



UV Illumination Studies at Trinity University

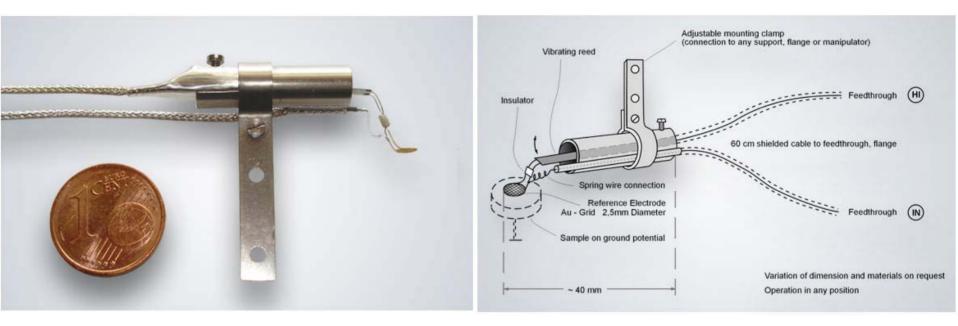
Dennis Ugolini, Mark Girard 2007 Workshop on Charging Issues July 27, 2007

- Calibration of Besocke Kelvin probe
- Discharge rates versus UV intensity, wavelength
- Addendum: International Cosmic Ray Conference



Besocke Kelvin Probe





- Vacuum-compatible probe from Besocke delta phi GmbH
- Probe + preamplifier = \$7,100
- Modulates probe electrode position with PZT
- Reported sensitivity of 0.1 mV



Probe Calibration



- Rub acrylic surface with viton O-ring
- Measure with probe and surface DC voltmeter (shown at right)
- Convert voltmeter reading to charge density via

$$\sigma = \frac{V_{meas}}{1kV} \frac{5.04 \times 10^{-7} \, C \, / \, m^2}{1 + (2L/D)^2}$$

L = meter-to-sample distance (2.5cm), D = diameter of sample (7.6cm)

 Find charge corresponding to probe noise level

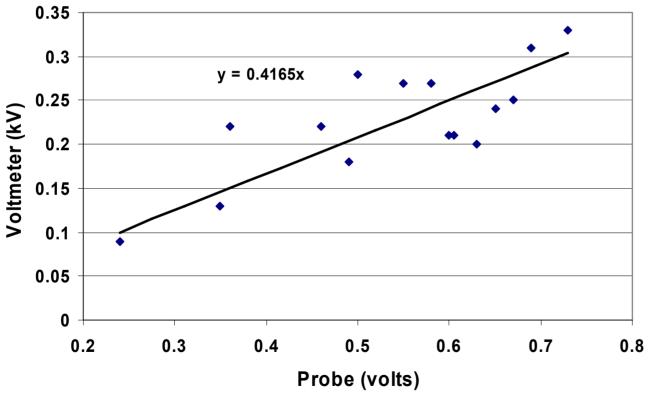




Calibration Results







For the fit shown, a probe signal of 1 volt

= 10⁻⁷ C/m²

= 8 × 10⁷ e⁻/cm²

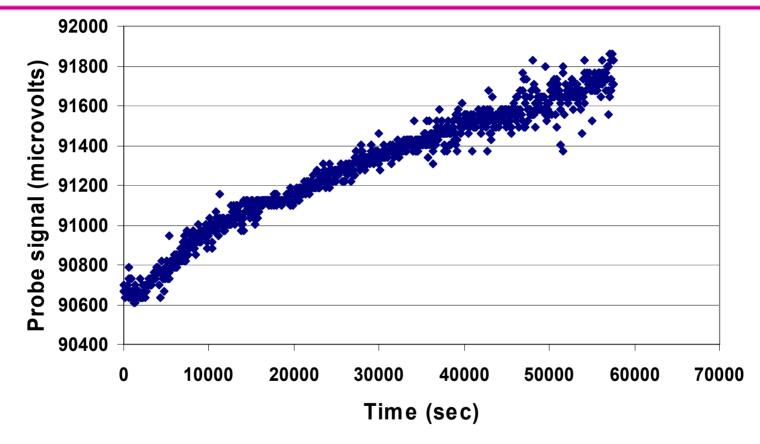
Since the probe noise level is measured to be +/- 0.1 mV, sensitivity

$$= 8 \times 10^3 e^{-1} cm^2$$

Sensitive to probe-tosample separation



Overnight Charging?



- Equivalent to ~ 10⁵ e⁻/cm²/day (sign known through UV test)
- Reduced when probe is powered off

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UV Light Source



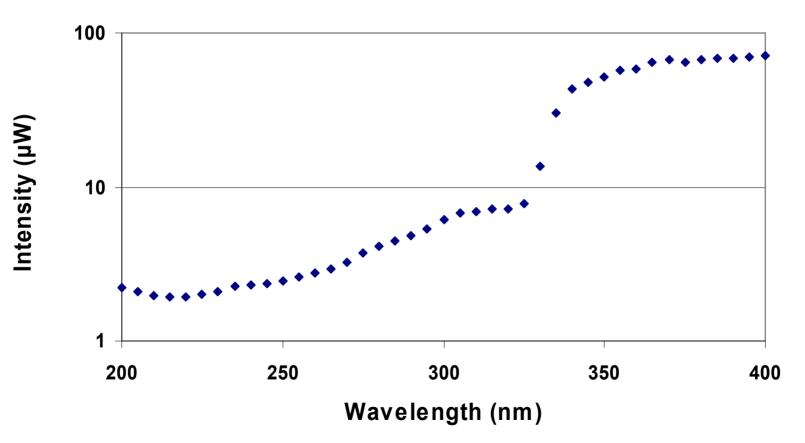


- 175W broadband Xenon lamp, 200nm 2,200nm
- 2400 lines/mm monochromator grating, 180nm-680nm range
- Intensity control through lamp knob, monochromator aperture





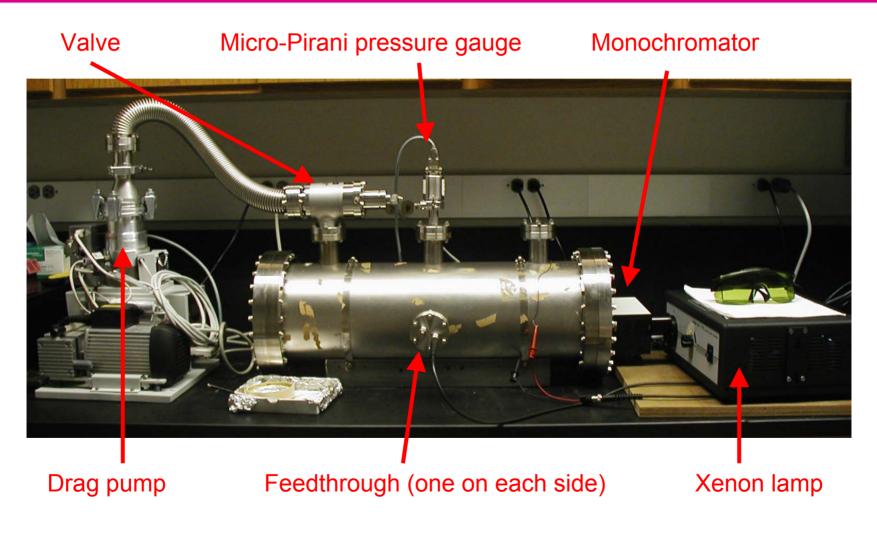
Intensity vs. Wavelength, 0.6mm Apertures



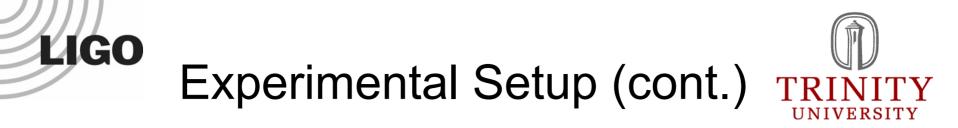


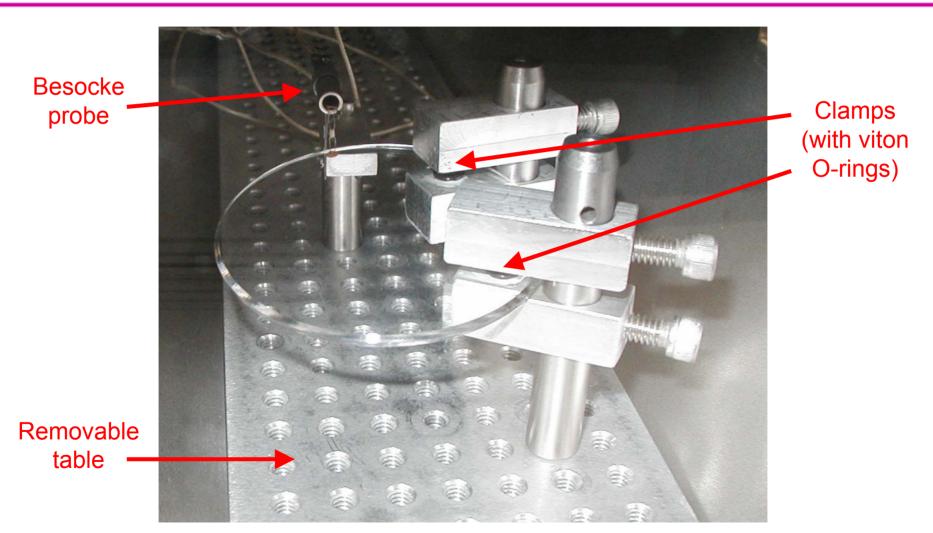
Experimental Setup





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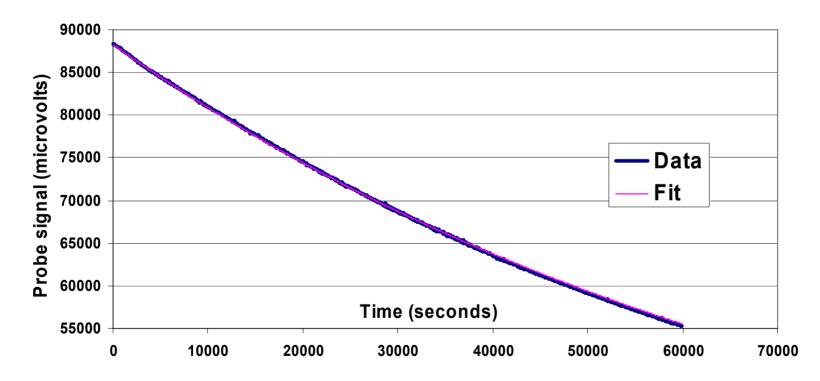




UV Discharge



Discharging at 215nm with fit curve

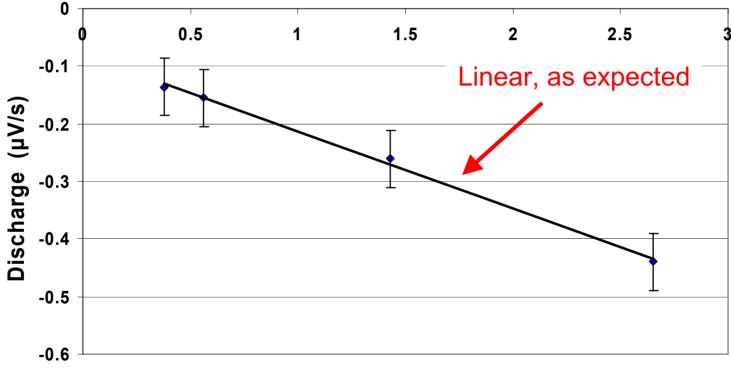


- Exp. fit implies linear relationship between charge level and discharge rate
- Gives correction for measurements taken at different charge levels



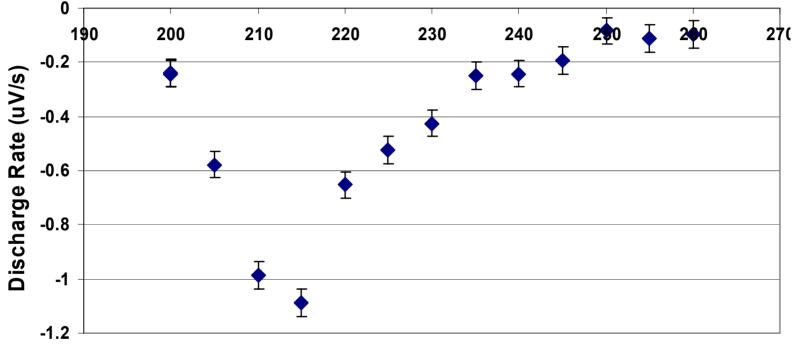


Intensity vs Discharge (varying apertures)



Intensity (µwatts)



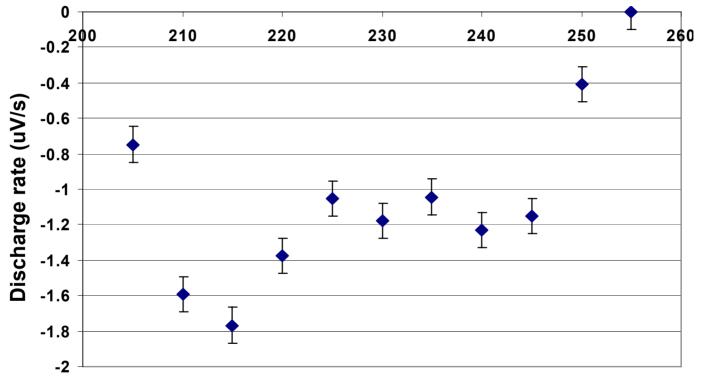


Wavelength (nm)

- Corrected for charge level, source intensity, viewport transmission
- Peak response at 215nm, surprisingly low at 255nm







Wavelength (nm)

- Greater response in 225-245nm range
- Concern same results for both sides of optic?

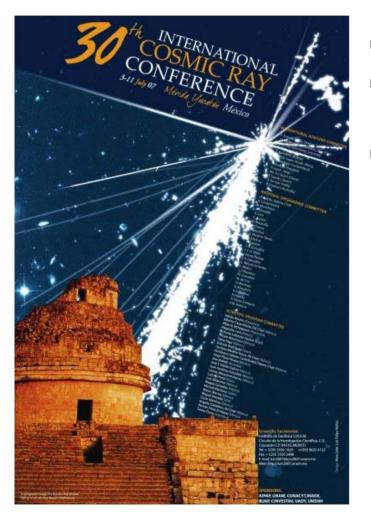




- Measure discharge rate, relaxation time, and spatial variation of charge for:
 - » Different optical coatings
 - » Different cleaning/handling techniques
- Testbed for other ideas as they come along

International Cosmic Ray Conference





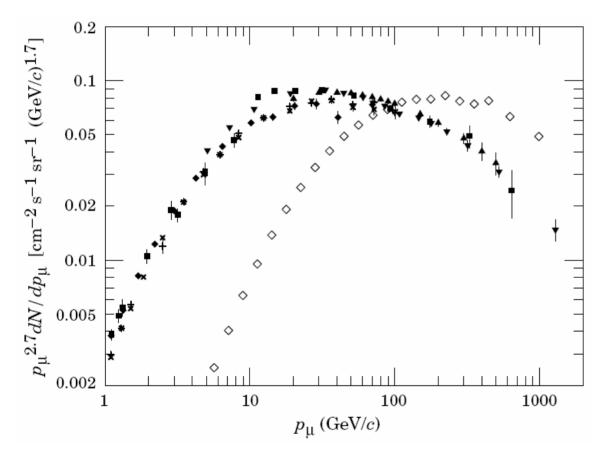
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- Charging due to cosmic rays "likely"
- GEANT model of vacuum chamber would be useful
- Until then, rough calculation of charging rate can be attempted as follows:
 - » Determine muon flux at ground (given in 2006 Review of Particle Properties)
 - » Measure thickness of vacuum chamber
 - » Estimate energy loss in iron (plots given in muon detector design documents)
 - » Convert to electrons via equation given in Braginsky et.al., Phys. Lett. **A350**, 1-4 (2006).
 - » Integrate to find total rate



Muon Flux at Ground



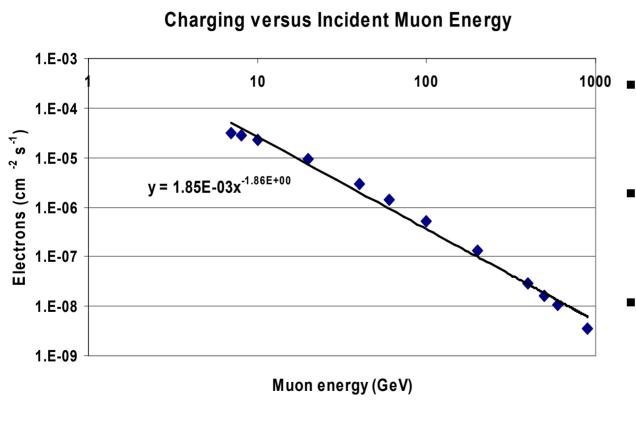


- Muon flux has cos²θ dependence
- Energy distribution also depends on angle:
 - » Solid points = 0°
 - » Empty points = 75°
- Gross assumption integrate over solid angle using distribution at zero degrees
 - » Not so bad most muons at low angles

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Rough Charging Estimate



- Assume muon passes through 1cm of iron
- Fractional energy transfer ~ 0.01
- Lower cutoff energy = E_{cr}/0.01 = 2.07 GeV
- Integration yields a charging rate of
 3 × 10³ e⁻/cm²/month,
 30 times less than
 Moscow State meas.

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