# Phase and alignment noise in grating interferometers





## All-Reflective Interferometry



Main optics do not use transmissive components

## **Grating Interferometers**



PER AD



## Phase and Alignment Noise

- Motivation: compute coupling coefficients related to the phase and alignment of the diffracted beams for the use in numerical simulations
- How does the motion of the grating (or the incoming beam) affect the direction of the outgoing beam?
- How does the motion of the grating affect the phase of the outgoing beam
- Idea: use only the geometry of the optical layout and ideal gratings (and compare to ideal transmissive optics)



## Grating Interferometers

The grating equation:

$$\sin\alpha + \sin\beta_m = m\lambda/d$$



... in 3D:

$$ec{q} imes ec{e_z} - ec{p} imes ec{e_z} = rac{m\lambda}{d} ec{e_y}$$

Note: Grating normal along z, grating grooves along y



## Alignment Effects

- Rotation around x- and y-axis has the same effects like with standard mirrors
- Rotation around z-axis (roll) is different from standard optics: causes approximately the same effect as tilt (rotation around x-axis), same type and amount of beam rotation
- Rotation of the grating has no effect on the phase of the light



## **Translation Effects**

- Translation of the grating or of the beam has no non-obvious effects on the alignment of the outgoing beam
- In the following we will describe the effects of translation on the phase of the diffracted beam





⋇











$$\xi_{\Delta x} = -\Delta x \left[ \frac{m\lambda}{d} \right]$$

≈1 for gratings with few diffraction orders

Optical phase change is proportional to beam translation!

Effect is proportional to *m*, especially it is zero for the zeroth order.

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## Cavity Alignment



$$\Delta L \approx R_c \,\Delta \gamma_1 \Delta \gamma_2 \qquad \longrightarrow \qquad \text{quadratic}$$
$$x'' \approx (R_c - L) \Delta \gamma_2 + R_c \Delta \gamma_1 \qquad \longrightarrow \qquad \text{linear}$$

⋇



## **Alignment Limits**

For a linear cavity (Virgo like) with a 2-port grating and typical values we can compute alignment limits:

 $\Delta L = 2R_c \gamma_1 \gamma_2$  yields:



$$\gamma_2 < 2 \cdot 10^{-16} \frac{\text{rad}}{\sqrt{\text{Hz}}} \left( \frac{h}{10^{-23}/\sqrt{\text{Hz}}} \right) \left( \frac{L}{3 \,\text{km}} \right) \left( \frac{3.5 \,\text{km}}{R_c} \right) \left( \frac{10 \,\text{nrad}}{\gamma_1} \right)$$

$$\Delta x = \frac{R_C}{\cos(\alpha)} \gamma_2 \quad \text{yields:}$$
  
$$\gamma_2 < 7 \cdot 10^{-24} \frac{\text{rad}}{\sqrt{\text{Hz}}} \left( \frac{h}{10^{-23}/\sqrt{\text{Hz}}} \right) \left( \frac{L}{3 \,\text{km}} \right) \left( \frac{\cos(\alpha)}{\cos(30^\circ)} \right) \left( \frac{3.5 \,\text{km}}{R_C} \right) \left( \frac{d}{\lambda} \right)$$



## Other Layouts (3-port, 4-port)

- Other diffractive layouts (using the same Virgo like parameters)
  - 3-port input coupler (cavity mode reflected at normal incidence): 10<sup>-21</sup> rad/sqrt(Hz)



 Diffractive beam splitter: 10<sup>-21</sup> rad/sqrt(Hz)





## **Compensating Phase Noise?**

Change of optical phase due to translation along z:





# Finding a Special Eigensystem

 Special eigenvector in the direction of the bi-section between incoming and diffracted beam:

$$\Theta = \tan\left(\frac{\alpha + \beta_m}{2}\right)$$

Grating motion orthogonal to this vector causes **no** phase noise





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## Summary

- Gratings introduce much more stringent alignment requirements
- Possible scenarios:
  - High-power high-frequency detectors
  - Alignment insensitive topologies?
- Grating roll has to be controlled and isolated as well as the other alignment degrees of freedom
- The sensitive/in-sensitive axes of motion are not parallel to the beams or the grating normal, suspension design should take this into account

Freise et al, LIGO-P070094-00-Z, to be published in NJP



#### ... end.

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