



Parametric Instability

Comparison between
Mesa beams and Gaussian beams
And
Modal misalignment suppression for
Advanced LIGO

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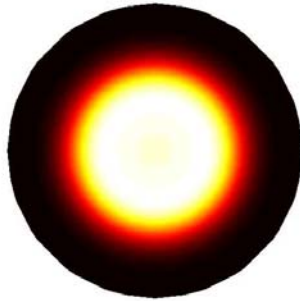
Comparison between Mesa beams and Gaussian beams



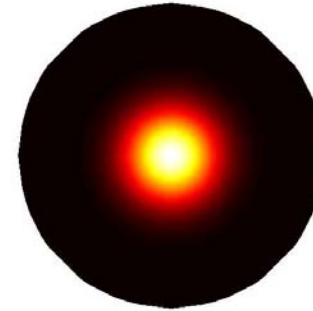
Cavity types



Mesa beam



Gaussian Beam



$$u_{00}^{MF}(D, \vec{r}) = \int_S \exp\left[\frac{-k(\vec{r} - \vec{r}')^2(1+i)}{2L}\right] d^2\vec{r}'$$

$$u_{00}^{GF}(R, \vec{r}) = \exp\left(-i\left[\frac{kL}{2} + \frac{k|\vec{r}|^2}{2R} - \arctan\left(\frac{2R-L}{L}\right)^{-1/2}\right]\right)$$

$$D = 4\sqrt{\frac{L}{k}}$$

Diffraction losses
 $\sigma = 19.5$ ppm

$$R = 2050.6 \text{ m}$$

Cavity length $L = 4000$ m $T_{ITM} = 2.3\%$ $T_{ETM} = 5.0$ ppm

Test mass: $r = 0.16$ m, $t = 0.13$ m, Sapphire M-axis, wedge 0.5deg



Parametric Instability- Simple cavity



$$R_j = \frac{4PQ_{mj}}{McL\omega_{mj}^2} \left(\frac{Q_{1i}\Lambda_{1i}}{1 + \Delta\omega_{1i}^2 / \delta_{1i}^2} - \frac{Q_{1ai}\Lambda_{1ai}}{1 + \Delta\omega_{1ai}^2 / \delta_{1ai}^2} \right)$$

Fixed parameters:

- circulating power $P = 830$ kW
- acoustic mode Q-factor $Q_m = 10^7$
- mass of the test mass $M = 40$ kg

Λ - overlapping parameter
(optical and acoustic mode shapes must be known)

Q_1, δ_1 – optical Q-factor and half line-width of the TEMs
(diffraction losses should be taken into account)

2 possible processes:

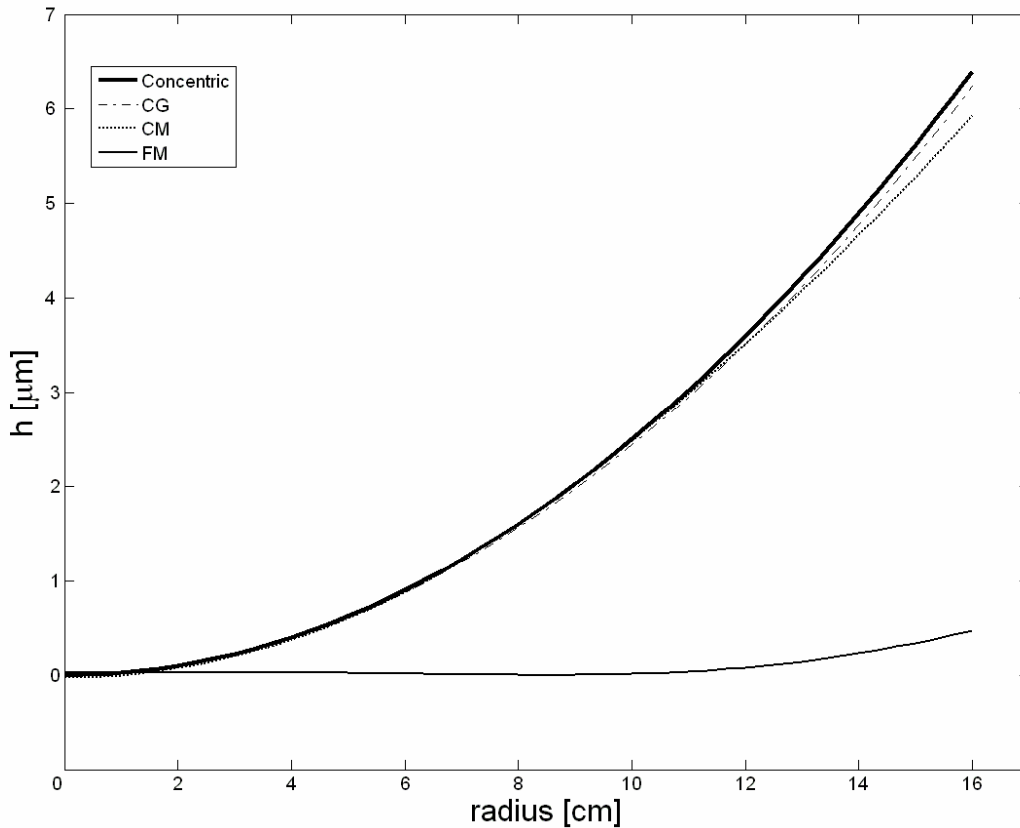
$$\omega_1 - \omega_0 = \sim \omega_m \quad : \text{damping} \rightarrow \Delta\omega_a \equiv \omega_1 - \omega_0 - \omega_m$$

$$\omega_0 - \omega_1 = \sim \omega_m \quad : \text{excitation} \rightarrow \Delta\omega \equiv \omega_0 - \omega_1 - \omega_m$$

optical $\omega_{0,1}$ and acoustic ω_m frequencies must be determined
→ cavity structure and acoustic mode calculation required



Cavity structure - Mirror shape



$$kh^F(\vec{r}) = \text{Arg}(u(D, \vec{r}))$$

$$h^F(\vec{r}) + h^C(\vec{r}) = \frac{r^2}{L}$$

Only near concentric cavities have small tilt instability thus only this cavity type is considered in PI analysis.



Cavity structure - Eigenvectors



The Fresnel-Kirchhoff propagation equation:

$$E_2(\vec{r}_2, L) = -\frac{ik}{2\pi} \iint_S G(\vec{r}_2, \vec{r}_1) E_1(\vec{r}_1, 0) F(\theta) d^2\vec{r}_1$$

For a fine meshgrid of the mirror surface the electric field E across element area d^2r becomes quasi-steady. The above equation can be written as an eigenequation.

$$\gamma E_1 = M_{21} M_{12} E_1$$

$$M_{12} = \frac{ik}{2\pi} \iint_S G(\vec{r}_2, \vec{r}_1) d^2\vec{r}_1$$

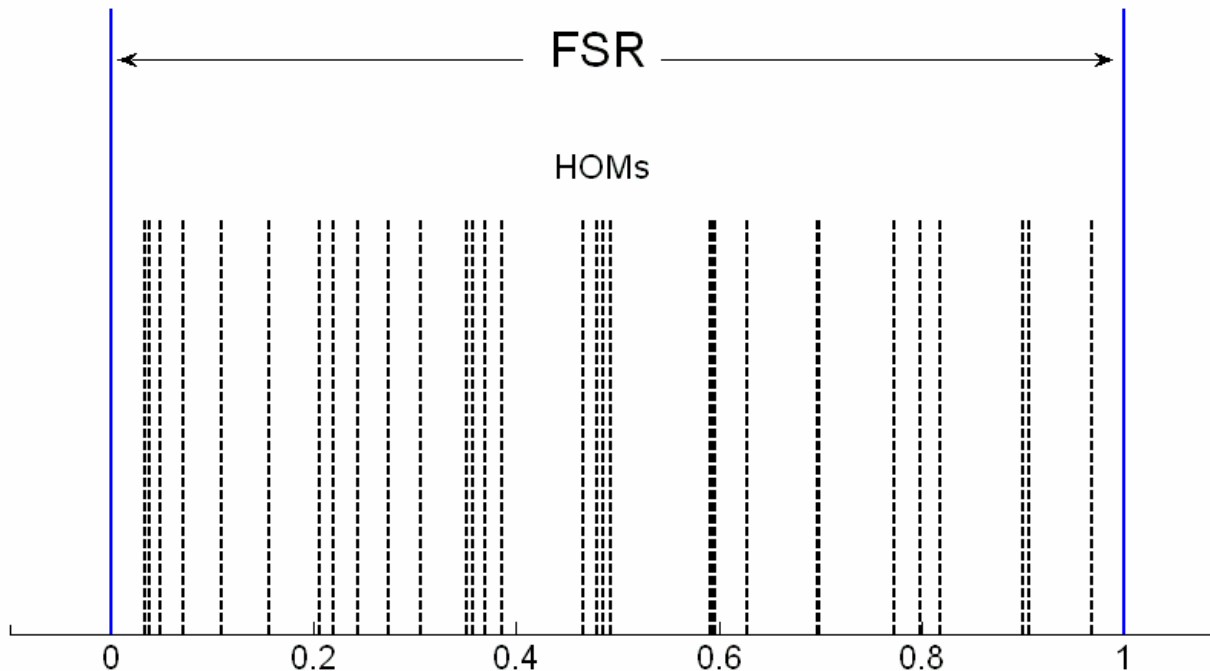
G – propagation kernel



Cavity structure – Gaussian beam



near concentric cavity (Gaussian)



$$\Delta \nu_{p0} = \frac{FSR}{\pi} (\arg(\gamma_p) - \arg(\gamma_0))$$

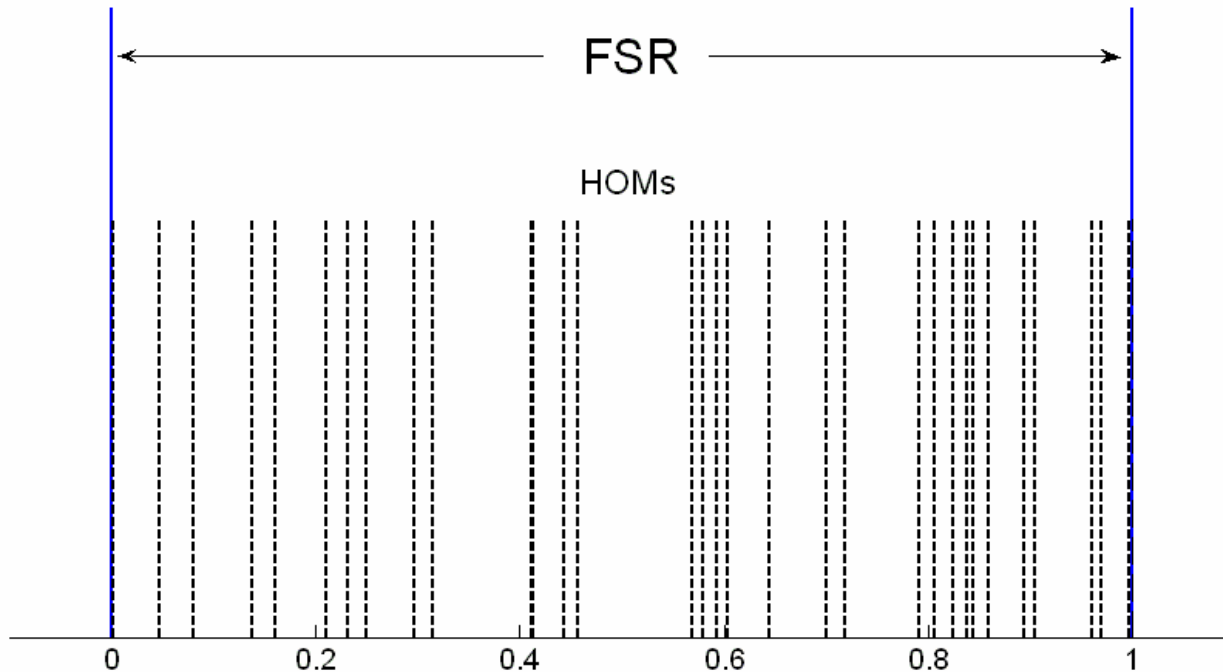


Cavity structure – Mesa beam



near concentric cavity (Mesa)

Near concentric cavity was obtained from near planar cavity using duality relation



Duality relation:

$$\gamma^C = e^{-4ikL} \left(\gamma^F \right)^* \rightarrow \Delta \nu^C = M \cdot FSR - \Delta \nu^F \quad \text{with} \quad M = 1$$



Acoustic modes analysis



$$M_{ab} \frac{\partial^2 u_i^b}{\partial t^2} + K_{aibk} u_k^b = 0$$

Global stiffness matrix:

$$K_{aibk} = \int_S C_{ijkl} \frac{\partial \varphi^a(\vec{\xi})}{\partial \xi_j} \frac{\partial \varphi^b(\vec{\xi})}{\partial \xi_l} dV$$

Global mass matrix:

$$M_{ab} = \rho \int_S \varphi^a \varphi^b dV$$

C – compliance tensor (due to the crystal symmetries can be reduced to the 6x6 matrix)

For sapphire M-axis:

$$C = \begin{pmatrix} 498.1 & & & & & \\ 110.9 & 496.8 & & & & \\ 110.9 & 163.6 & 496.8 & & & \\ 0 & 0 & 0 & 153.1 & & \\ 0 & -23.5 & 23.5 & 0 & 147.4 & \\ 0 & 0 & 0 & -23.5 & 0 & 147.4 \end{pmatrix} \quad \text{Symmetric}$$

assuming $u = \mu_i \cos \omega_i t$
we obtain an eigenequation

$$|K - \omega_i^2 M| = 0$$

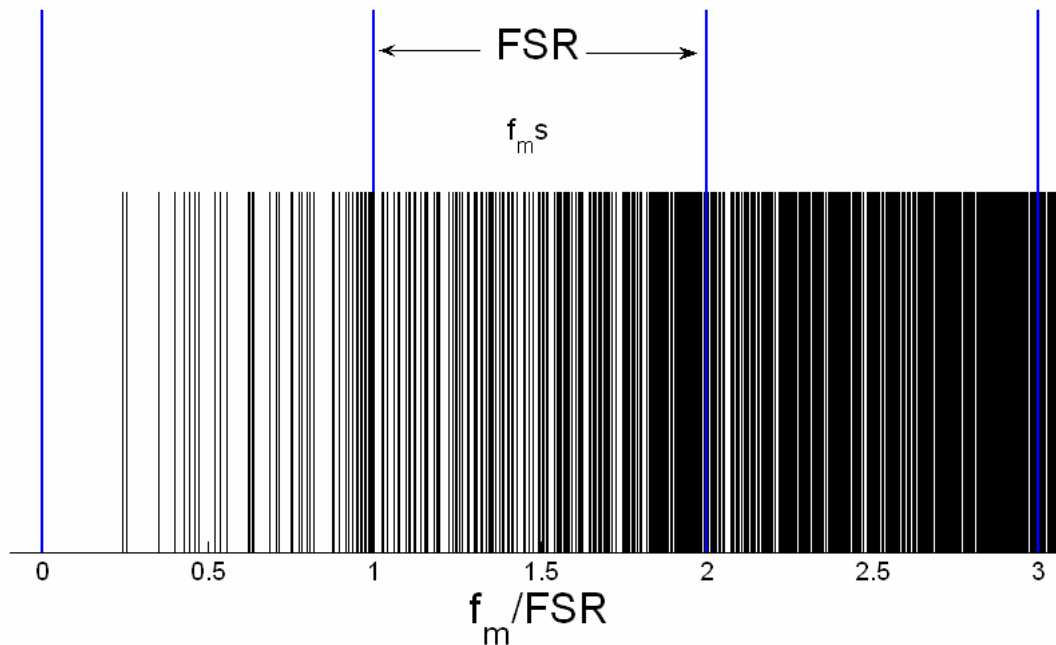
Solved with ANSYS – 140k nodes



Acoustic modes frequencies



acoustic modes M-axis Al_2O_3



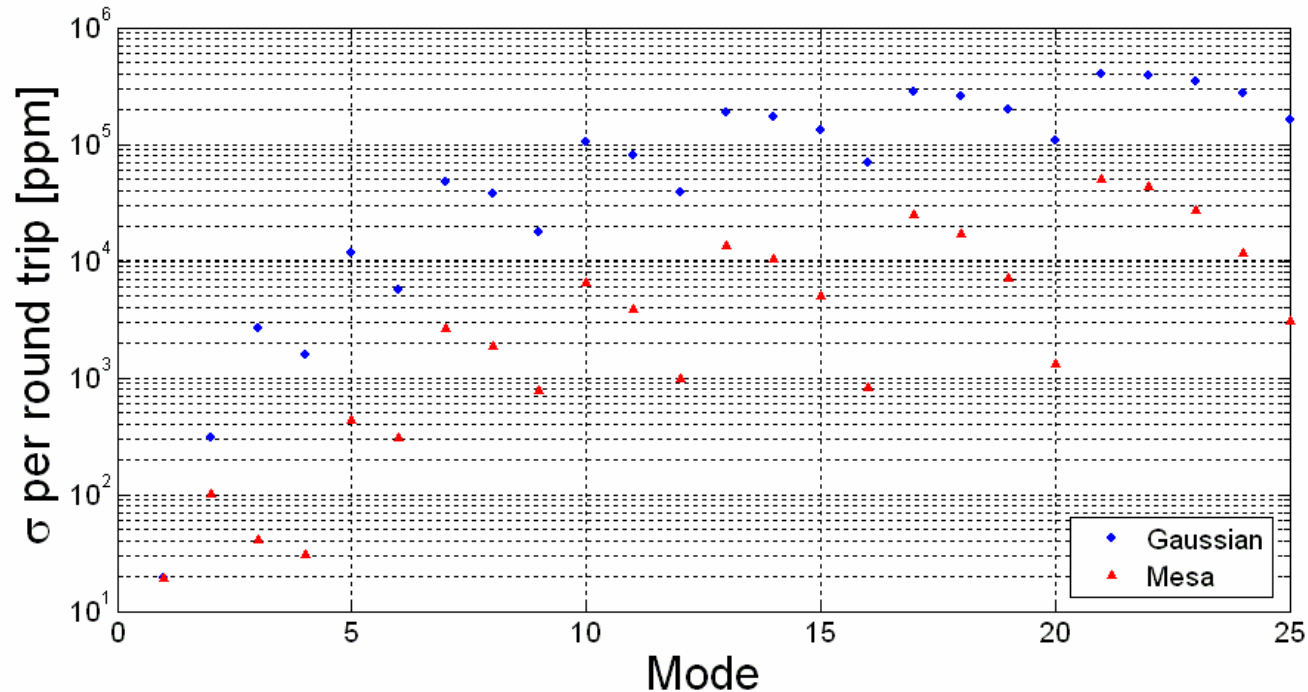
800 modes: 8.9 kHz – 114.6 kHz



Diffraction losses



$$\sigma_p = 1 - |\gamma_p|^2$$

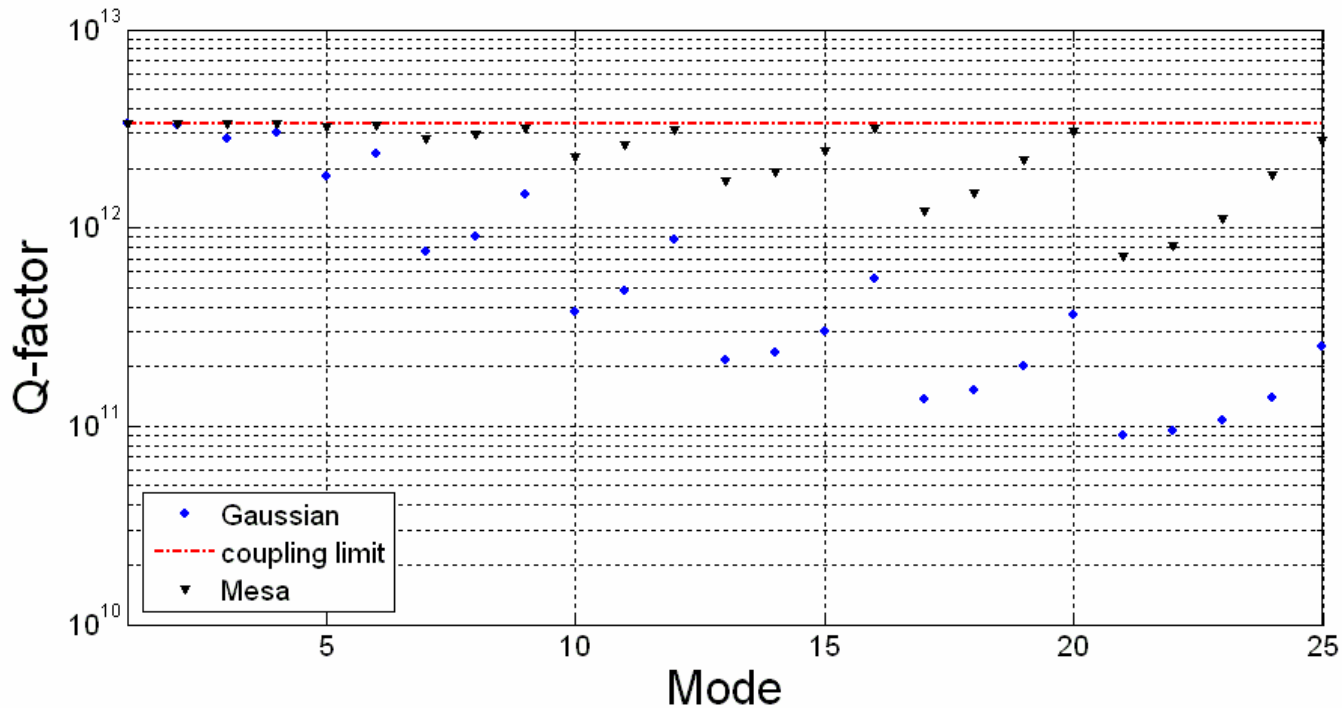




Optical modes Q-factors



$$Q_p = \frac{-2\pi\nu_p}{FSR \ln[(1 - \sigma_p)R_1R_2]}$$



Q-factors of the mesa TEMs are close to the coupling limit. Diffraction losses have substantially smaller effect on Q in comparison to the Gaussian cavity.



Overlapping parameter



Parametric gain R is also set by the overlapping parameter Λ . The spatial match between HOM and fundamental mode must be conserved.

$$\Lambda_{ij} = \frac{V \left(\int (\vec{E}^0 \circ \vec{E}^i) \mu^j_{\perp} d\vec{r}_{\perp} \right)^2}{\int |\vec{E}^0|^2 d\vec{r}_{\perp} \int |\vec{E}^i|^2 d\vec{r}_{\perp} \int |\vec{\mu}^j|^2 d\vec{r}}$$

- spatial match is defined as an integral $\int (\vec{E}^0 \circ \vec{E}^{HOM}) \mu_{\perp} d\vec{r}_{\perp}$ The higher the integral value the better is a match between modes.

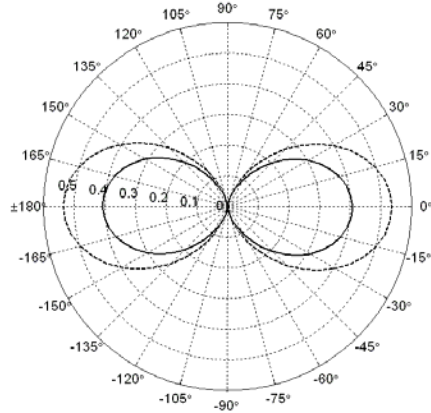
- $V / \int |\vec{\mu}|^2 d\vec{r}$ is a mass ratio of the mass of the test mass and the effective mass of the acoustic mode. The effective mass refers to the excitation susceptibility of an acoustic mode.



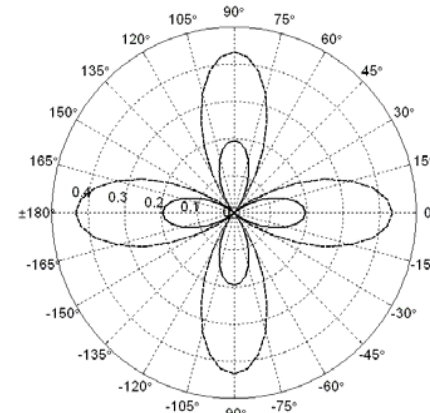
Optimal overlapping parameter



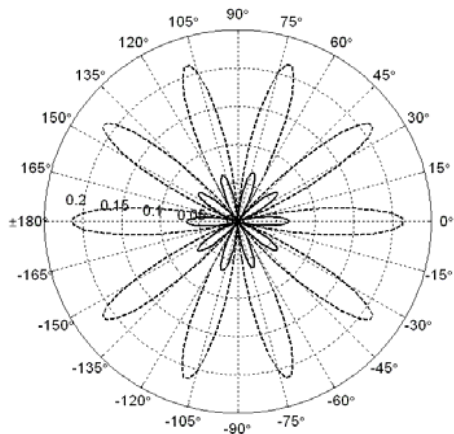
Λ_{MAX} → rotation of the acoustic mode on the optical mode



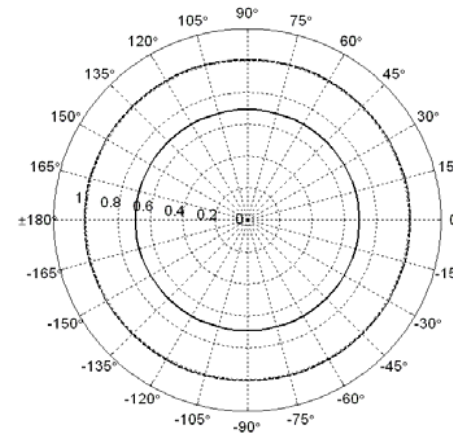
TEM 01



TEM 02



TEM 18



TEM 10



Overlapping parameter

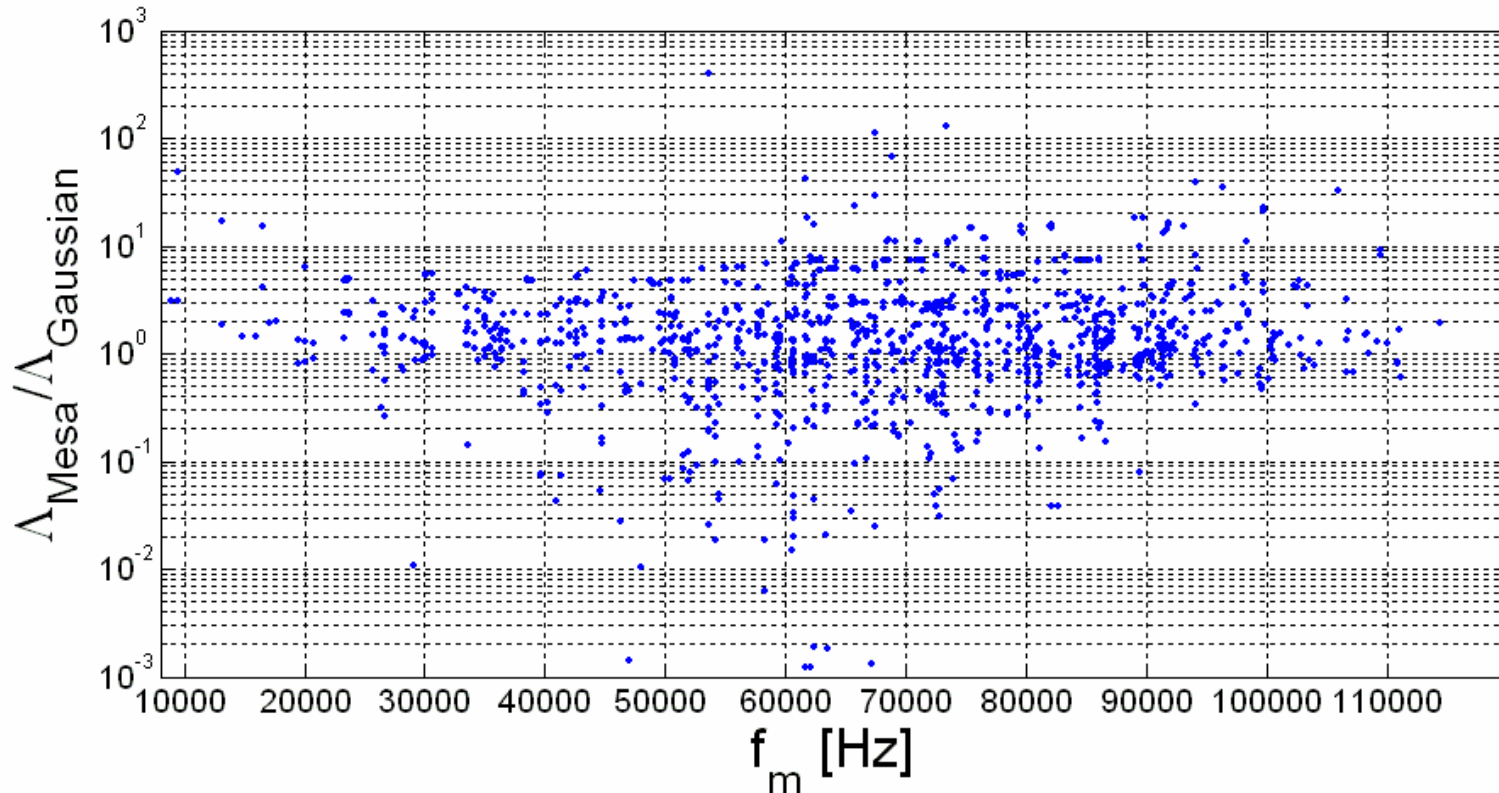


Λ selection criteria: $\Lambda \geq \omega_m^2 / g Q_m$, $g = \frac{4PQ_o}{mLc}$ (only opto-acoustic interactions with such Λ can be dangerous)

- $\Lambda_{\text{Mesa}} > \Lambda_{\text{Gaussian}}$: **772**

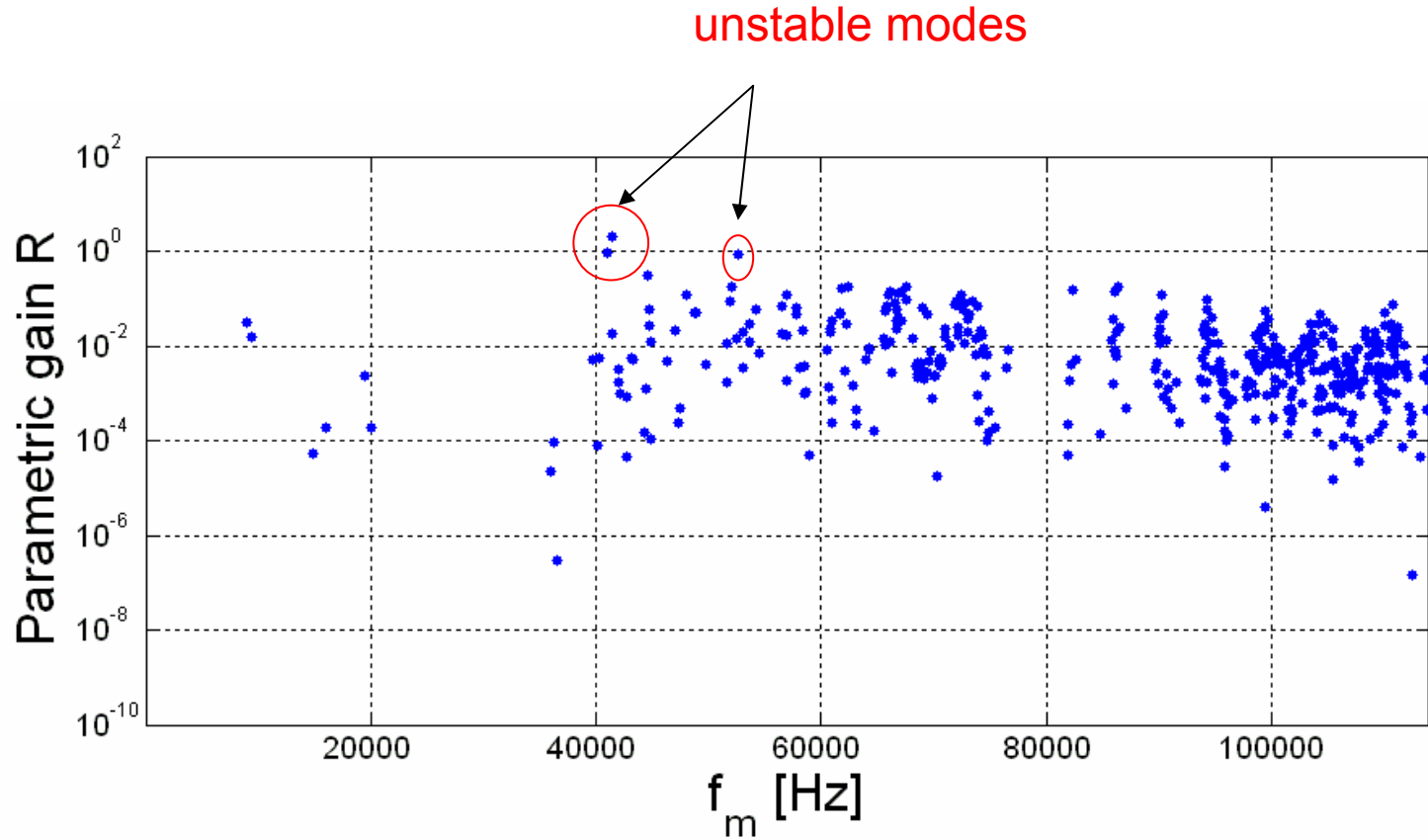
- $\Lambda_{\text{Gaussian}} > \Lambda_{\text{Mesa}}$: **404**

→ **~2 times more overlaps for mesa beam**



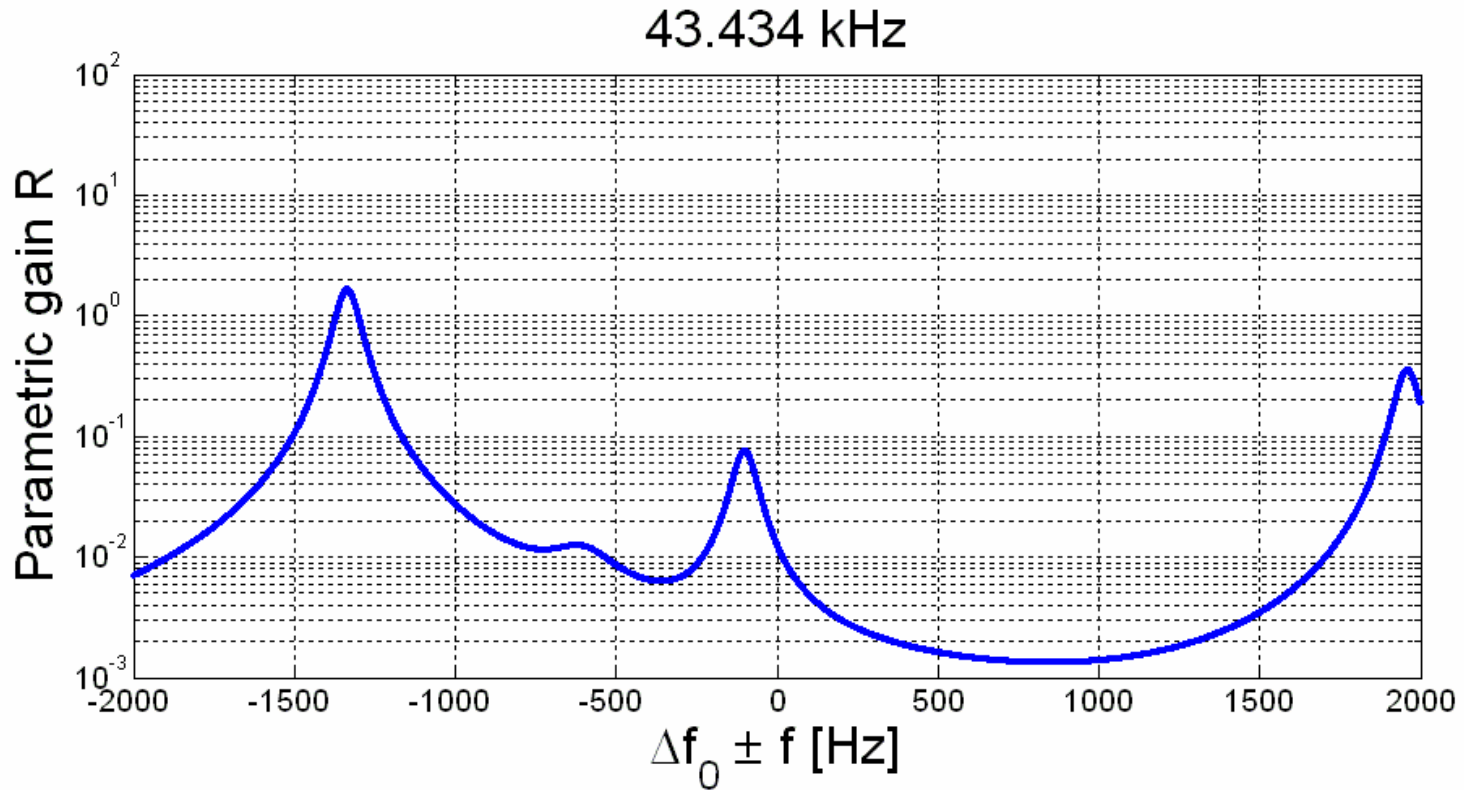


R-value for Gaussian cavity



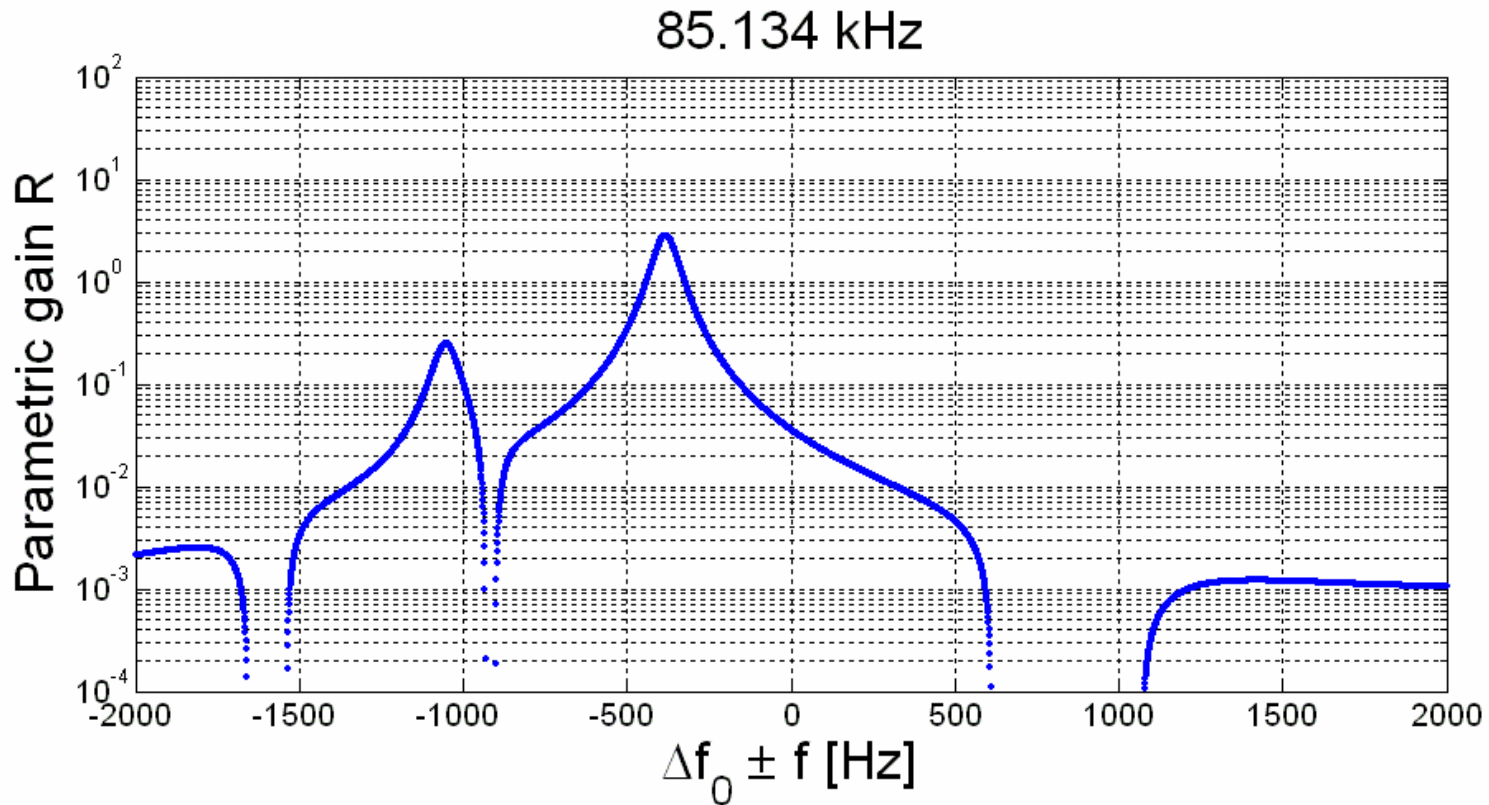


R-value vs. $\Delta\omega$ uncertainty



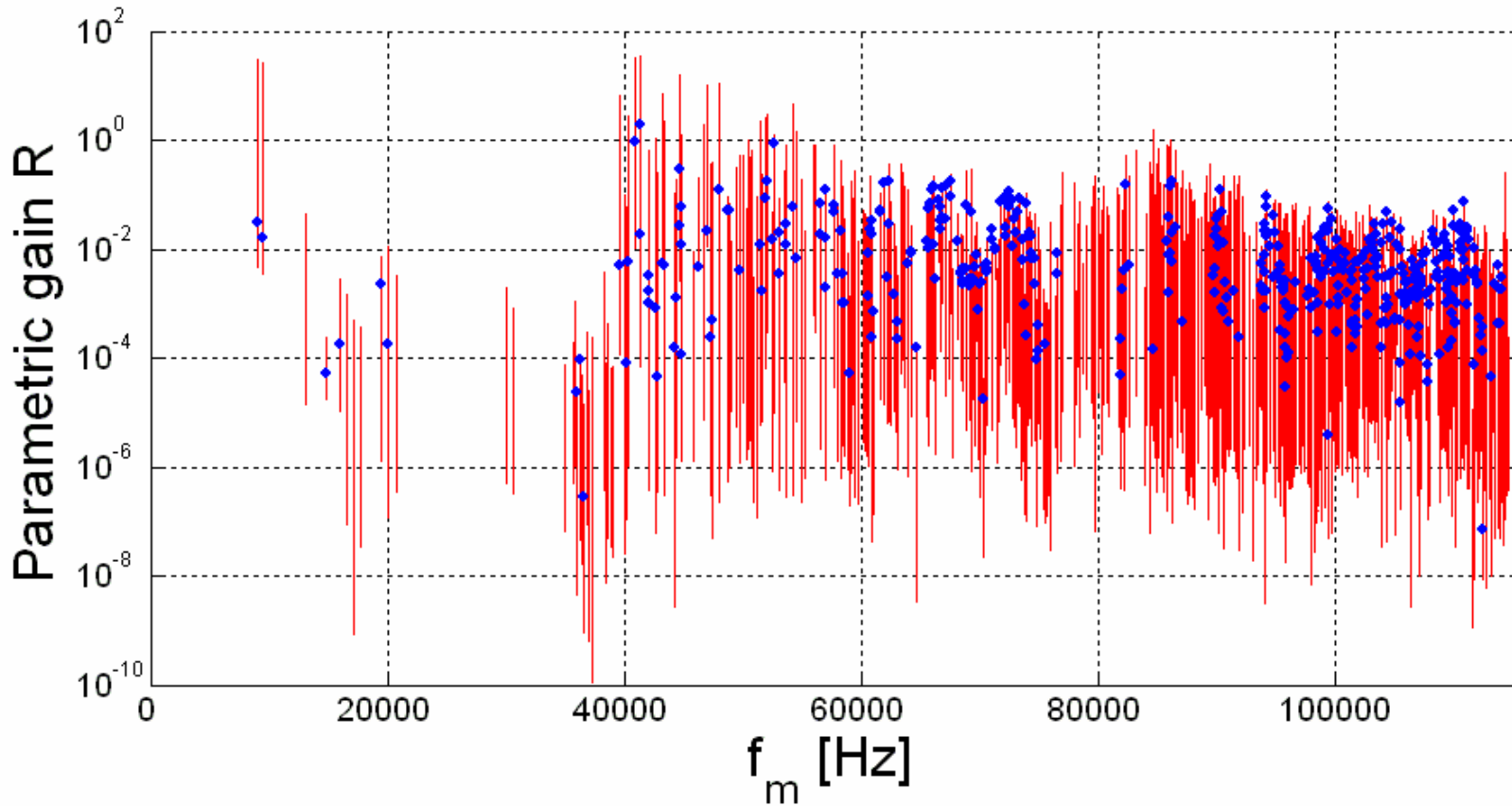


R-value vs. $\Delta\omega$ uncertainty





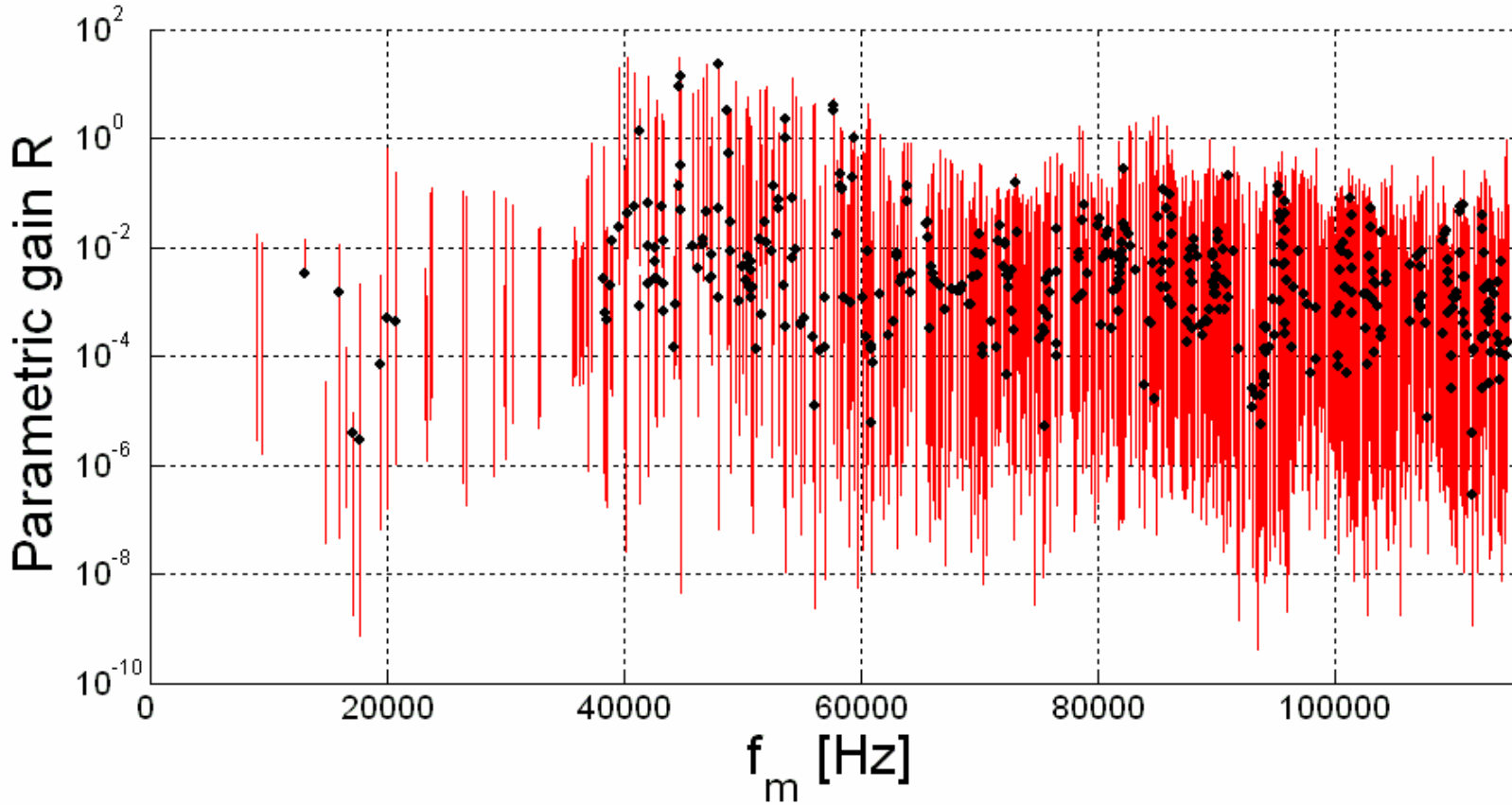
R-value for Gaussian cavity



possible unstable modes: 18



R-value for mesa cavity



possible unstable modes: 46



Conclusions



- Diffraction losses of the mesa HOMs lower than for the Gaussian HOMs by a factor of ~ 15 .
- HOMs Q-factor in the mesa cavity set by coupling losses.
- The mesa cavity has ~ 2 times more overlaps with value bigger than for Gaussian cavity.
- The mesa cavity is ~ 2.6 times more susceptible to parametric instability than the Gaussian cavity.



Modal misalignment suppression for Adv. LIGO



Parametric Instability



$$R = \frac{2PQ_m \Lambda \omega_s}{McL \omega_m^2} \operatorname{Re} \left[\frac{1}{(1 + \chi^2)} \left(\frac{-1}{i\Delta\omega + \lambda_1} + \frac{-\chi^2}{i\Delta\omega + \lambda_2} \right) \right]$$

Optical mode frequency:

$$\omega_s = \frac{1}{2} (\omega_{s1} + \omega_{s2})$$

Arm tuning parameter:

$$d = \frac{1}{2} (\omega_{s1} - \omega_{s2})$$

IFO configurations:

- 1) IFO with PRM and symmetrical arms
- 2) IFO with PRM and asymmetrical arms
- 3) IFO with PRM, SRM and symmetrical arms
- 4) IFO with PRM, SRM and asymmetrical arms

Model limitations: PRC unstable, no diffraction losses allowed



PARAMETERS



$T_{ITM} = 1.4\%$
 $T_{ETM} = 5\text{ppm}$
 $T_{PRM} = 2.5\%$

$SRM = 20\%$
 $\delta = 11 \text{ deg}$
(see T070247-01)

Wavelength = $1.064\text{e-}06 \text{ m}$

Arm length = 3994.75 m

Mirror radius = 0.17 m (coated $0.16.8 \text{ m}$)

Optical modes: up to order 12

Total number of optical modes = 60

Estimated numerically

Axial mode order = 5

Total number of acoustic modes = 5507

(range: $5.8 \text{ kHz} - 140 \text{ kHz}$)

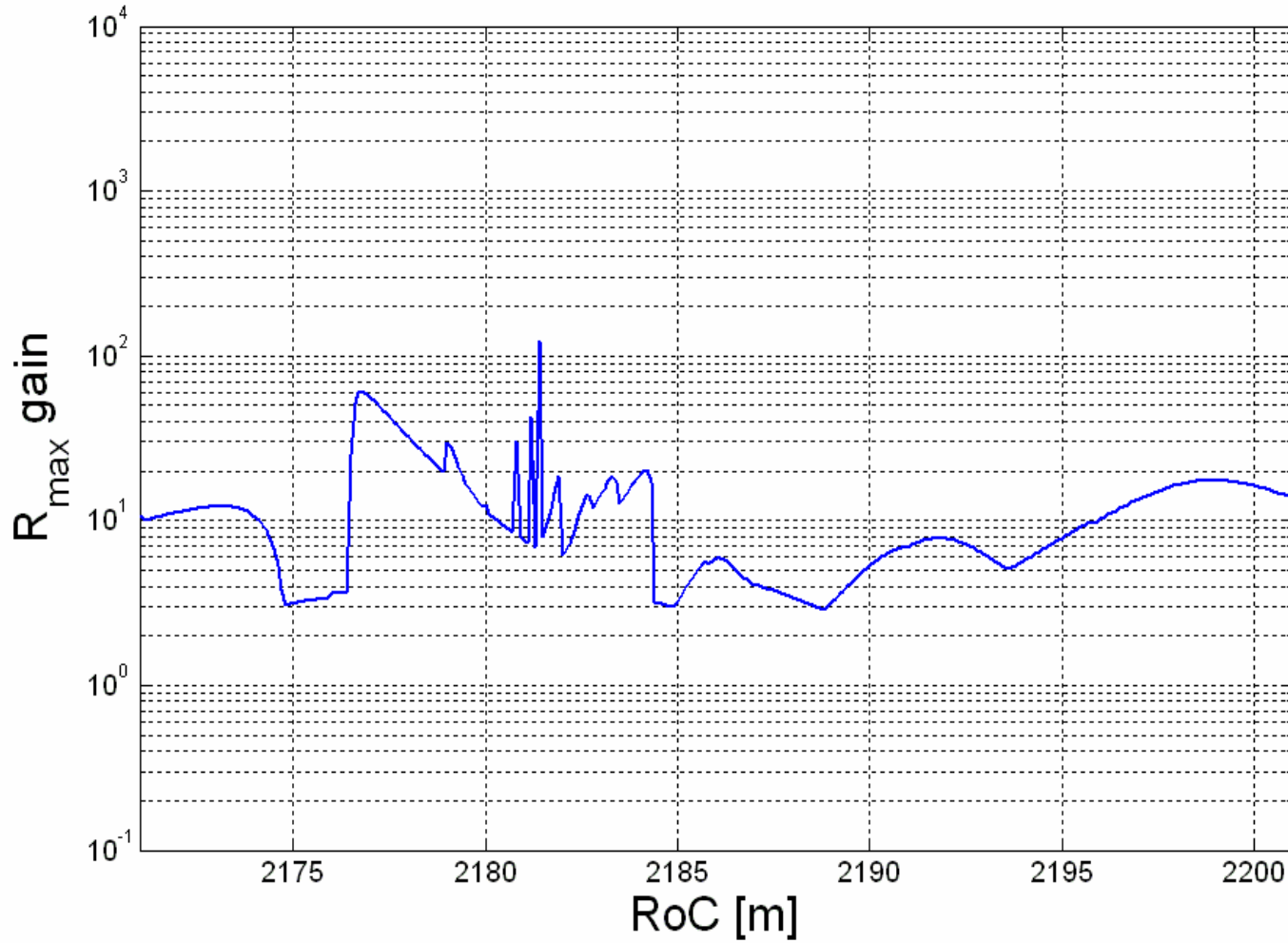
RoC step: ***0.1m***

$RoC_{ITM} = 1971 \text{ m}$

$RoC_{ETM} = 2191 \text{ m}$

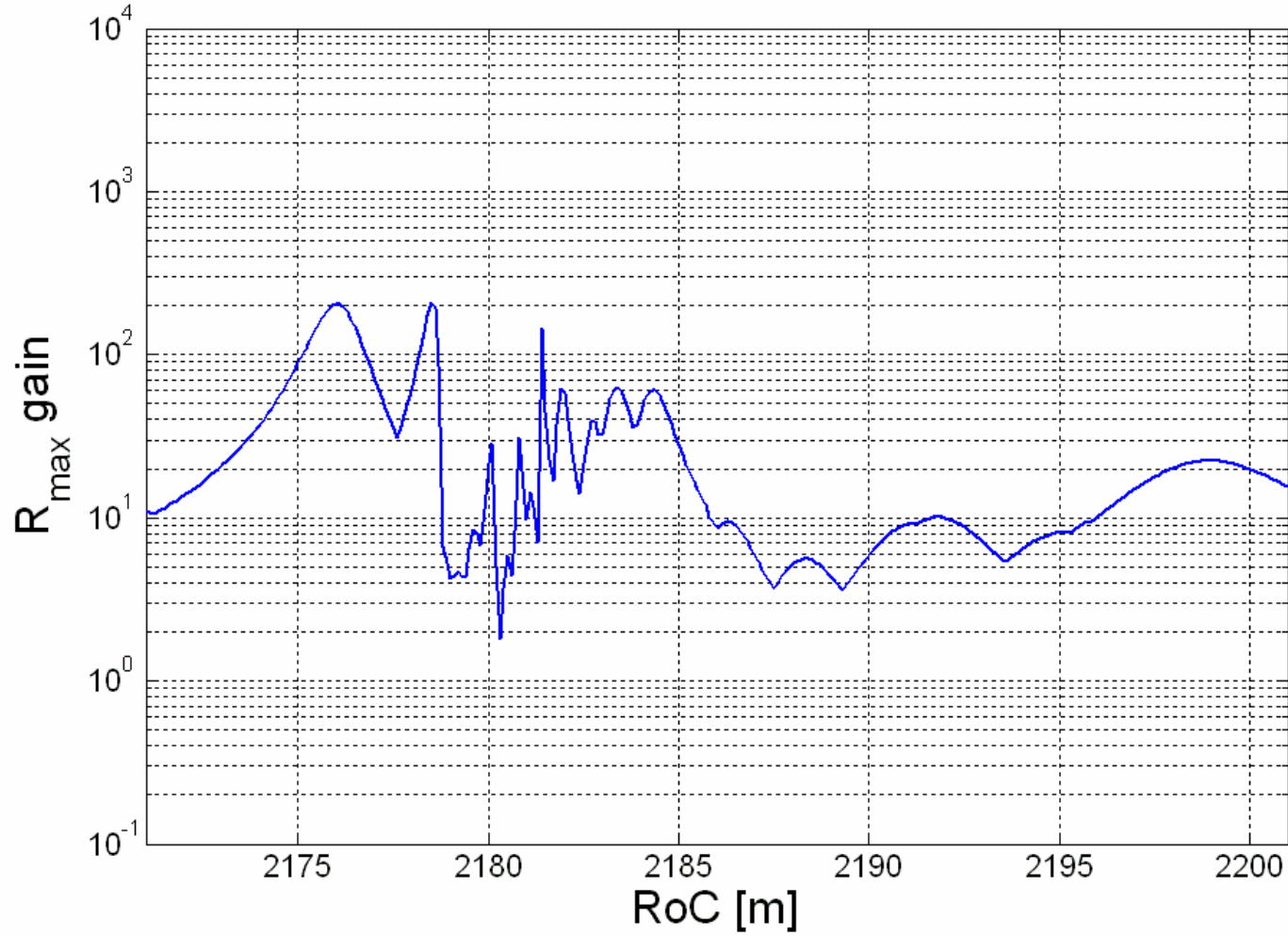


IFO with PRM



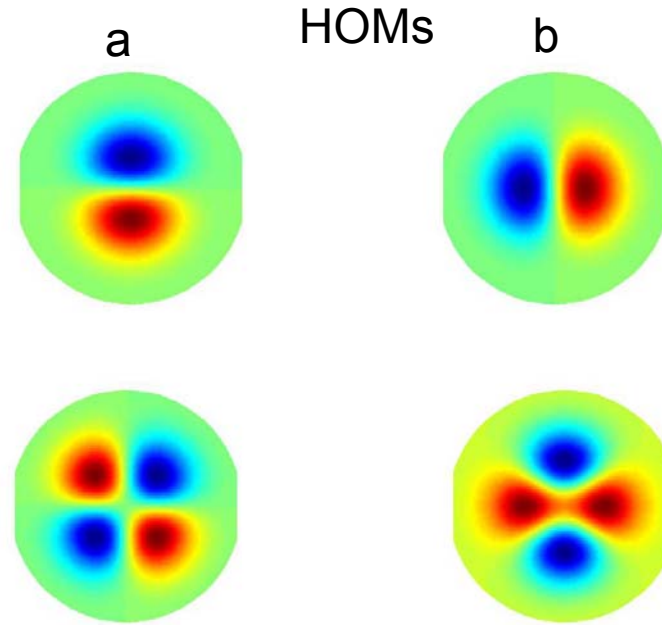
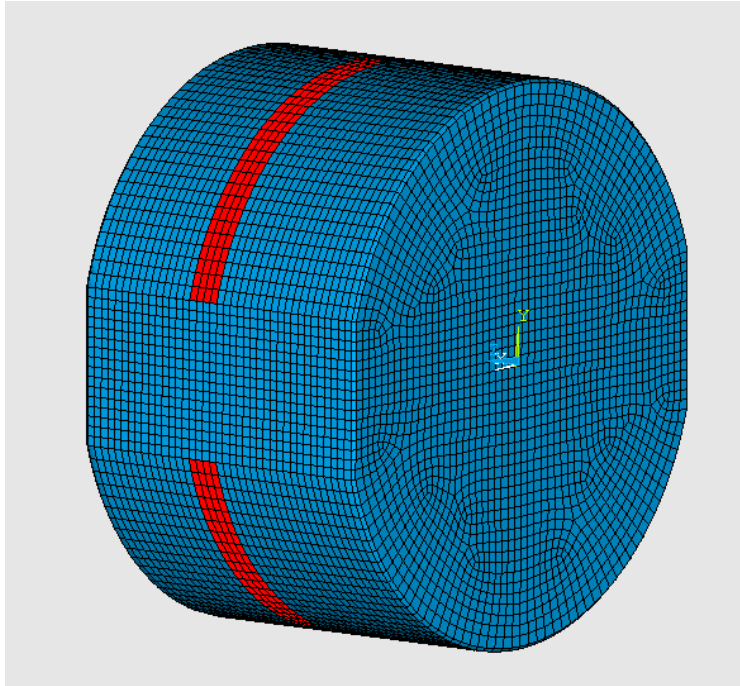


IFO with PRM and SRM



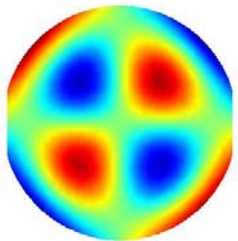


Optical mode orientation

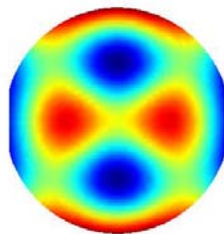


Elastic modes

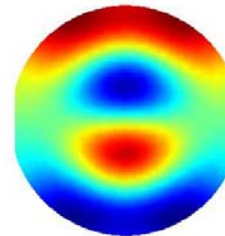
19287 Hz



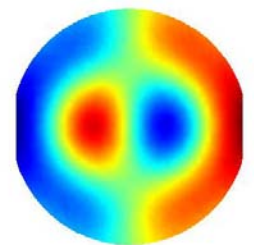
19476 Hz



20438 Hz

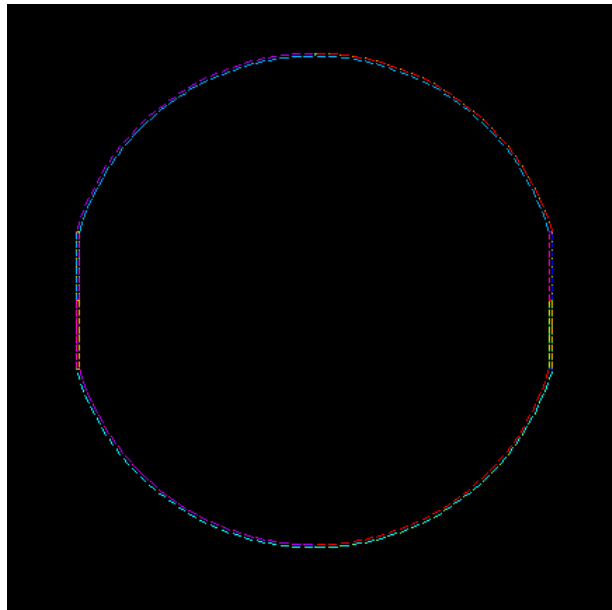


20514 Hz

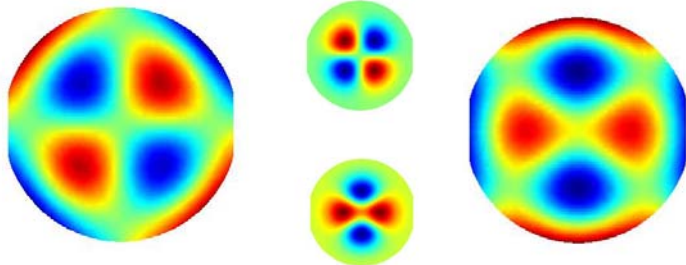
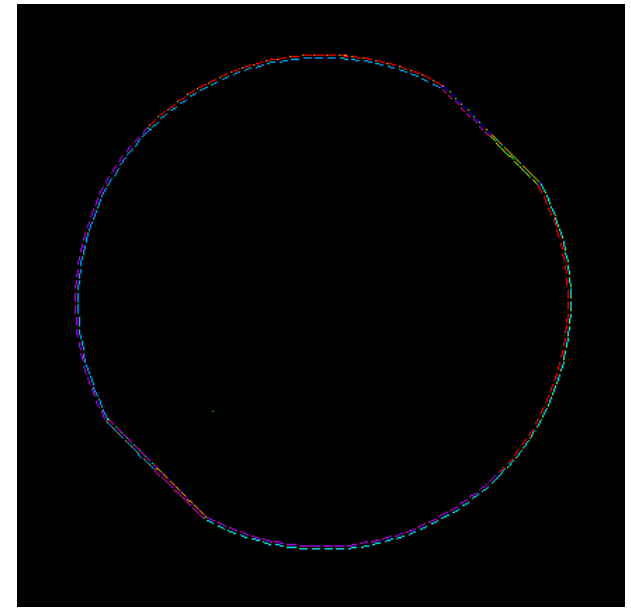
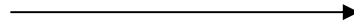




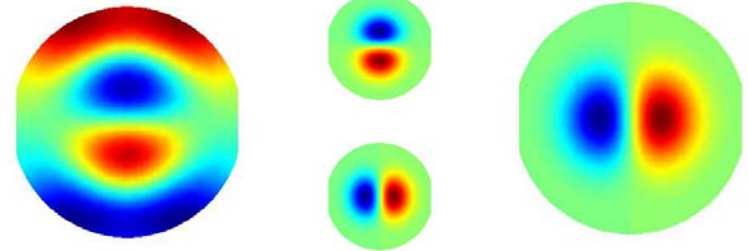
Coating rotation



45°



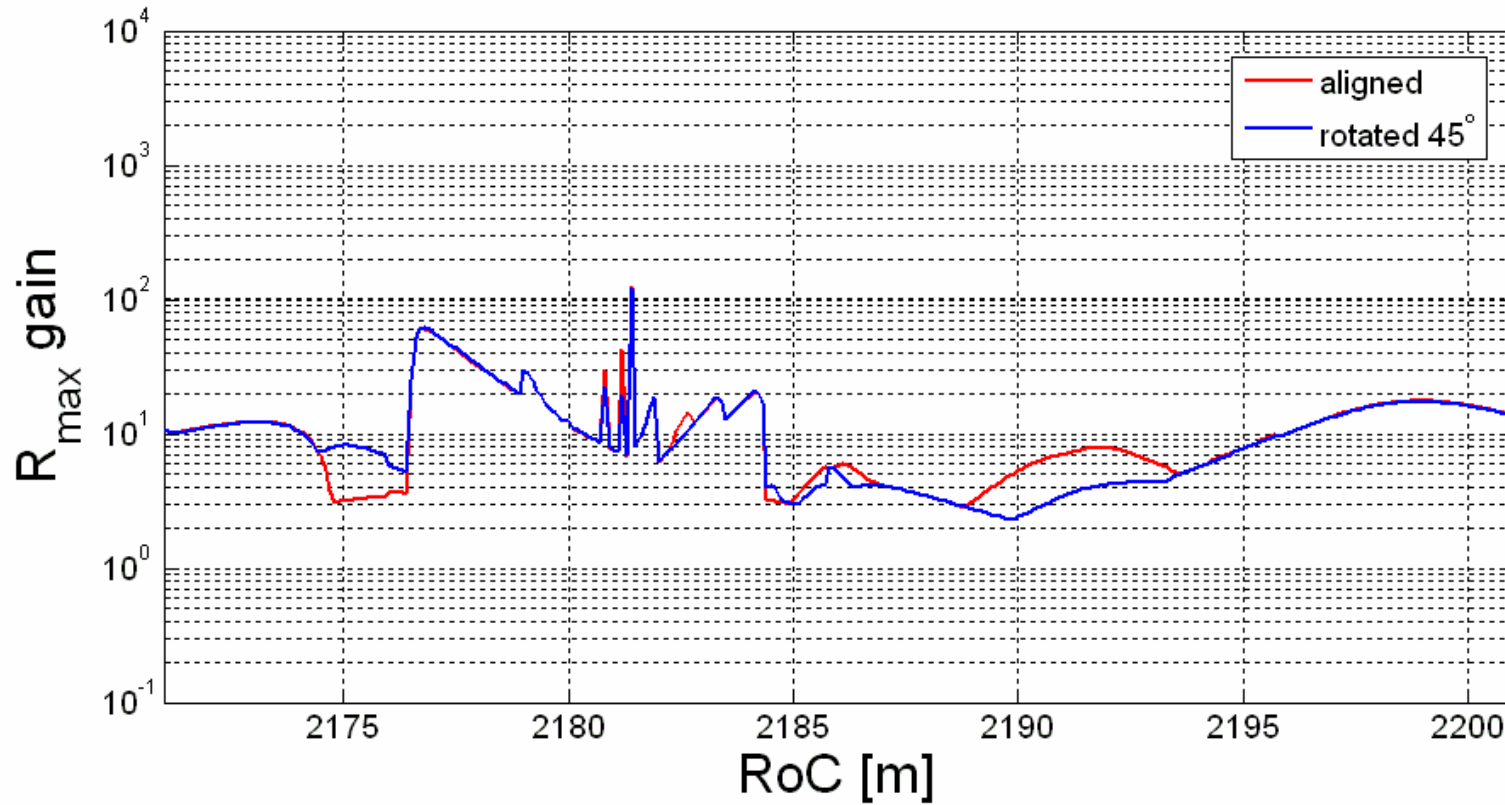
does not work



work $\rightarrow \Lambda$ reduced

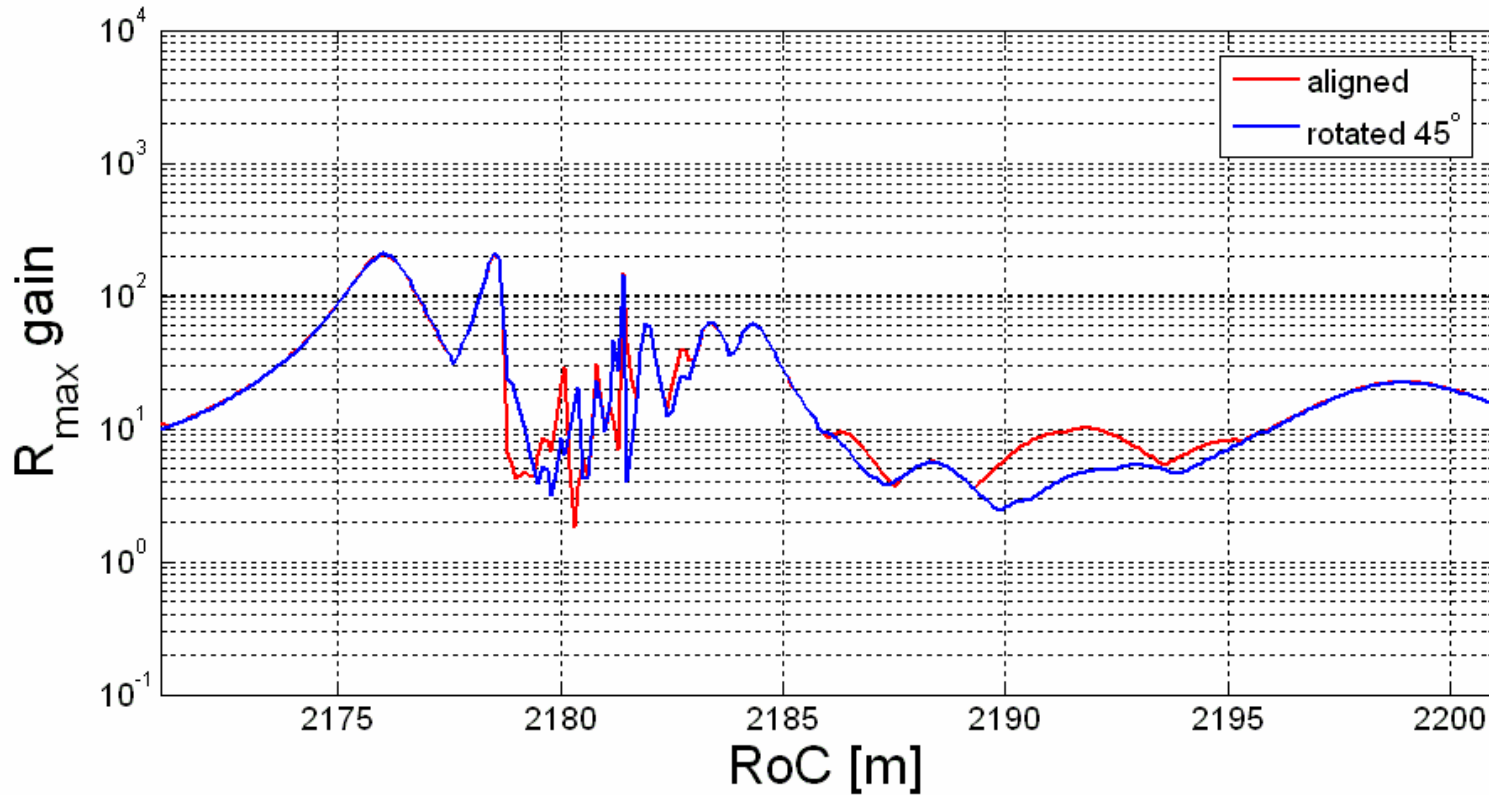


Modal misalignment –IFO with PRM





Modal misalignment – IFO with PRM and SRM





Conclusions



- Arm asymmetry substantially lower parametric gain, $R \sim 10-20$ in large RoC range.
- Modal misalignment has some effect on reduction of R- value.
- Further R reduction expected from diffraction losses (only for HOMs above 3rd order), unfortunately we do not have a mathematical model for it.