



UNIVERSITY OF
BIRMINGHAM

Advanced LIGO UK Production BOSEM Status and Interferometric OSEM Development

Stuart Aston
University of Birmingham

LSC-Virgo Meeting

Amsterdam (Netherlands)
20th - 26th September 2008

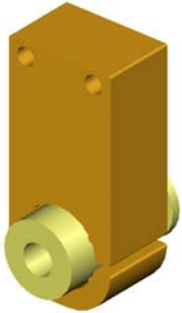
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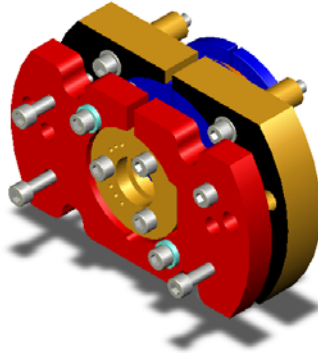


- **BOSEMs** (*S. Aston, R. Cutler, D. Lodhia, A. Vecchio*)
 - Configuration
 - Quantities
 - Characterization
 - Noise Prototype Status
 - Full Production Status
- **Interferometer** (*C. Speake, S. Aston, F. Peña Arellano, D. Hoyland*)
 - Design Motivation
 - Fringe Interpolation Method
 - Design Realisation
 - Prototype Fabrication
 - Optical Development
 - Electronics Development
 - Characterization
- **Summary**

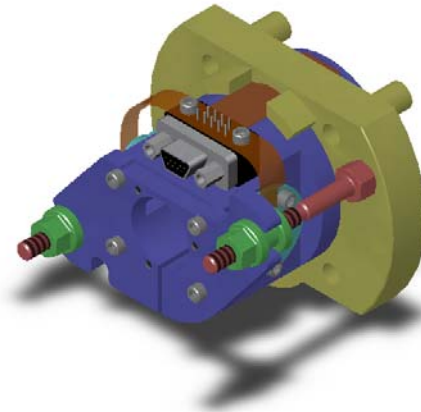




Initial LIGO
(OSEM)



Advanced LIGO
Controls Prototype (Hybrid OSEM)



Advanced LIGO
Noise Prototype & Final
Production (BOSEM)

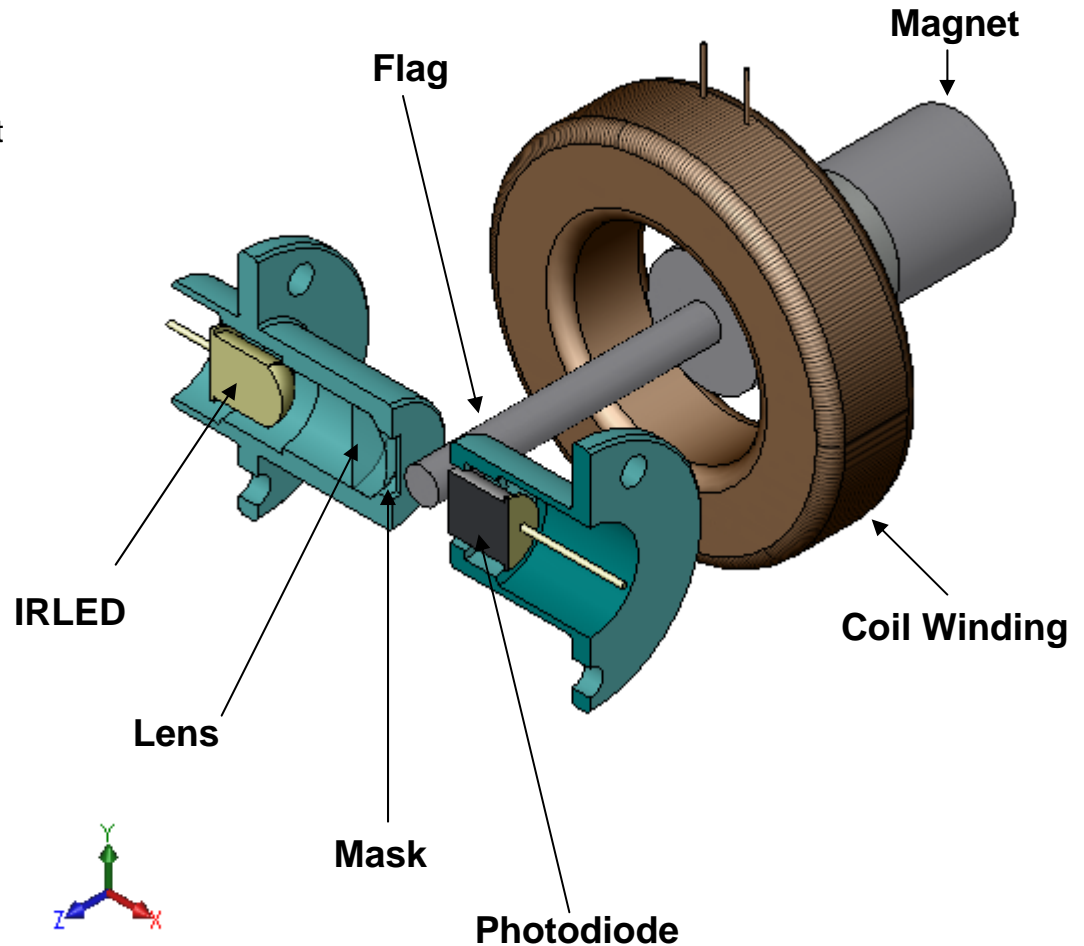
1st Generation

2nd Generation



- Key sensor and actuator components (see right):-
- Significant changes from 1st generation Initial LIGO OSEM:-

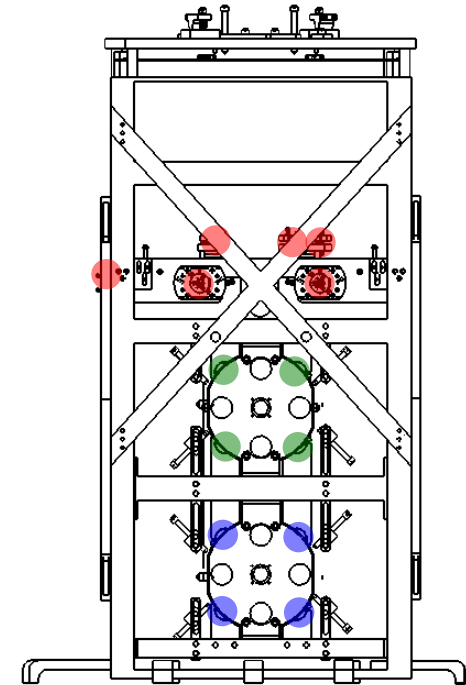
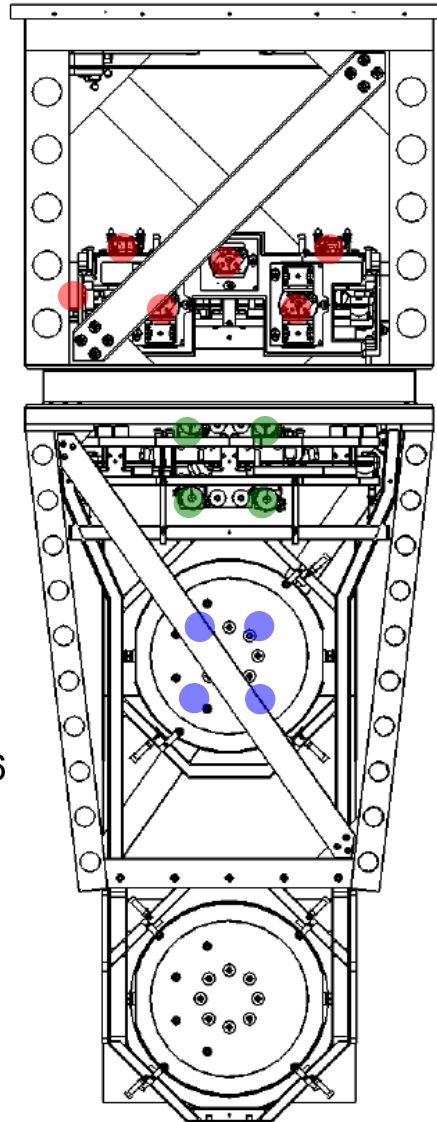
- Stronger actuator force
50mN -> 500mN
 - Higher coil current
 - More coil windings
 - Larger magnets
- No Ceramabond used in fabrication process
- Standard leaded device packages (i.e. not surface mount)
- Commercial-of-the-shelf connectors (sub-D to μ -D)
- Flexi-circuit for interconnections





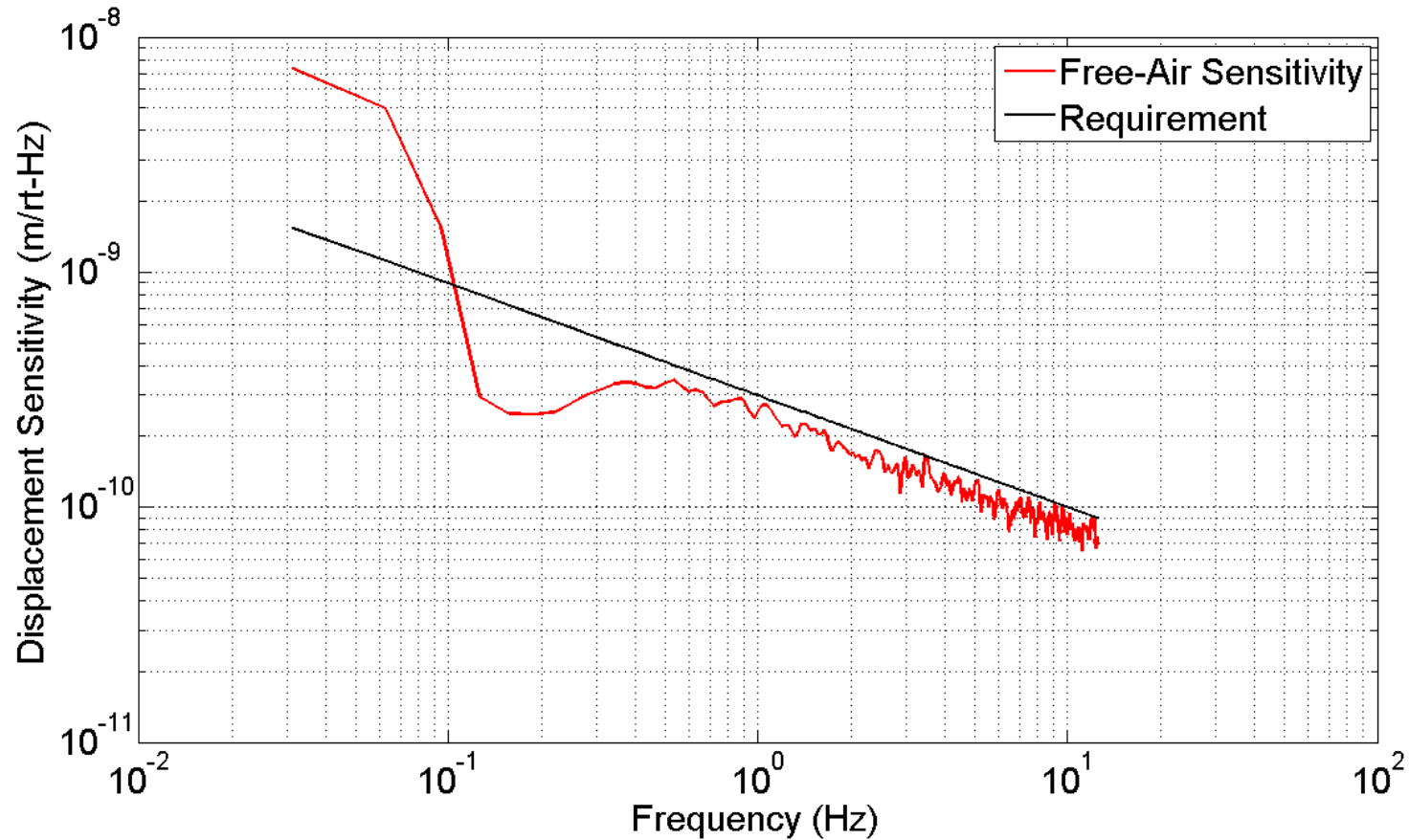
BOSEM:- Quantities

- BOSEM counts (not including spares):-
- Quad suspensions:-
 - End Test Mass (ETM) ~20
 - Input Test Mass (ITM) ~20
- Triple suspensions:-
 - Input Mode Cleaner (IMC) ~14
 - Recycling Mirror (RM) ~14
 - Folding Mirror (FM) ~10
 - Beam Splitter (BS) ~10
- Double suspensions:-
 - Output Mode Cleaner (OMC) ~6
- Single suspensions:-
 - Steering Mirror (Tip/Tilt) ~4
- Delivery Quantities:-
 - Noise Prototypes = 79 OSEMs
 - Advanced LIGO = 654 OSEMs





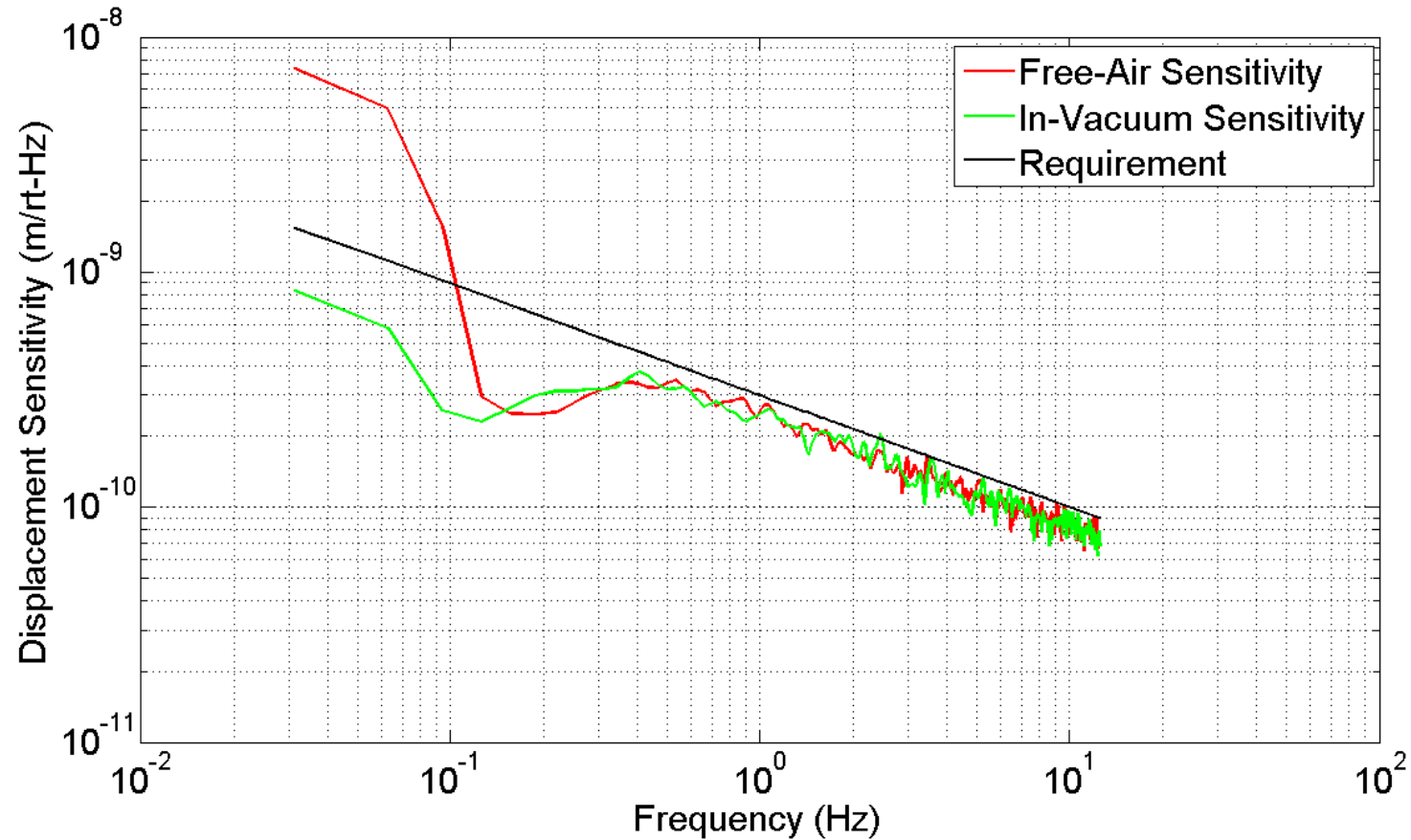
- Requirements^[1]:-
 - Sensitivity @ 1Hz = $3 \times 10^{-10} \text{m}/\sqrt{\text{Hz}}$ and @ 10Hz = $1 \times 10^{-10} \text{m}/\sqrt{\text{Hz}}$
 - Operating Range 0.35mm (peak-peak)



^[1] K.A. Strain "Input to the OSEM selection review decision". LIGO-T040110-01-K



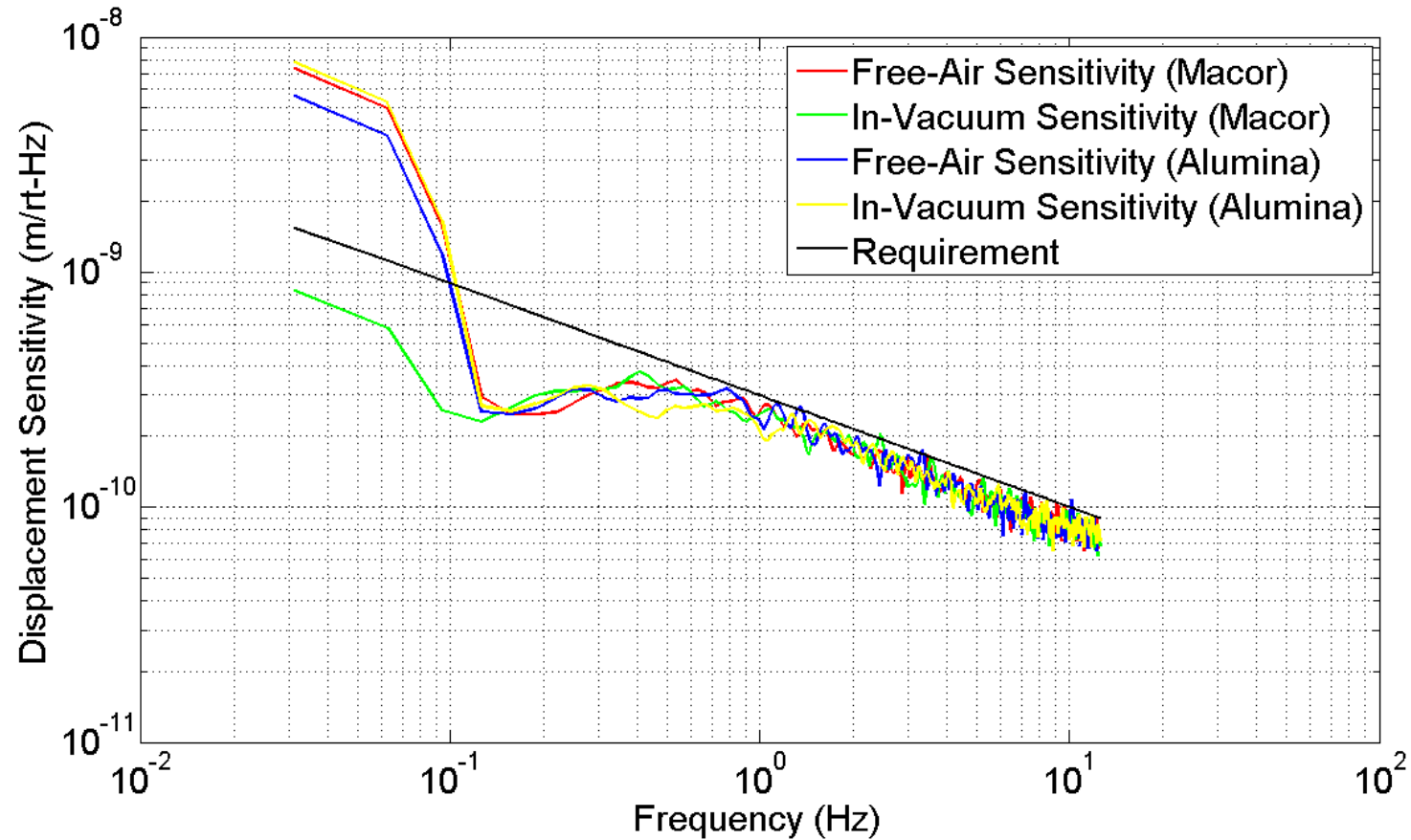
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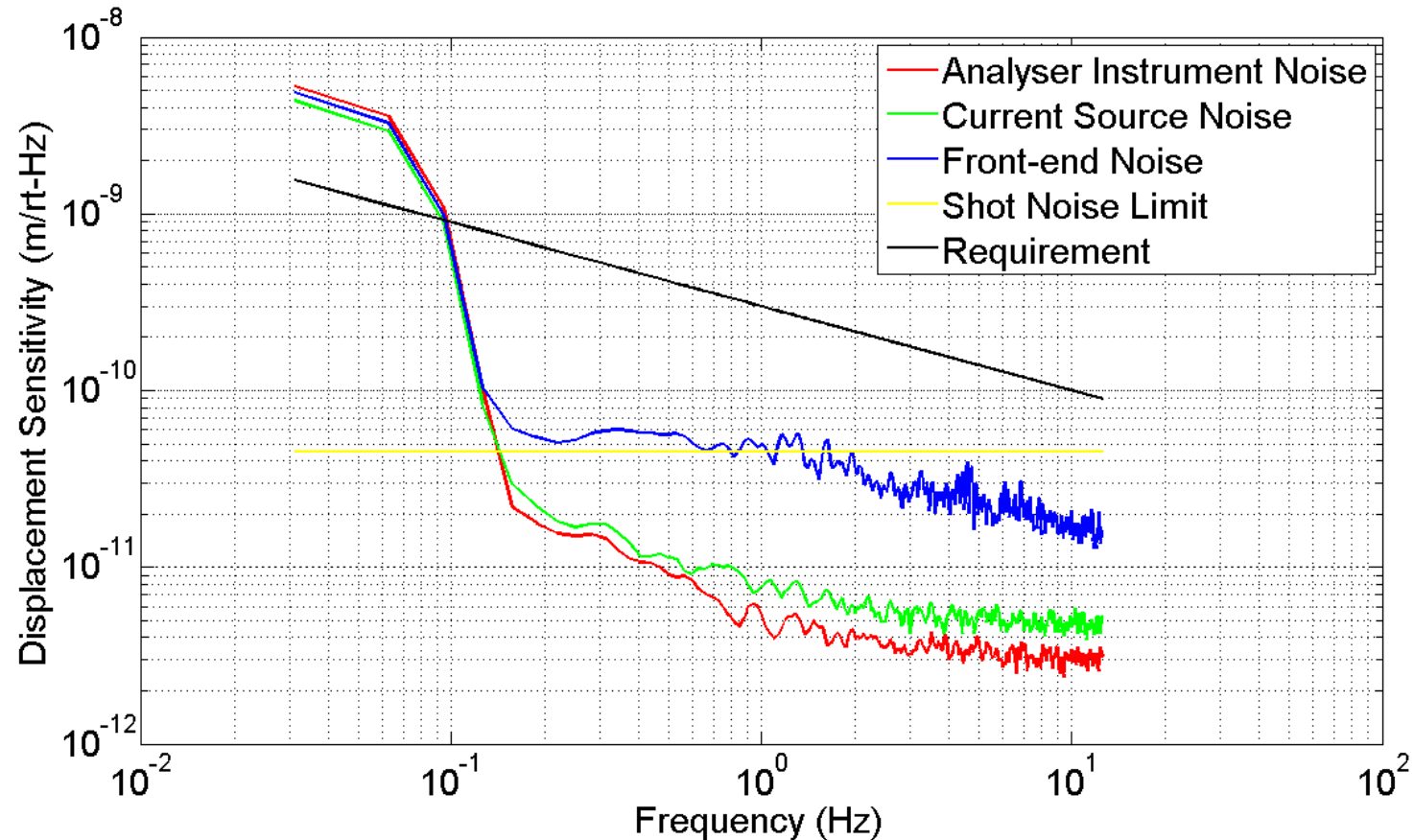


- In search of a factor of 3 improvement @ 1Hz:-
 - Switched Macor Sleeve (1.46 W/mK) with Alumina Sleeve (35 W/mK)
 - Applied thermal compound at key interfaces (no observable improvement)





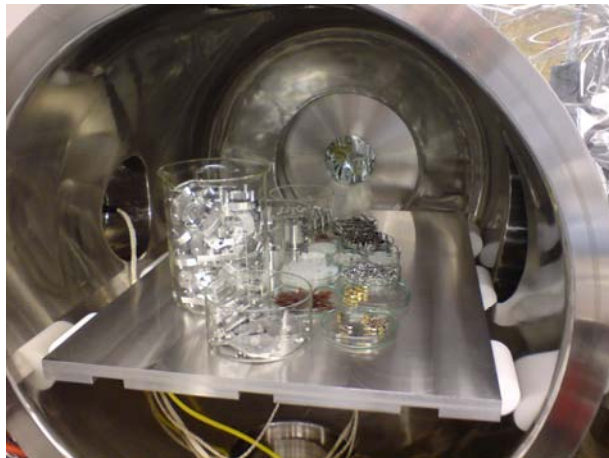
- Sanity check of electronics noise sources:-
 - Do not appear to be restricting us (n.b. these have been end-to-end measurements)
 - Further investigation of noise sources is ongoing (plus other stabs-in-the-dark)





BOSEM:- Noise Prototype Status

- Noise Prototype deliveries are complete! 😊 (Spring 08)
 - Comprised of:-
 - 79 units in total delivered to USA
 - Units were assembled, tested and shipped as LIGO UHV compliant parts
 - Cleaning and baking carried out in Birmingham facilities
 - LIGO UHV qualification (e.g. RGA scans) undertaken at Caltech
 - Current Status:-
 - Units employed in Quad at LASTI, OMC, and Tip/Tilt suspensions
 - Operational experience gained has been fed back into final production BOSEM



Pre-assembly bake out



Dedicated clean-room assembly suite



Assembled BOSEMS at testing station



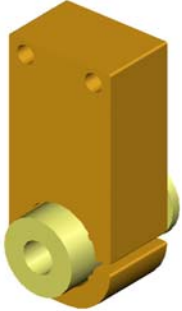
- Full production of 654 BOSEMs (inc spares and Noise Prototypes) underway:-
 - Rough Breakdown:-
 - ~6 months parts procurement and fabrication (March 08 – August 08)
 - ~6 months assembly and testing (September 08 – February 09)
 - Key Milestones:-
 - 1st delivery, 28th November 2008 (training exercise) ~24 Units
 - Final delivery, 1st June 2009



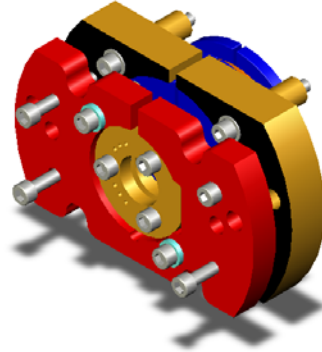
Received opto-electronic parts



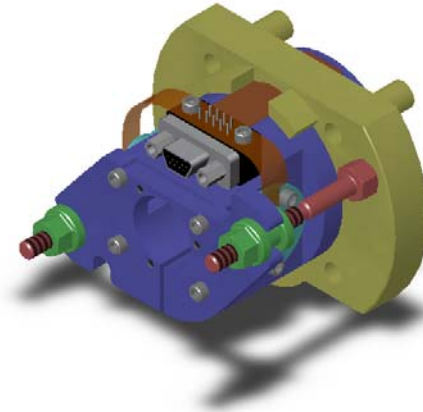
Mechanical parts awaiting inspection



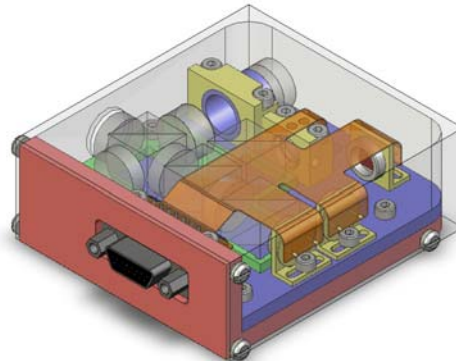
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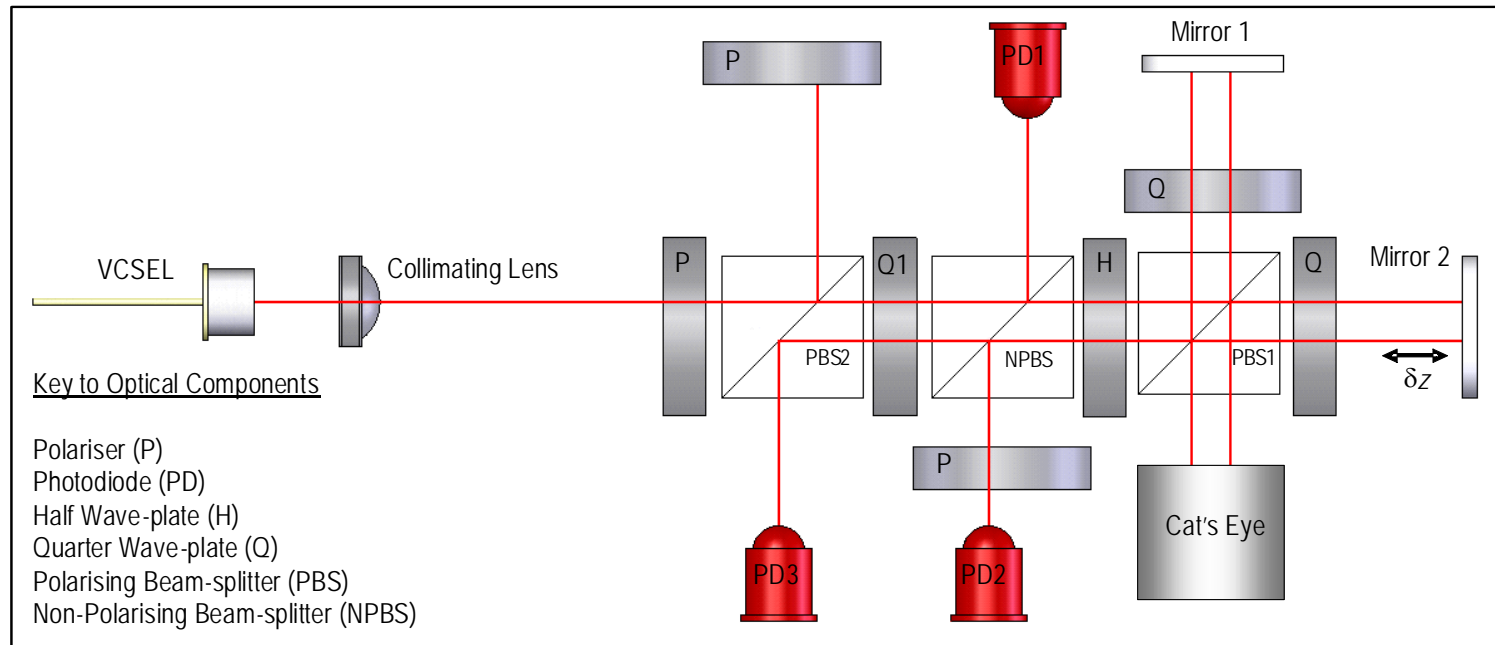


Advanced / Ultra LIGO
(Interferometric Sensor (EUCLID))



Interferometer:- Design Motivation

- OSEMs offer the best you can hope for with a shadow sensor configuration
- To ensure good low frequency stability we needed to avoid active parts that can age, thermally expand, generate heat, exhibit hysteresis, e.g. piezos, AOMs, EOMs etc. This naturally led to a Homodyne Interferometer
- Required to be, compact / portable, and robust against misalignment. This has led us to develop a compact interferometric sensor (EUCLID)

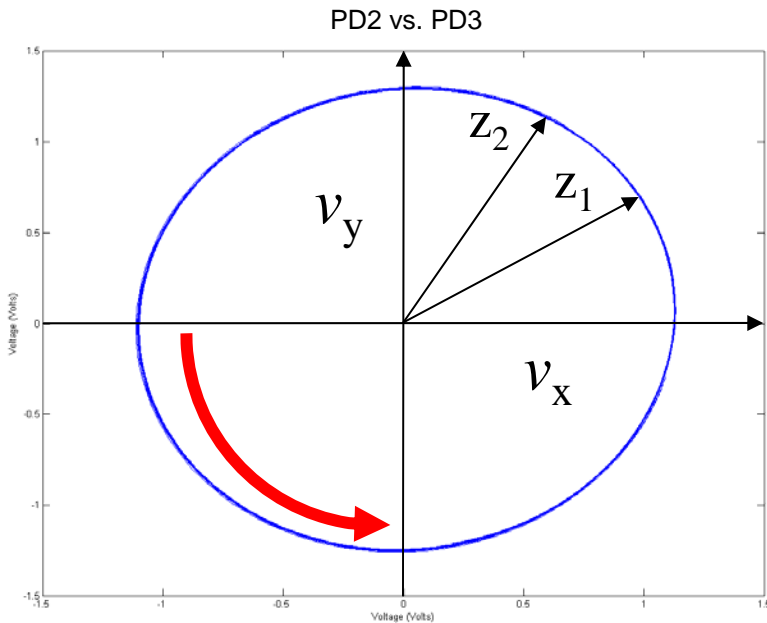


Optical Layout [2]

[2] C. C. Speake and S. M. Aston "An interferometric sensor for satellite drag-free control". IOP, Class. Quantum Grav. 22 (2005)



- Two fringe intensities I_2, I_3 are 90° out of phase (PD2 and PD3)
- Target mirror motion (Mirror 2) generates a circular Lissajous figure with I_2, I_3 plotted as v_x, v_y



- Phase calculation:-

$$\Delta\phi = \arctan\left(\frac{\begin{pmatrix} v_{y_2} & -v_{y_0} \end{pmatrix}}{\begin{pmatrix} v_{x_2} & -v_{x_0} \end{pmatrix}}\right) - \arctan\left(\frac{\begin{pmatrix} v_{y_1} & -v_{y_0} \end{pmatrix}}{\begin{pmatrix} v_{x_1} & -v_{x_0} \end{pmatrix}}\right)$$

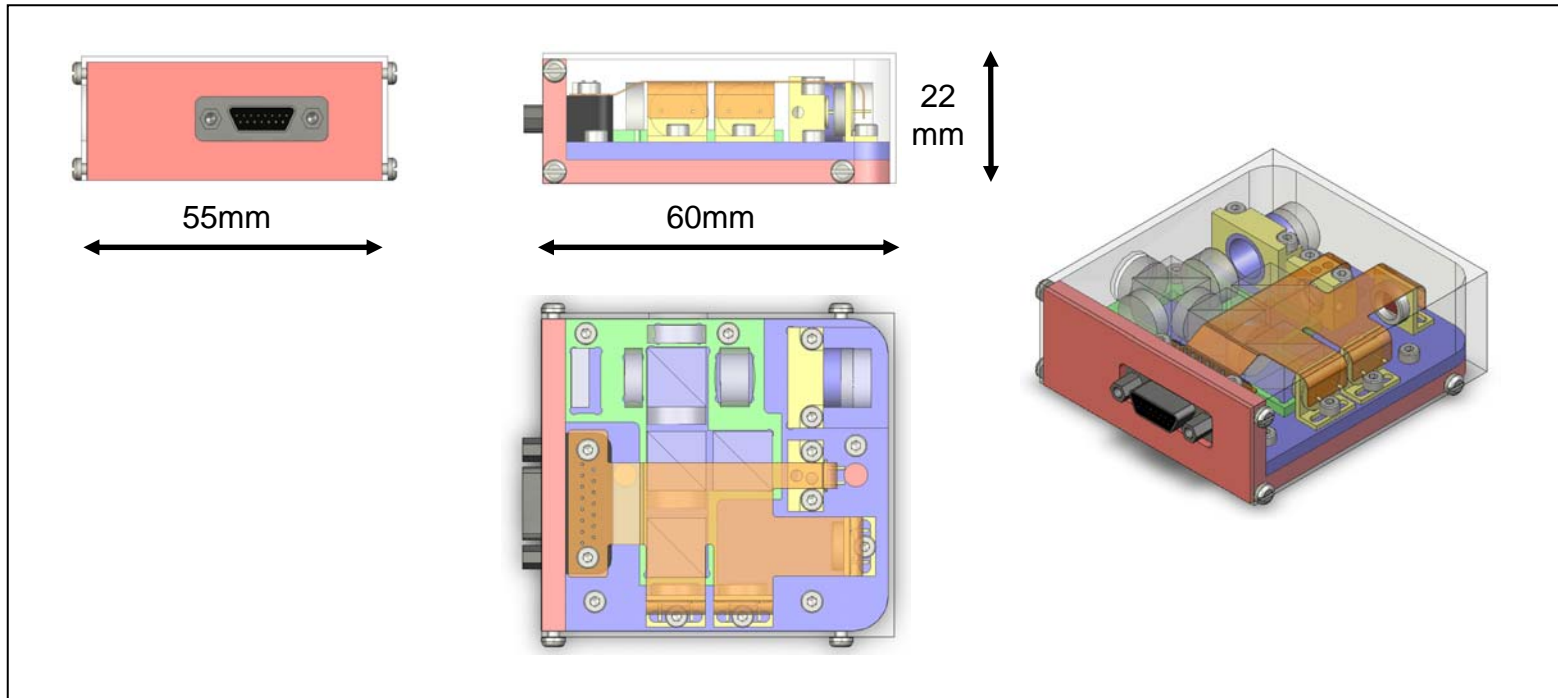
- Displacement calculation:-

$$\Delta z = \frac{\Delta\phi}{2\pi} \cdot \frac{\lambda}{4}$$



Interferometer:- Design Realisation

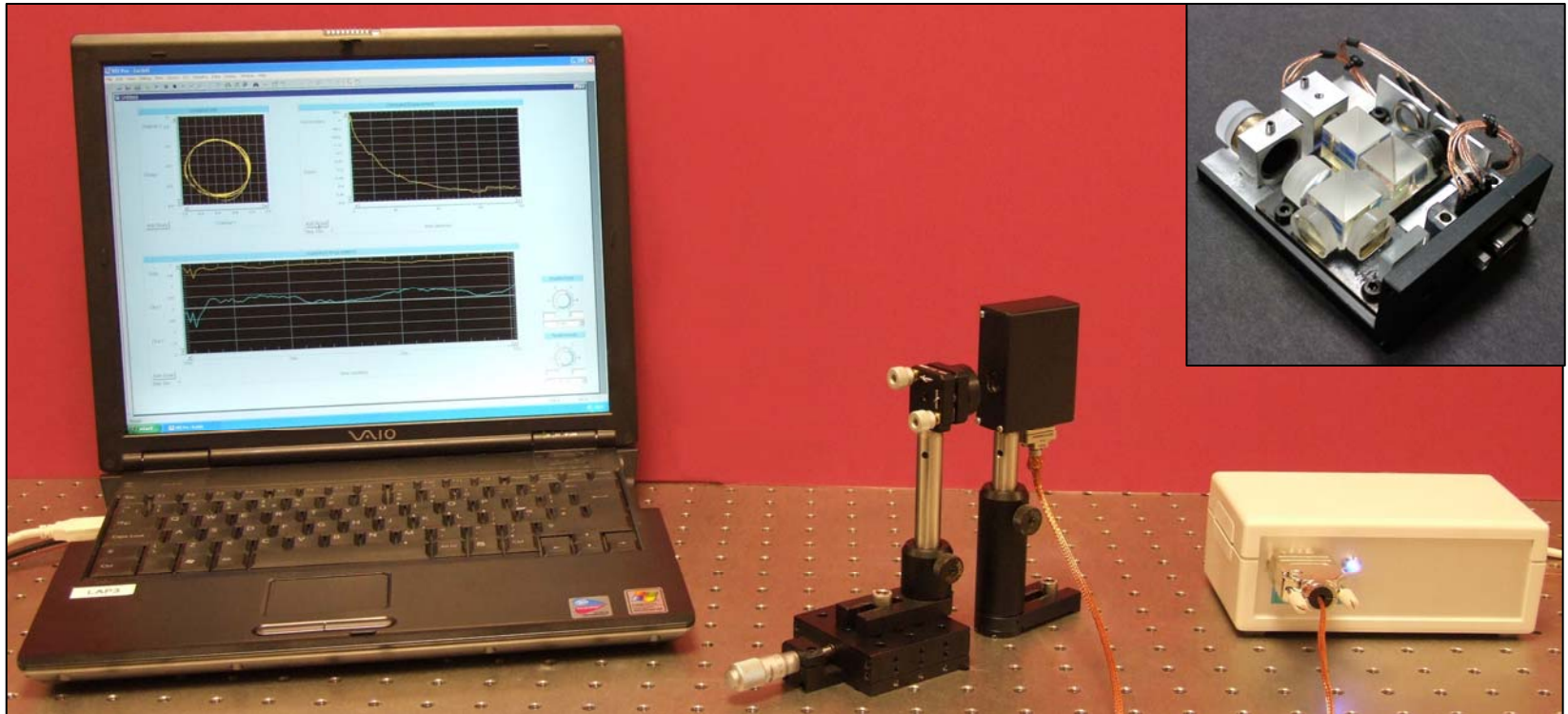
- Compact dimensions of 60mm x 55mm x 22mm
- Robust against misalignment +/- 1°
- Resolution of 1 pm/ $\sqrt{\text{Hz}}$ over a large working range > 2mm
- Can be constructed to be LIGO UHV compliant
- Incorporates VCSEL with known high MTBF



3D CAD Model



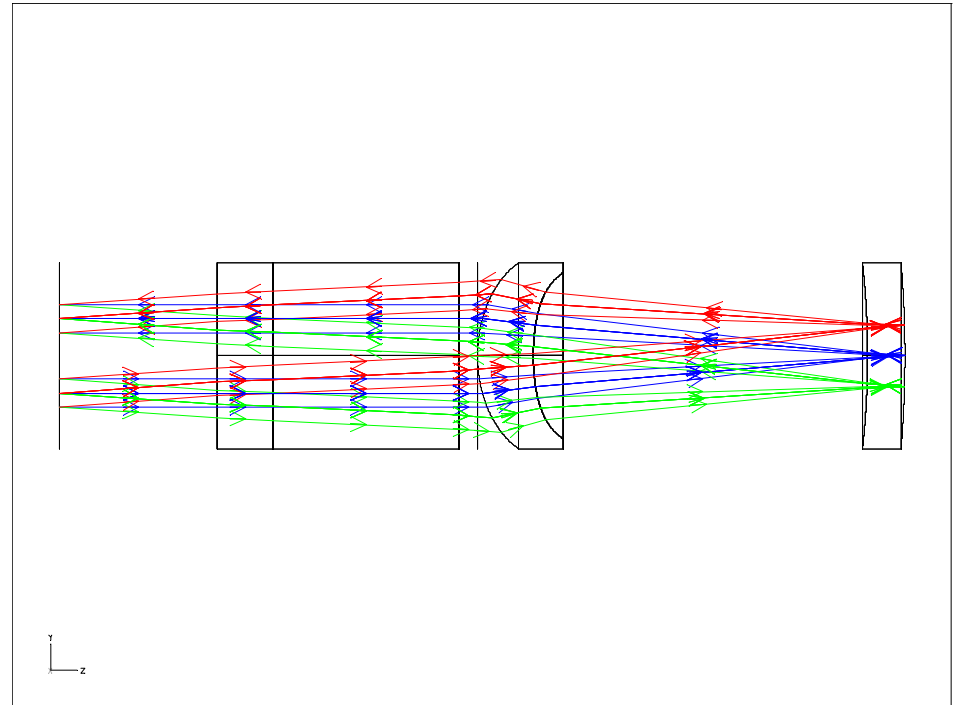
- We have designed and fabricated a bench-top prototype for evaluation
 - Also Includes:-
 - Data processing software
 - Support electronics
- Characterization is ongoing



Prototype and support equipment. (Inset: with cover removed)



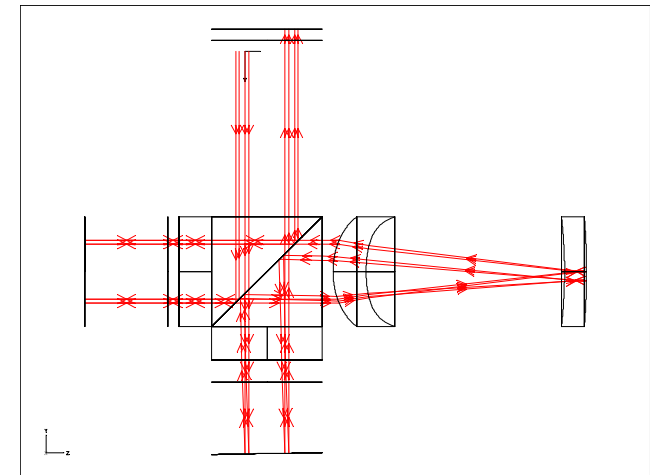
- Optimisation of the lens shape with Zemax
 - Fabián Peña Arellano
- The effect of aberrations on the wavefront is minimised
- The parameters are also adjusted using other configurations in order to achieve the best mirror tilt immunity possible



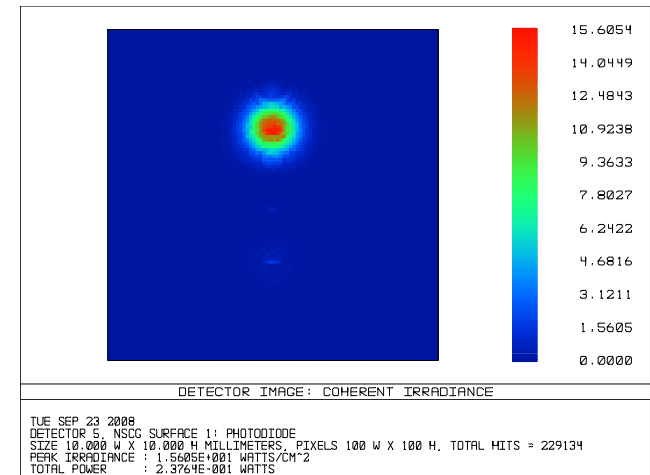
OPTO CAD Model



- Calculation of the fringe visibility
 - *Fabián Peña Arellano*
- Achromat. Preliminary quantitative agreement
 - For 1° mirror tilt:-
 - ± 1 mm displacement range
 - 50% change in visibility
- Aspheric. Preliminary prediction
 - For 1° mirror tilt:-
 - ± 5.7 mm displacement range
 - 50% change in visibility



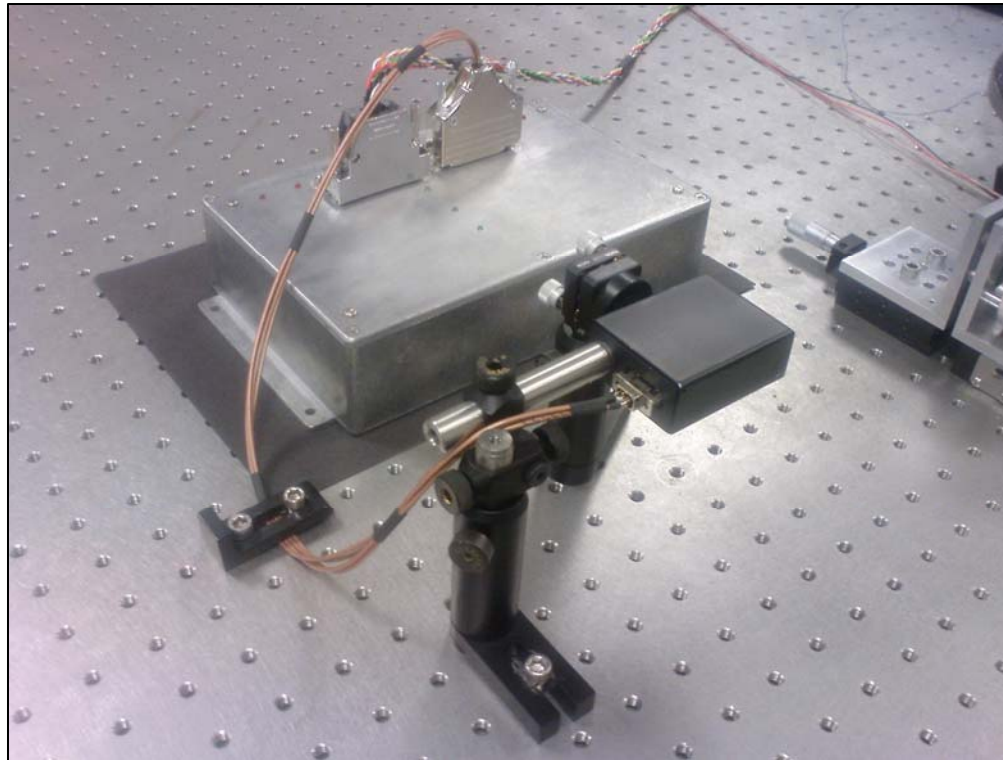
OPTO CAD Model



OPTO CAD PLOT



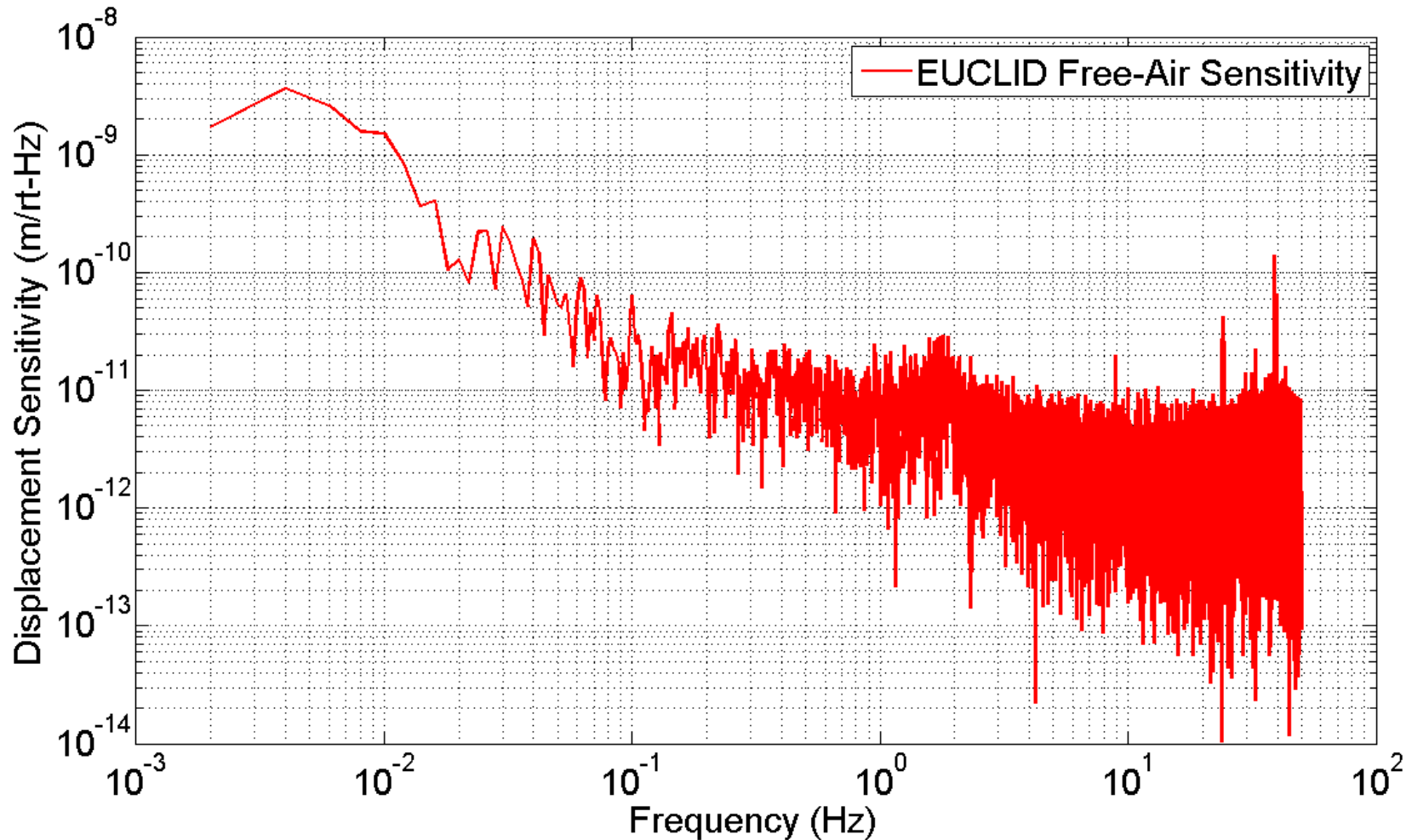
- Design and fabrication of support electronics:-
 - *David Hoyland*
- Modified BOSEM Satellite Box:-
 - Utilises single channel (of 4 available) current source
 - Utilises 3 channels (of 4 available) of trans-impedance amplifiers



Support Electronics and Interferometer Prototype

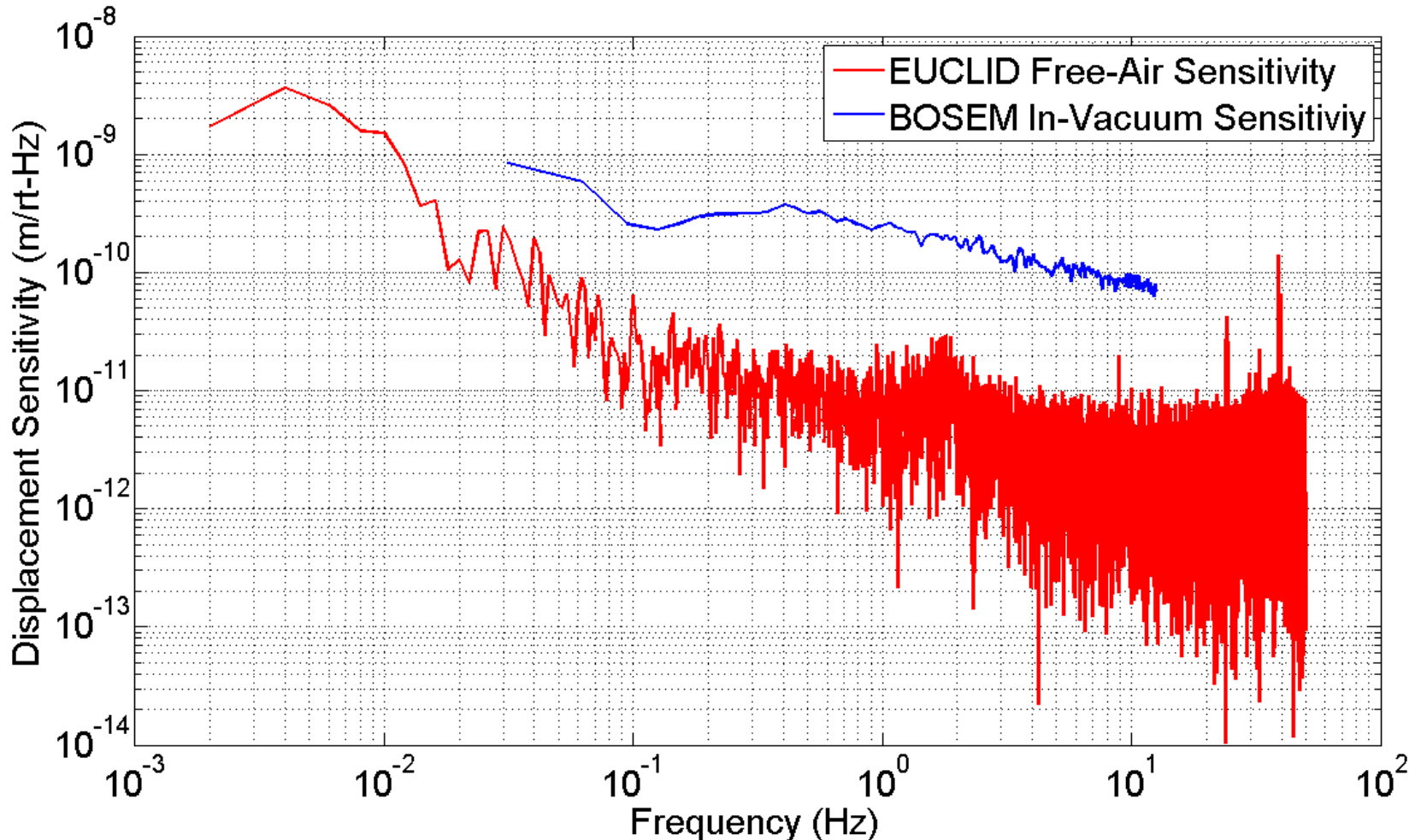


- Results from end-to-end test driving EUCLID with Satellite Box





- Comparison between Interferometric OSEM and BOSEM performance
 - n.b. we plan to go in-vacuum and try temperature controlled VCSEL diodes





- Continue with our task to deliver and support in-vacuum BOSEMs and Harnesses for Advanced LIGO
- Deliver and support electronics coil drivers for Advanced LIGO
- Build upon the experience gained with BOSEMs and apply to interferometer development. For example:-
 - Transfer knowledge from manufacturing study and clean-assembly processes
- Continue to develop and characterise interferometric OSEM (EUCLID)
 - Produce additional prototypes
 - Hope to supply a quantity of non-UHV and UHV samples to LIGO for evaluation
- We seek feedback on interferometric OSEM specifications. For example:-
 - What features would be desirable for potential applications?
 - Sensitivity
 - Working range
 - Dimensions
 - Tilt-immunity
 - Cost



Thank you for your attention