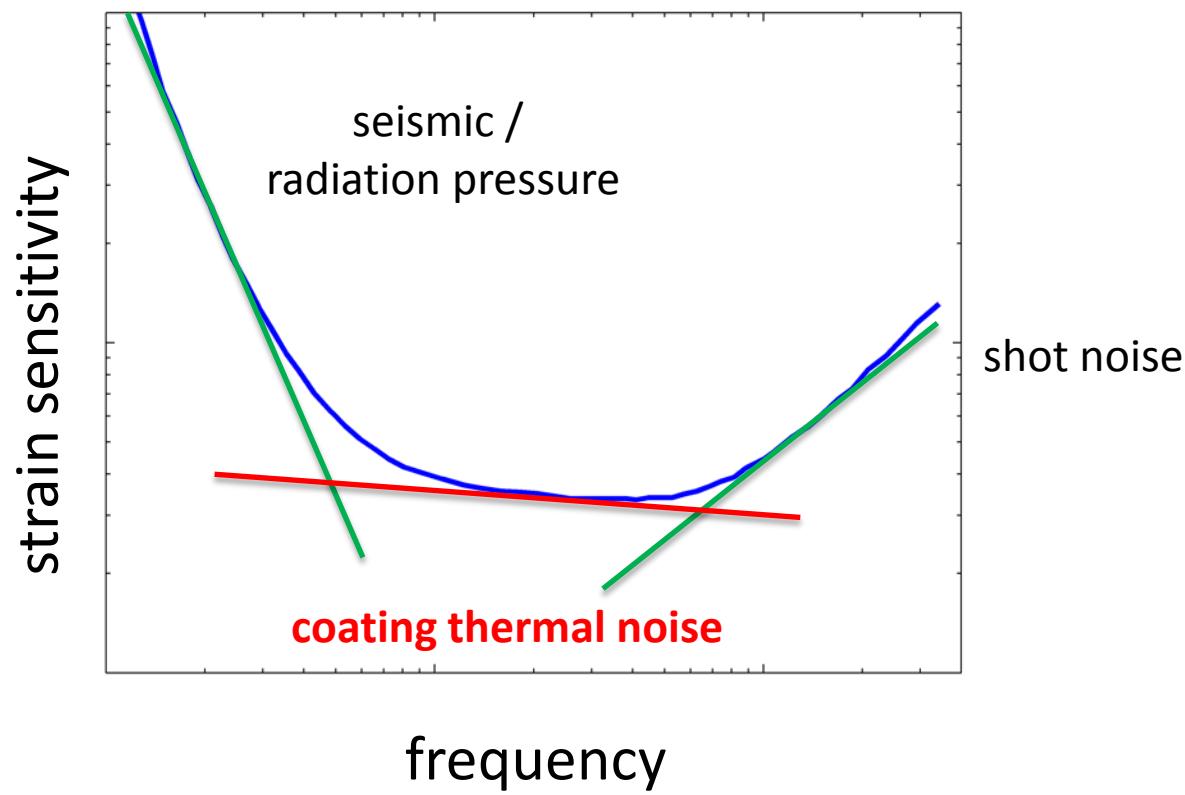


Outline

- **RWGs as mirrors**
- **Advanced grating concepts**
- **Fabrication**
- **Optical losses**

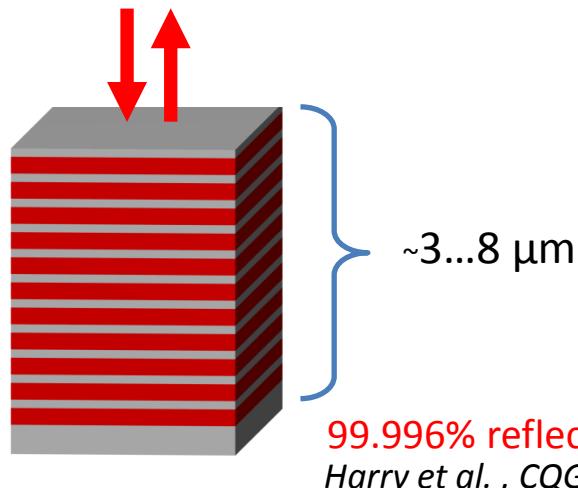
- **Summary & Outlook**

Introduction – coating thermal noise



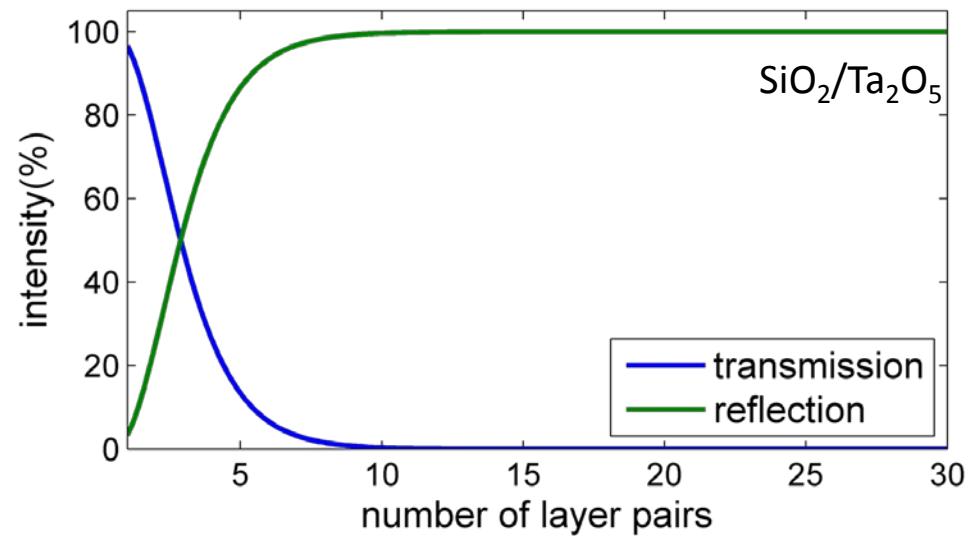
Introduction – reducing coating thermal noise

conventional: dielectric **crystalline**
multilayer stack



99.996% reflectivity
Harry et al., CQG (2002)

multiple beam interference:

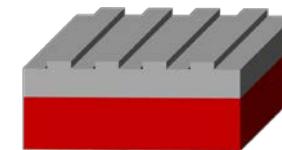


reflectivity  Brownian coating thermal noise

1. Possibility- optimization materials: crystalline coatings (Al_xGaAs)
See talk by G. Cole!

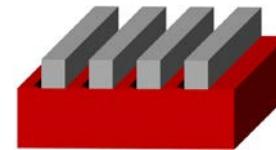
2. Possibility- alternative optical concepts + material optimization: Resonant waveguide gratings (RWGs)

Origin: weakly modulated high-index grating
on low-index substrate (narrow band)

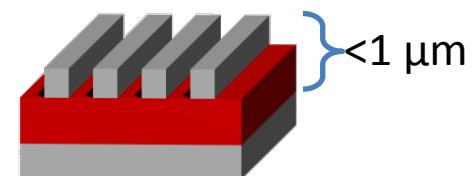


Popov et al., Opt. Comm. (1985)

Stronger modulation: larger bandwidth



Replace low-index substrate



Replace low-index material



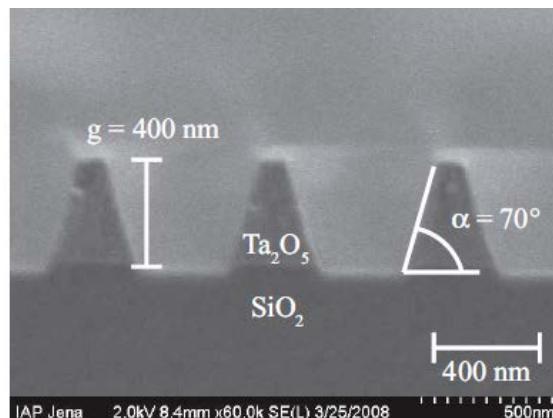
high-index

low-index

Brückner et al., PRL (2010)

RWGs as mirrors – so far

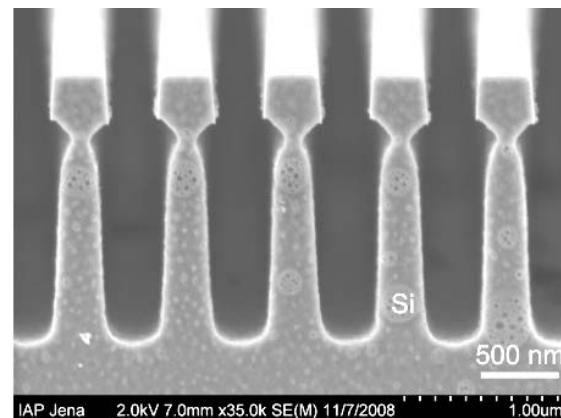
Tantala grating on silica for 1064 nm



Brückner et al., Opt. Express (2008)

$$R(1064 \text{ nm}) = (99.08 \pm 0.05)\%$$

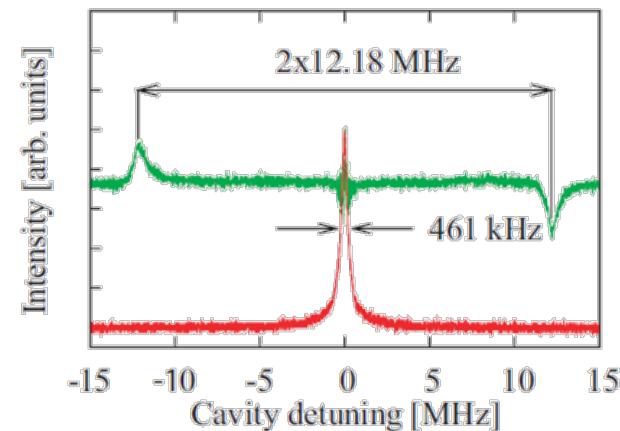
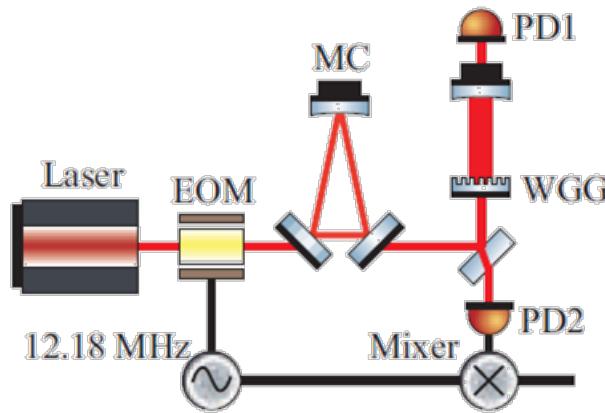
Monolithic silicon gratings for 1550 nm



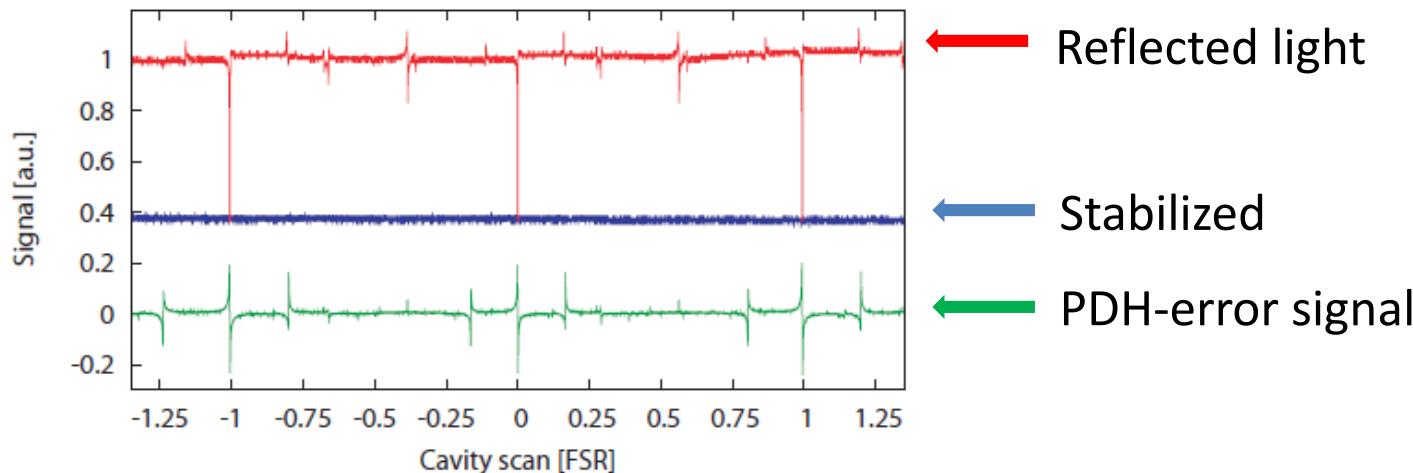
Brückner et al., PRL (2010)

$$R(1550 \text{ nm}) = (99.77 \pm 0.01)\%$$

reflectivity
measurements
in cavity

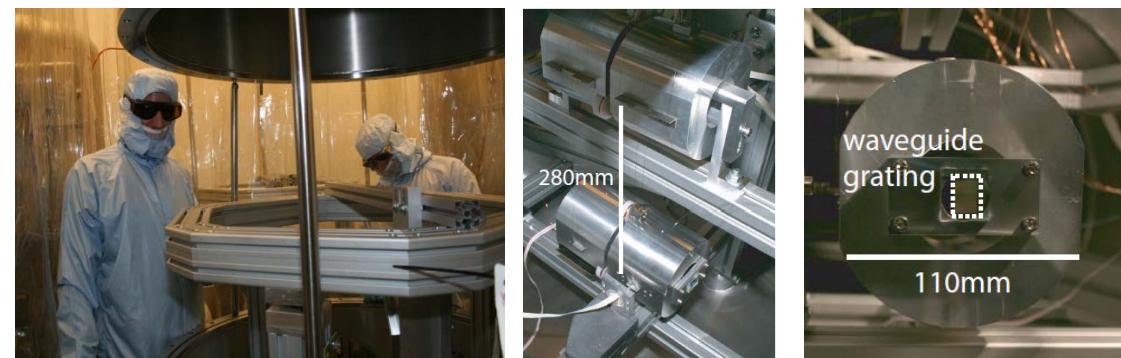


Tantala RWG



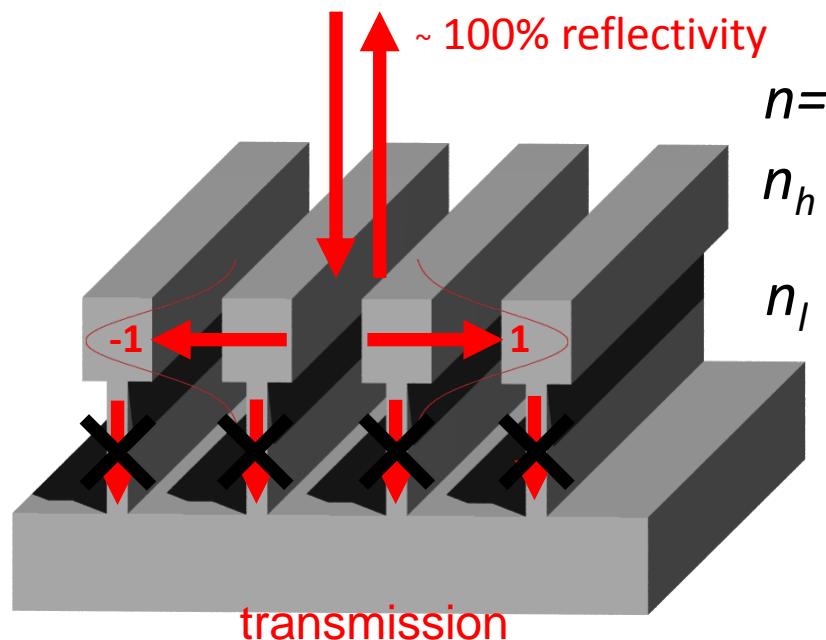
- Finesse = 790 (± 100) $\rightarrow R \geq 99.2(\pm 0.1)\%$
- Cavity stabilization with standard PDH-technique

D.Friedrich et al. Opt. Express 19, 14955 (2011)





Functionality of RWGS – horizontal modes



silicon as high-index material
($n_h=3.48$ @ $\lambda=1550$ nm)

1st diffraction orders
in grating

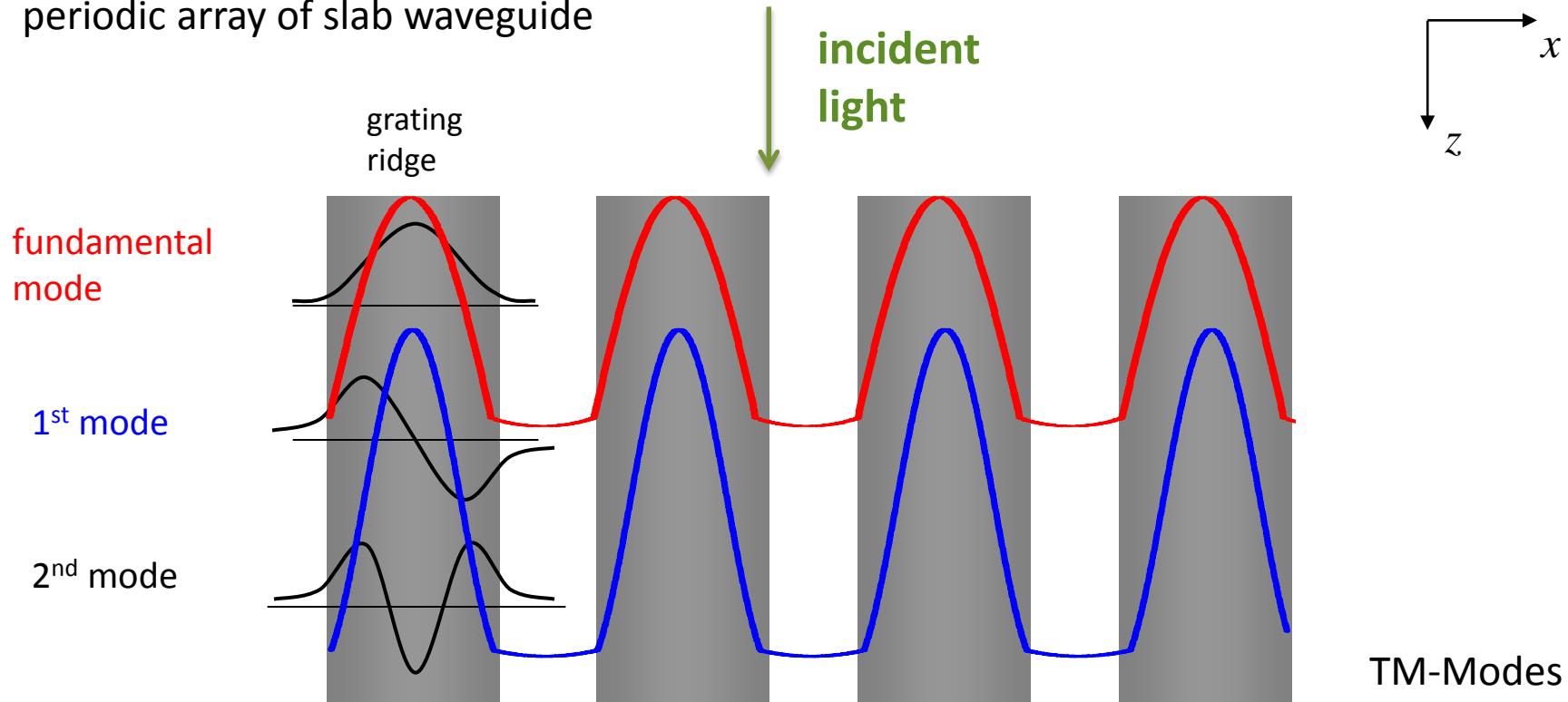
$$\lambda/n_h < p < \lambda/n_l$$

p... grating period

0th diffraction order
outside grating

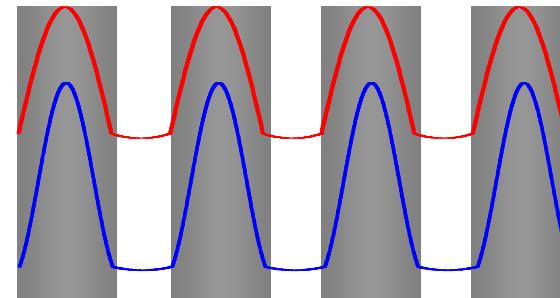
Functionality of RWGS – vertical (Bloch-) modes

periodic array of slab waveguide

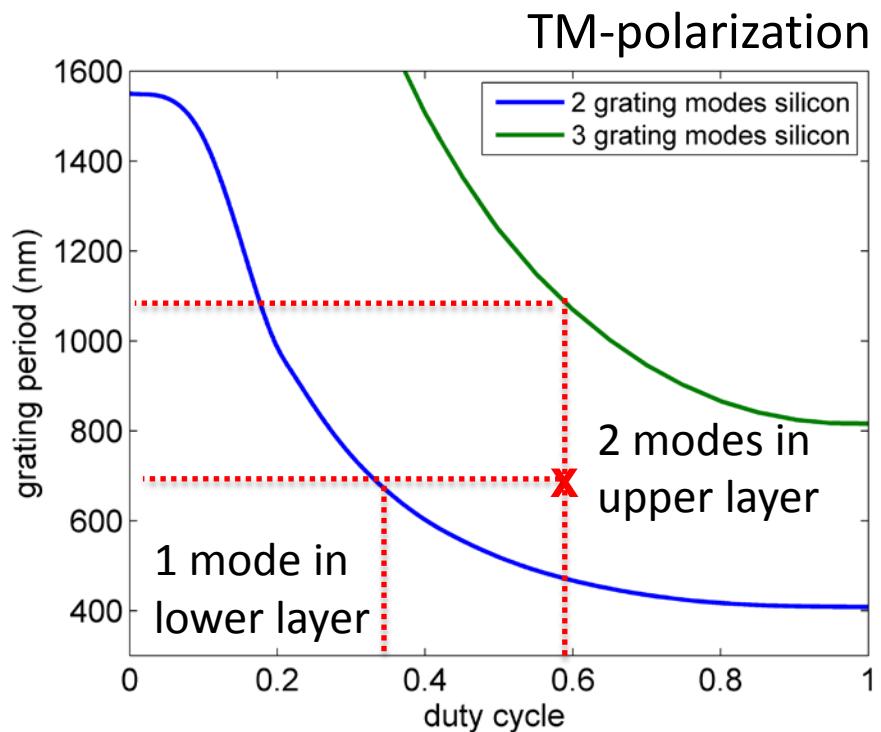
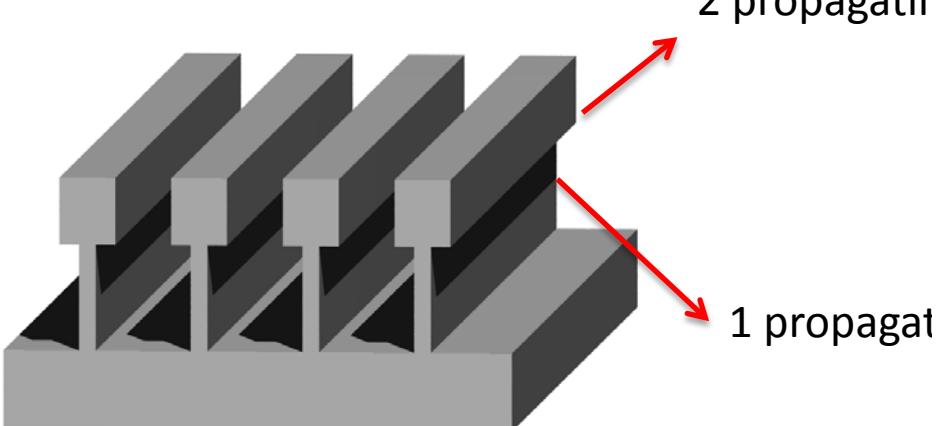


Effective index dependent on:

- fill factor
- period/ wavelength
- polarisation
- angle of incidence

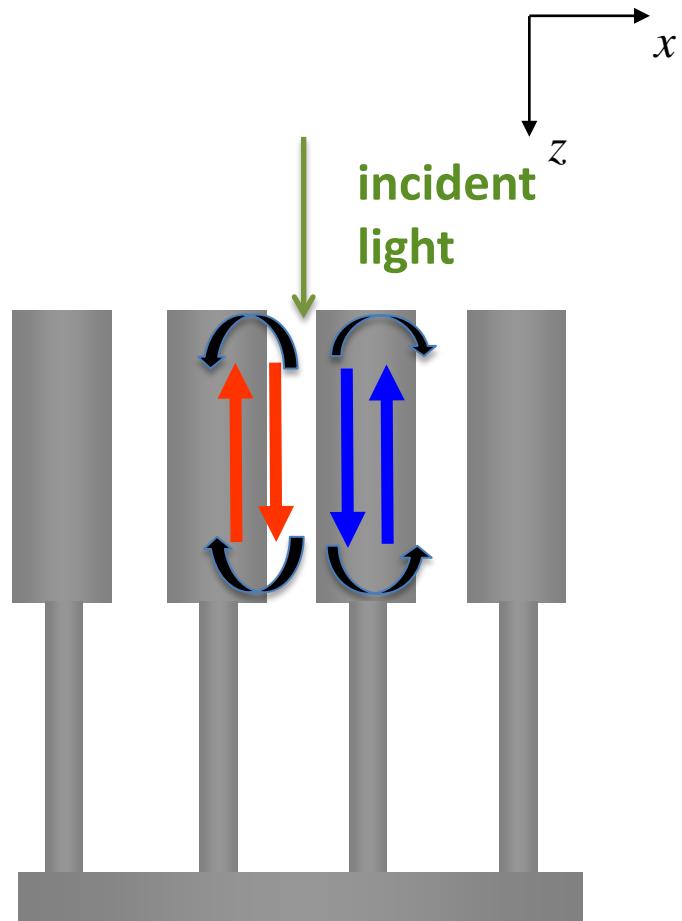


High efficiency:



mechanism for R=100%:

- phase difference due to different effective indices (different k_z),
- reflection/ interference at boundaries
- no additional phase in lower grating
- Amplitudes of grating modes need to cancel out



exact intensities and phases + physical insight

Lalanne , J. Lightwave Tech. (2006)

Karagodsky et al., Opt. Express (2009)



... interaction of light with surface structure...



S. Vyatchanin



S. Hild



Friedrich-Schiller-Universität Jena

D. Heinert

R. Nawrodt

S. Kroker



東京大学
THE UNIVERSITY OF TOKYO

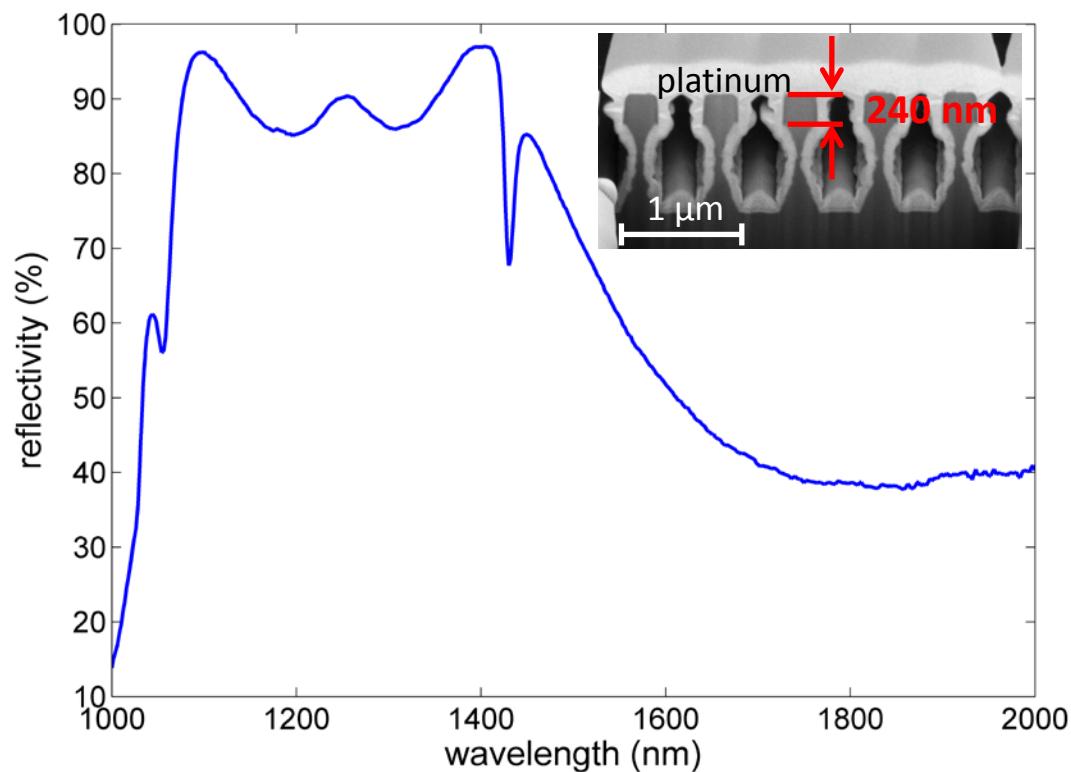
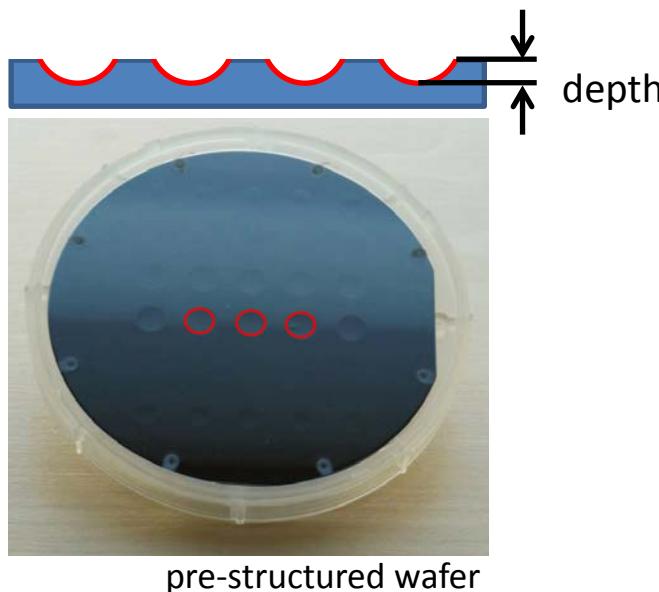
D. Friedrich

K. Yamamoto

Feel free to join!



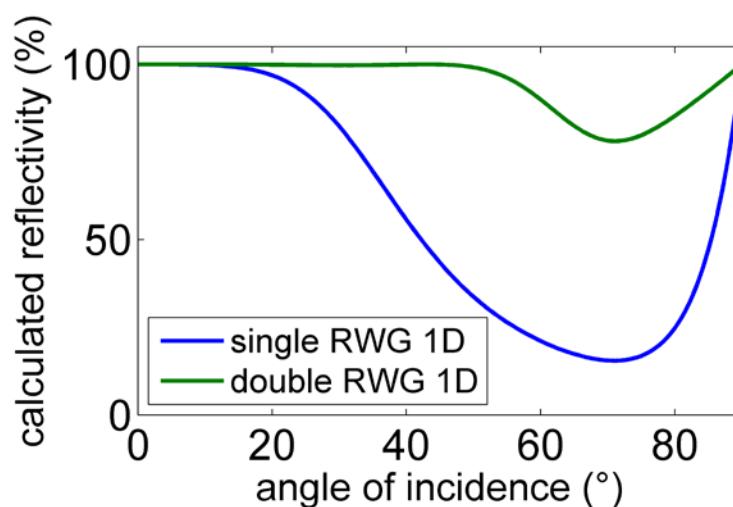
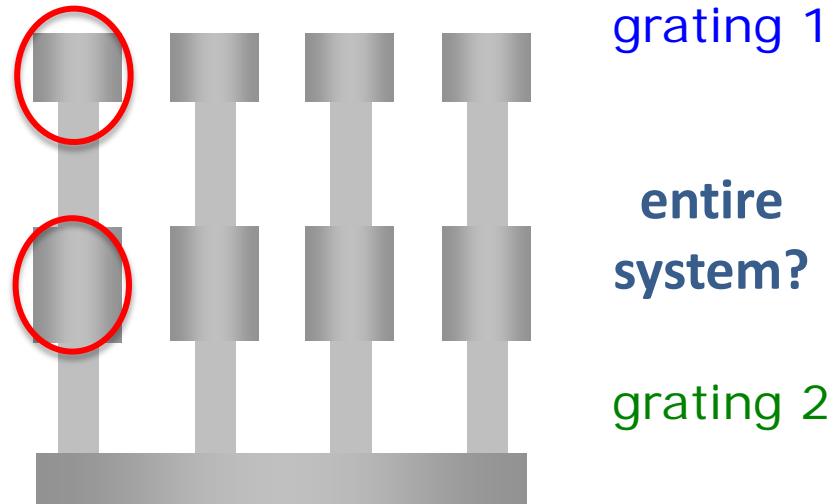
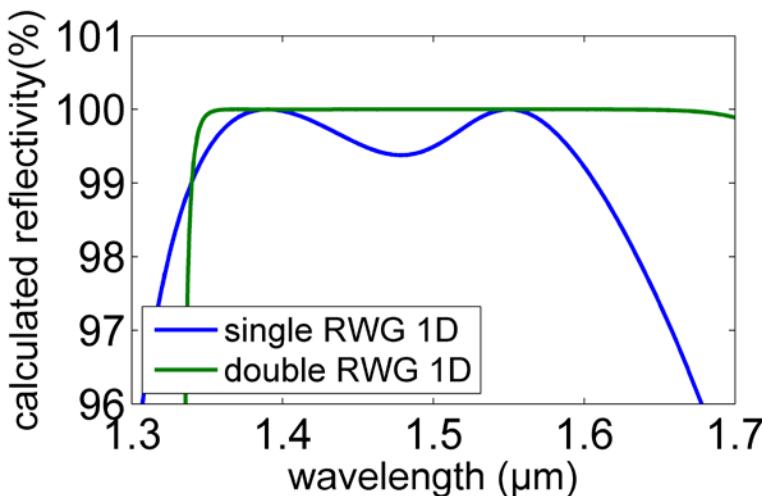
Mirrors on Curved Substrates



	grating	LIGO
ROC (m)	0.6	~2000
r_{Mirror} (m)	0.0035	0.17
depth (μm)	~10	~7

maximum reflectivity : 97% ($\lambda=1396 \text{ nm}$, $\theta=5^\circ$)
resonance to larger wavelength with larger period/thickness...

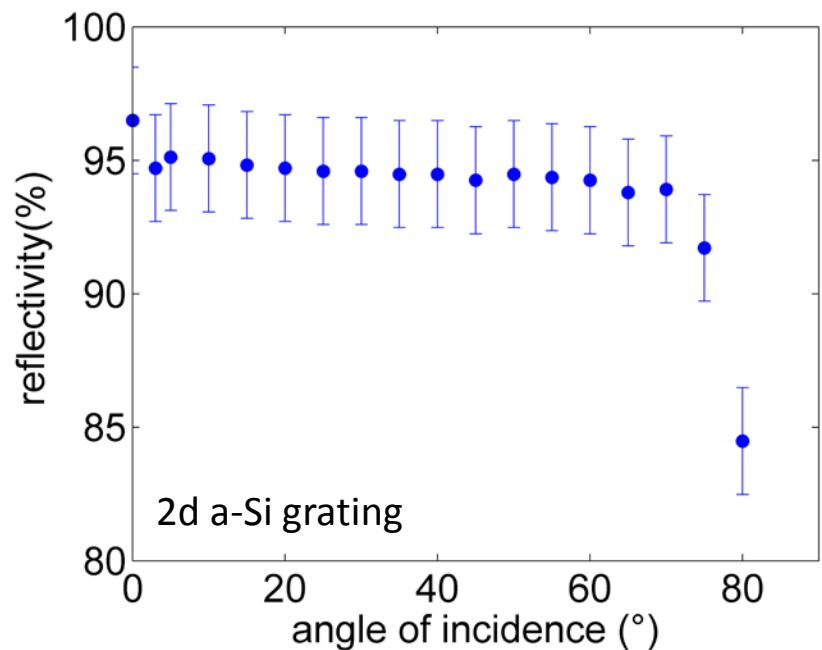
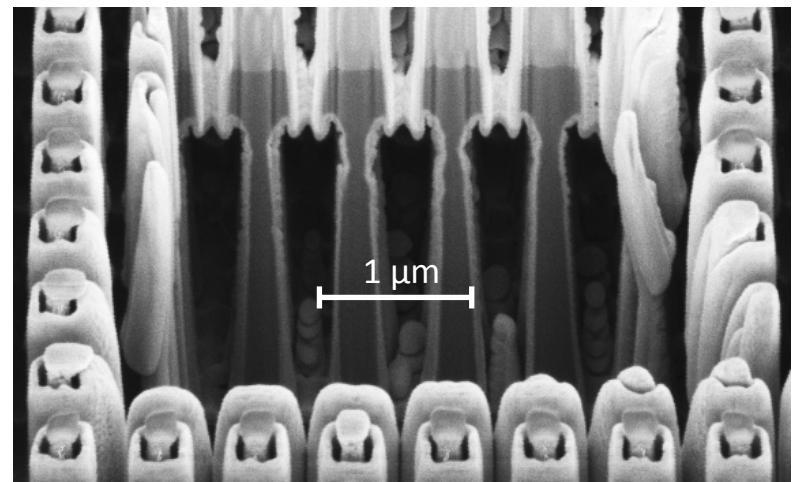
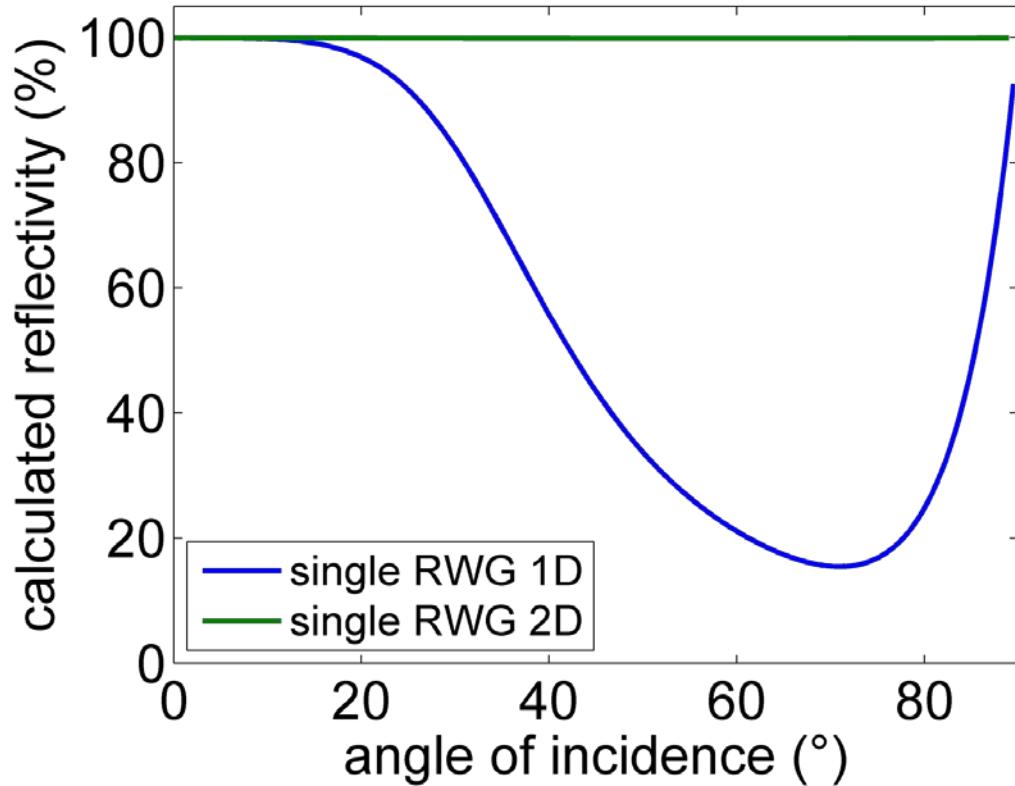
1. stacked RWGs



Thickness determines wavelength/ angle of max. reflection

- enhance angular/ spectral bandwidths
- enhance experimentally feasible reflectivity

2. RWGs with 2D periodicity

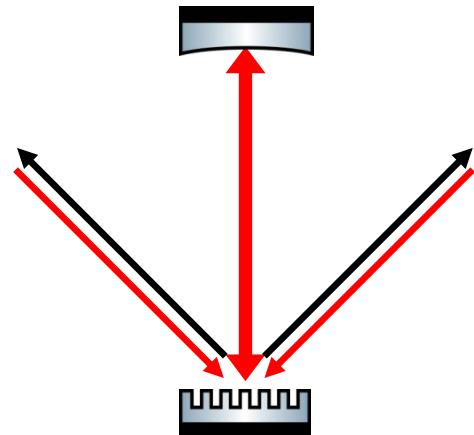


basis for flat focusing mirrors

Combination von highly efficient mirrors with diffraction grating

Highly efficient mirror

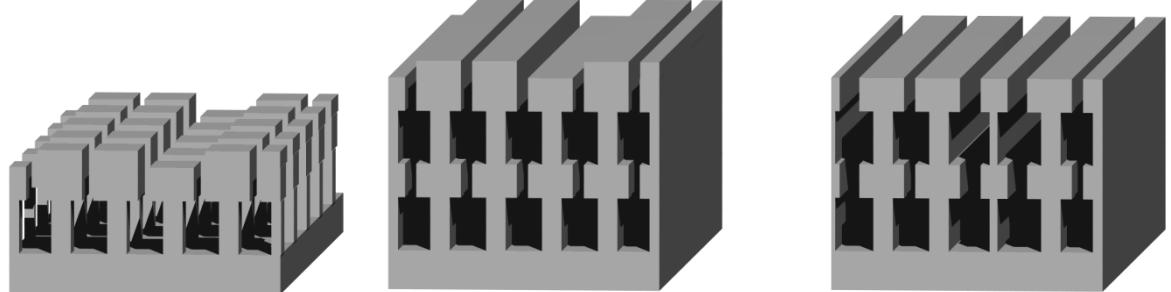
$p_{mirror} < \lambda$
only 0th order



high Finesse
→ only weak perturbation

Diffraction grating

$p_{diff} = \lambda \dots 2\lambda$
0th, ±1st orders

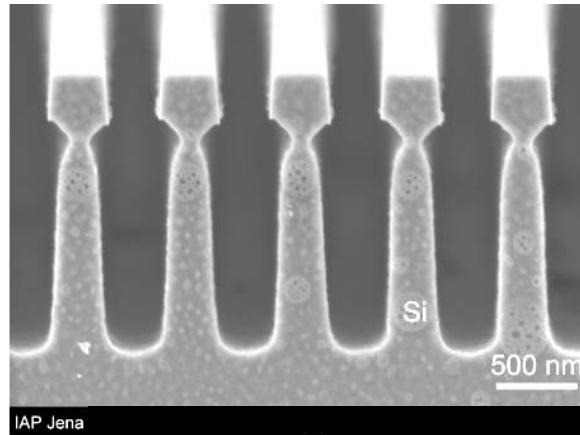
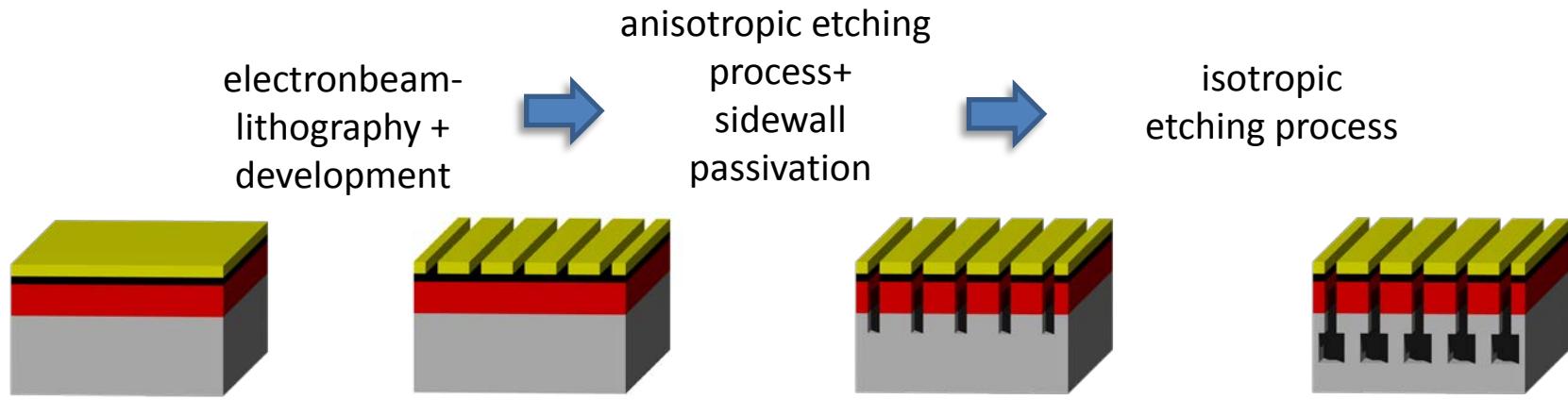


transversal
modulation:
Grating depth

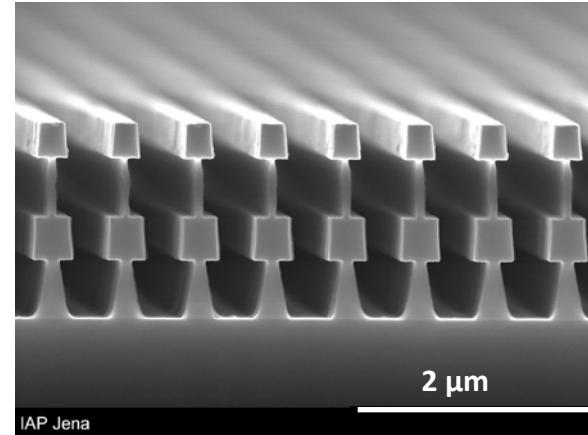
lateral
Modulation:
Ridge width/
-position

High angular tolerance of reflector necessary

Fabrication



More complex
structures?



silicon



SiO_2



chromium



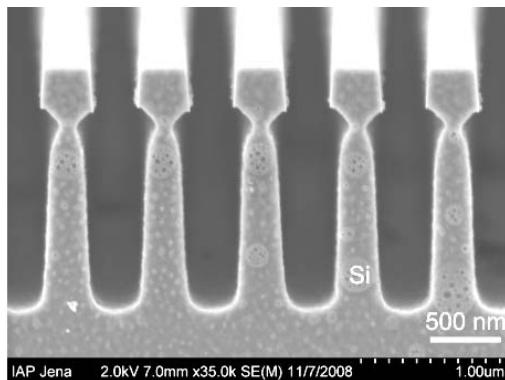
electron-
beamresist



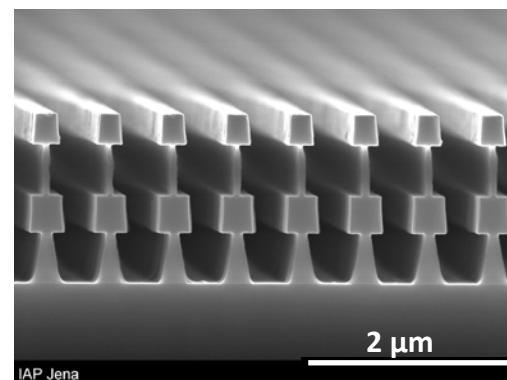
Where is the missing light?

$$R = (99.8 \pm 0.01)\%$$

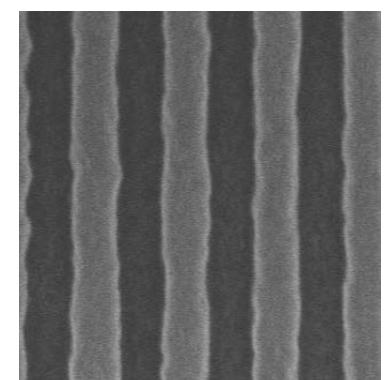
absorption
material properties
See talk by J. Komma!



transmission
structure deviations
systematic
(ridgewidth/-depth)

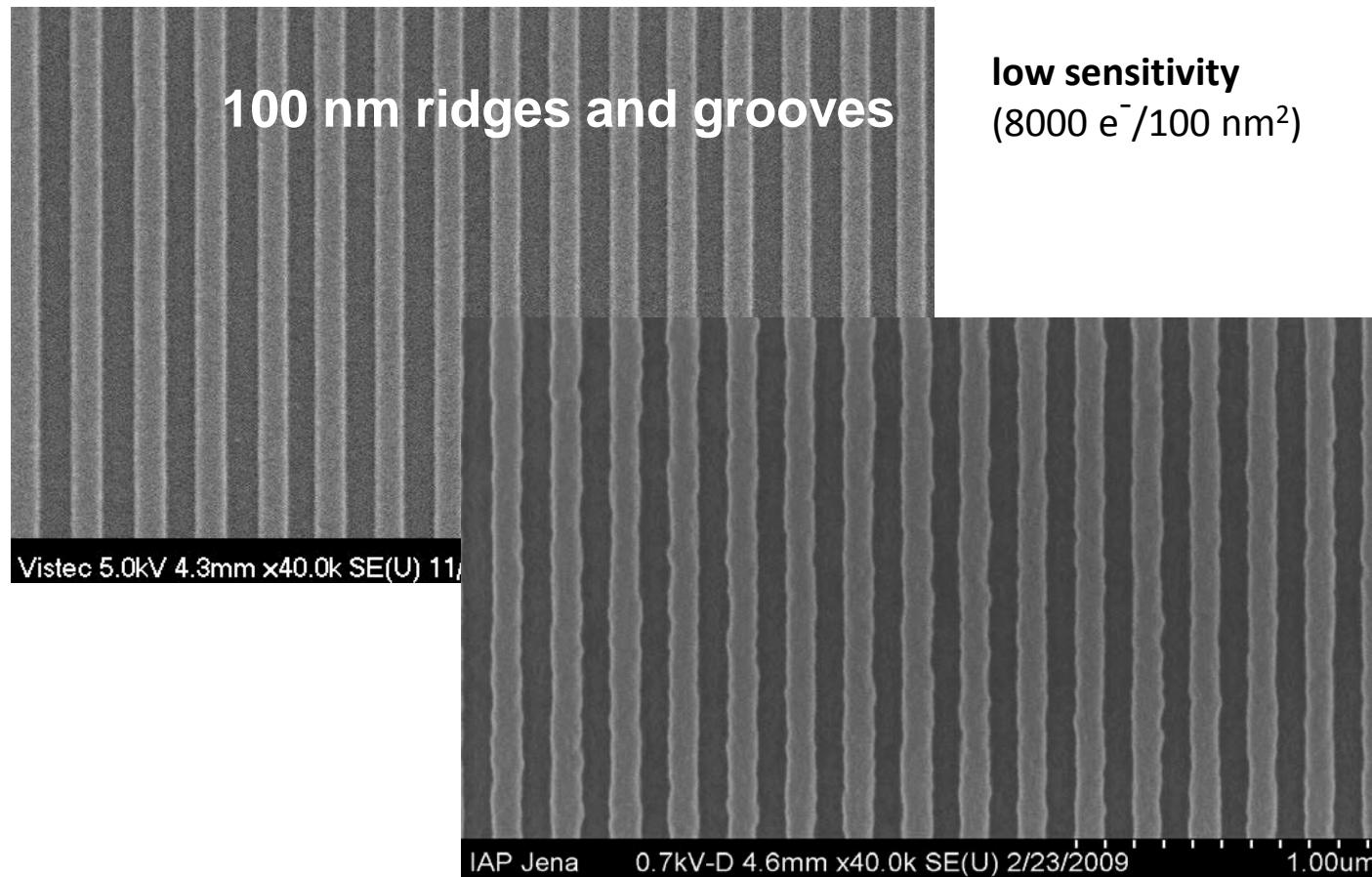


scattered light
stochastic
(line edge roughness)

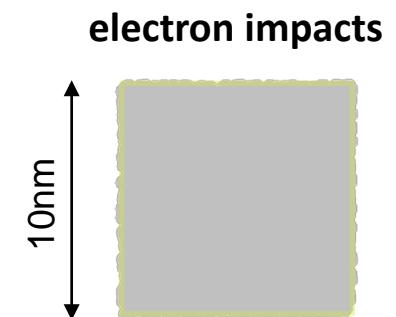




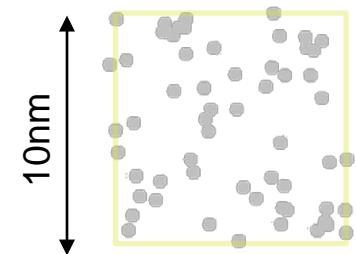
Line edge roughness due to particle statistics in the lithographic process



low sensitivity
($8000 \text{ e}^-/\text{100 nm}^2$)

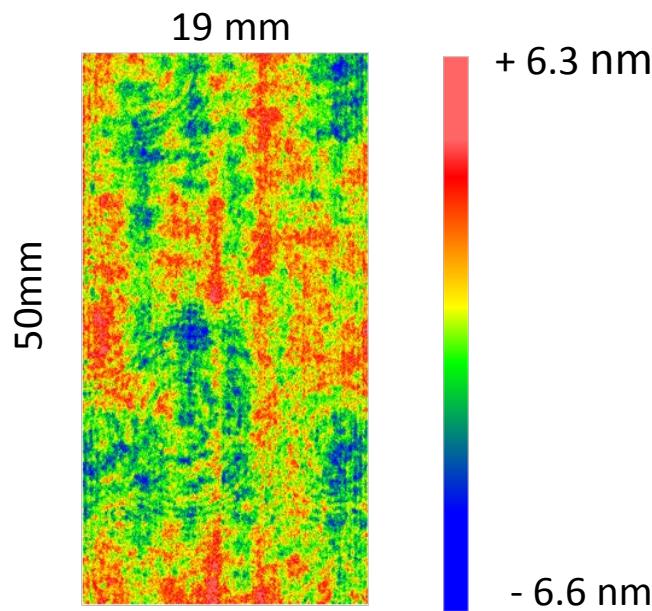


high sensitivity
($60 \text{ e}^-/\text{100 nm}^2$)



Grating accuracy

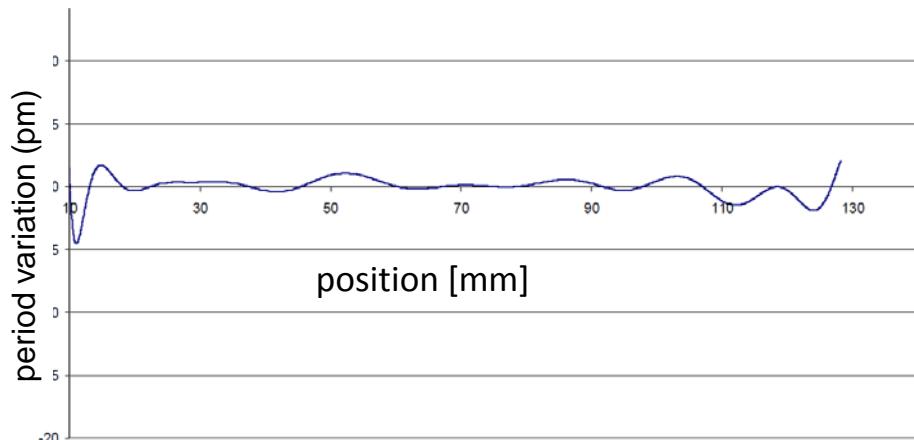
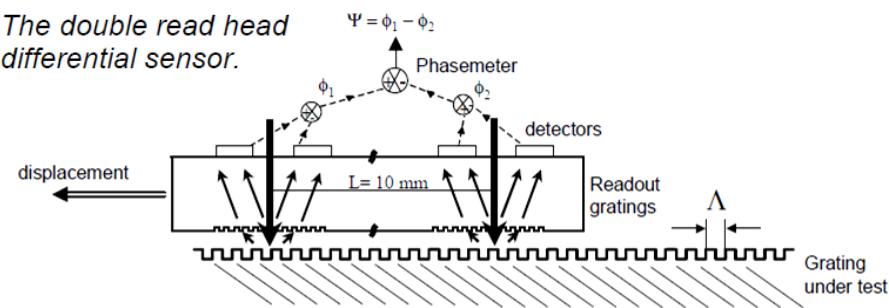
wave-front measurement (1 μm period grating +
technology, Littrow-Mount, $\lambda=633 \text{ nm}$)



	wavefront	placement
PV	12.8 nm	< 10.3 nm
rms	1.4 nm	< 1.1 nm

significantly better than interferometric gratings

Report on the grating writing analysis Laboratoire Hubert Curien CNRS - Fraunhofer IOF



Sophisticated device for measurement of scattered light ALBATROSS
at Fraunhofer Institute IOF



Measurements of scattered light at 1550 nm set up right now!

Summary

- RWGs capable of providing high reflectivity
- Experiments for higher R running
- Origin of optical losses to be identified
- Thermal noise to be investigated