



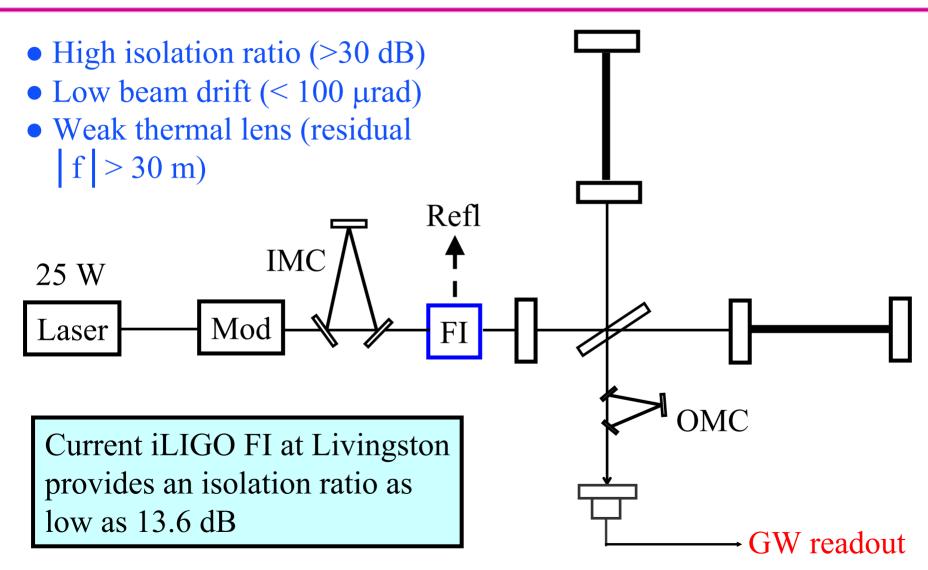
eLIGO/advLIGO FI update

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FI for eLIGO



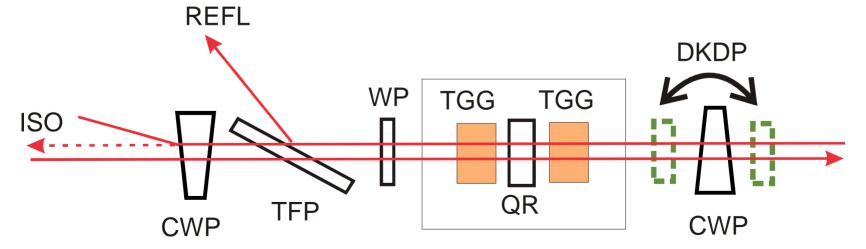




eLIGO/advLIGO FI configuration



- Faraday Rotator
 - Two 22.5° TGG crystals with a reciprocal 67.5° QR
- Calcite wedge polarizers (superior to TFPs)
- Thermal lens compensation with dn/dT material (deuterated dihydrogen phosphate, or 'DKDP')





Problems we have encountered

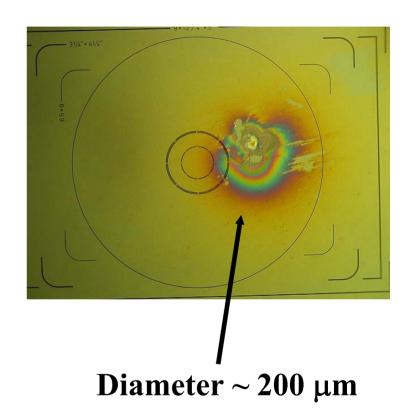


- TGG and DKDP crystals:
 - surface damage
 - improper AR coatings

SOLUTION: New vendor (NG)

Magnet

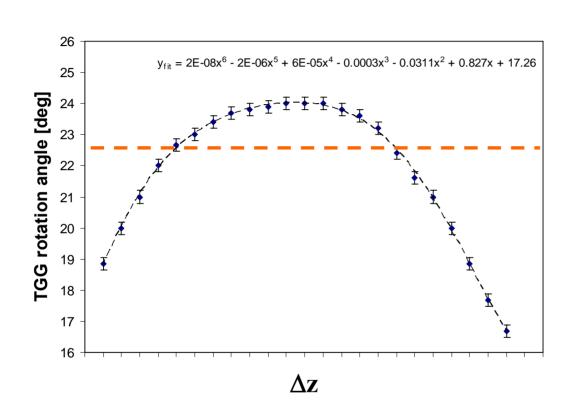




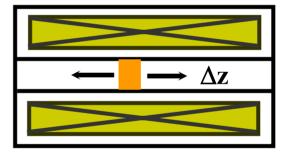
FI magnet ring failure

LIGO Characterization of individual **TGGs**





Incident laser power: 1 W



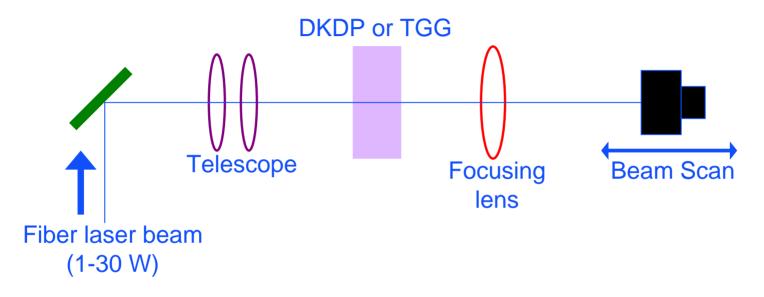
Result from 6th order polynomial fit:

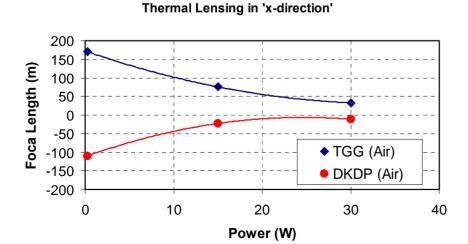
Max TGG rotation angle (crystal #2): 24.03 deg +/- 0.01 deg



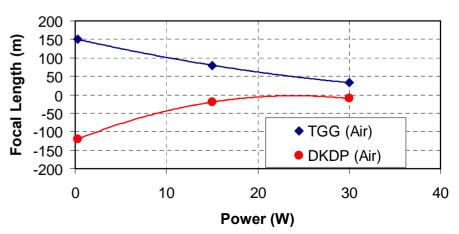
Thermal lens created by TGG and DKDP









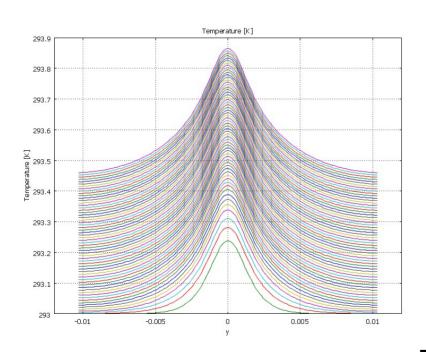


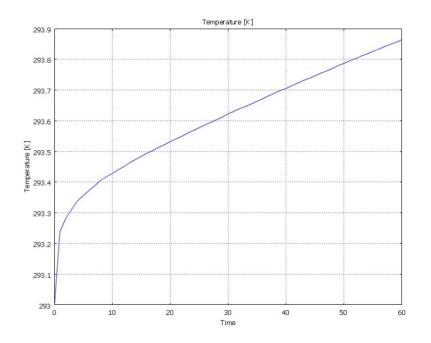
LIGO

COMSOL Simulation (radiation enabled on faces and edges)

LSC

30 W incident power Time from $\underline{t = 0s \text{ to } t = 60s}$ with a line every 1s Temperature in the center of the pumped face (point [0,0,0])





Parameters:

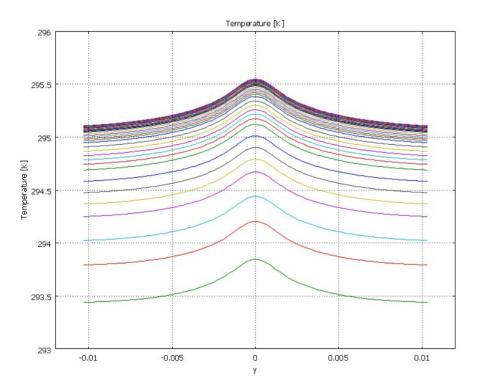
TGG radius = 10.3 mm
Laser beam radius = 1.5 mm

The time constants of thermal lensing are in the order of seconds

LIGO COMSOL Simulation (radiation enabled on faces and edges)

LSC

30 W incident power Time from t = 0s to t = 3600s (60min) with a line every 60s (1min)



294.5 293.5

Temperature [K]

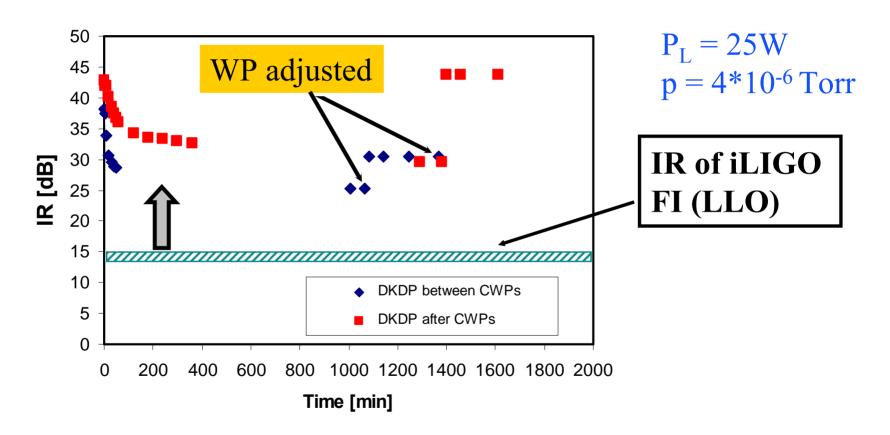
Parameters:

TGG radius = 10.3 mm Laser beam radius = 1.5 mm Temperature in the center of the pumped face (point [0,0,0])

LIGO

Faraday Isolator performance Isolation ratio (CWPs)





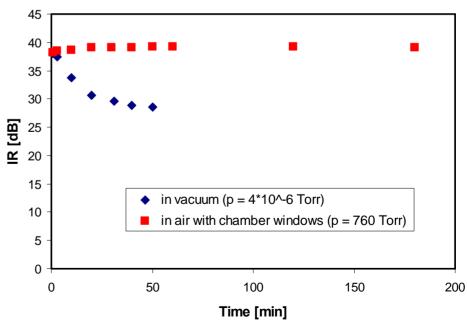
- Decrease in isolation ratio under vacuum as time went on
- Could be recovered by < 2 degree rotation of waveplate

LIGO Faraday isolator performance Effect of vacuum



- $P_L = 25 \text{ W}$
- DKDP between CWPs

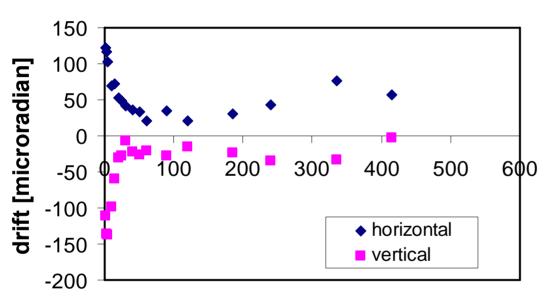




LIGO Faraday isolator performance thermal drift (CWPs)



Refl beam thermal drift intrinsic to FI

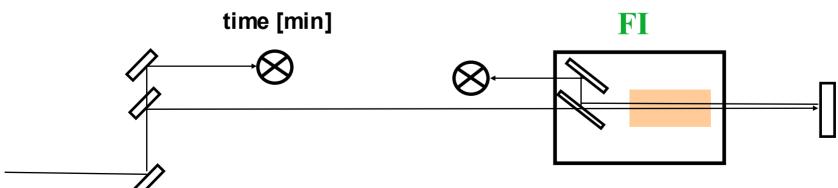


$$P_{\rm L} = 25 \text{W}$$

p = 4*10⁻⁶ Torr

Result:

Over the course of six hrs the horizontal and the vertical intrinsic drifts are < 50 microradians







ELIGO FI Schedule

LLO

- Assembly, final test and measurements of the FI will be done in the clean optics lab at the LIGO Livingston Observatory
- Installation of the FI for LLO will start on Nov 13, 2007.
- Integration and commissioning of the FI and Input Optics with PSL is planned to start on June 3, 2008.

LHO

- Installation of the FI at LHO is planned to start in December 2007 or January 2008.
- Integration and commissioning of the FI and Input Optics with PSL is planned to start on Feb. 19, 2008.





Conclusions

- Initial LIGO FI measurements reveal that isolation requirements for E-LIGO are very relaxed
 - \rightarrow 20 dB will work well
- Obtained 43 dB isolation ratio with CWPs in vacuum Decrease in isolation ratio under vacuum should be recovered by < 2 degree rotation of waveplate. Intrinsic thermal drift of the reflected beam was
 < 50 microradians for the horizontal and vertical axes
- Performance is expected to improve with Northrop Grumman crystals (TGG/DKDP) and new precision mount for DKDP with tip/tilt positioning