

Progress report from 40m team for the Advanced LIGO optical configuration

LSC meeting

August 15, 2005

O. Miyakawa, Caltech and the 40m collaboration

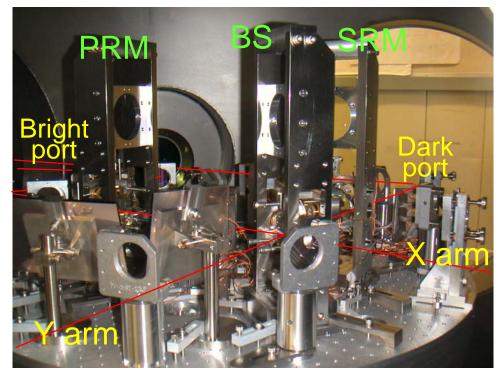
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LIGO Caltech 40 meter prototype interferometer

Objectives

- Develop lock acquisition procedure of detuned Resonant Sideband Extraction (RSE) interferometer, as close as possible to AdLIGO optical design
- Characterize noise mechanisms
- Verify optical spring and optical resonance
- Develop DC readout scheme
- Extrapolate to AdLIGO via simulation



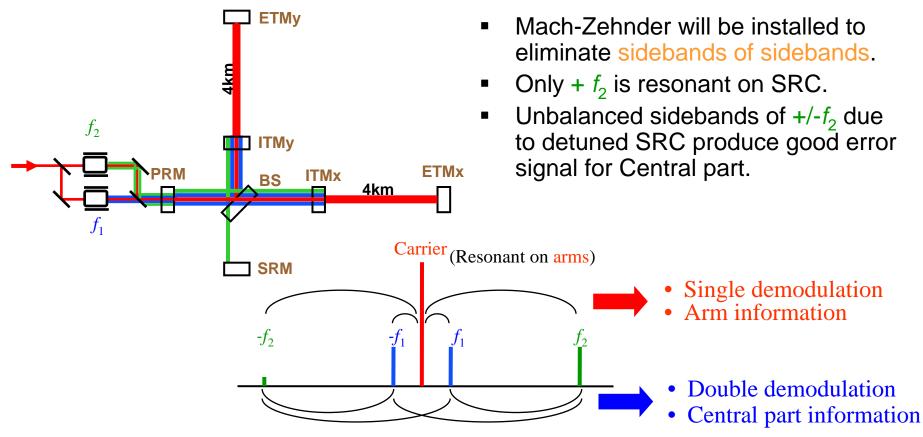


Differences between AdvLIGO and 40m prototype

- 100 times shorter cavity length
- Arm cavity finesse at 40m chosen to be = to AdvLIGO (= 1235)
 - » Storage time is x100 shorter.
- Control RF sidebands are 33/166 MHz instead of 9/180 MHz
 - » Due to shorter PRC length, less signal separation.
- LIGO-I 10-watt laser, negligible thermal effects
 - » 180W laser will be used in AdvLIGO.
- Noisier seismic environment in town, smaller stack
 - » ~1x10⁻⁶m at 1Hz.
- LIGO-I single pendulum suspensions are used
 - » AdvLIGO will use triple (MC, BS, PRM, SRM) and quad (ITMs, ETMs) suspensions.



AdLIGO signal extraction scheme



- Arm cavity signals are extracted from beat between carrier and f_1 or f_2 .
- Central part (Michelson, PRC, SRC) signals are extracted from beat between f_1 and f_2 , not including arm cavity information.



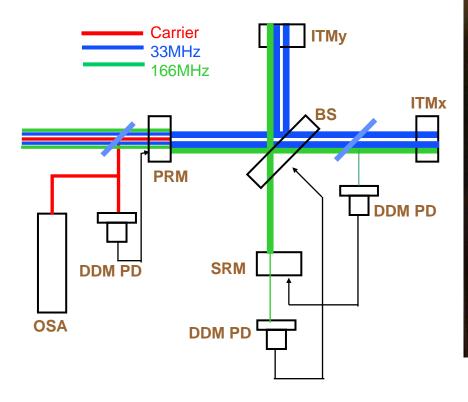
DRMI lock using double demodulation with unbalanced sideband by detuned cavity

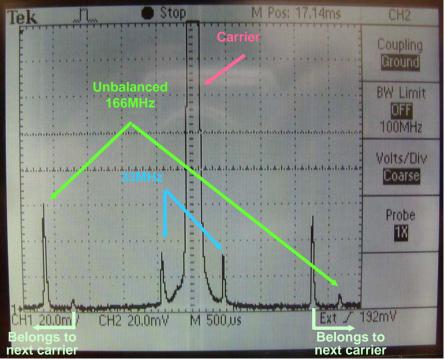
August 2004

DRMI locked with carrier resonance (like GEO configuration)

November 2004

•DRMI locked with sideband resonance (Carrier is anti resonant preparing for RSE.)





Typical lock acquisition time : ~10sec Longest lock : 2.5hour



All 5 degrees of freedom under controlled with DC offset on L_{+} loop

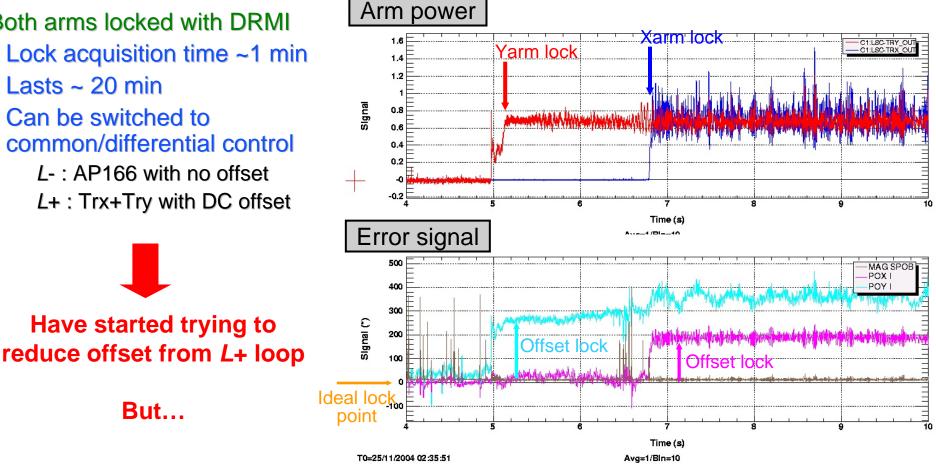
Both arms locked with DRMI

- Lock acquisition time ~1 min
- Lasts ~ 20 min
- Can be switched to common/differential control
 - L-: AP166 with no offset

Have started trying to

But...

L+ : Trx+Try with DC offset





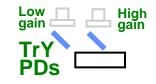
Progress in last 6 months

- For the last 6 months, we have been able to control all 5DOF, but with CARM offset.
- Reducing the CARM offset has been made difficult by technical noise sources. We have spent last 6months reducing them;

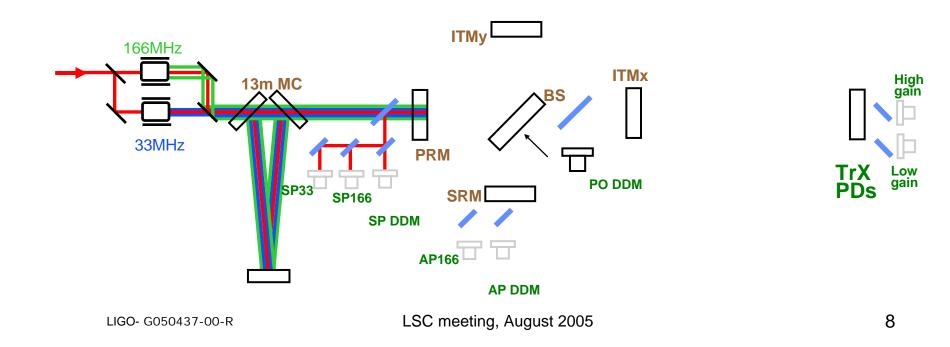
»suspension noise, vented to reduce couplings in ITMX
»improved diagonalization of all suspensions
»improved frequency noise with common mode servo
»automation of alignment and lock acquisition procedures
»improved DC signals and improved RF signals for lock acquisition.

 We can now routinely lock all 5 DOFs in a few minutes at night.

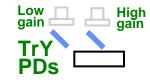




Start



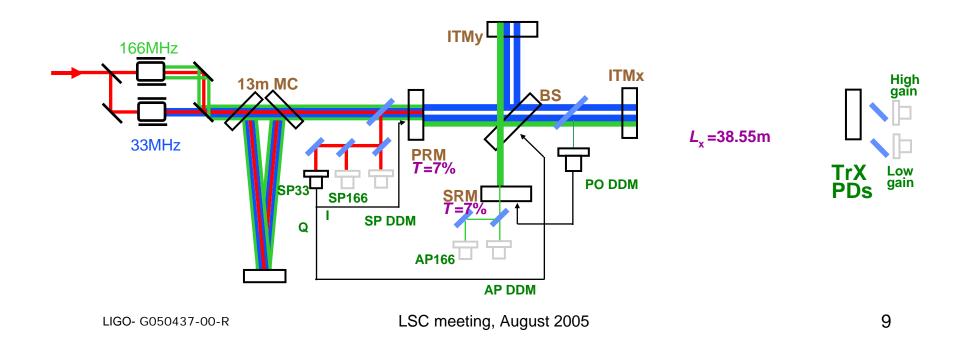




done every 10sec

DRMI

L_y=38.55m

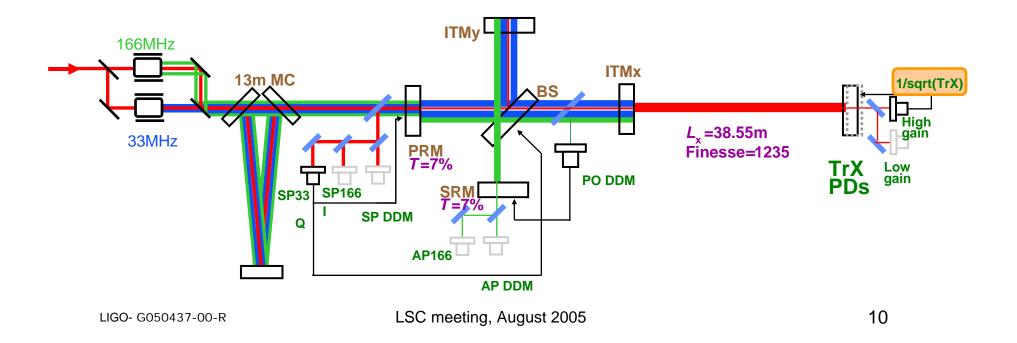




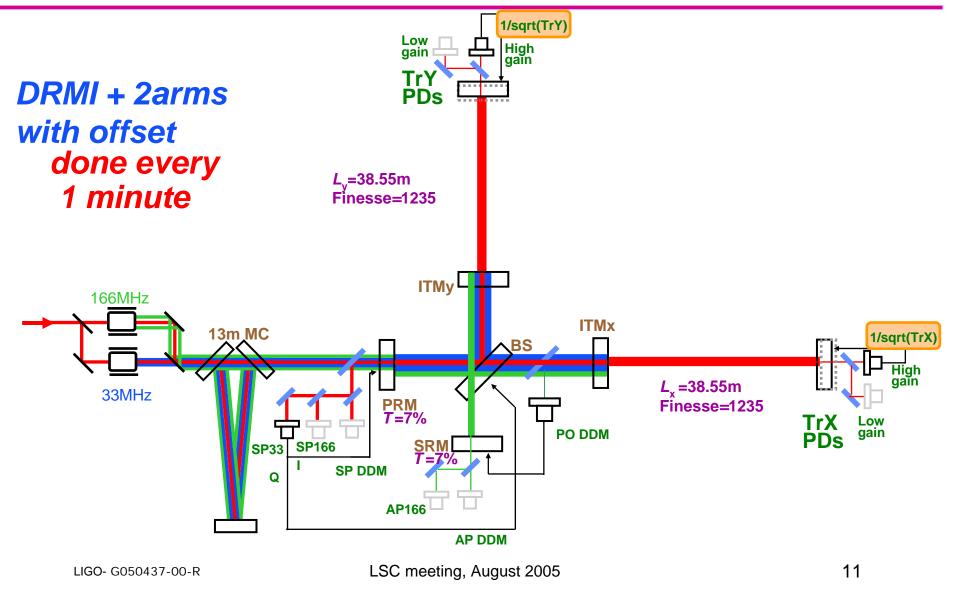
DRMI + single arm with offset done every 30 seconds



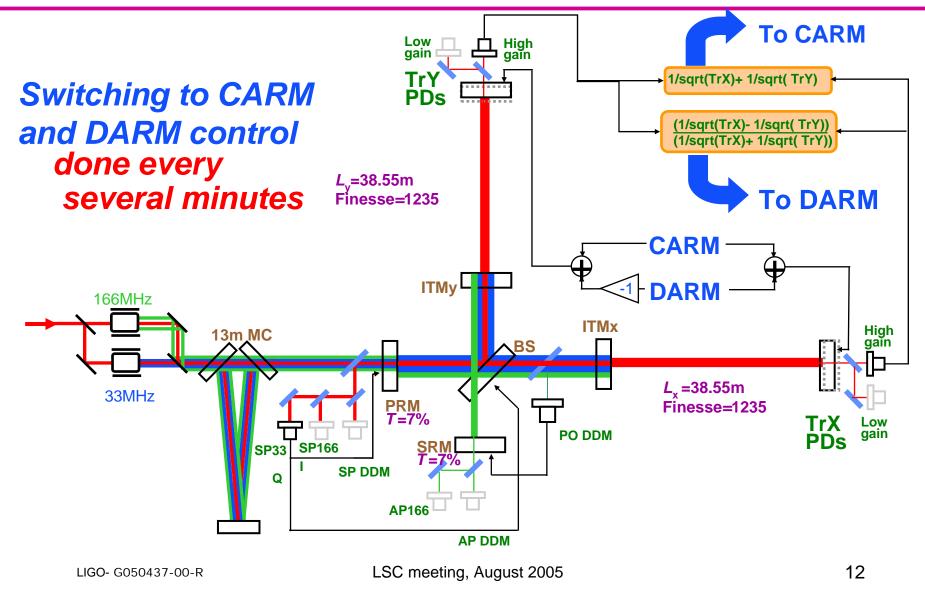
L_y=38.55m Finesse=1235



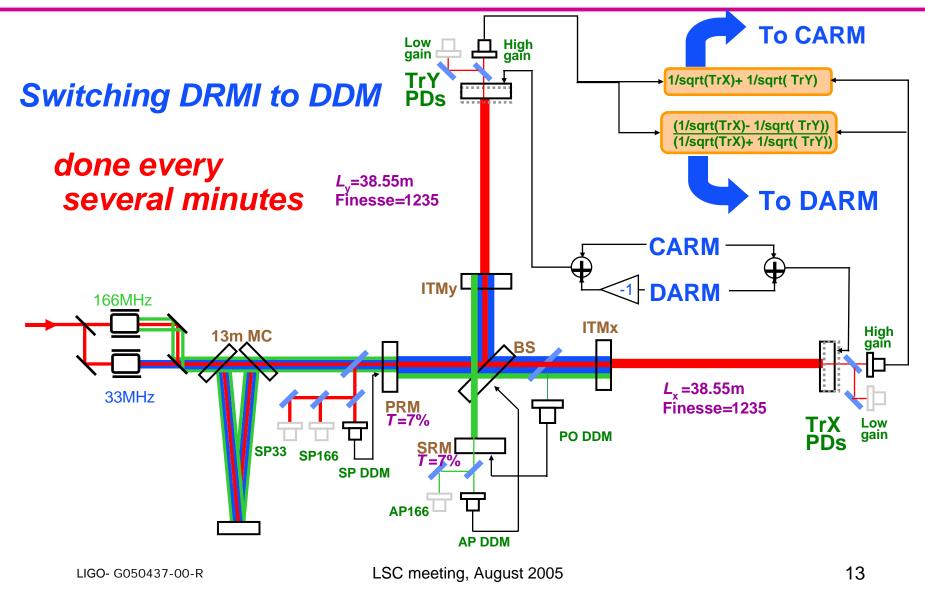




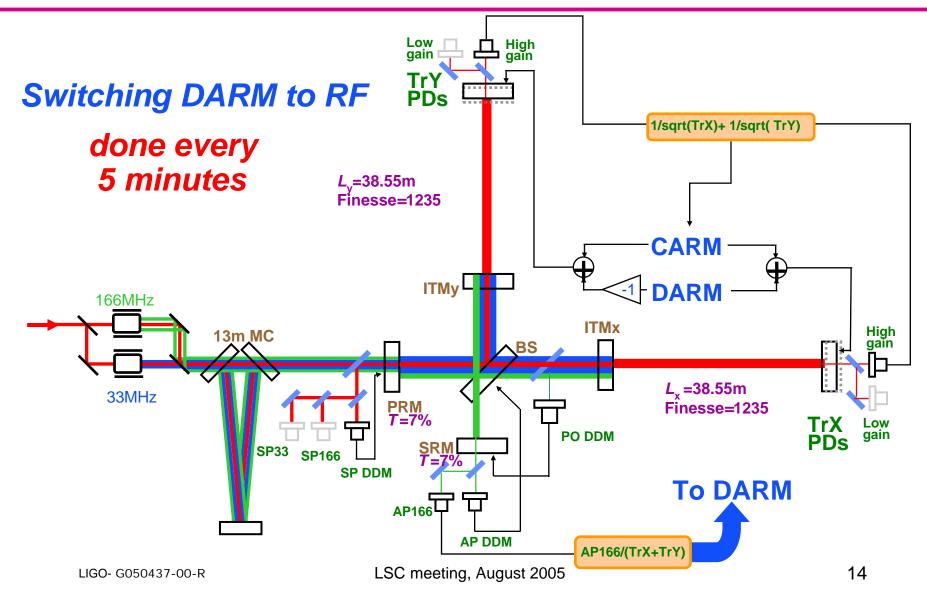




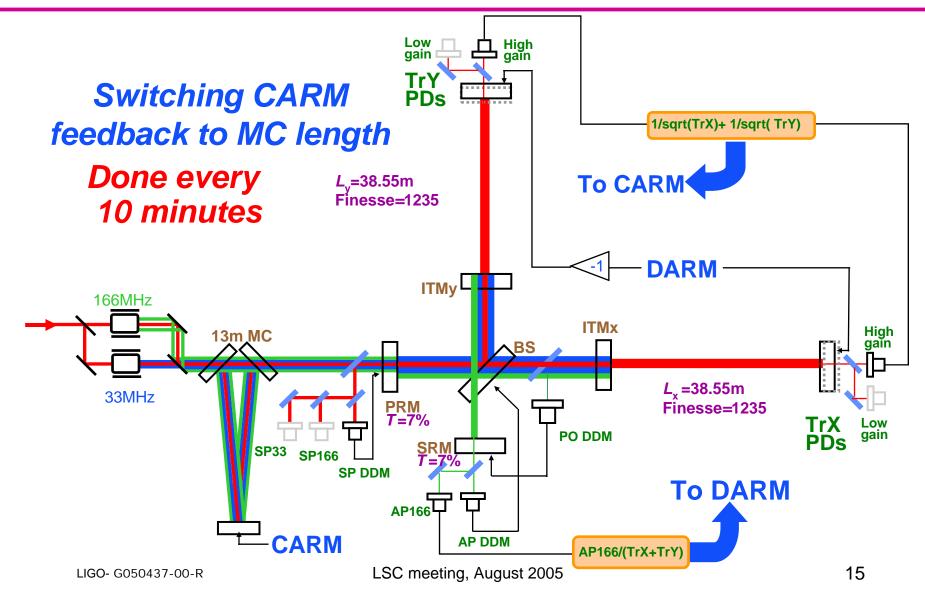




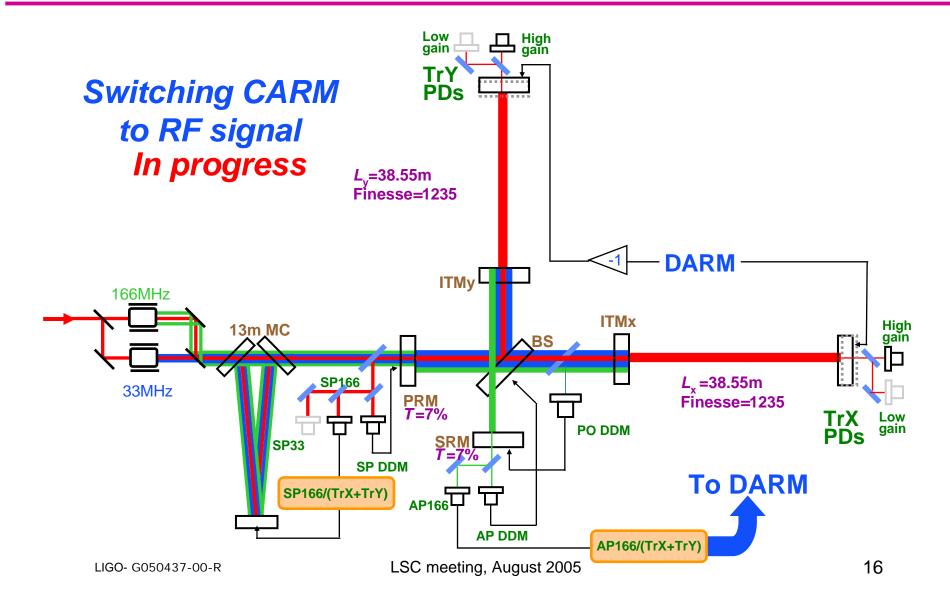






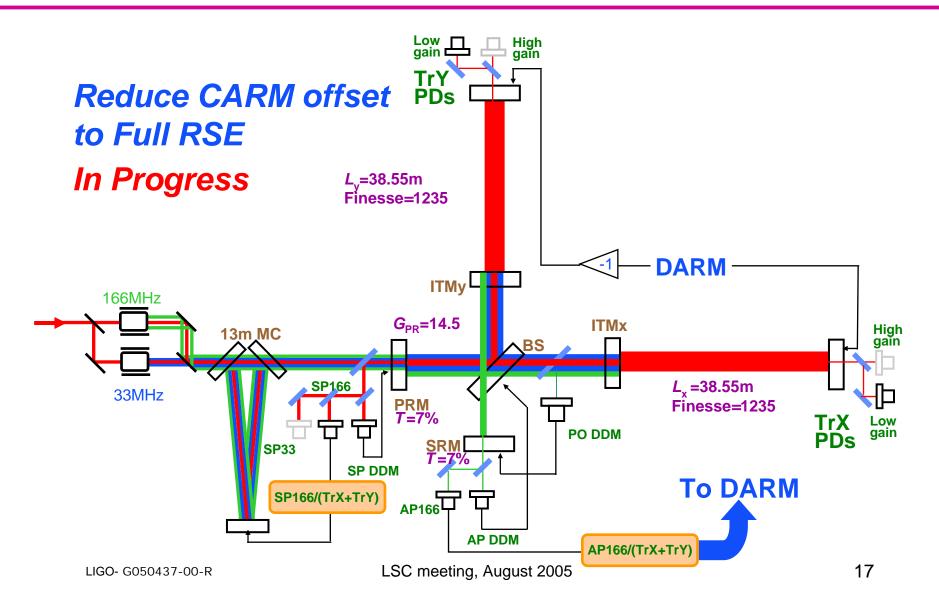








Lock acquisition procedure towards detuned RSE Ly=38.55m Finesse=1235





Differential mode of Arm : RSE peak! Common mode of Arm : small offset exists

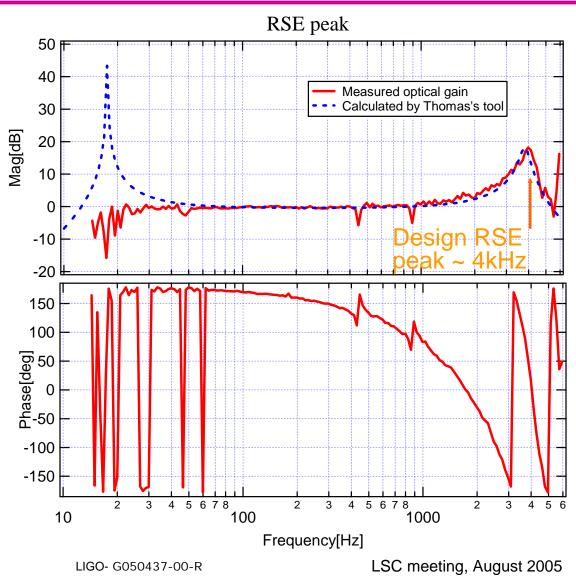
Effectively the same as low power recycling (G=1.4) RSE

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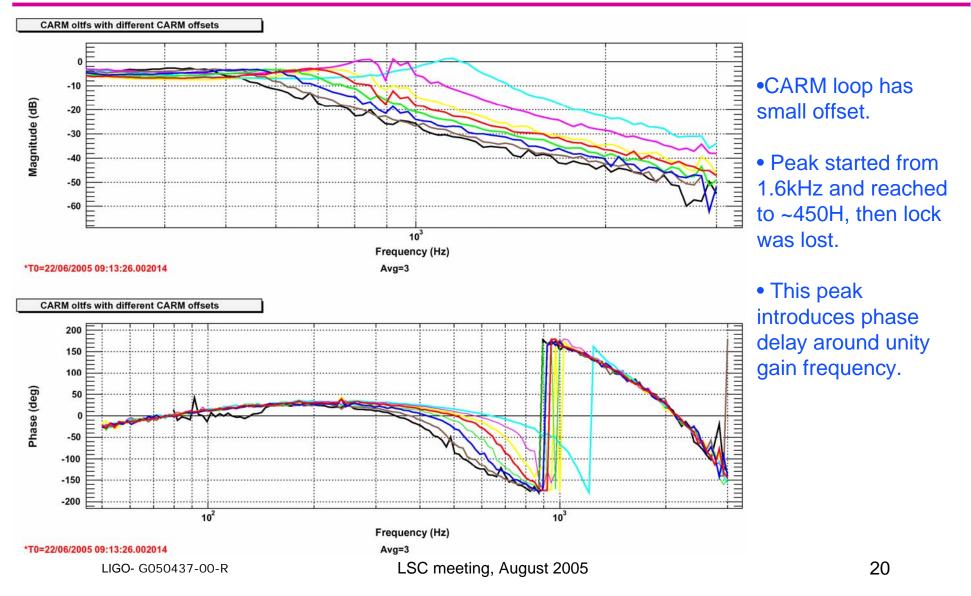
RSE peak!

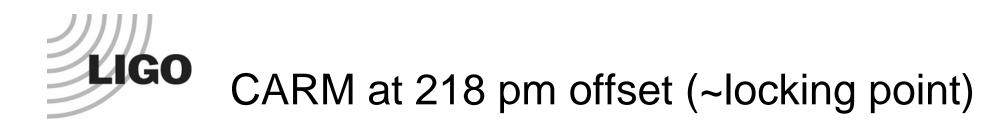


- Optical gain of *L* loop
 DARM_IN1/DARM_OUT divided by
 pendulum transfer function
- No offset on *L*-loop
- ~60pm offset on L+ loop
- Phase includes time delay of the digital system.
- Optical resonance of detuned RSE can be seen around the design RSE peak of 4kHz.
- Q of this peak is about 7.
- Effectively the same as Full RSE with GPR=1.4 with 1W input laser.
- Model was calculated by Thomas's tool.
- We will be looking for optical spring peak.

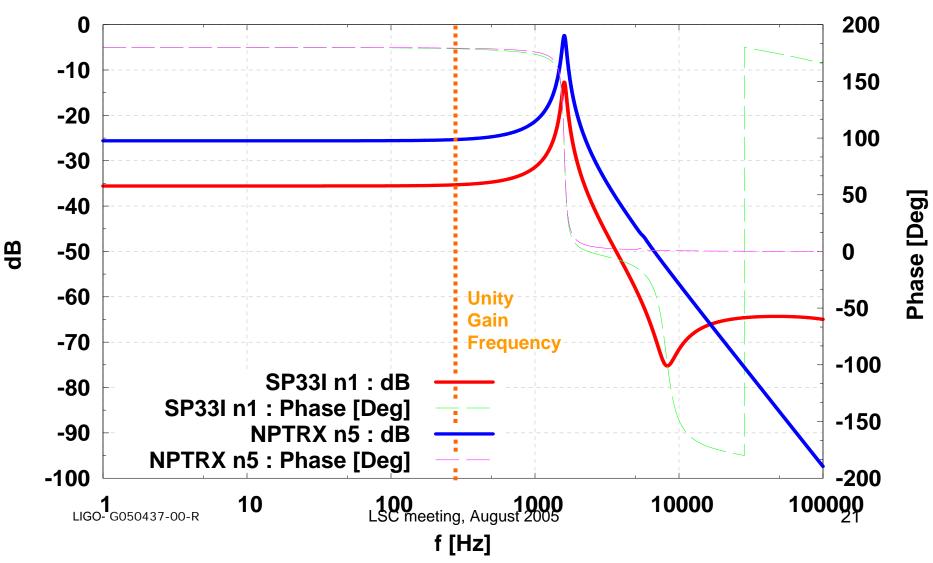


Peak in CARM loop





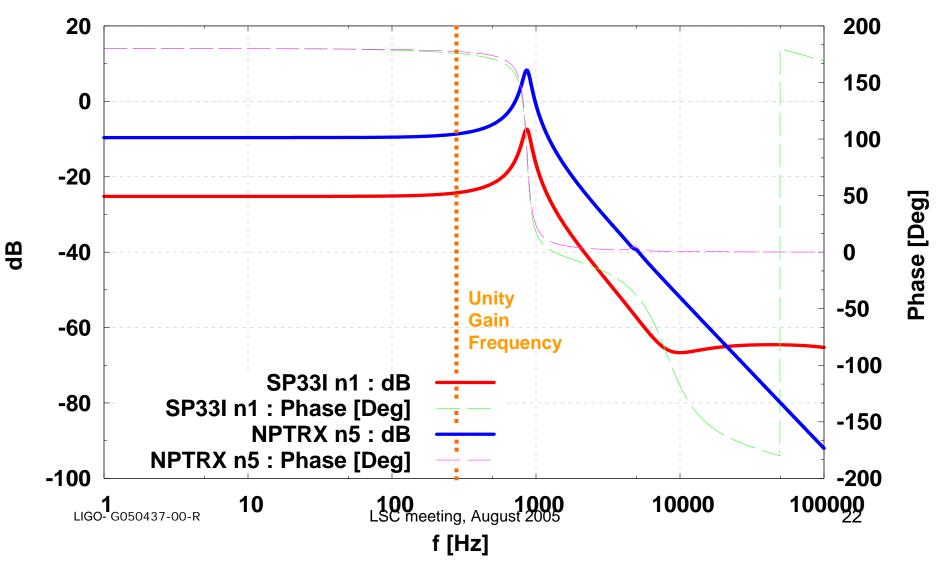
CARM optical response, CARM offset at 218 pm





CARM at 118 pm offset

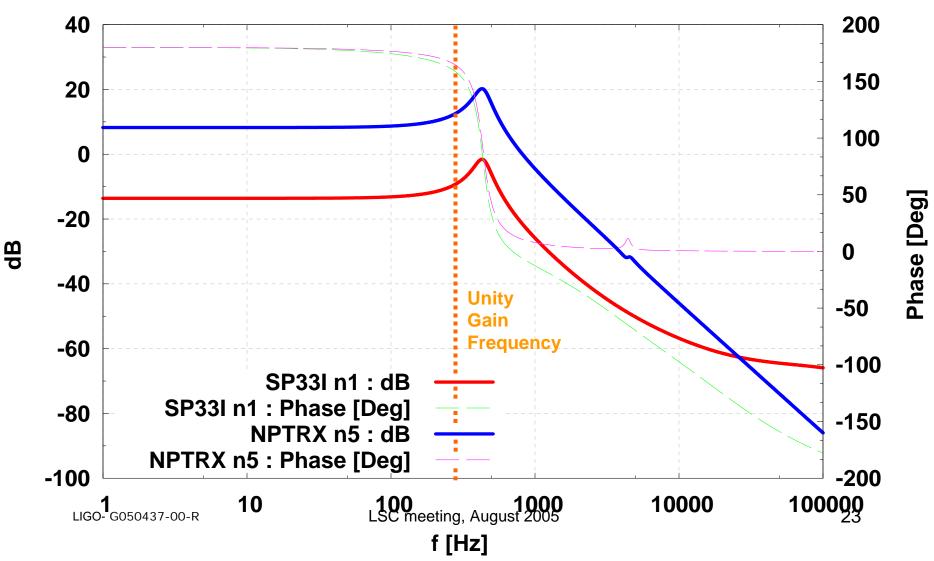
CARM optical response, CARM offset by 118 pm





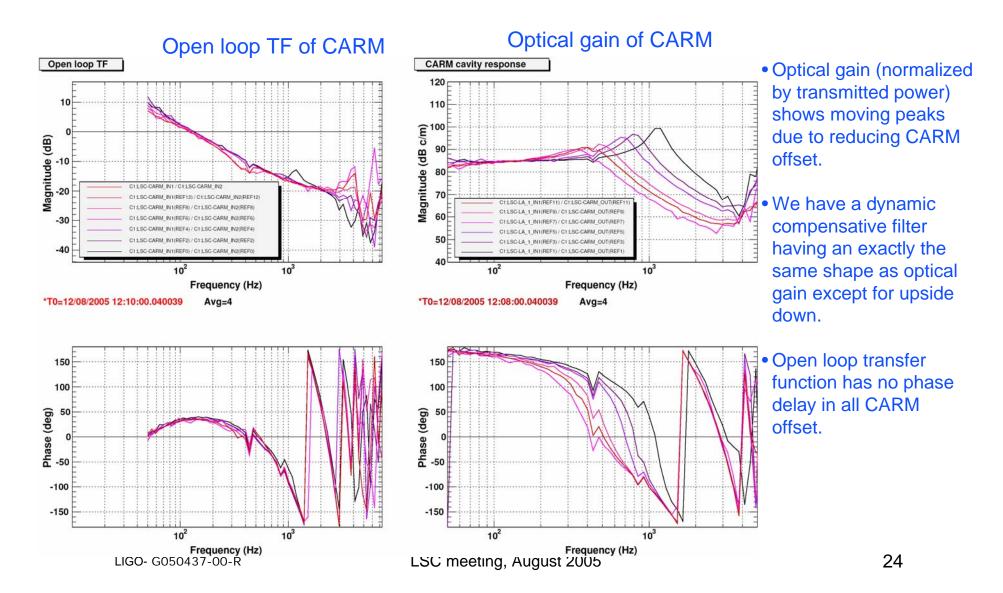
CARM offset at 59 pm (losing lock)

CARM optical response, CARM offset by 59 pm





Dynamic compensative filter for CARM servo by Rob Ward





How large is the CARM offset? Evaluation of power recycling gain

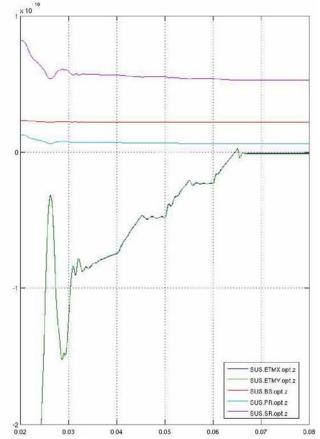
	Design	Measured(calculated)	
Cavity reflectivity	93%	85%(X arm 84%, Yarm 86%)	
PRM reflectivity	93%	92.2%	
Loss in PRC	0%	2.3%	
Achievable PRG	14.5	5.0	
Coupling	Over coupled	Under coupled	

 Estimated actual power in arms with 1W input and smallest CARM offset achieved is
 0.5kW of ~2kW(25% of the way to full resonance). LIGO

CARM switching from DC to RF signal

E2E simulation by Matthew Evans in June 2005

- CARM moved to RF signal
 - » Not yet done at 40m
 - » REFL port HF demod (a.k.a., SP166)
 - » Normalized by arm power
 - » Offset and gain matched at hand over
 - » Offset swept to zero slowly
 - » Coupled-Cavity pole compensation required
 - Pole (actually more complicated) moves down as resonance is approached
 - Compensation filter uses sum to make a "moving zero"
 - More detail may be required for 40m



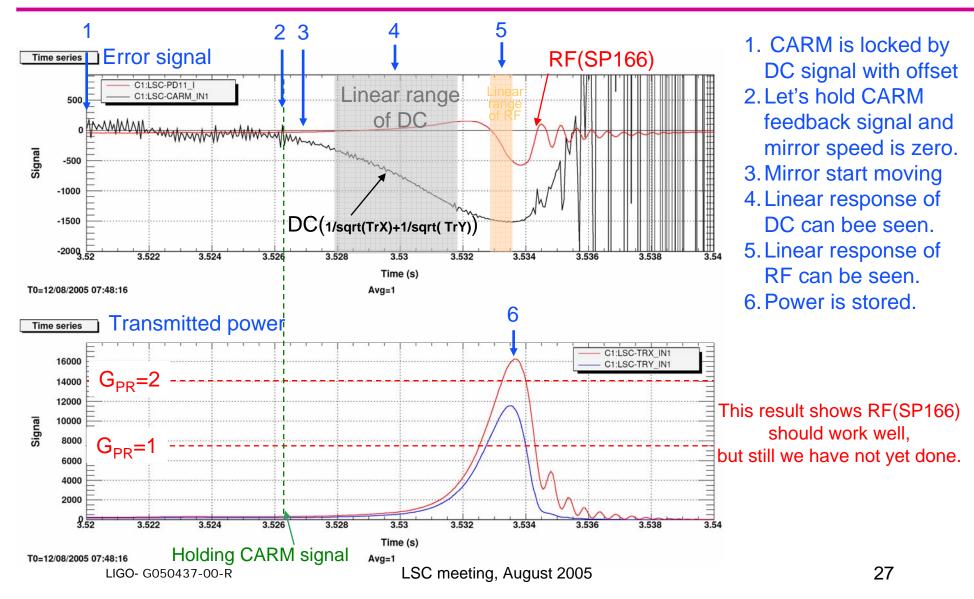
Simulation verifies that controlled reduction of CARM offset should work.

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CARM switching from DC to RF signal



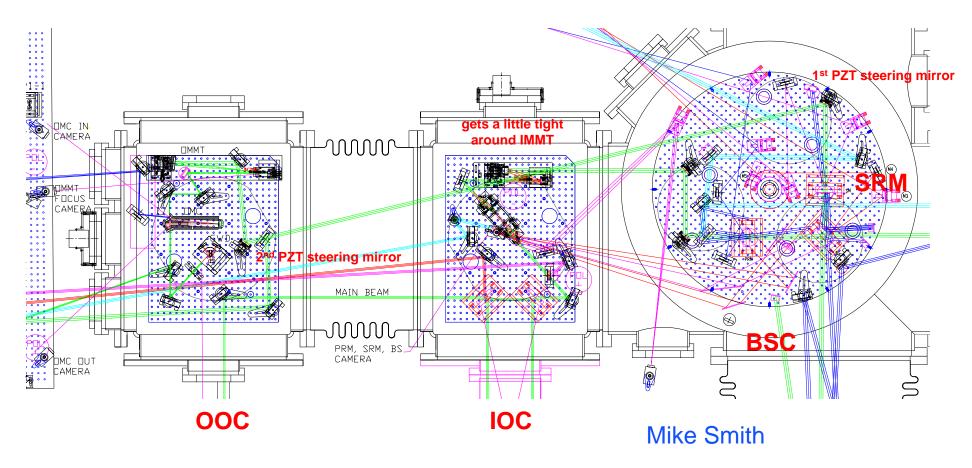


DC Readout at the 40m

- DC Readout eliminates several sources of technical noise (mainly due to the RF sidebands):
 - Oscillator phase noise
 - Effects of unstable recycling cavity.
 - The arm-filtered carrier light will serve as a heavily stabilized local oscillator.
 - Perfect spatial overlap of LO and GW signal at PD.
- DC Readout has the potential for QND measurements, without major modifications to the IFO.
- We may not be able to see shot noise at low frequency, given our noise environment. We may not even see any noise improvements, but we might!
- The most important thing we will learn is : How to do it
 - » How to lock it?
 - » How best to control the DARM offset?
 - » What are the unforeseen noise sources associated with an in-vacuum OMC?
 - » How do we make a good in-vac photodiode? What unforeseen noise sources are associated with it?



Output Optical Train

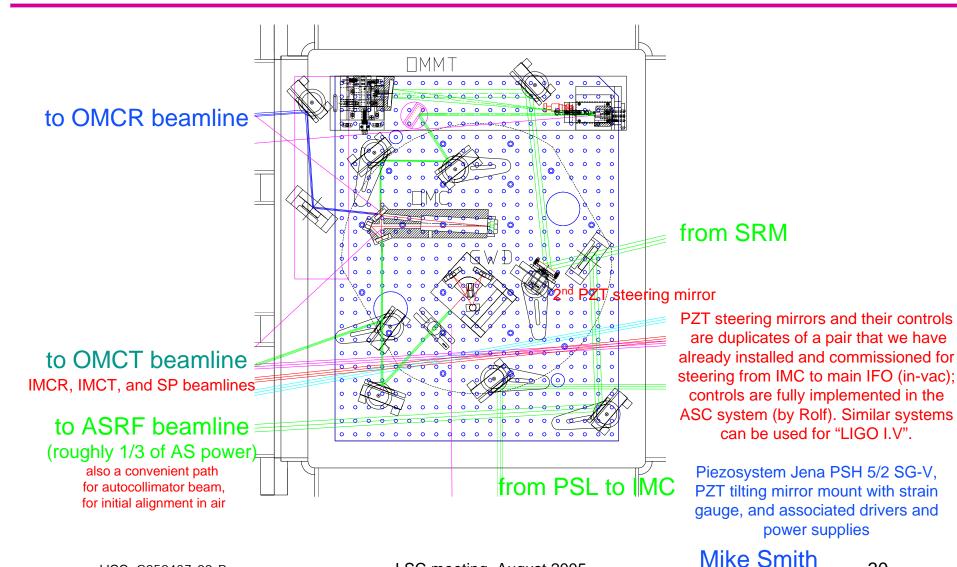


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Output Optic Chamber



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30

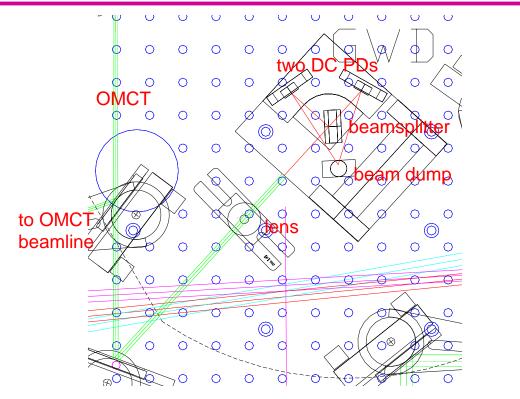


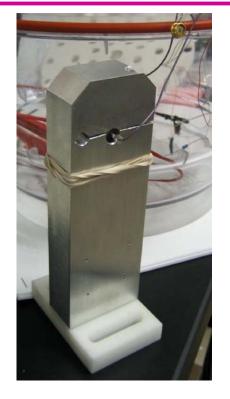
OMMT layout

Primary radius of curvature, mm	618.4	
Secondary radius of curvature, mm	150	
Defocus, mm	6.3	
Input beam waist, mm	3.03	
Output beam waist, mm	0.38	
Make mirror(s) by coating a co to get larger selection of RC	Mike Smith	
		01



Two in-vac DC PDs



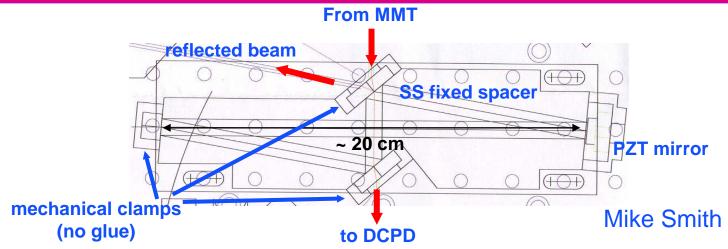


- Ben Abbott has designed an aluminum stand to hold a bare photodiode, and verified that the block can radiate 100 mW safely.
- A small amplifier circuit will be encased in the stand, and vacuum-sealed with an inert, RGA detectable gas.
- Two such assemblies will be mounted together with a 50:50 beamsplitter to provide in-loop and out-of-loop sensors.

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- Mirrors mounted mechanically, on 3 points (no glue)
- curved mirror: off-the-shelf CVI laser mirror with ROC = 1 m \pm 0.5%
- Fixed spacer should be rigid, vented, offset from table
- OMC length signal:
 - Dither-lock? >> Should be simple; we'll try this first.
 - PDH reflection? >> There's only one sideband, but it will still work.
- Servo:

LIGO

- Will proceed with a simple servo, using a signal generator and a lock-in amp.
- Feedback filters can easily be analog or digital.
 - Can use a modified PMC servo board for analog.
 - Can use spare ADC/DAC channels in our front end IO processor for digital.
- PZT actuation

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