# AGENDA FOR THE LIGO INTEGRATION MEETING

9 February 1995

<b>Facilities</b>
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0900 - 0930

0930 - 1030

• Ground motion measurements at the sites

Foundation motions and alignment

Lisa Sievers

Mike Gamble

### **Beam Tube**

• QT Status & Baffle design issues

• Tube motion analysis

Beam tube scattering measurements & baffle material options

• Synopsis of the Baffle Review Meeting

Charts only

ho oral presentation Larry Jones

Mike Gamble

Albert Lazzarini

Rai Weiss

**BREAK** 

Vacuum Equipment

• Deferral of getter pump procurement for initial interferometer

Procurement status & update

1030 -1045

Mike Zucker

John Worden

### Detector

· Length control modeling

1130 - 1245

Lisa Sievers

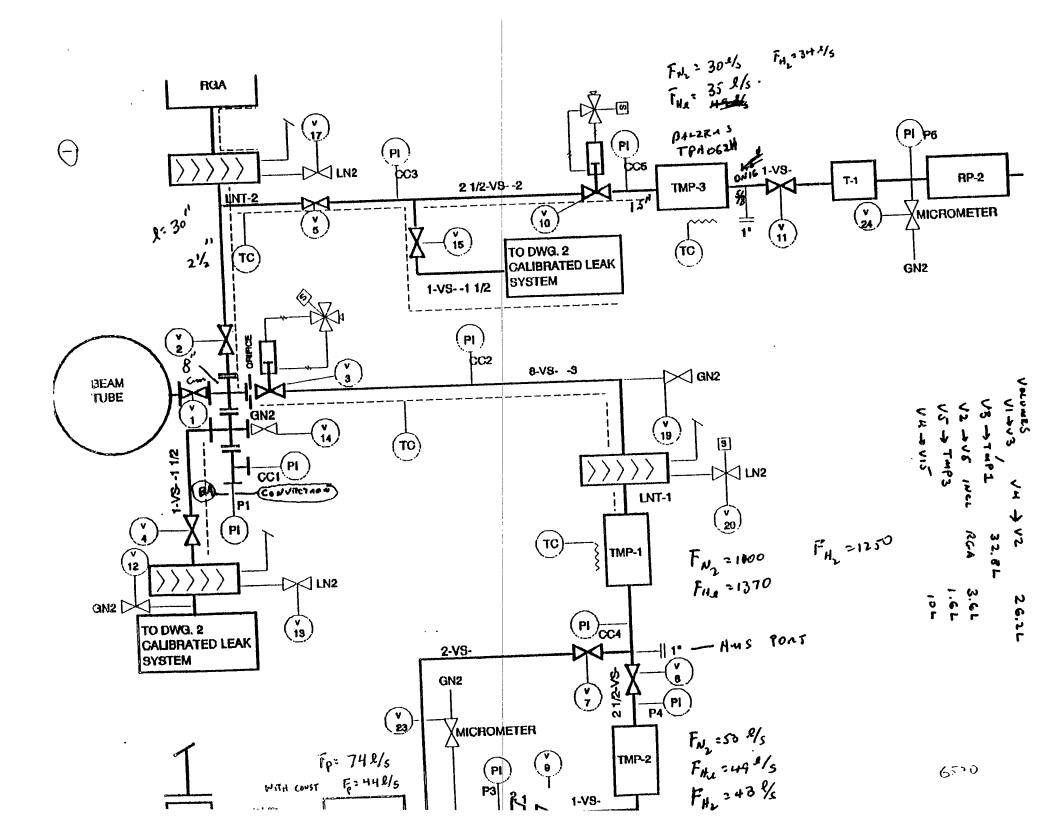
Jordan Camp

Dave Redding (JPL)

Yaron Hefetz

IFO configuration definition

6950127-0013



He ON FATTE TUBE

4.7 ×10-10 cc-atus/sec

ALL LRAKS IN CC-ATM/SAC

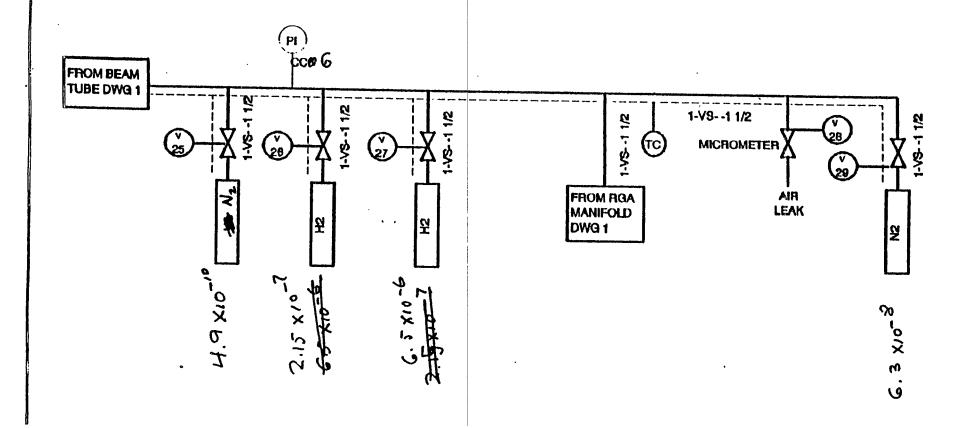
BRAM TUBR AARA

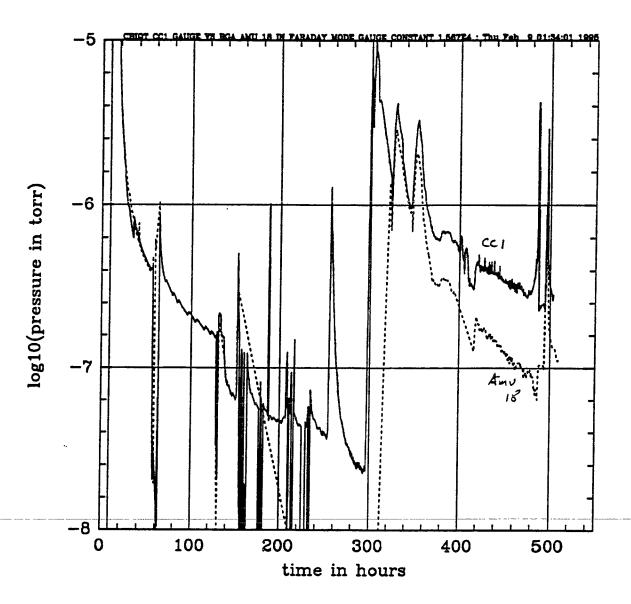
1.71 X/6 Cm²

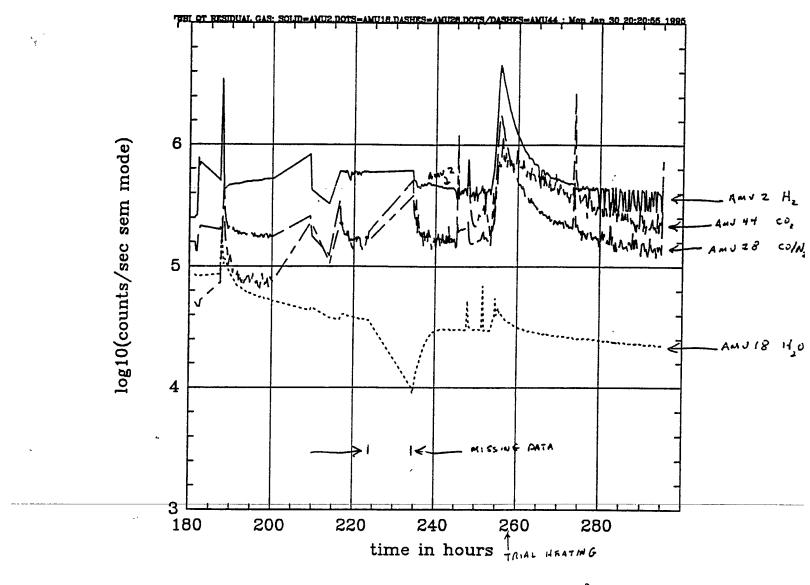
BRAM TUBR JOLUMA

5.124 X 104 Liter

Restauration of the second







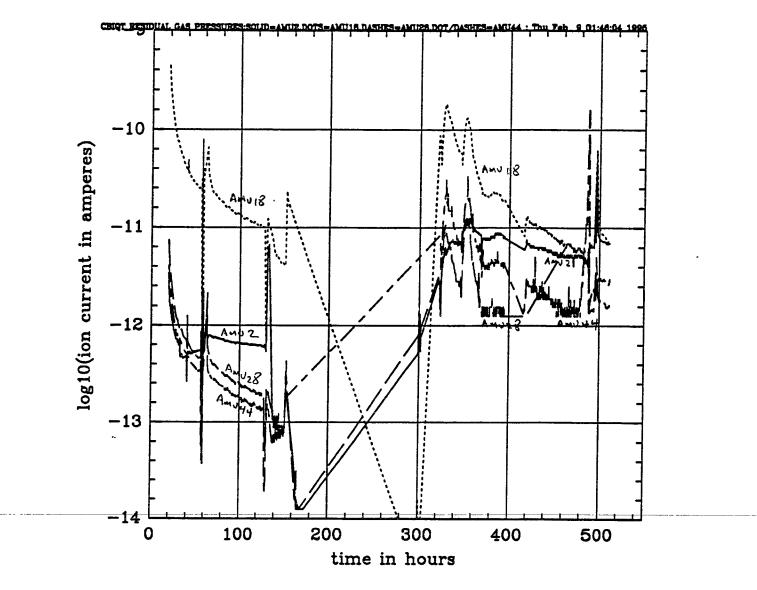
COUNTER / SEM MODE LNT! LNT2 AT 770K

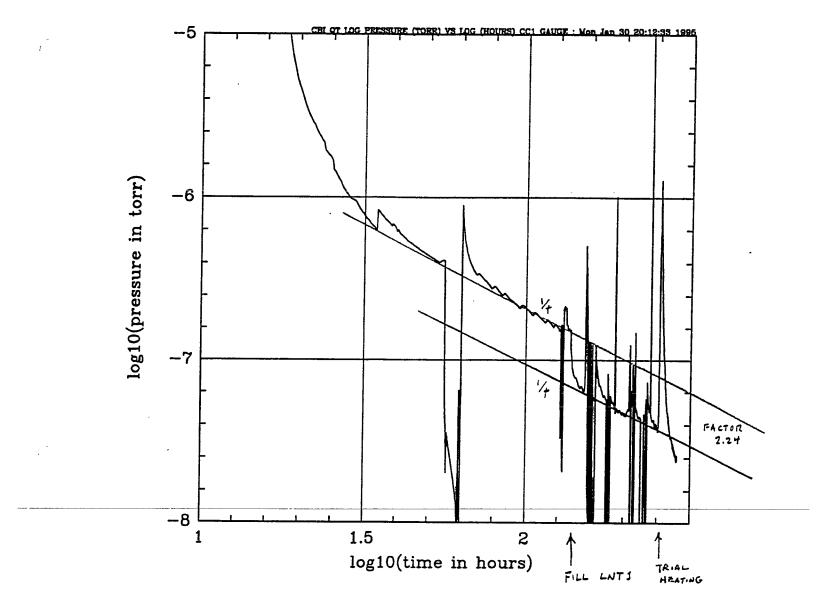
COUNTRY GAJGE CONSTANTS tors/court/SEC

3.2 ×10 -15 H2

1.5 ×10-15 (?) N2

NOTE: CC GAUGES ON

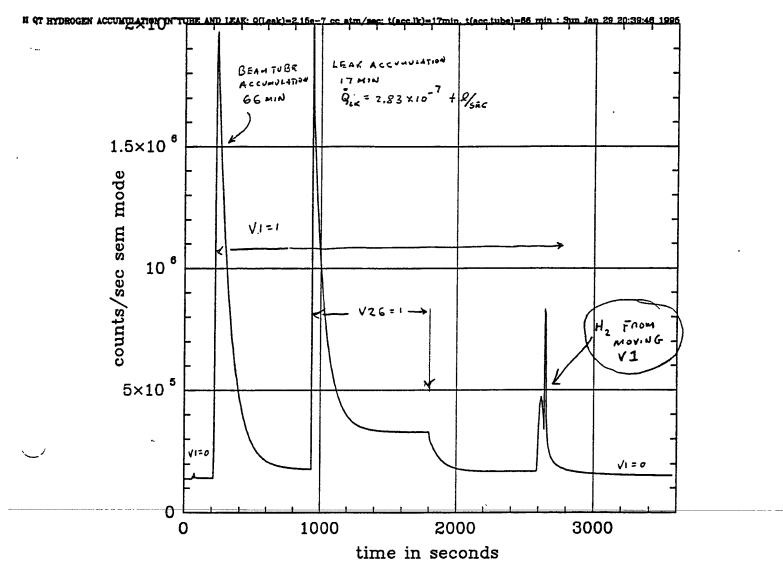




$$\frac{\text{WATER OUT GASSING}}{\text{J}_{40}(t)} = \frac{1.15 \times 10^{-8}}{\text{J}_{(hours)}} + \text{orr liters/sec cm}^2$$

$$\frac{U_{SING}}{F_{H_{20}}} = 311 \pm 10 \quad liters/sec$$

$$F_{H_{20}} = 696 \pm 23 \quad liters/sec \quad LNT1 \quad AT \quad 77^{\circ}K$$



HYDROGEN MEASUREMENT: LNTI, LNTZ AT 77°K.

RGA COUNTER MORE

CC GAUGES OFF

RESULTS!

H<sub>2</sub> Pumping stree 
$$611 \pm 5$$
 liters/sic  $3_{H_2} = 84$  Dec

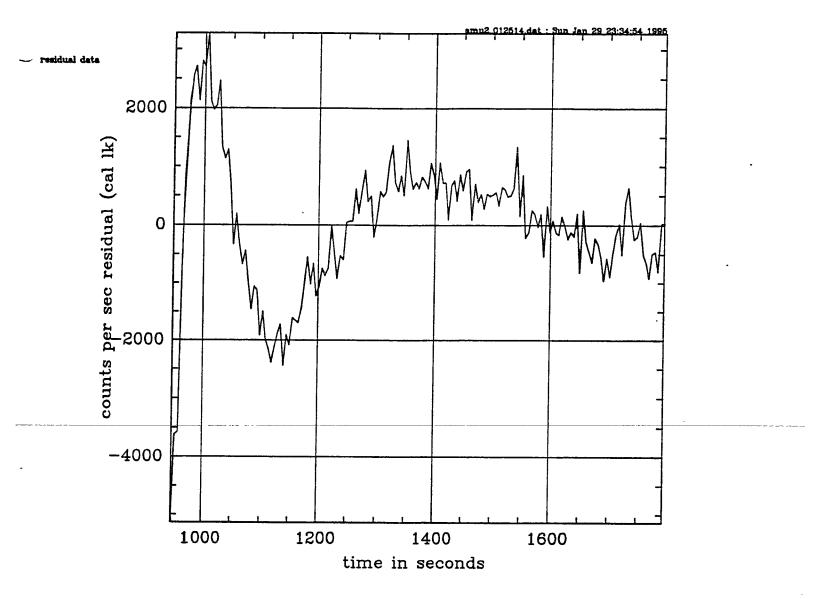
 $J(H_2) = 5.1 \pm .3 \times 10^{-14}$  for liters/sic cm<sup>2</sup> Transient

 $= 3.8 \pm 1. \times 10^{-14}$ 

H<sub>2</sub> Caige constant  $3.1 \times 10^{-15}$  Torr/cont/sic  $576.401$  Strans state

 $= 3.8 \pm 1. \times 10^{-15}$  Torr/cont/sic  $576.401$  Strans state

amuz -012514 .ps



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<ul> <li>IFO configuration definition</li> </ul>		Yaron Hefetz



## LIGO Beam Tube Vibration Study

(Dynamic Response of the Tube to the Hanford Site Ambient Vibration Spectrum)

Presented at the

LIGO Integration Meeting, 9 February 1995

by

M. Gamble of the LIGO DETECTOR GROUP



### **OBJECTIVE**

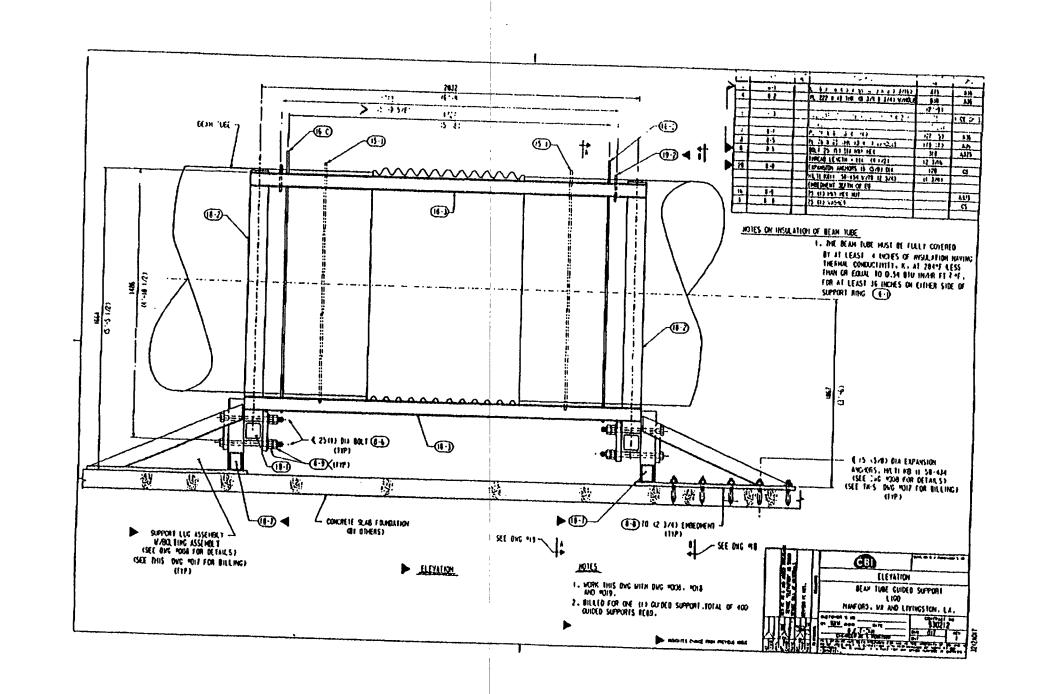
Determine the response of the LIGO beam tube to the standard LIGO input spectrum (Hanford site ambient spectrum) and infer the transfer function matrix elements, as functions of frequency, describing the support hardware.

### **METHOD**

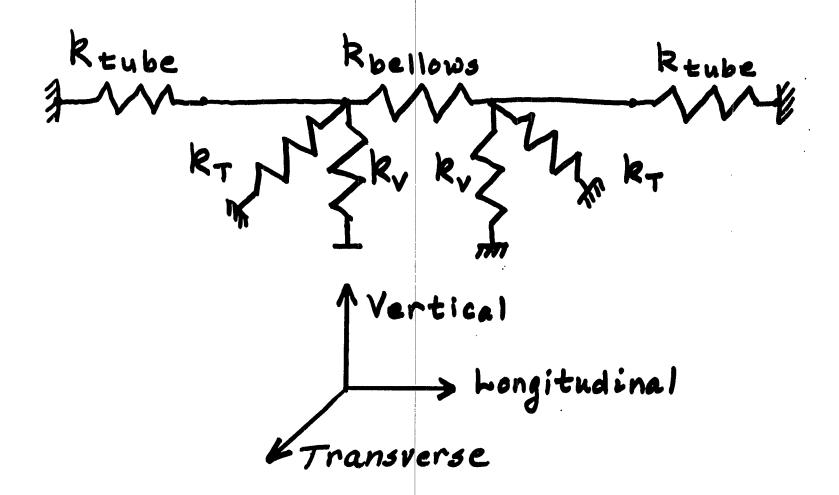
A general purpose finite element code used to determine the response of tube extension and bending modes (using 3D, kirchoff beam elements) and tube ovalization modes (using thin, isoparametric shell elements) supported by grounded hardware (modeled using 2D spring elements).

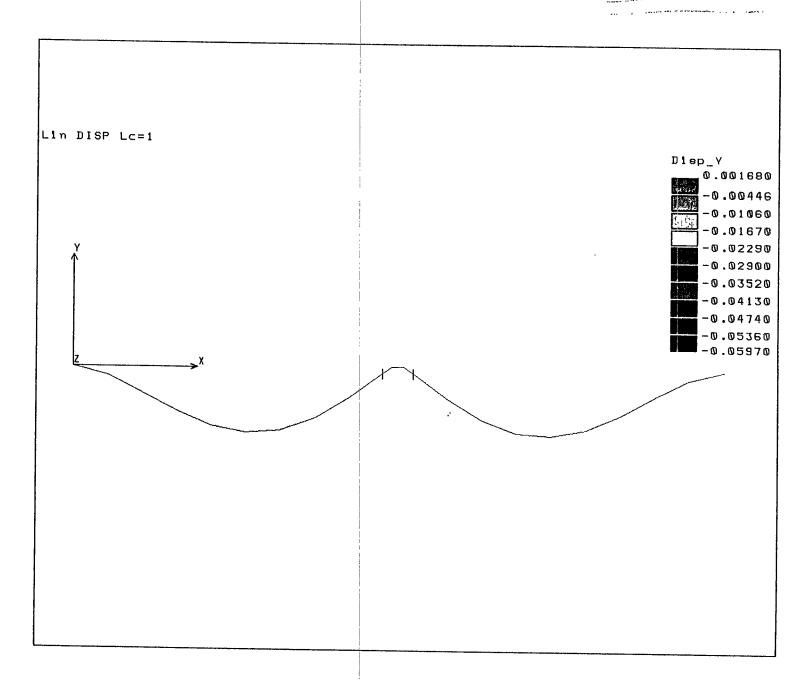
### **ASSUMPTIONS**

- Neglect the effects of soil / structure interactions (apply input PSD directly to the support hardware base nodes).
- The tube's vibrational behavior can be characterized by modeling only 2 rigid pipe supports and a single guided pipe support located at an arbitrary bellows.
- Diagonal matrix elements are much larger than off diagonal elements.
- Small strains and linear / elastic materials.

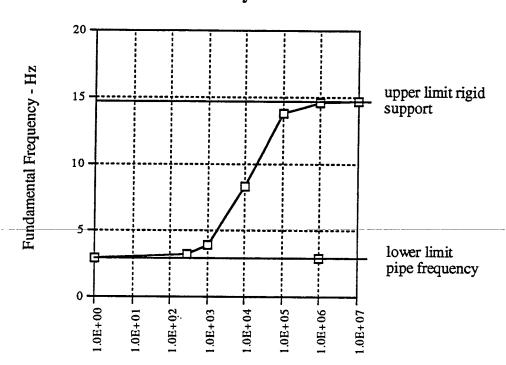


## 3D Finite Element Model Schematic



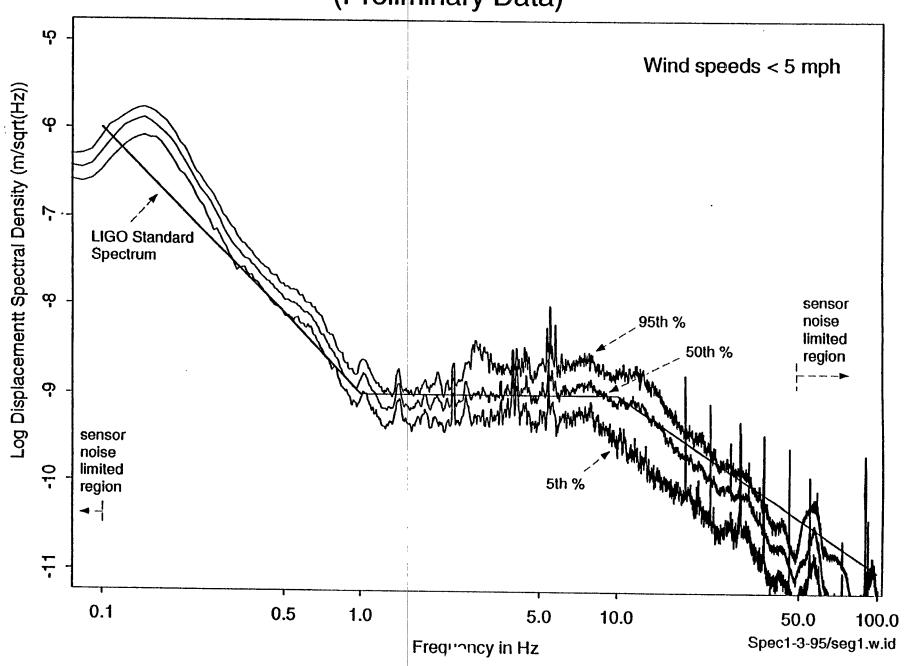


#### Bellows Horizontal Stiffness Trade Study



Horizontal Stiffness - Kz

Hanford Corner Station SW Arm Axis, Morning Traffic December 13,1994 (Preliminary Data)





### INPUT / OUTPUT QUANTITIES

#### **INPUT POWER SPECTRUM**

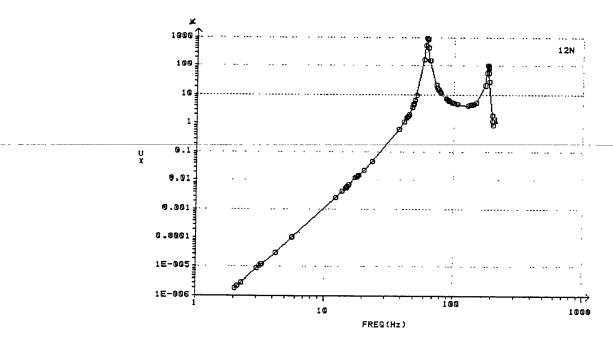
- Hanford site ambient ground spectrum
- Displacement power
- Standard power units [m<sup>2</sup>/Hz]

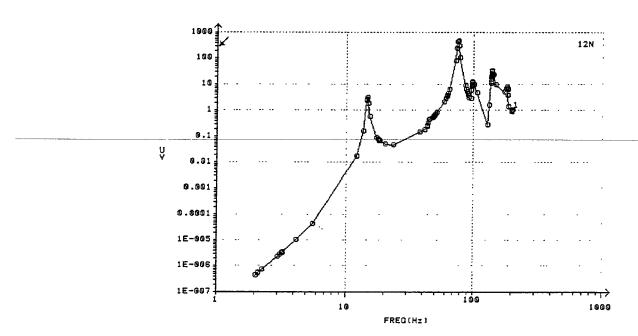
### **TRANSFER FUNCTION**

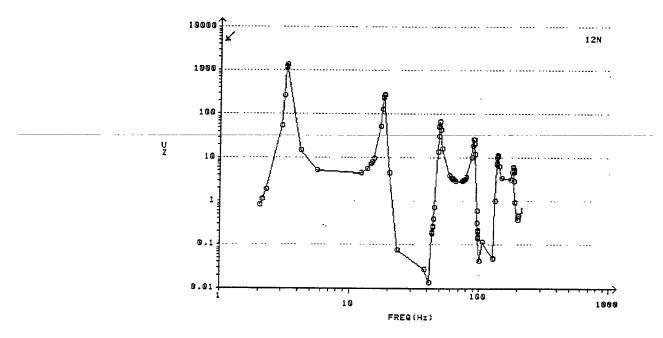
- RELATIVE frequency response function, H(f)
- Plotted values are /H(f)/2

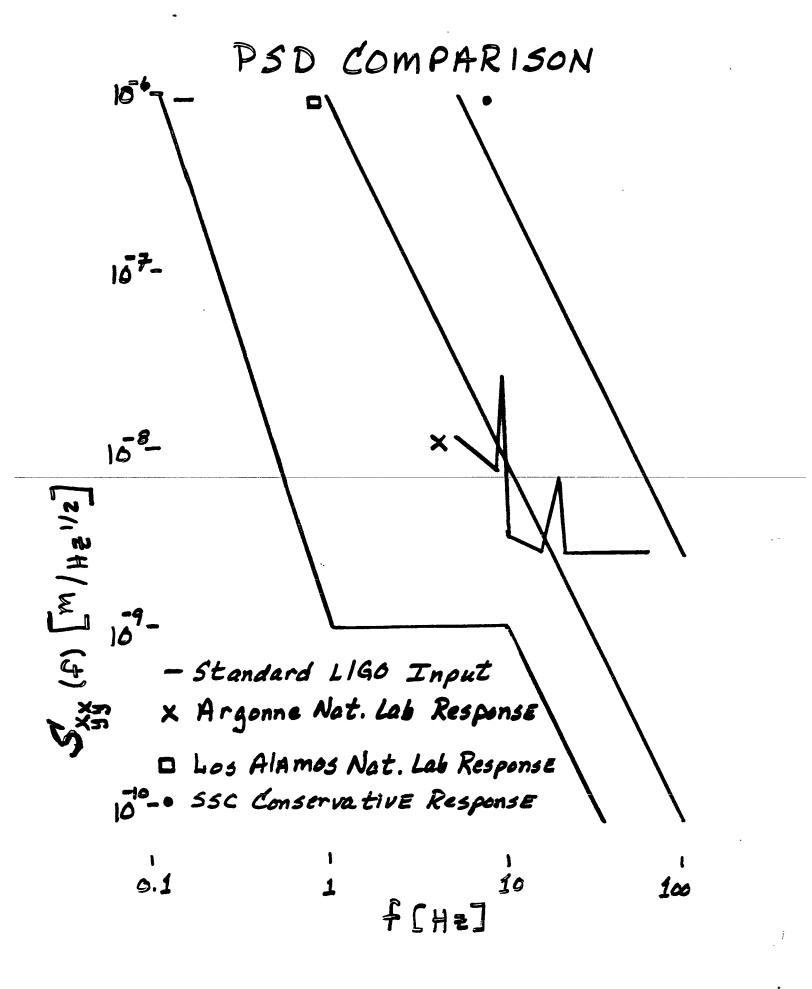
### **OUTPUT POWER SPECTRA**

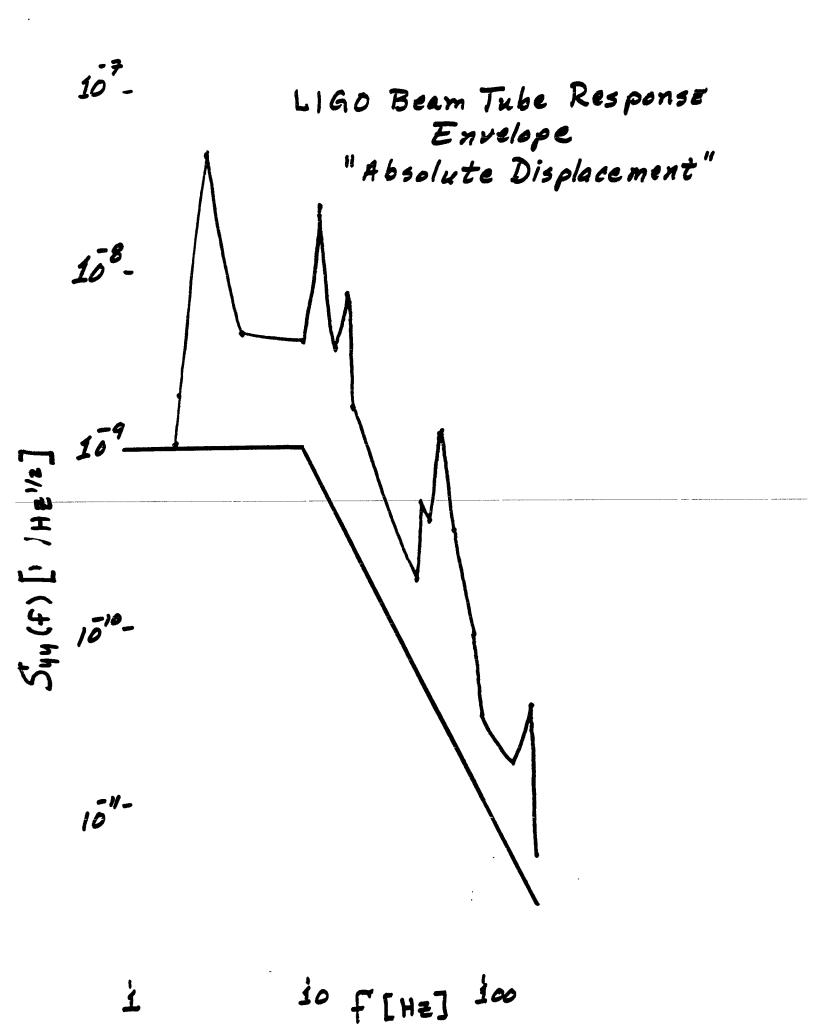
- RELATIVE displacement power { Szz(f) = Syy(f) Sxx(f) }
- ABSOLUTE displacement power  $\{Syy(f) = /H(f)/2 Sxx(f)\}$
- "Linearized" units [m/Hz<sup>1/2</sup>]













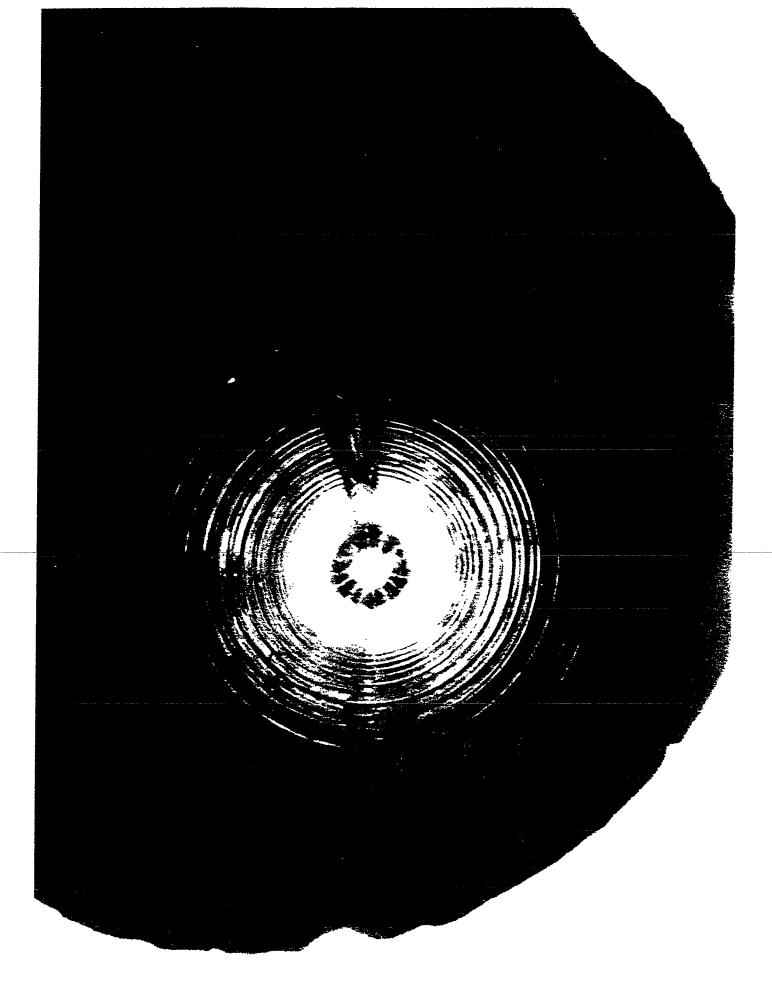
### **RESULTS**

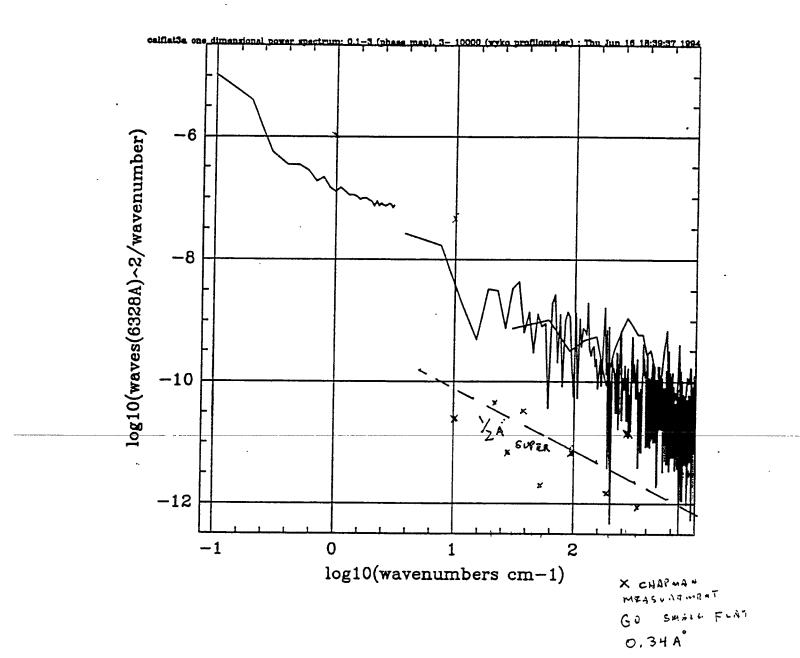
- The gravity-induced deflection of the tube was well modeled by CBI and has a magnitude of approximately 0.0015 m.
- The fundamental mode of vibration of the tube is a strong function of the lateral stiffness of the guided tube support and has a magnitude of approximately 3.2 Hz, when modeled as described by CBI mechanical drawings. The second mode of vibration agrees with CBI calculations and has a magnitude of approximately 14 Hz.
- Gain factors for the tube / support system reach a maximum of approximately 30 in the longitudinal, transverse, and vertical directions. The system is generally stiffest in the longitudinal direction and most flexible in the transverse direction.
- The frequency of the tube's first ovalization mode is approximately 35 Hz and demonstrates little participatory mass.
- Beam tube responses are well above the input PSD at frequencies below 20 Hz. Large peaks, also well above the input PSD, are found near 60 Hz, within the LIGO operations bandwidth.

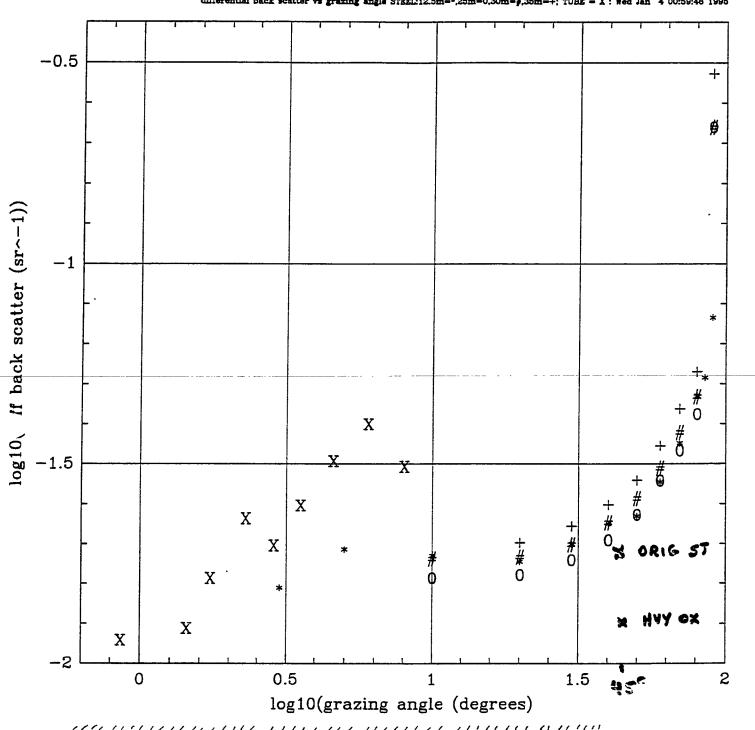


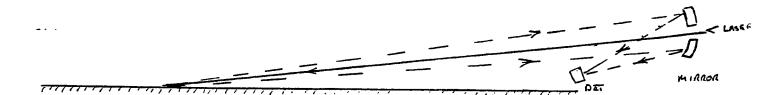
### **CONCLUSIONS and RECOMMENDATIONS**

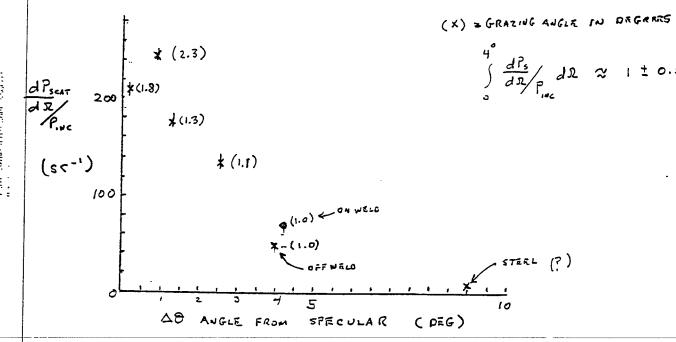
- The LIGO beam tube, as currently envisioned, will undergo large vibration-induced deflections within the gravity wave observation bandwidth.
- The response estimates provided today neglect other sources of motion enhancement such as soil / structure resonance, acoustic noise sources, and baffle / tube motion enhancement,
- As a minimum, the current guided support design should be reconsidered and potentially stiffened in the transverse direction.
- Visco-elastic material should be considered for the interfaces of the support hardware and their mating surfaces fixed to the ground.
- An estimate of the response behavior of the baffle, excited by the beam tube, using an accurate shell model of the tube, should be executed.

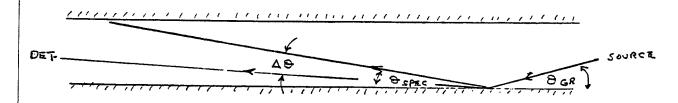








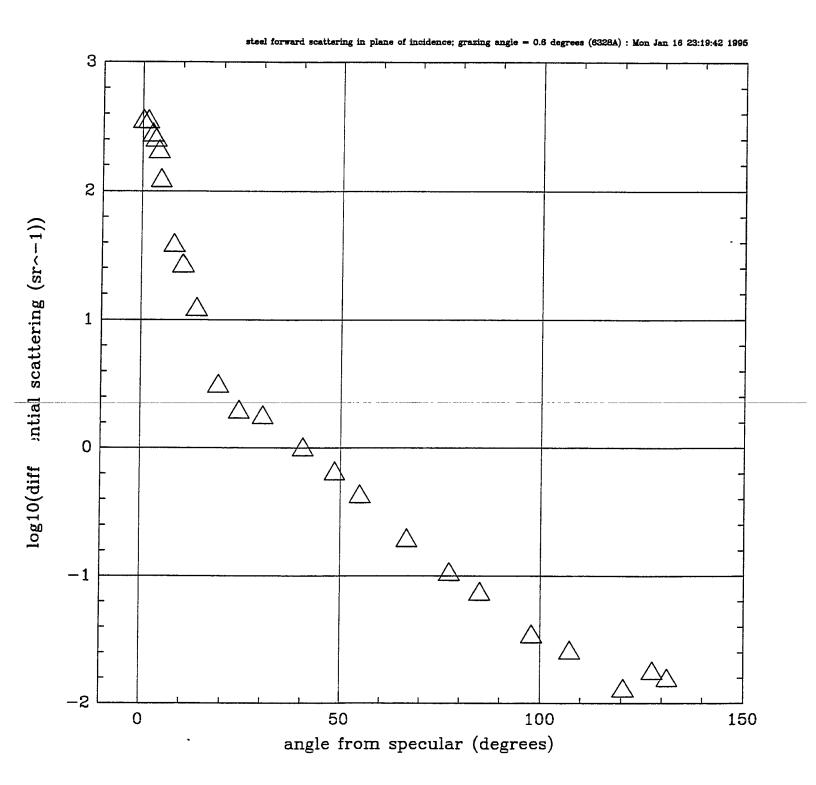


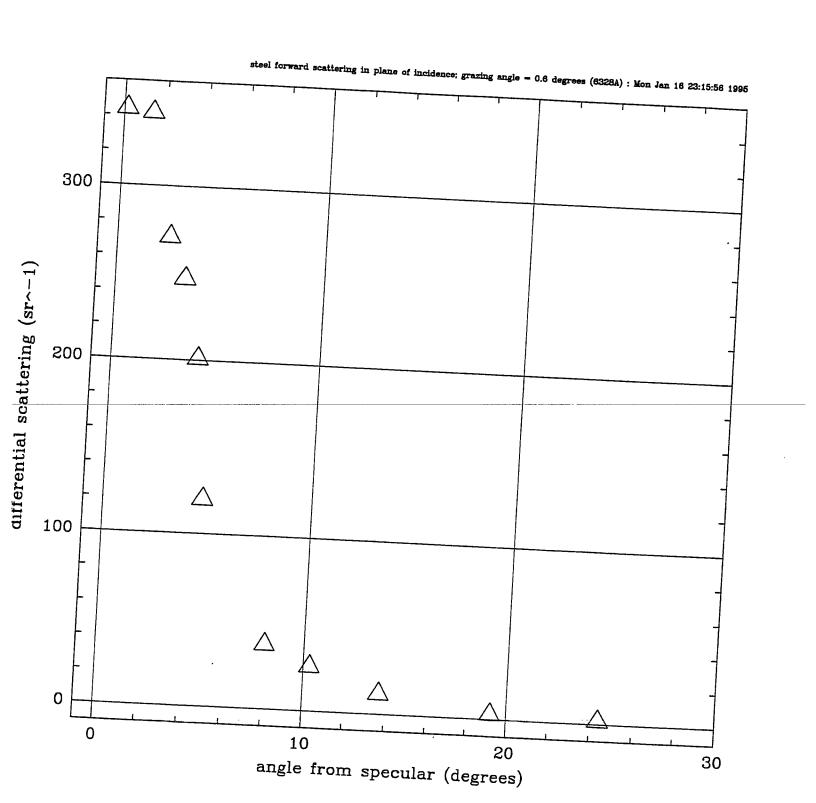


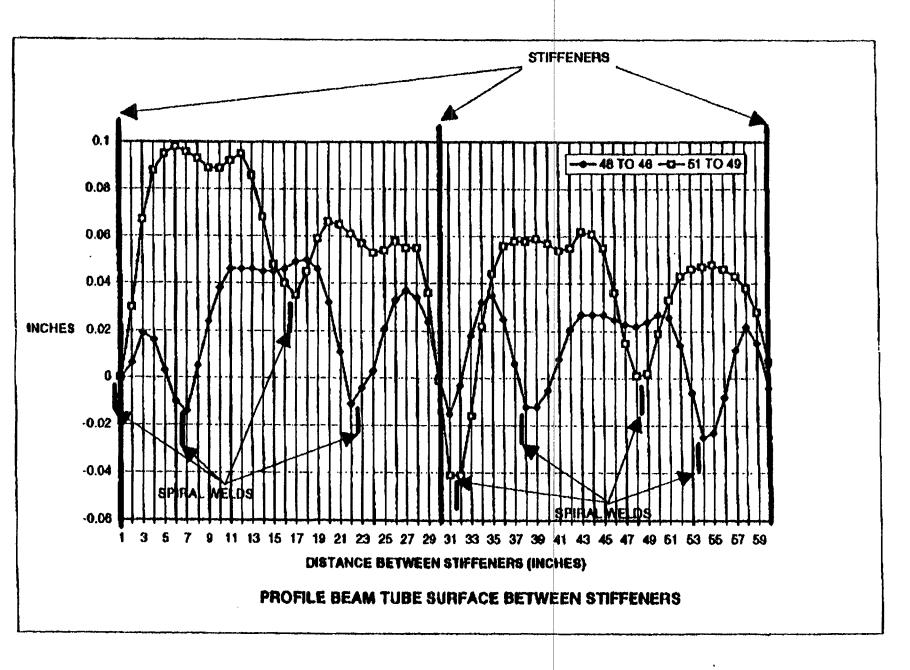
LASER BEAM DIVERGENCE (FULL ANGLE) 1.65 × 10-3 RABIANS

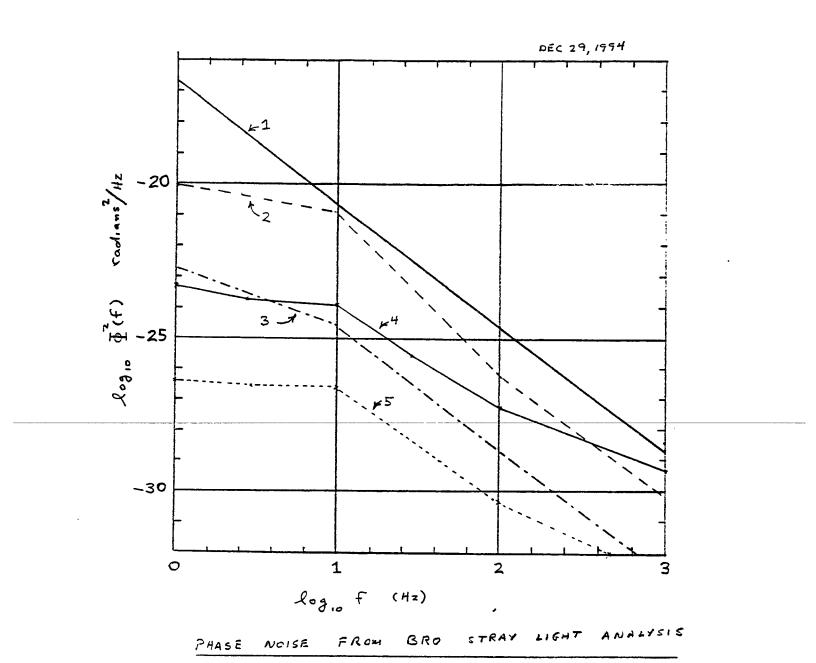
NEED TO MAKE FURTHER MEASUREMENTS

- 48 > 5 04 1) STEEL SAMPLE
- 2) TOTAL SCATTER
- 3) LOSS









1 1/10 (AMPLITURE) S GL 10 g ms 4 Km

2 ALL PATHS, CAVITY MIRROR RECOMBINATION FO = 10 HZ

3 NO BACK SCATTER PATHS, CAVITY RECOMBINATION FO = 10 HZ

4 ALL PATHS, DETECTOR RECOMBINATION TO SEUS ~ 10-3

5 NO BACK SCATTER PATHS, DETECTOR RECOMB.

BAFFLE AND STREL SURFACE DIFFERENTIAL SCATTERING  $\frac{dP_{SCAT}}{dR} = 1.25 \times 10^{-2} \text{ Cod } \Theta \text{ SC}^{-1}$ 

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### SOURCES OF SCATTERED-LIGHT INDUCED PHASE NOISE

Phase Noise Contribution	Source(s)	Comments
$\left(\frac{1}{P}\frac{dP}{d\Omega}\right)_{\pi}$ (backscatter) $\left(\frac{1}{P}\frac{dP}{d\Omega}\right)_{0}$	Baffle materials - BRDFs  Tube wall - BRDF  Tube wall - BRDF	Use diffuse (black) materials Use absorptive (black) glass (VIRGO) Avoid glints (handling/fabrication)  Baffle to avoid seeing walls  Baffle to avoid seeing walls
(forward scatter) $ \left(\frac{1}{P}\frac{dP}{d\Omega}\right)_0 $ (diffraction)	Edges	Reduce number of baffles by appropriate baffle spacing Reduce coherence effects by serrating & randomizing serration amplitudes
A[f]	Beam tube/baffle motion amplification	Preliminary effect indicates amplifications up to 30x possible Shell model for beam tube; radial motions - Rod model; translational motions
σ <sub>recombination</sub>	Mirror surface roughness at high v [cm <sup>-1</sup> ]	Address coating and polishing techniques - Future effort not immediately addressable

- MEETING WAS SPONSORED BY CALTECH (CHAIRED BY STAN WHITCOMB, ORGANIZED BY KIP THORNE) ON FRIDAY, SATURDAY 6,7 JANUARY.
- ATTENDED BY:
  - LIGO
    - BARRY BARISH, VLADIMIR BRAGINSKY, RON DREVER, EANNA FLANNIGAN, MIKE GAMBLE, LARRY JONES, ALBERT LAZZARINI, GARY SANDERS, RICK SAVAGE, GERRY-STAPFER, ROBBIE VOGT, RAI WEISS, JOHN WORDEN, HIRO YAMAMOTO
  - VIRGO
    - JEAN-YVES VINET
  - GEO
    - WALTER WINKLER
  - BREAULT RESEARCH ORGANIZATION -- CONSULTANT
    - BOB BREAULT
  - US NAVAL SURFACE WEAPONS CENTER, CHINA LAKE -- CONSULTANT
    - HAL BENNETT

### VIRGO BAFFLE AND BEAM TUBE CONCEPTUAL DESIGN

- BEAM TUBE WILL BE EITHER LIGO-LIKE, STRAIGHT TUBE OR OF A CORRUGATED DESIGN, WITH CORRUGATIONS HAVING 90 cm PITCH, 25 mm AMPLITUDE
  - BOTH CONCEPTS BEING EVALUATED
- BAFFLES ARE MADE OF BLACK GLASS IN A TRUNCATED-CONE (45°) PLUS CYLINDRICAL HOOP CONFIGURATION
  - 3 REFLECTIONS OFF (AR COATED) ABSORPTIVE GLASS PROVIDE  $T = 2 3 \times 10^{-7}$
  - DIRECT BACKSCATTER IS VERY LOW DUE SURFACE FINISH OF THE GLASS:  $BRDF \approx 1 \times 10^{-6}$
  - NO SERRATIONS ARE PLANNED [AT THIS TIME]
  - BAFFLE HOOPS ARE PLANNED TO BE INCLINED TO OPTICAL AXIS  $\theta \approx 10^{\circ}$
- BAFFLE SPACING WILL BE A GEOMETRIC PROGRESSION:  $Z_n = Z_0 (1.15)^n$
- BAFFLE DESIGN TO BE COMPLETED BY MARCH 1995; DECISION ON DESIGN BY APRIL 1995;
- UNRESOLVED ISSUES:
  - INSERTION OF BAFFLES INTO TUBES NOT YET ADDRESSED;
  - BAFFLES EDGES WILL NEED TO BE TAPERED TO KNIFE-LIKE CHARACTERISTIC TO REDUCE GLINTS
  - NEED TO SERRATE BAFFLES TO BE ADDRESSED

### GEO BAFFLE AND BEAM TUBE CONCEPTUAL DESIGN

- BEAM TUBE NEEDS TO BE 1/2 LIGO/VIRGO DIAMETER
- BAFFLES ARE OF A CONFIGURATION SIMILAR TO VIRGO:
  - CONE ANGLE IS26° -- 5 REFLECTION'S REQUIRED TO EXIT VERTEX OF CYLINDER-CONE
  - TWO SUCH LIGHT TRAPS AT EACH BAFFLE, ARRAGNED BACK-TO-BACK (LOOKING AT OPPOSITE MIRRORS)
  - BAFFLE HEIGHT IS 3 cm; BAFFLE IS SPACING LINEAR, 6 m
  - MATERIAL EITHER POLISHED STAINLESS STEEL OR BLACK-GLASS COATED SS
  - DIRECT BACKSCATTER IS VERY LOW DUE SURFACE FINISH OF THE STEEL/GLASS:  $BRDF \approx 1 \times 10^{-6}$
  - ISSUE OF SERRATIONS WAS NOT ADDRESSED [YET]
- BAFFLE DESIGN IS VERY PRELIMINARY

### LIGO BASELINE BAFFLE DESIGN

- BAFFLES ARE MADE OF A SINGLE BENT SHEET INCLINED AT ANGLE OF 55 AWAY FROM NEAREST TEST MASS -- PROJECTION ALONG THE RADIAL DIRECTION IS 6 cm
- BAFFLE HAS SAME OPTICAL PROPERTIES AS BEAM TUBE (OXIDIZED) STAINLESS STEEL
- REGULAR SERRATIONS ALONG THE EDGE ARE PRODUCED BY COMPUTER CONTROLLED WATERJET CUTTING -- 2 mm DEPTH, ALONG RADIAL DIRECTION; UNIFORM IN PERIOD AND AMPLITUDE
- BAFFLE DIAMETER IS SLIGHTLY OVERSIZED TO PROVIDE A "SPRING LOADED" PRESS FIT
- BAFFLE INSERTED WITH A HELICAL BIAS OF  $\cong$  30 cm/TURN, WITH  $\cong$  30 ° OVERLAP
- BAFFLES ARE UNIFORMLY SPACED ALONG TUBE FROM END TO MID-POINT; BEYOND MID-POINT, THE SEQUENCE IS REVERSED:

0<1<100 m; NO BAFFLING

100 m < L < 250 m; BAFFLES EVERY 7 m, EVERY 3rd IS AT A BEAM TUBE SUPPORT 250 m < L < MID-POINT; BAFFLES EVERY 21 m, LOCATED AT SUPPORTS

#### **OUTCOME OF THE REVIEW**

- LIGO IS NOT READY TO FREEZE BAFFLE DESIGN(!)
- THE BEAM TUBE BACKSCATTER, MEASURED BY RAI, WAS MUCH HIGHER THAN WAS ORIGINALLY ASSUMED IN THE BAFFLING ANALYSIS AND DESIGN.

BRDF FOR NEAR GRAZING INCIDENCE (1-2 °) WAS ≅100x GREATER THAN EXPECTED BACKSCATTER FROM TUBE SURFACE IS THE LARGEST SOURCE OF PHASE NOISE IN THE PRESENT BASELINE DESIGN.

• THE BEAM TUBE FORWARD SCATTER WAS MEASURED TO BE MORE BENIGN THAN WAS ORIGINALLY ASSUMED.

THERE IS NO WELL-DEFINED SPECULAR COMPONENT, BUT RATHER, A BROAD MAXIMUM ALONG THE SPECULAR DIRECTION, WITH  $\theta_{rms} \cong 3^{\circ}$ .

• MIKE GAMBLE'S PRELIMINARY ANALYSIS OF COUPLING OF GROUND MOTION TO THE BEAM TUBE WALL INDICATES AMPLIFICATIONS AS GREAT AS 30x MAY OCCUR AT STRUCTURAL RESONANCES, WITH TYPICAL AMPLIFICATIONS OF 2x OFF RESONANCE IN 50 - 200 Hz BAND:

ANALYSIS WAS FOR A ROD MODEL OF THE BEAM TUBE -- NO SHELL DOFs.

LARGEST AMPLIFICATION IS TRANSVERSE TO THE BEAM AXIS, IN THE PLANE PARALLEL TO THE GROUND -- AFFECTS DIFFRACTION CONTRIBUTION.

AMPLIFICATION TRACED TO THE DESIGN OF THE GUIDED SUPPORT AT EXPANSION JOINTS

• A NUMBER OF MODIFICATIONS TO THE LIGO BAFFLE DESIGN WERE PROPOSED:

TIGHTEN DESIGN GOAL TO PROVIDE FOR A LIMITING CONTRIBUTION TO THE DETECTOR NOISE FLOOR CORRESPONDING TO 0.1 SQL FOR A 1-TON TEST MASS

REPLACE BARE OXIDIZED STAINLESS STEEL MATERIAL WITH A DIFFUSIVE/ABSORPTIVE COATING TO REDUCE BACKSCATTER

BAFFLE THE 1st 100 m OF THE BEAM TUBE

REMOVE UNNEEDED BAFFLES FROM NEAR MID-STATION TO REDUCE DIFFRACTING EDGES

• USE SPACING SEQUENCE CORRESPONDING TO GEOMETRICAL PROGRESSION RATHER THAN LINEAR ONE -- VIRGO APPROACH

FURTHER REDUCE DIFFRACTION BY RANDOMIZING SERRATION AMPLITUDES BY  $\pm 0.5\,$  mm PLACE BAFFLES AT OR NEAR FIXED BEAM TUBE SUPPORTS TO REDUCE STRUCTURAL AMPLIFICATION OF GROUND MOTION

CONSIDER BETTER FASTENING METHOD FOR BAFFLE WITHIN TUBE

ADDITIONAL MEASUREMENTS AND CALCULATIONS ARE NEEDED:

FOLD RAI'S BEAM TUBE BRDF MEASUREMENTS INTO FORWARD PROPAGATION OF SCATTERED LIGHT

REDO MONTE CARLO ANALYSIS WITH MEASURED SCATTERING DISTRIBUTIONS COMPLETE THE STRUCTURAL ANALYSIS OF BEAM TUBE MOTION AMPLIFICATION

## LIGO ACTIONS RESULTING FROM THE RECOMMENDATIONS OF THE BAFFLE REVIEW MEETING

- A REVIEW OF OUR SCHEDULE INDICATES WE HAVE UNTIL DECEMBER 1995 TO PRODUCE A FINAL BAFFLING DESIGN
- LIGO WILL COMPLETE THE SCATTERING ANALYSIS:

COMPLETE THE FORWARD PROPAGATION ANALYSIS

**COMPLETE BRDF ANALYSIS (DONE)** 

REFINE THE BAFFLE MOTION ESTIMATE

UPDATE PERFORMANCE CURVES AND SAFETY MARGIN ESTIMATES

- LIGO CONTRACTED BREAULT RESEARCH ORGANIZATION (BRO) TO PROVIDE A LIST OF CANDIDATE DIFFUSE AND/OR ABSORPTIVE MATERIALS WITH THE LEVEL OF OPTICAL PERFORMANCE REQUIRED FOR THE BAFFLES
- LIGO WILL SELECT TWO CANDIDATE COATINGS;

FIRST CANDIDATE (i.e., BEST OPTICAL PERFORMER) IS EXPECTED TO BE AN ORGANIC DYE-BASED MATERIAL, WHICH WILL BE EVALUATED FOR CONTAMINATION EFFECTS ON OPTICS;

SECOND CANDIDATE WILL BE THE BEST INORGANIC MATERIAL

• BOTH MATERIALS WILL BE EVALUATED FOR OUTGASSING CHARACTERISTICS

## LIGO ACTIONS RESULTING FROM THE RECOMMENDATIONS OF THE BAFFLE REVIEW MEETING (continued)

• LIGO WILL WORK (WITHIN THE SCHEDULE CONSTRAINTS -- 3-4 MONTH SETUP TIME; 3-4 MONTH RINGDOWN TESTING) TO SET UP A RINGDOWN FACILITY AND TO TEST THE ORGANIC DYE-BASED CANDIDATE FOR CONTAMINATION OF OPTICS <u>AT THE LEVEL REQUIRED BY LIGO</u>:

IF THE RINGDOWN TESTING FALLS BEHIND SCHEDULE... OR
THE ORGANIC DYE-BASED CANDIDATE FAILS THE RINGDOWN TESTS... OR
IF THE RESULTS ARE INCONCLUSIVE...
BAFFLE COATING SELECTION WILL DEFAULT TO THE BACK-UP MATERIAL

• LIGO WILL PURSUE, IN PARALLEL, THE MECHANICAL DESIGN OF THE MODIFIED BAFFLES TO INCLUDE:

MOUNTING CONCEPTS

SERRATION DESIGN

BAFFLE CONFIGURATION AND FABRICATION

THOSE ASPECTS OF DESIGN THAT ARE NOT DEPENDENT ON COATING AND COATING PROCESS SELECTION WILL BE CARRIED FORWARD, MAINTAINING THE OPTION TO SELECT EITHER COATING MATERIAL AND PROCESS - FINAL DESIGN WILL AWAIT MATERIAL SELECTION