

Thermal Compensation System for Virgo+

Alessio Rocchi

INFN Roma Tor Vergata

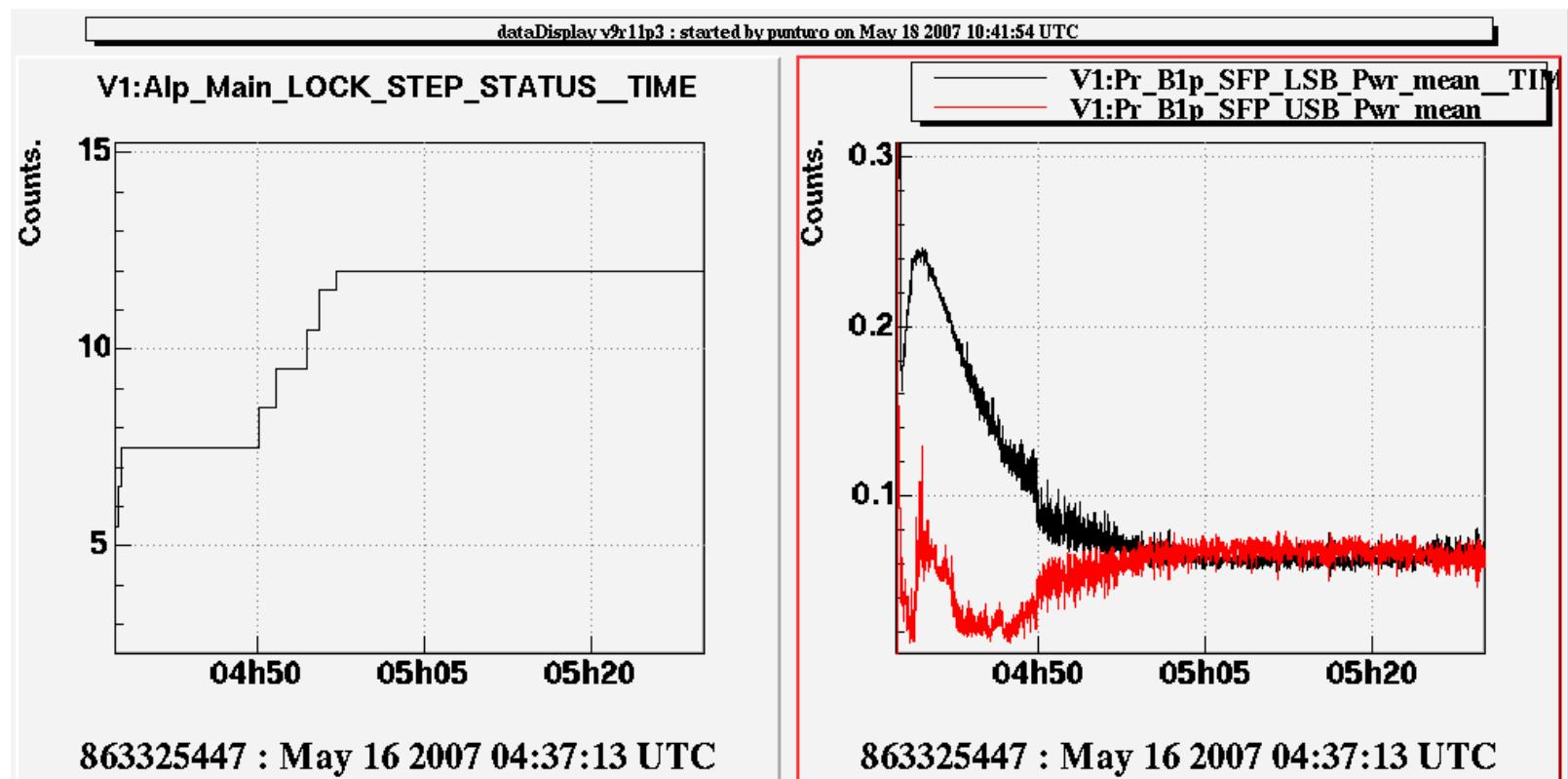


LIGO-G070734-00-Z



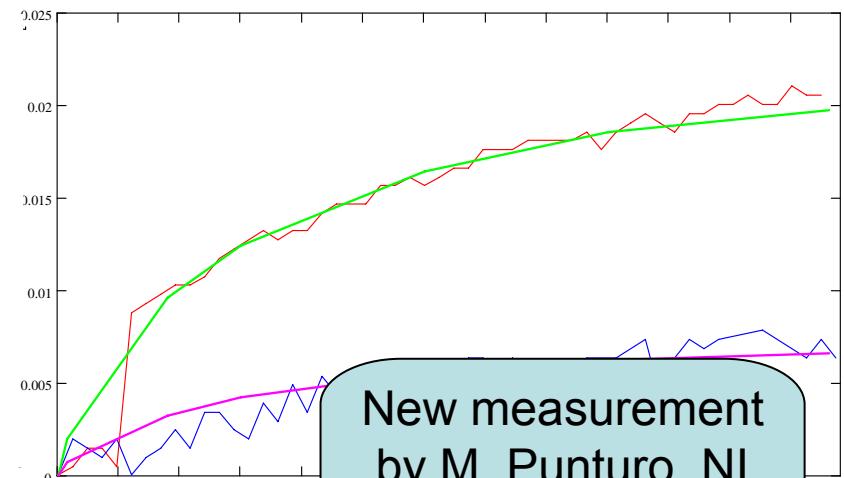
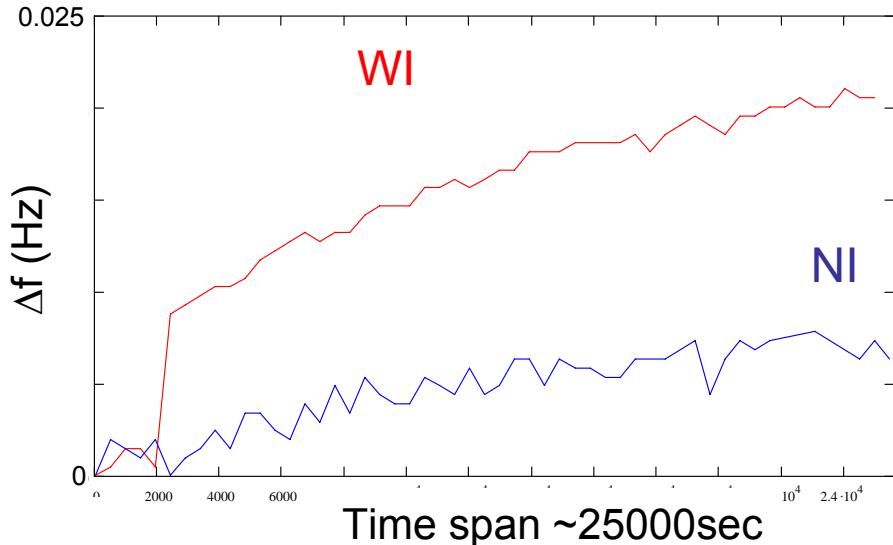
Thermal Lensing

- Observed in Virgo through reduction of the sidebands gain.



ITMs absorptions evaluated through coupled thermal-modal FEM with Ansys

Red and blue curves experimental frequency shifts green and pink FEM results



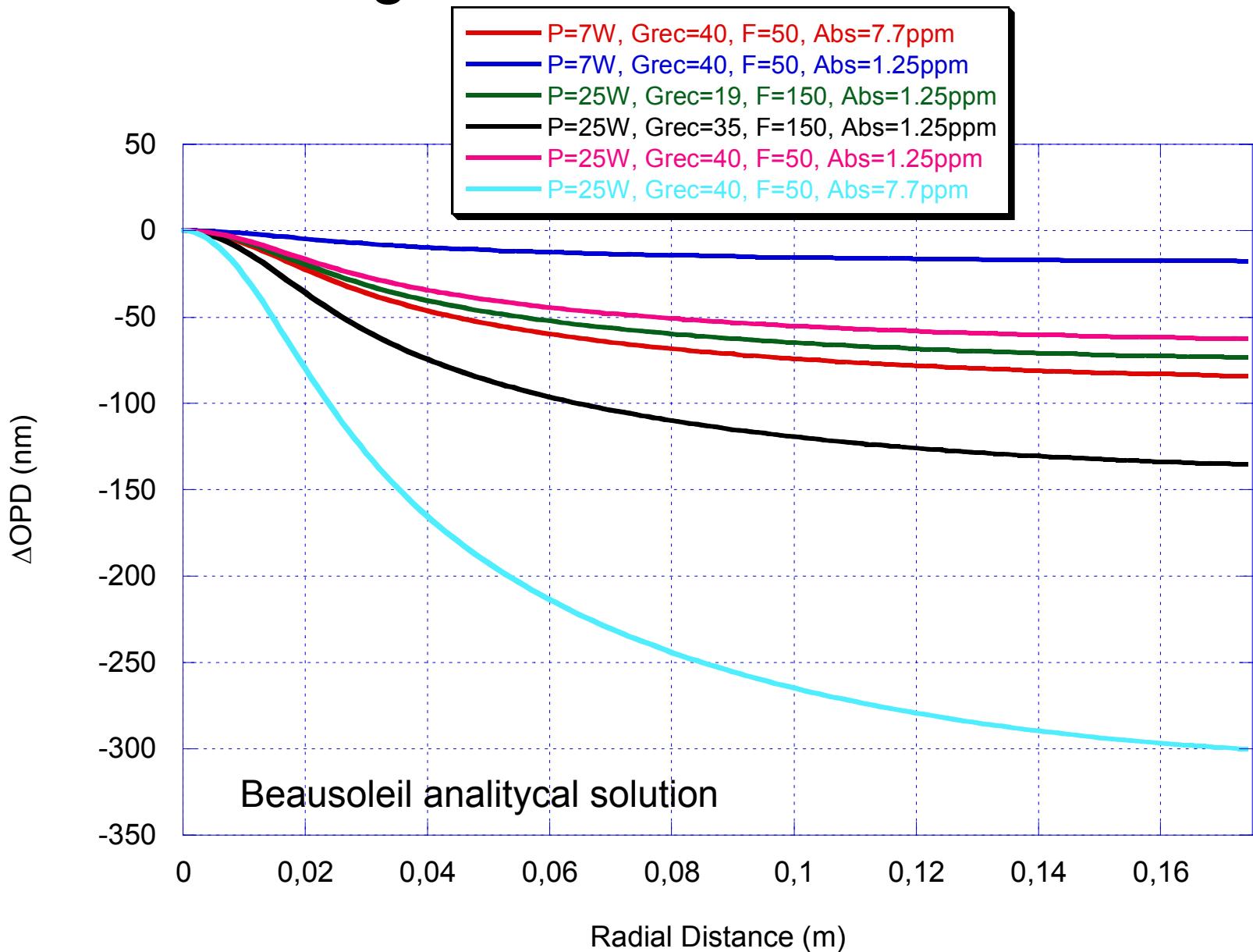
ITMs effective absorptions

WI (7.8 ± 1.1) ppm

NI (2.3 ± 0.3) ppm

Errors come from dependency of the model from uncertainty on elastic parameters dependency on temperature

Thermal lensing evaluation



Thermal Compensation System

- TCS for Virgo+: annular heating of ITMs with dc CO₂ laser
- “Annular” profile
 - from simulations: reasonable parameters to compensate inner radius~2.5 cm, outer radius~11÷14 cm

Desirable properties:

- It is easily adaptable as new understanding of the ITF is realized
- It does not require a significant vacuum incursion to install as this would lead to significant down time to the instrument

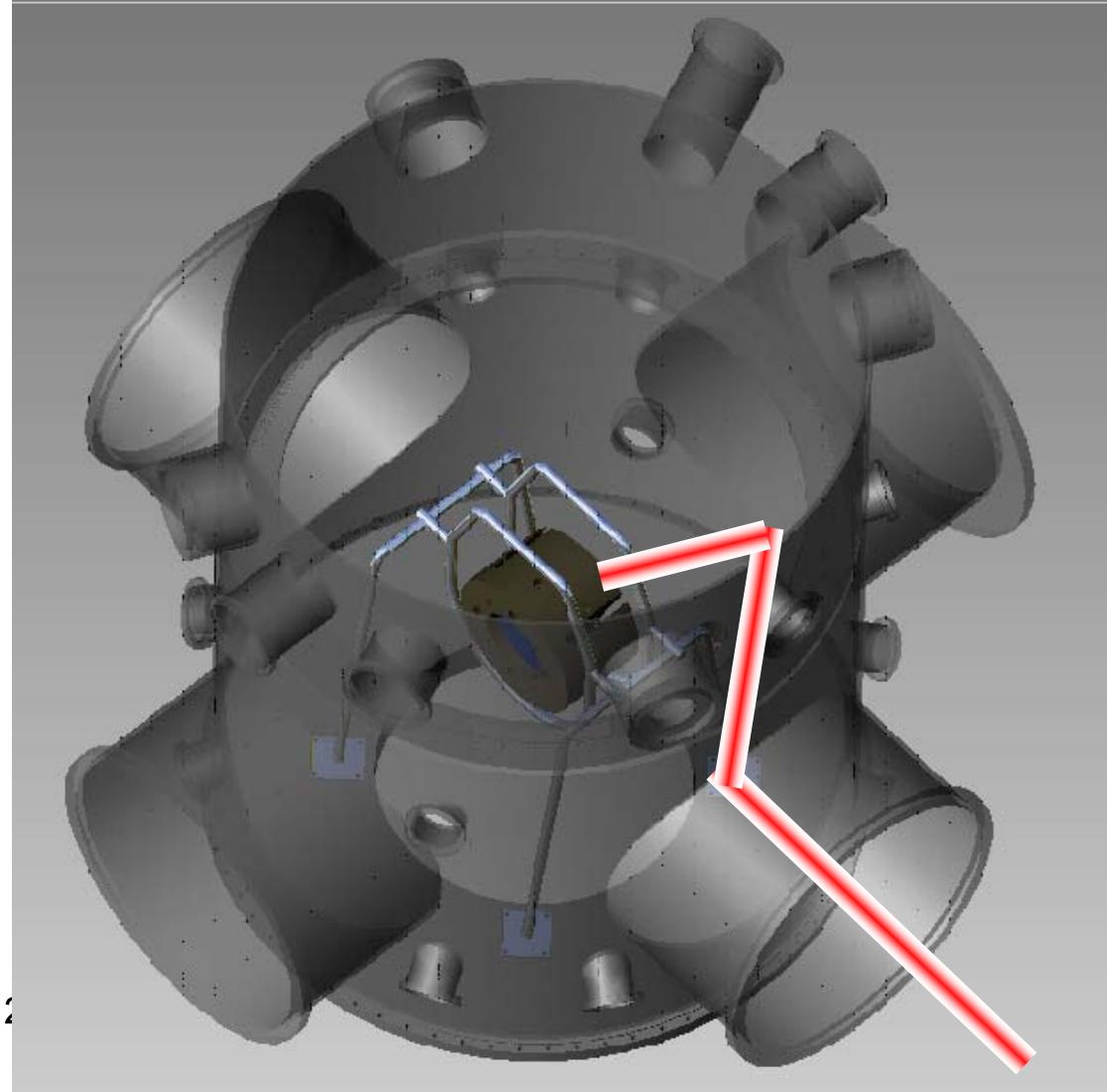
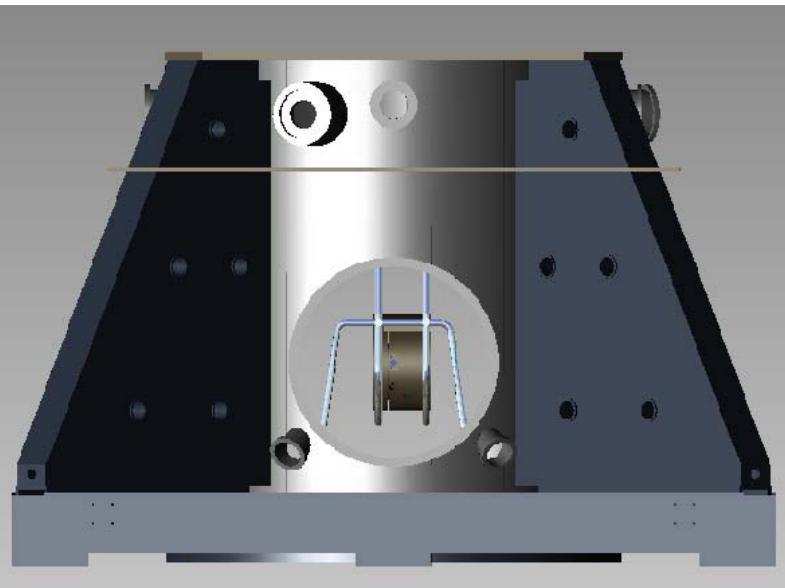
Two phases installation:

1. Feb 2008, system without power stabilization control loop
2. May-June 2008, installation of the power stabilization control loop

TCS Power requirements

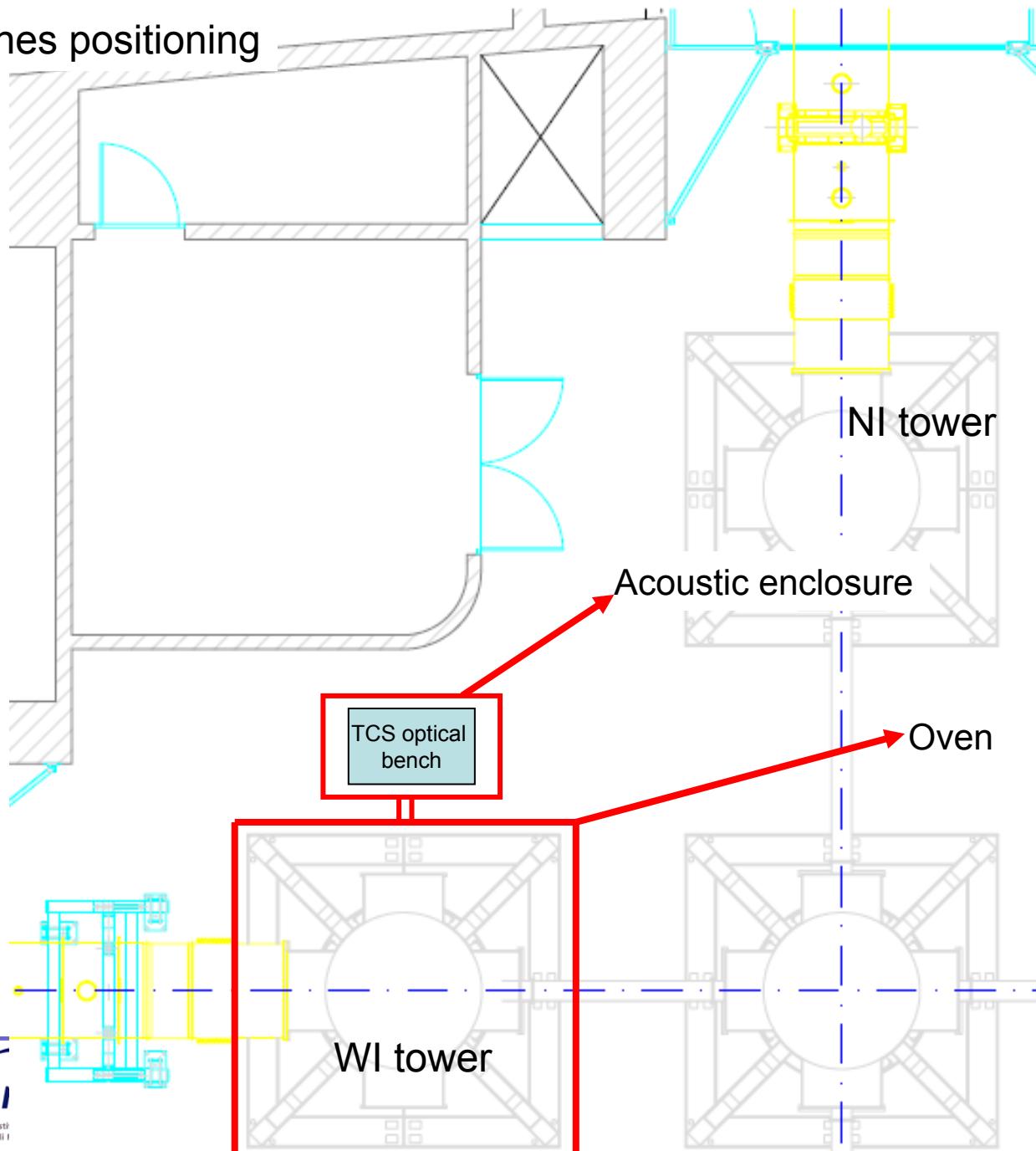
Virgo on WI mirror (absorption 7.7ppm)	~1.8 W
Virgo (nominal absorption)	~0.4 W
Virgo on WI mirror (absorption 7.7ppm, $P_{\text{las}}=25\text{W}$)	~6.2 W
Virgo (nominal absorption, $P_{\text{las}}=25\text{W}$)	~1.3 W
Virgo+, ($F=150$, $G_{\text{rec}}=19$) (nominal absorption)	~1.5 W
Virgo+ ($F=150$, $G_{\text{rec}}=35$) (nominal absorption)	~2.8 W

Optical layout



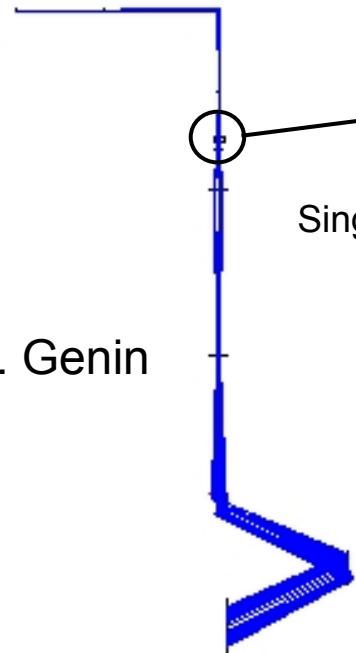
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Optical benches positioning

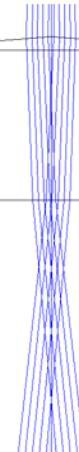


Final design of the Optical Imaging System

P. La Penna & E. Genin



Single axicon to get the “annular” profile



100% efficiency

MON OCT 15 2007

3D LAYOUT

VIRGO_L1-A
CONFIG

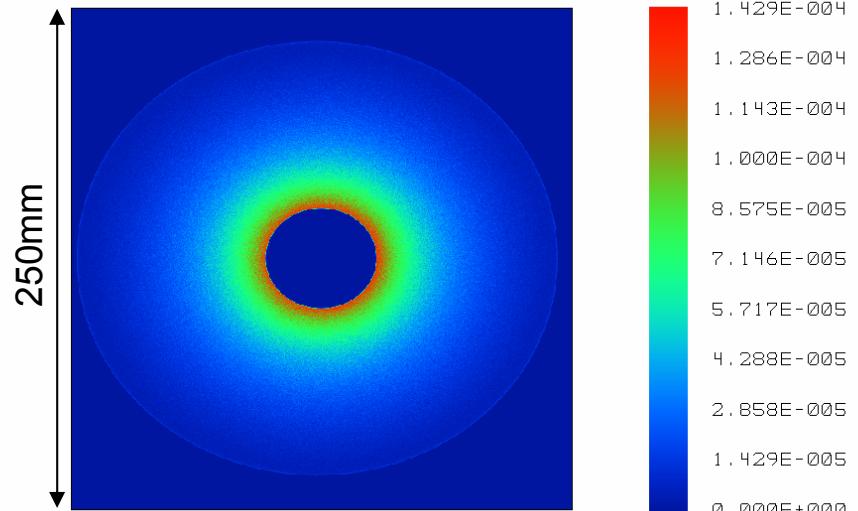
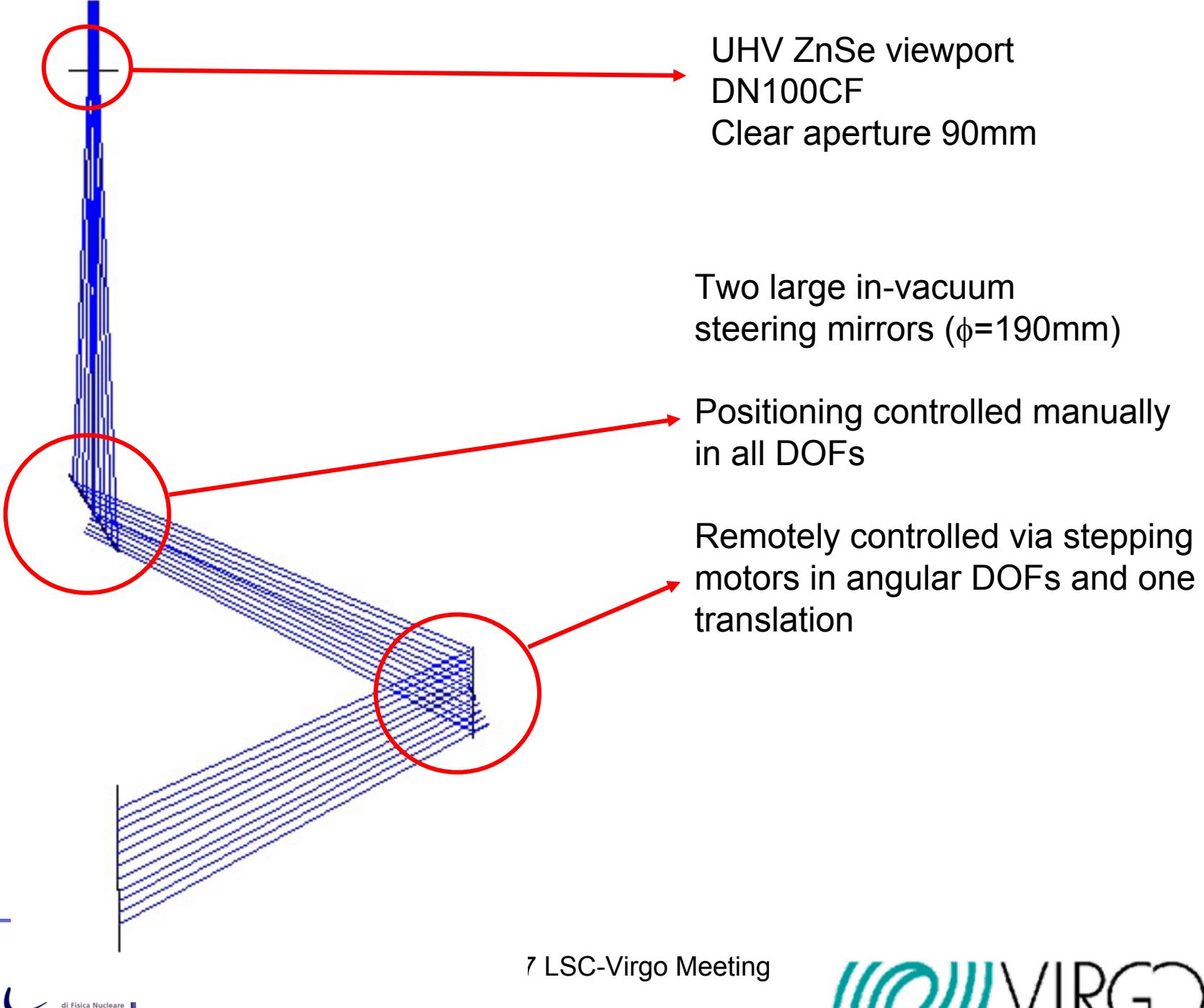


IMAGE DIAGRAM

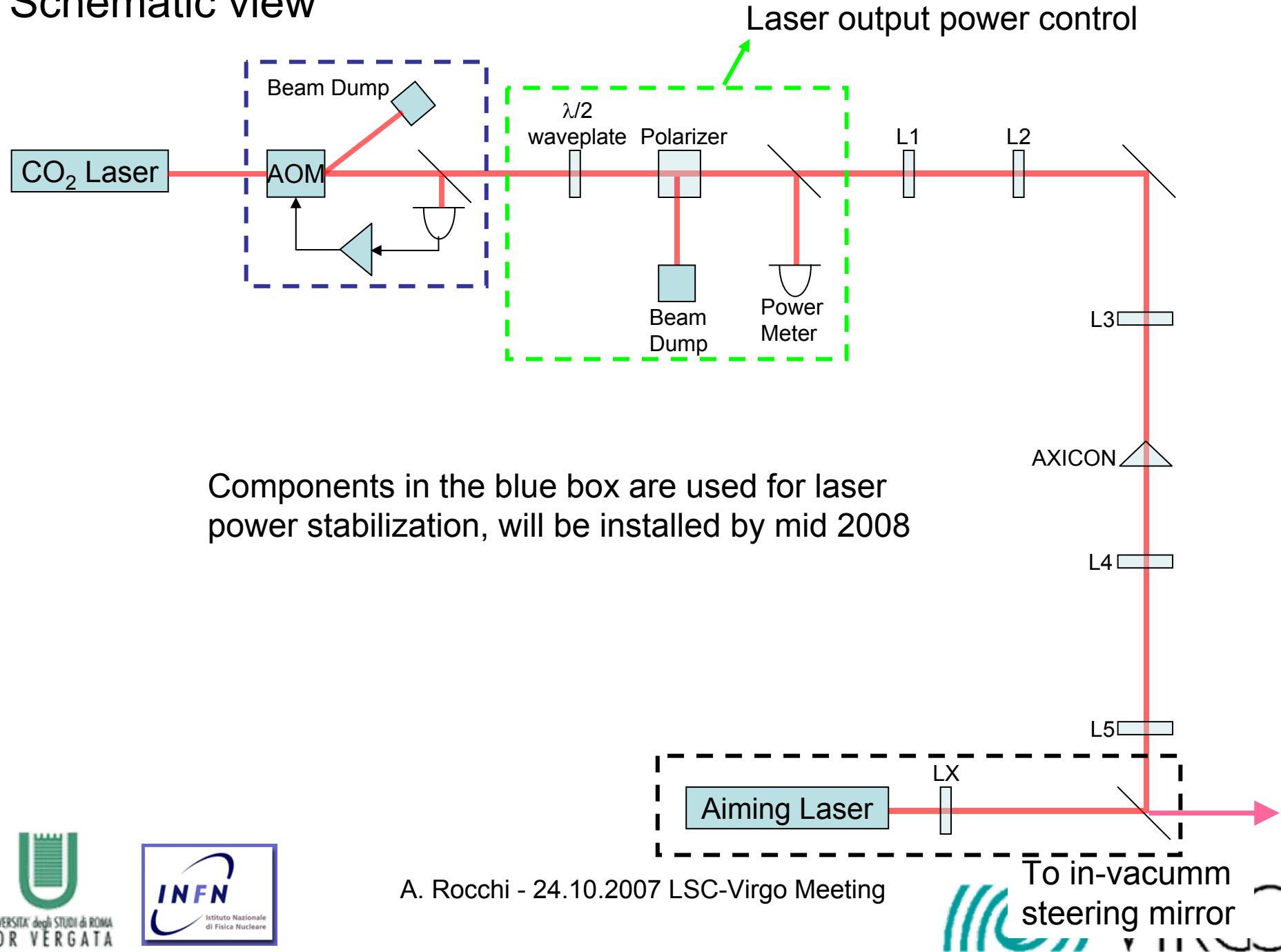
MON OCT 22 2007
IMAGE WIDTH = 250.0000 MILLIMETERS, 600 X 600 PIXELS
FIELD POSITION: 0.0000 0.0000 DEG
PERCENT EFFICIENCY: 100.000%, 1.000E+000 WATTS
SURFACE: 34, UNITS ARE WATTS PER MILLIMETERS SQUARED.

VIRGO_L1-AXICON_DEFMIRR_LAST3.ZMX
CONFIGURATION 1 OF 1

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Schematic view



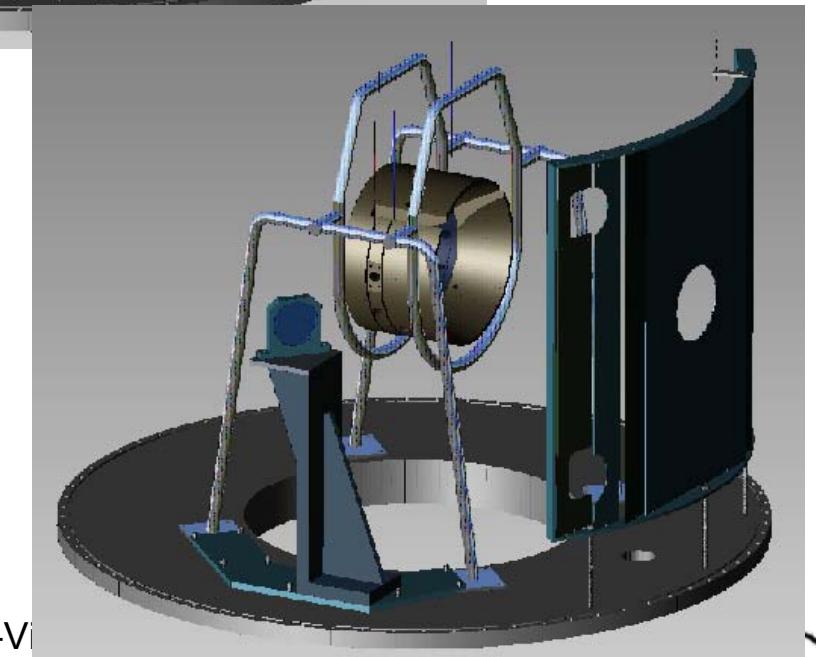
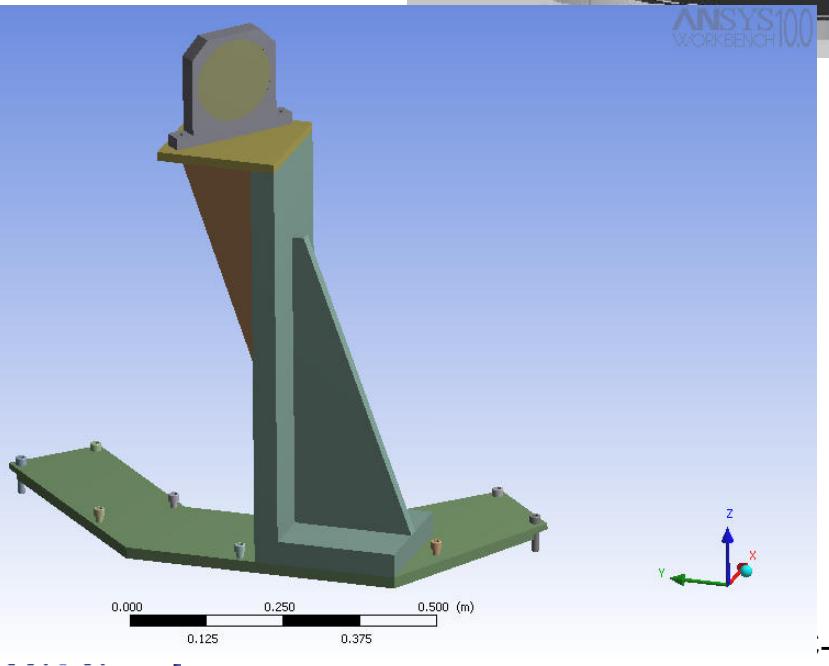
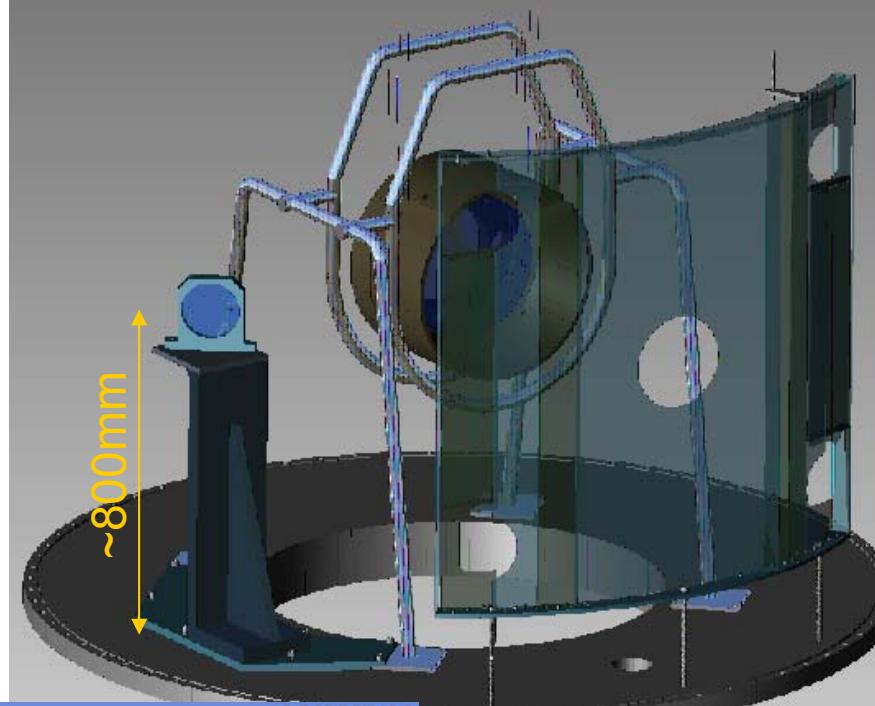
Tests in Tor Vergata Labs before installation on Virgo

- Measure the RIN of the laser
 - So we can have an estimate of the injected noise in the first phase installation
- Check the efficiency of the Optical Imaging System

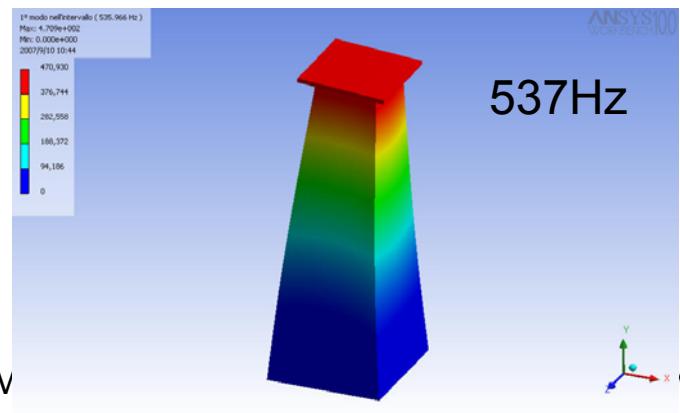
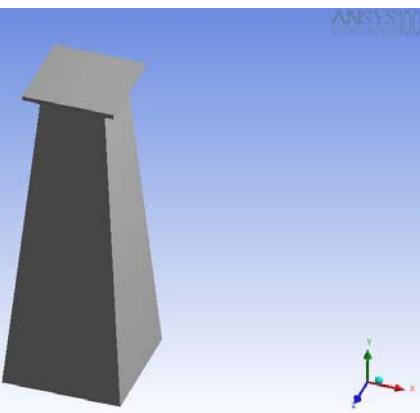
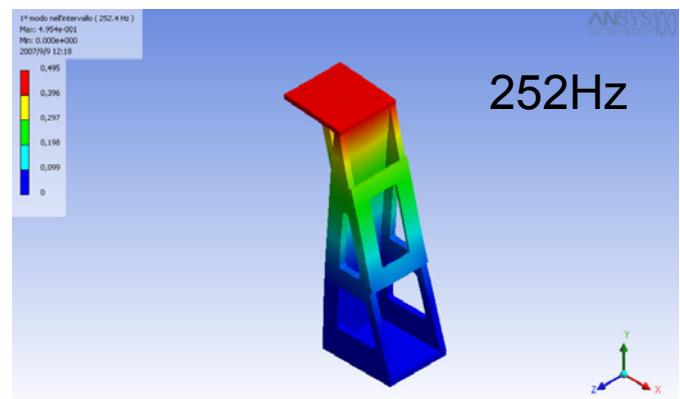
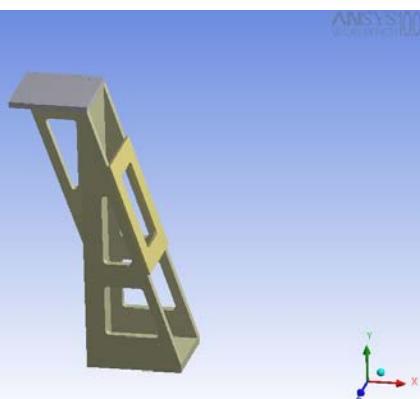
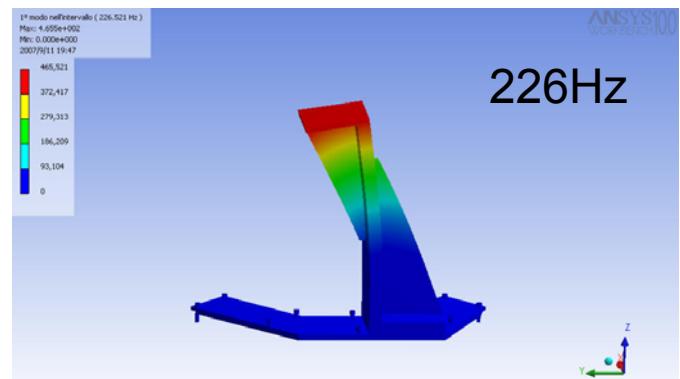
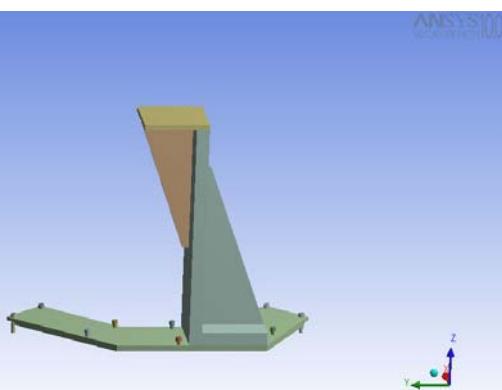
Preliminary design of the first in-vacuum mirror

Material: Aluminum

T. Zelenova



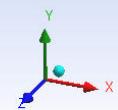
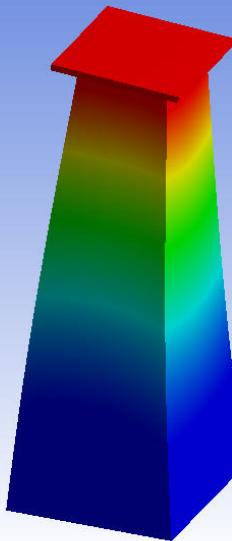
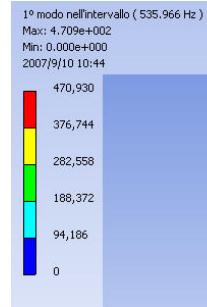
Exploring different designs for the in-vacuum optics holders, with T. Zelenova



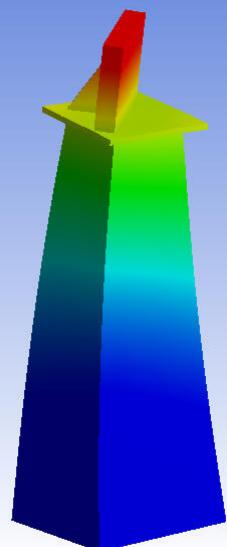
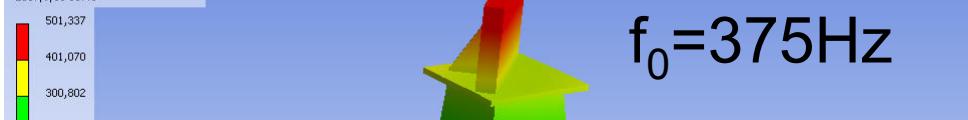
chi - 24.10.2007 LSC-Virgo M

$f_0 = 537\text{Hz}$

Unloaded



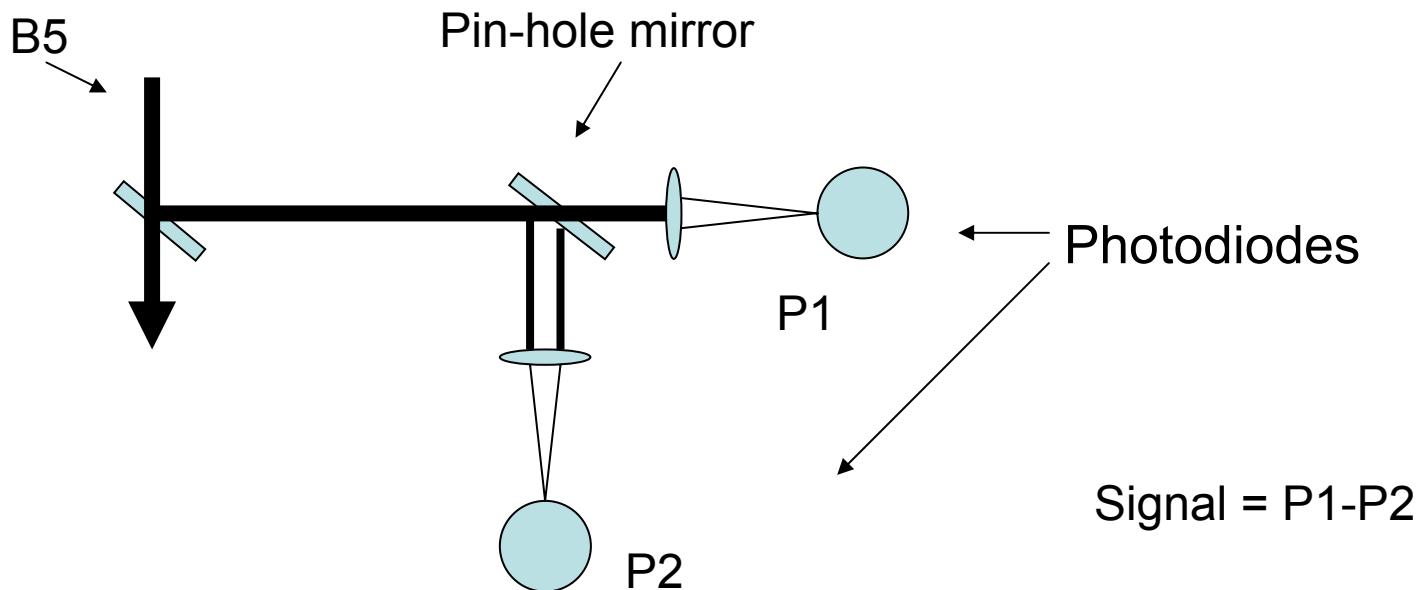
1° modo nell'intervallo (375.181 Hz)
Max: 5.013e+002
Min: 0.000e+000
2007/9/10 10:40

 $f_0 = 375\text{Hz}$ 

Loaded with mirror (2kg)

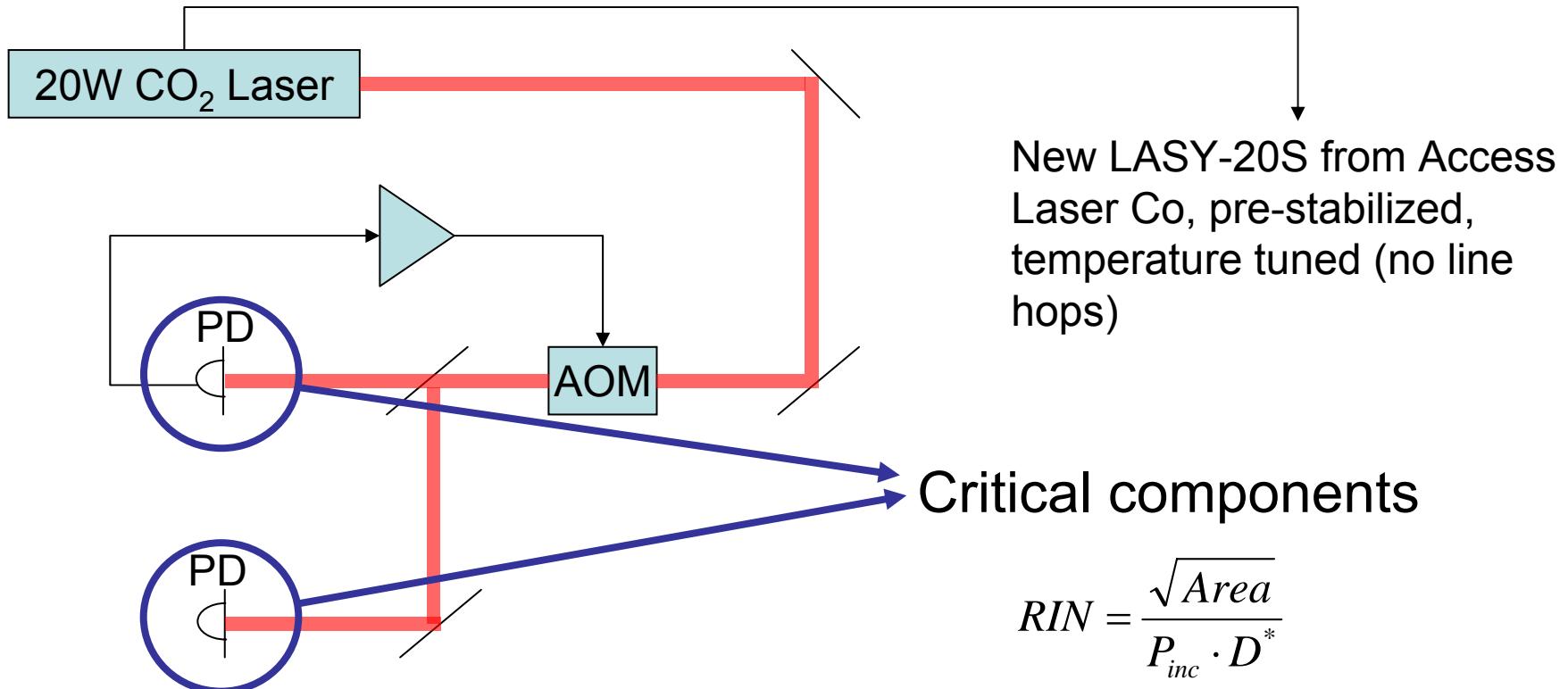
Error signal generation

- E. Calloni & M. Laval demonstrated that Phase camera 0 + DarkF → tools to “measure” TCS
- Pine-hole photodiodes → error signal generation
- Phase camera 1 → an upgraded tool to understand sidebands and TCS



This is only one error signal, but two are needed. Investigating the possibility to use the signal from B2

CO_2 laser power stabilization

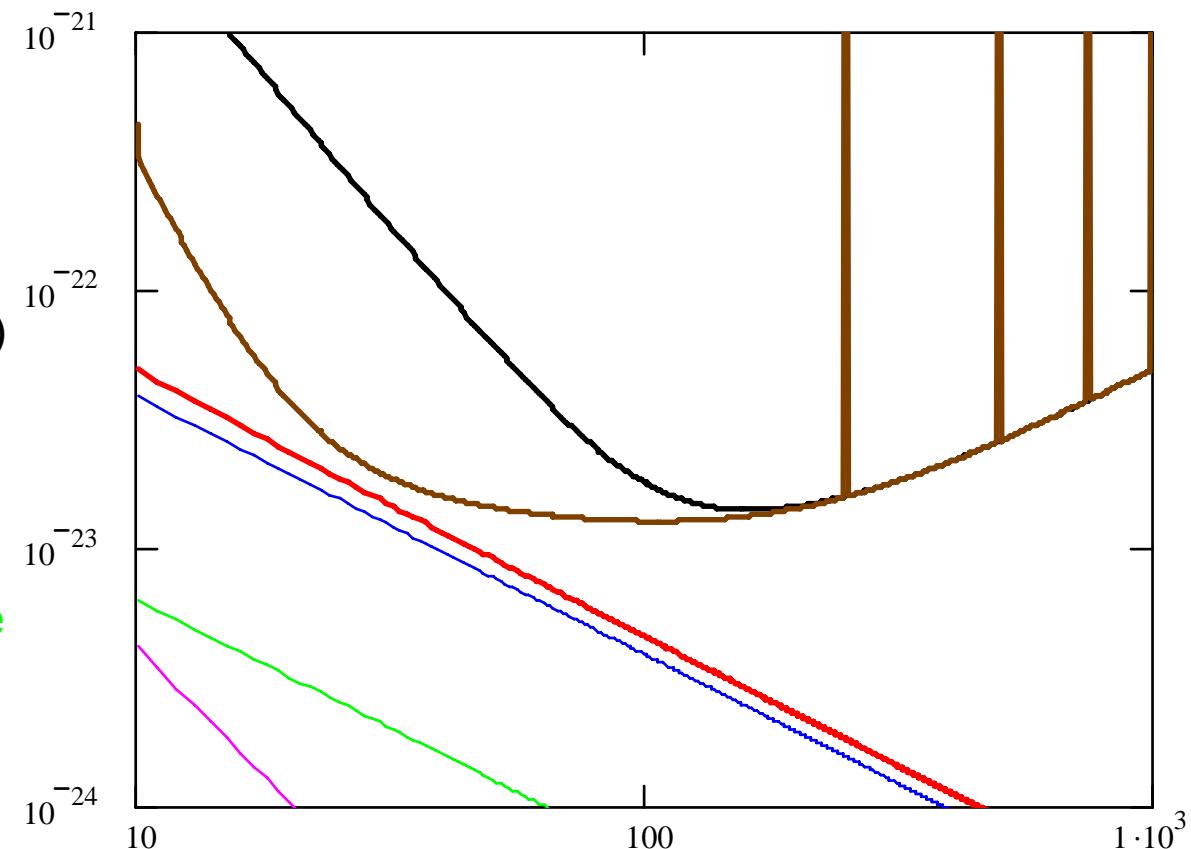


	Detectivity ($\text{cmHz}^{1/2}/\text{W}$)		Area	
	TE	Room Temp.	TE	Room Temp.
VIGO photodiodes	$10^8 - 10^9$	$10^7 - 10^8$	$0.5 \times 0.5 - 2 \times 2 \text{ mm}^2$	$2 \times 2 - 4 \times 4 \text{ mm}^2$
Coherent photodiodes		$2.2 \cdot 10^8$		$5 \times 5 \text{ mm}^2$

Noise budget

Evaluation of TCS noise injected in Virgo+, if RIN is stabilized @ $3 \times 10^{-7}/\sqrt{\text{Hz}}$, flat over the frequency range 10 Hz to 1000 Hz, TCS power=1.5W.

Virgo+ sensitivity
(without monolithic suspensions)
Virgo+ sensitivity
(with monolithic suspensions)
Total
Flexural
Thermoelastic+Thermorefractive
Radiation pressure



The End

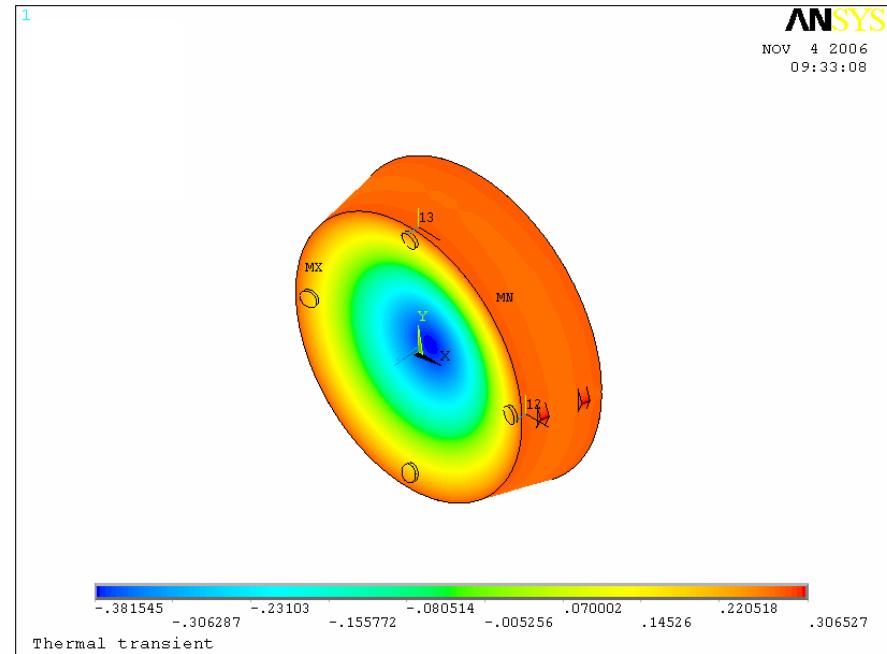
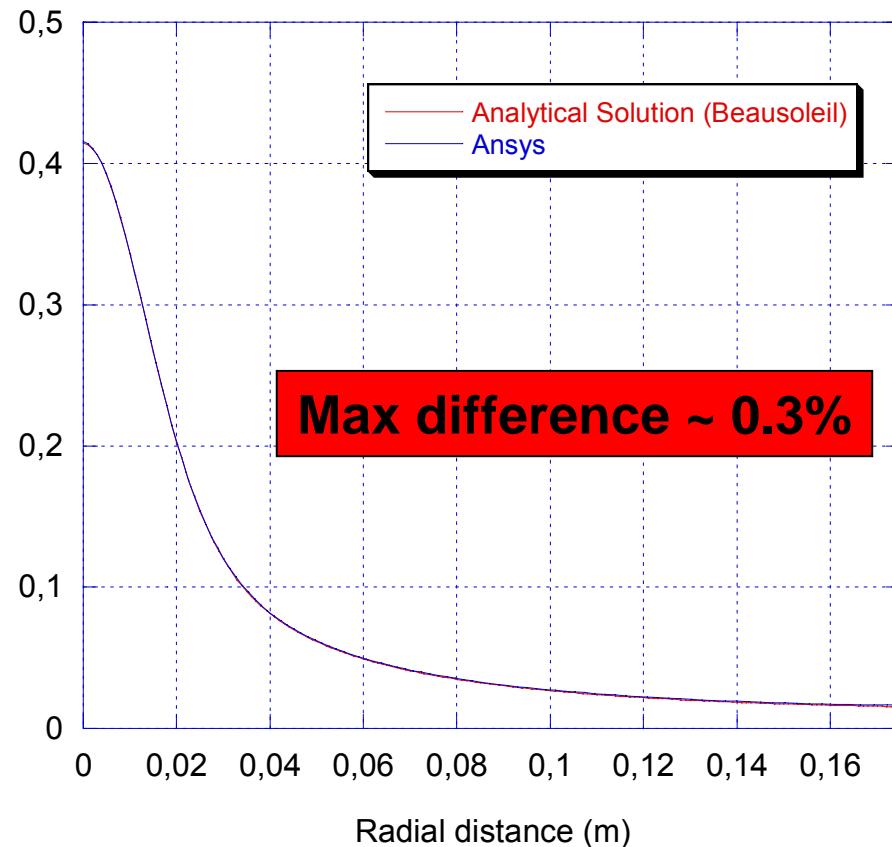


A. Rocchi - 24.10.2007 LSC-Virgo Meeting



Building and validating the FE model

Comparison with analytical solution
Beausoleil ([LIGO-P020026-00-D](#))



- DRUM MODE
- ANSYS FEM $f = 5584.15$ Hz
- Experimental value $f = 5584.7$ Hz

Difference is 0.55Hz
less than 0.01%

Possible optical layout

AXICON

