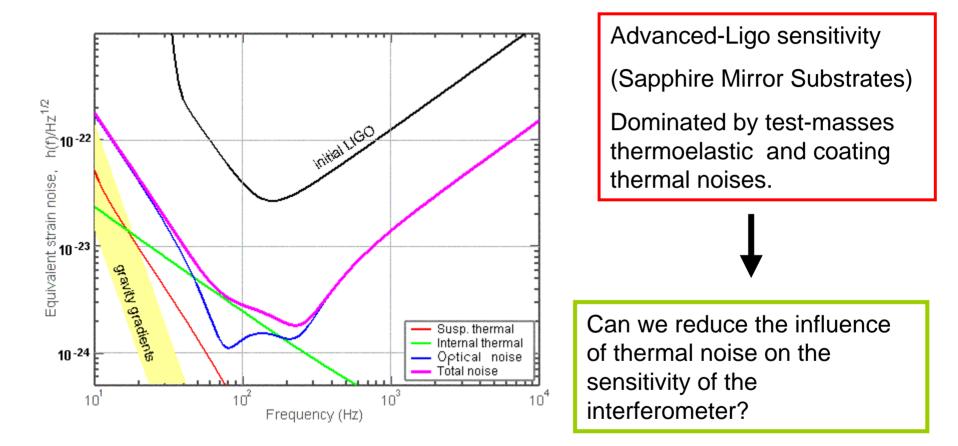


# Flat-Top Beam Profile Cavity Prototype

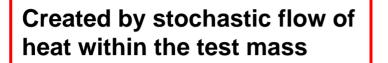
J.Agresti, **E.D'Ambrosio**, R. DeSalvo, J.M. Mackowsky, M. Mantovani, A. Remillieux, B. Simoni, P. Willems

LIGO-G040412-00-D

### **Motivations for a flat-top beam:**



#### **Thermoelastic Noise:**



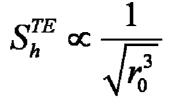
Fluctuating hot spots and cold spots inside the mirror

Expansion in the hot spots and contraction in the cold spots creating fluctuating bumps and valleys on the mirror's surface

Mirror surface Fluctuating bumps

Interferometer output: proportional to the test mass average surface position, sampled according to the beam's intensity profile.

#### Gaussian beam



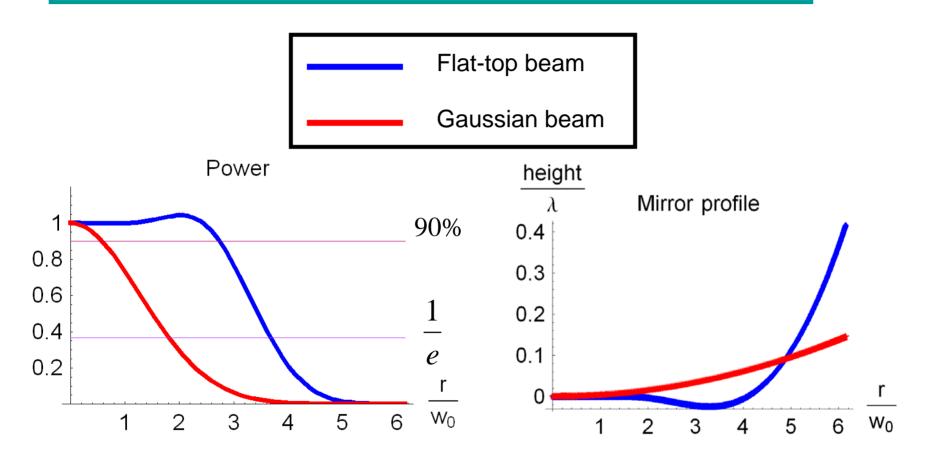
Mirror surface

*r*<sub>0</sub> As large as possible (within diffraction loss constraint).

The sampling distribution changes rapidly following the beam power profile

Flat-Topped Beam

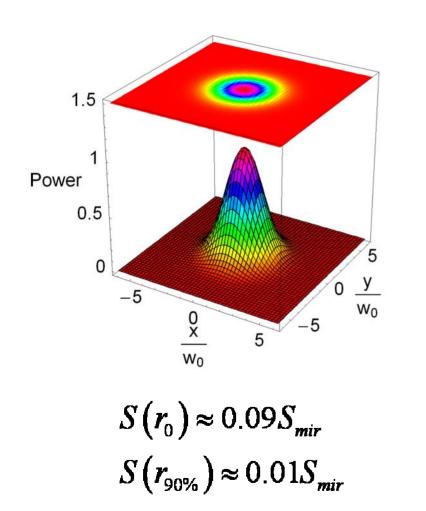
Larger-radius, flat-top beam will better average over the mirror surface. Diffraction prevents the creation of a beam with a rectangular power profile...but we can build a nearly optimal flat-top beam:

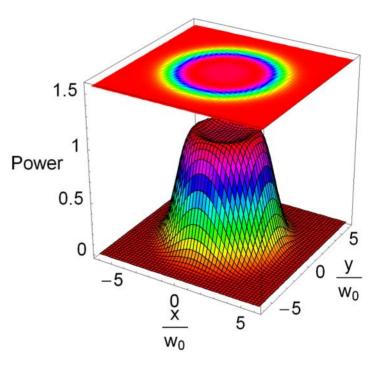


•The mirror shapes match the phase front of the beams.

#### Comparison between the two beams

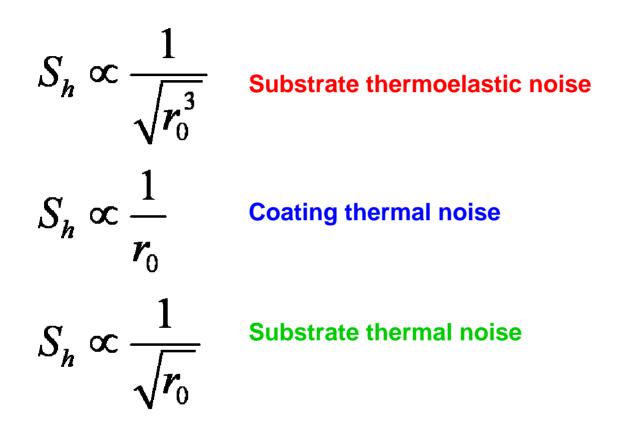
(Same diffraction losses)





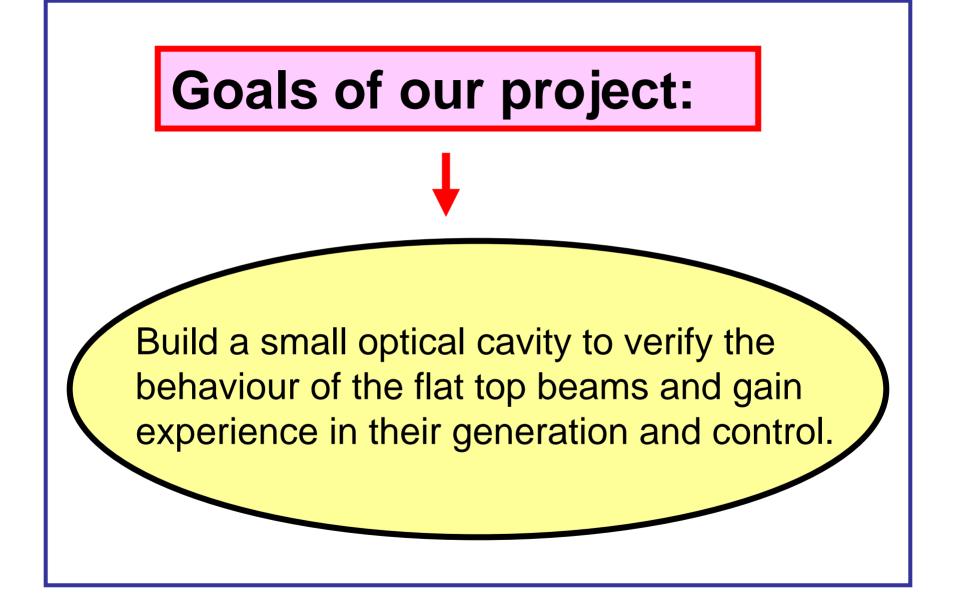
$$S(r_0) \approx 0.36S_{mir} \qquad R = 4$$
$$S(r_{90\%}) \approx 0.20S_{mir} \qquad R = 20$$

#### Indicative thermal noise suppression trends

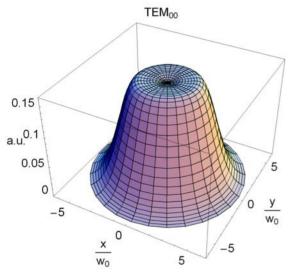


Exact results require accurate information on material properties (Q-factors)

Expected gain in sensitivity ~ 3

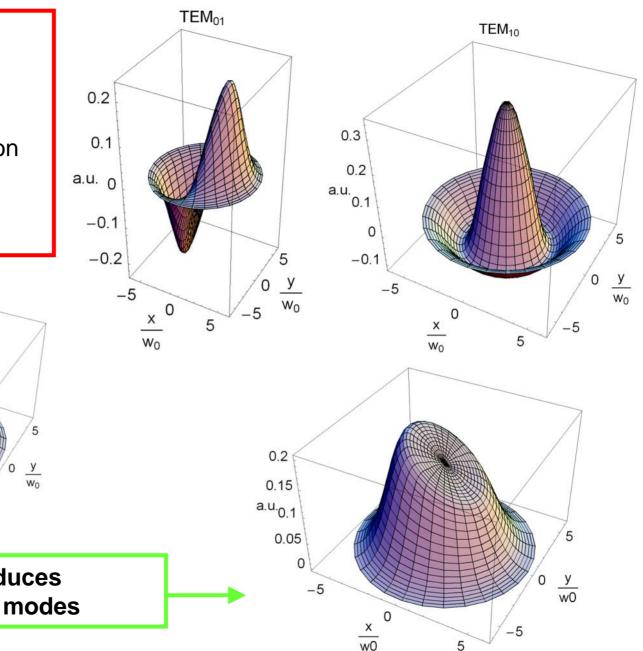


We will investigate the **modes structure** and characterize the **sensitivity to perturbations** when non Gaussian beams are supported inside the cavity.



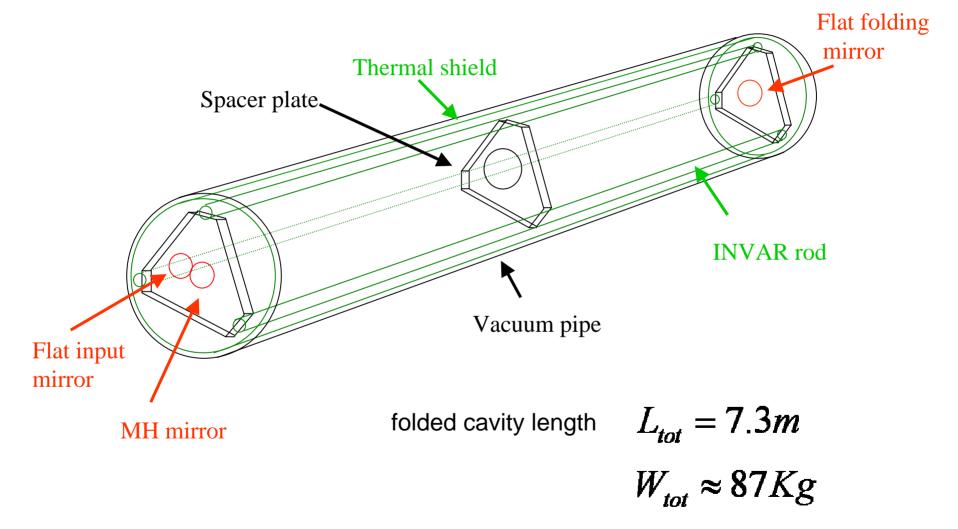
Misalignment produces coupling between modes

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#### Design of the test cavity : Rigid cavity suspended under vacuum



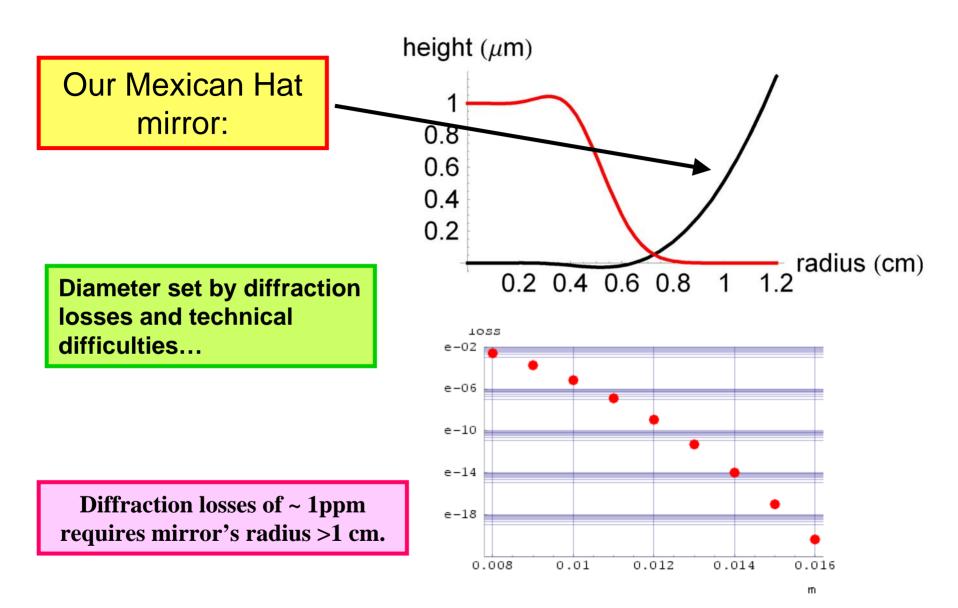
#### **Optical and mechanical design:**

- Injection Gaussian beam designed to optimally couple to the cavity.
- Required finesse  $\mathcal{F}$  = 100 to suppress Gaussian remnants in the cavity.

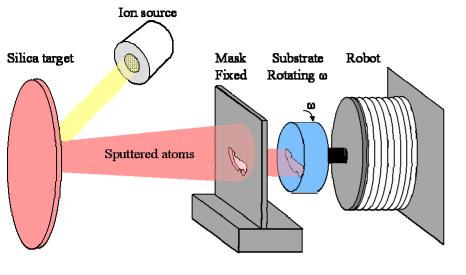
Length stability: ~ 5 nm

- INVAR rods (low thermal expansion coefficient).
- Stabilized temperature.
- Vacuum eliminates atmospheric fluctuations of optical length.
- Ground vibrations can excite resonance in our interferometer structure: suspension from wires and Geometrical-Anti-Spring blades.

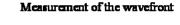
Mirror's size constrained by beam shape and diffraction losses

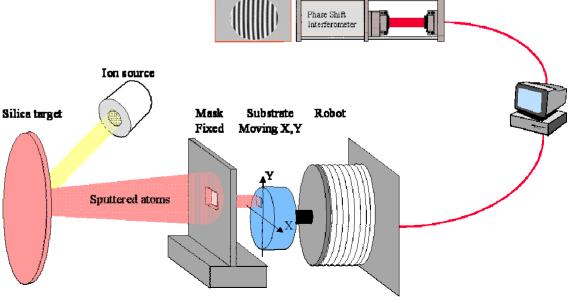


#### LMA's Technique to build Mexican Hat mirrors



- Rough Shape Deposition:
- Coating the desired Mexican Hat profile using a pre-shaped mask
- Achievable precision ~60nm Peak to Valley



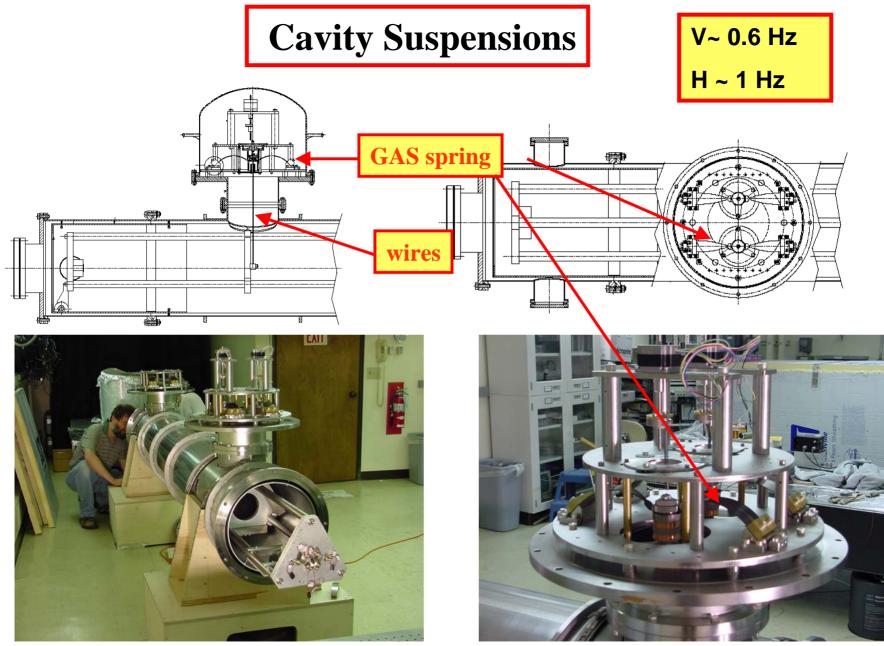


• Corrective coating:

- Measurement of the achieved shape
- Coating thickness controlled with a precision <10 nm.

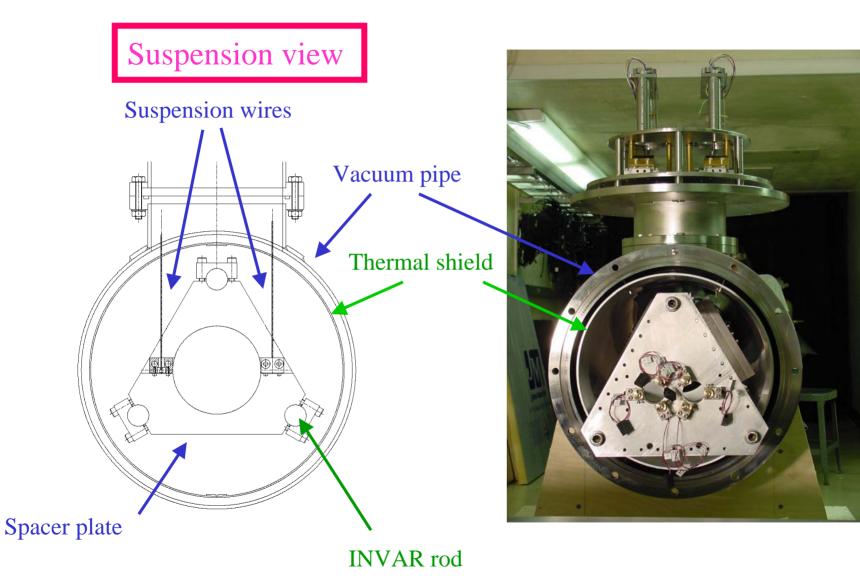
Maximum slope ~ 500nm/mm

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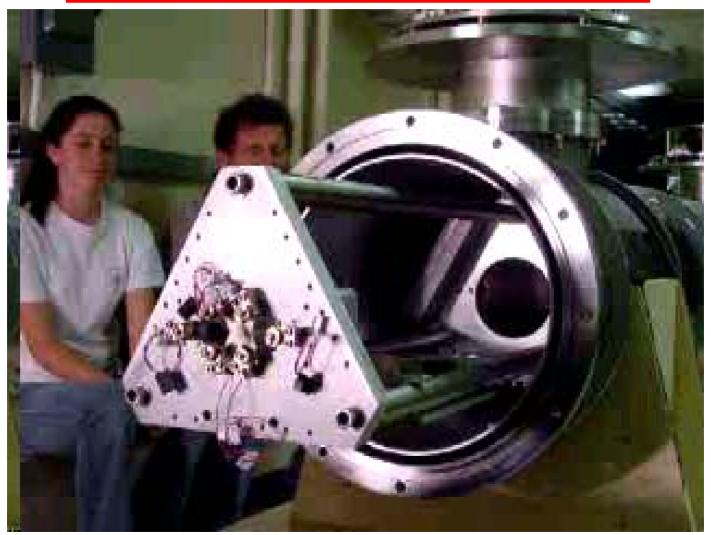


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### Cavity Vacuum & Thermal Shield



# Suspension at work!



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## Schedule for the future

- Complete building the cavity including the optics and the electronics
- Lock the cavity with spherical mirror (test the apparatus)
- Switch to Mexican-Hat mirror as soon as available
- Characterization of Flat-top beam modes and misalignment effects

Next possible developments

Flat topped beam inside a nearly-concentric cavity: same power distribution over the mirrors but less sensitive to misalignment.

Overcome the technical limitation on the slope of the coating...not impossible.