

# PCI Technique for Thermal Absorption Measurements

Ashot Markosyan, Roger Route, Helena Armandula, Martin Fejer *and many others from the LIGO team* 

# Absorption Loss Measurements on LIGO Optics



### • Photothermal Common-Path Interferometer

## Surface Absorption Measurements on Coated Optics

Transversal scans through mirror: point data.

Longitudinal seams: surface inspection

Time scans



- Photothermal Common-Path Interferometer
- Surface Absorption Measurements on Coated Optics
- Transversal scans through mirror: point data
- Longitudinal seams: surface inspection
- Time scans



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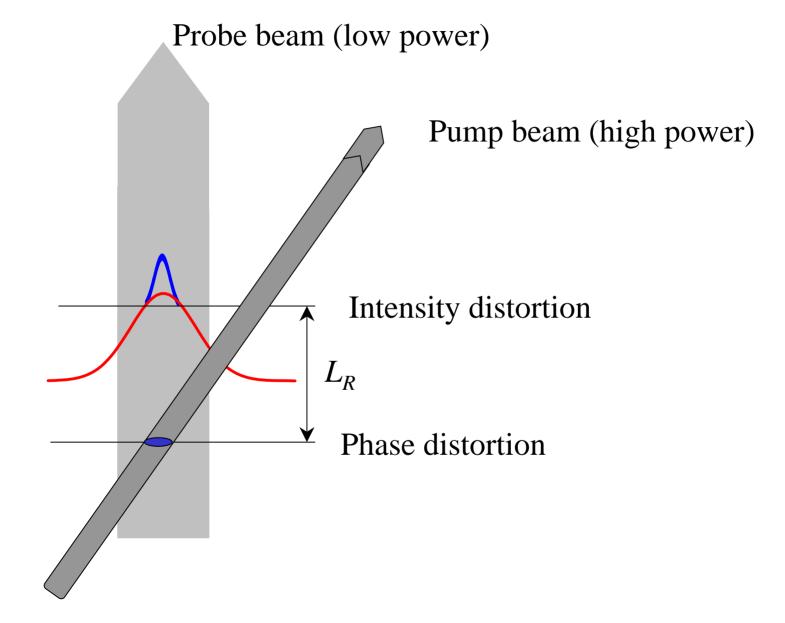
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# <u>PCI Method</u> (by Alex Alexandrovsky)





The probe phase is distorted due to heating:

$$\Delta \varphi = kL \frac{\partial n}{\partial T} \Delta T$$

 $\Delta \phi$  transforms into an intensity distortion. At the Rayleigh length the intensity modulation is:

$$\Delta I/I_0 = k\delta = (2\pi/\lambda)\Delta nL$$

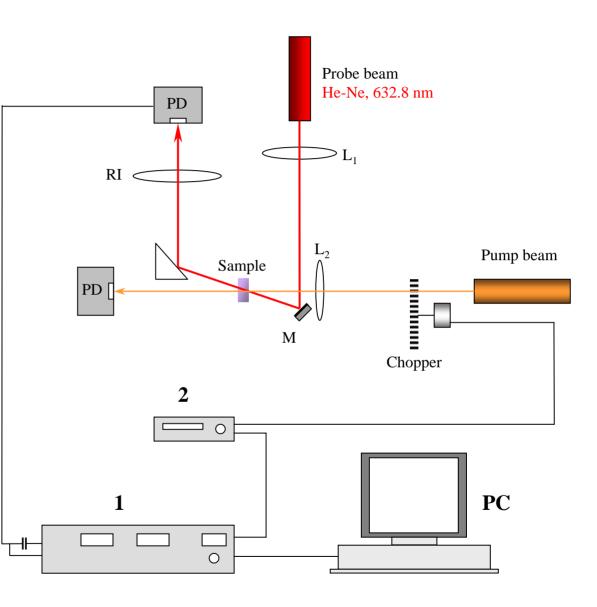
$$\Delta n = (\partial n / \partial T) \Delta T$$

The simplified formula for  $\alpha$  with a "chopped" pump beam of power  $W_p$  is:

 $\alpha \approx C(\partial n / \partial T)(\lambda / L)(w_0^2 / W_P)(\Delta I / I)f$ 



- **RI** R-imaging System
- **PD** Photodetectors
- Lenses
- M Mirror
- 1 Lock-in Amplifier SR830 DSP
- 2 Chopper Controller SR340
- PC with GPIB-connection (LabVIEW software)





## **PCI Method**

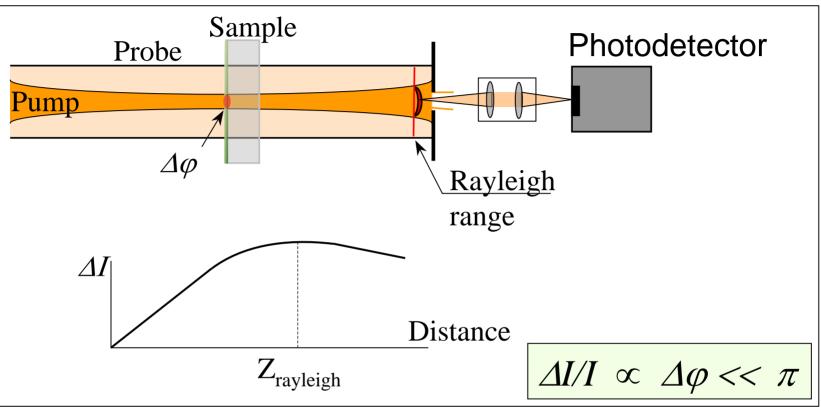


Fig.1.  $w_{pump} < w_{probe}$ . When the pump is on, the probe experiences phase distortion  $(\Delta \varphi)$  in the heated area. At some distance,  $\Delta \varphi$  transform into  $\Delta I$  and an effective interference occurs between the central spot (distorted) and outer ring area (undistorted) of the probe

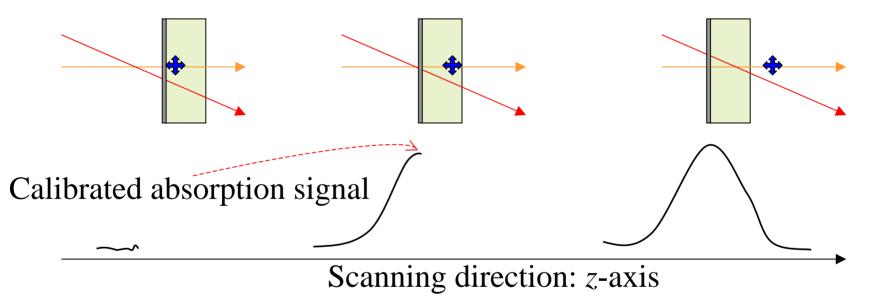


# Measurement technique: transversal *z*-scan (through surface)

Start position

Maximum Signal

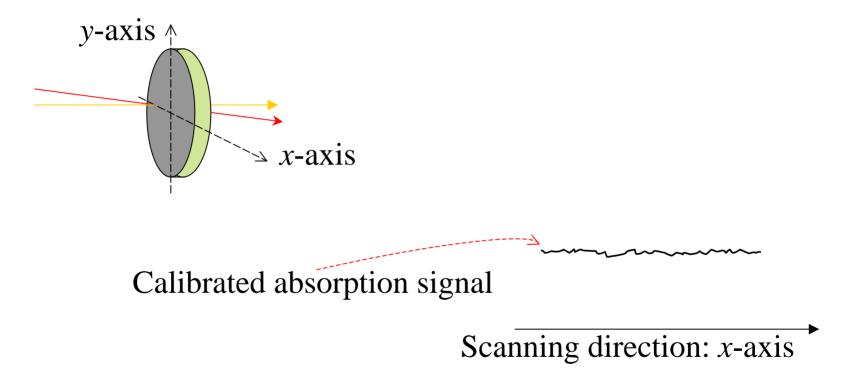
End position



 $\Rightarrow$  > signal reading point (the focus of the optical system)

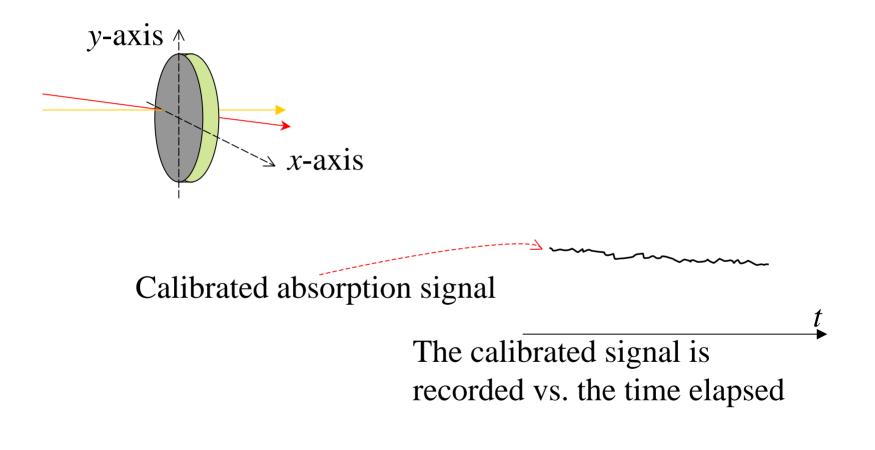
# **Measurement technique: longitudinal** *x***-scan** (across surface)

The heated point is positioned at the coated surface

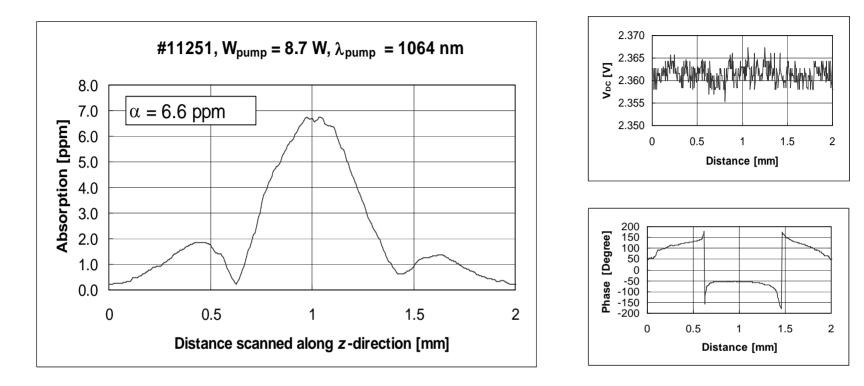


# Measurement technique: time scan

The heated point is focused at the coated surface. The calibrated signal is recorded vs. time elapsed



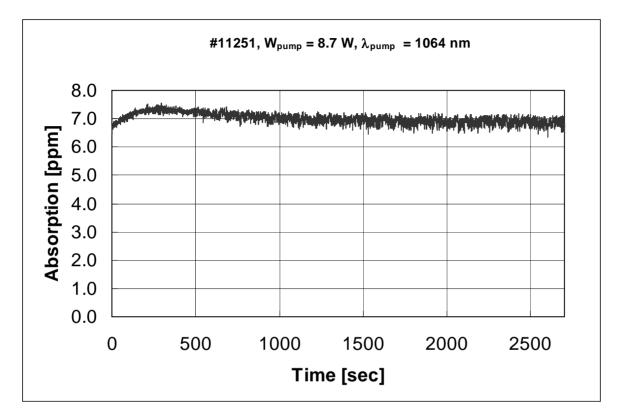




A transversal *z*-scan of sample #11251.  $\alpha = 6.6$  ppm. Additionally monitored parameters : *probe beam stability, phase delay with respect to chopper* 



### HfO<sub>2</sub>-SiO<sub>2</sub>/SiO<sub>2</sub> coating, $\emptyset = 1$ inch



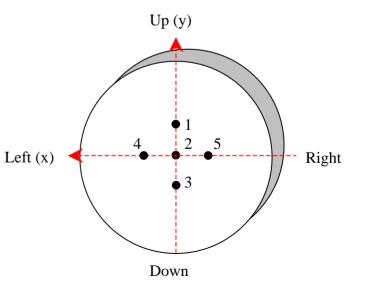
A characteristic time dependent 45 min *t*-scan of sample #11251.  $\alpha$  shows a "bump" (7.3 ppm) at about 5 min.



### Tabulated data on series of HfO<sub>2</sub>-SiO<sub>2</sub>/SiO<sub>2</sub> coated samples

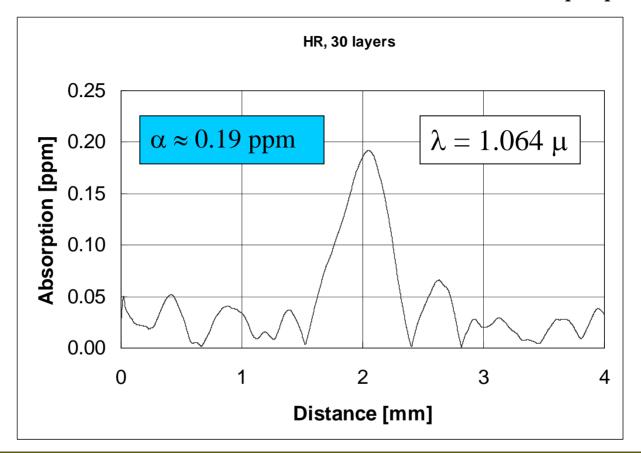
Sample #	Absorption (ppm)					
	3 mm above	Center	3 mm below	3 mm left	3 mm right	Average
014-01	7.2	7.4	7.1	7.1	7.2	7.2
018-01	8.0	8.0	7.9	8.1	7.9	8.0
024-01	6.9	6.7	6.9	6.8	6.9	6.9
11251	6.8	6.4	6.6	6.6	6.6	6.6

Position of measured points



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### Method's resolution: down to 0.1 ppm with $W_{pump} < 10$ W

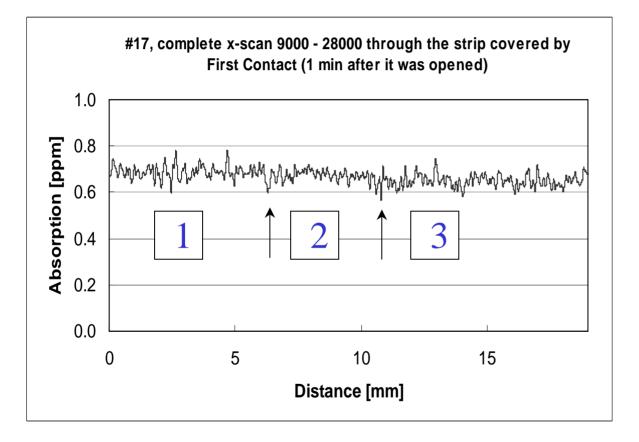


A transversal *z*-scan of a 30 layer HR mirror on a FS substrate.  $W_{pump} = 8.8 \text{ W}, \lambda = 1.064 \text{ } \mu. \text{ <u>Absorption measured} = 0.19 \text{ } ppm.$ </u>



### A coated 3 inch disk marked #17, $\alpha = 0.65$ ppm.

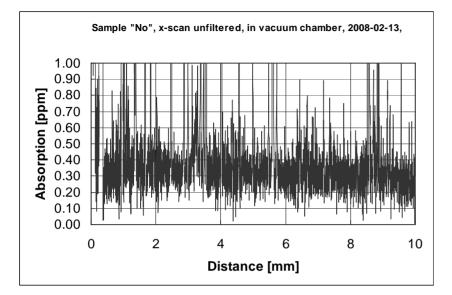
#### Covered by a First Contact strip for 1 week



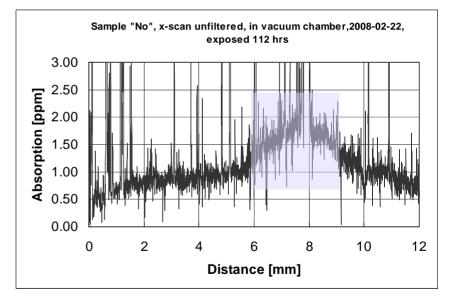
Testing the "First Contact" effect.  $\alpha = 0.65$  ppm.

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### Study of the UV illumination effect on LIGO coated mirrors



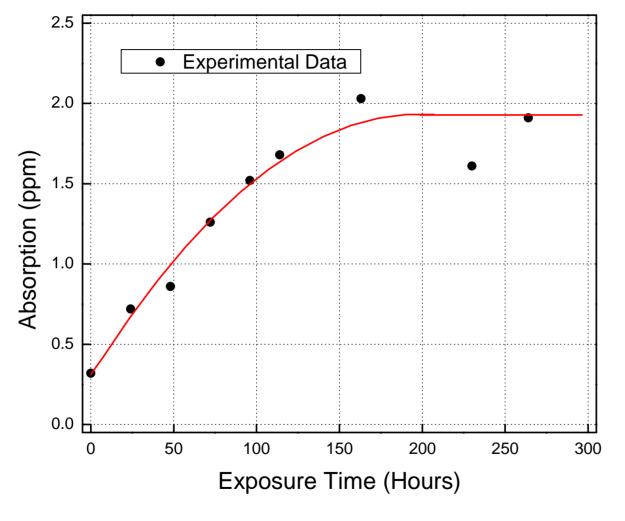
Characteristic *x*-scans before and after UV illumination (*see Ke-Xun Sun's talk for details*)





### Study of the UV illumination effect on LIGO coated mirrors

Sample # "No"



Variation of the absorption loss vs. Exposure Time



### Summary

# The PCI method is capable to measure thermal absorption losses as low as 0.1 ppm with 8 W incident power

Point by point as well as continuous scan across the coating are possible

The lowest Absorption value measured on LIGO mirrors till now were 0.19 ppm