Minimizing the Resonant Frequency of MGAS Springs for Seismic Attenuation System in Low Frequency Gravitational Waves Interferometers

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Maddalena Mantovani, Juri Agresti, Erika D'Ambrosio, Riccardo De Salvo, Barbara Simoni

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Motivations



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Ground Based Interferometer are Limited by Newtonian Noise under 10 Hz Seismic Attenuation was Designed to Match that Limit

Recent Calculations Show the Possibility of Suppression of Newtonian Noise by Suspending the Test Masses in Deep Caves

LIGO Cella Suppression of Newtonian Noise



Suppression of NN by a Factor of 10⁻⁶ in Amplitude, 30 in Frequency Seismic Attenuation Must be Redesigned to Match the New Limit

Horizontal Achievable with Longer Wires in Wells Vertical Attenuation Requires New Development



Horizontal Attenuation



Inverted Pendula





Existing MGAS Spring





Spectrum Analyzer



Spectrum Analyzer









Characterization Work

- Determination of the Working Point
- Circuit Calibration
- Frequency behavior as a function of the Gain
- Q Factor Analysis
 - Impulse Response
 - Frequency Response

After fixing the vertical height to the minimum frequency point we change the Gain to find the frequency behavior vs. gain

	Frequency [Hz]	
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gain [turns]

Q Factor Analysis; Two Methods

Fitting the Damped Exponential in Free Oscillations

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 $Q = \tau \cdot v$

Fitting the Resonant Width in the Transfer Function







Q factor



(filter body to payload)



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Transfer Function with Different Gain values



Lowering the system stiffness

As the Transfer Function is shifted to lower frequencies, The Q factor decreases



Frequency [Hz]

Q Factor Analysis



Both methods show a quadratic behavior for the Q factor

Implies structural, non viscous, damping

Q factor

Need for Thermal Compensation



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$$\frac{dY}{dT}: -\frac{g}{L\omega_y^2}\alpha_E$$

Where α_E is the thermal expansion coefficient given by the Young's modulus

Virginio Sannibale et al.

The large thermal dependence of the blade working point with temperature, at low stiffness, makes it practically impossible to obtain very low resonant frequencies

Overnight Blade Tip Position stability Integrated on 10 Period

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These residuals are mainly the r.m.s. 100 mHz oscillator resonance excited by the seismic activity

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Power Considerations



The Power consumption is of the order of "10mW

LIGO Two Anomalies at Low Frequency are Discovered

1. Slower Slope in the Lower Frequency Transfer Function Measurements



Transfer Function Measurements with different Excitation Amplitude

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The slope does not change

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Transfer Function Slope for Different Integration Times



Transfer Function with Different Integration Times



The Slope does not depends on the integration time

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Frequency [Hz]

Slope for Different Vertical Positions



Second anomaly founded



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Resonat Frequency [mHz]

• So far the working point to frequency dependence has always been observed to be hyperbolic, getting narrower at lower frequency

• At lower frequency we find a departure from the hyperbolic behavior (the fit in the bottom data set is a fourth power polynomial)

• We suppose that higher order effect may be taking over when the anti-spring gets close to fully neutralize the spring constant

• The newly observed behavior is coincidental with the departure from 1/f2 of the attenuation transfer function

• We do not know yet if the two anomalous effects are correlated.

Is this useful for LIGO??



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the LIGO Deep Fall Back solution prototype for one-pier preattenuators

DFBS prototype







Real time



DFBS Prototype

So far driven down to 120mHz despite the additional springs (4/3 of the equivalent bellow stiffness) in fully passive configuration

Expect \leq 30mHz if tuned with electromagnetic anti-spring

attenuation plateau at 10⁻³ proven on preceding prototypes (expect similar performance)

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Conclusions

• This development of Electromagnetic Correction Springs was designed to boost the Low Frequency attenuation performance of Geometric Anti Springs for future Underground Low Frequency Gravitational Wave Interferometric Detectors (Horizontal Attenuation was always easily available)

• This development allows for further depression of seismic perturbations, beyond the performance of the Superattenuators, reaching down to 1 Hz and even below.



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Conclusions

• The retrofitting of this technology on existing system like the LIGO deep fall back pre-isolator solution may allow the introduction of attenuation factors as large as one thousand for frequencies above 1 Hertz

• Sizeable attenuation at the micro seismic peak at 150 mHz can be obtained as well



Next Steps

• Find the source of the less then $1/f^2$ behavior

- Repeat the Measurements with Accelerometers
- Study and Reduce the Control Circuit Noise Improving some of its Performances (Yanyi Chen)
- Make a Mechanical Thermal Compensation
- Use this Seismic Attenuation System on the Mexican Hat Interferometer (Barbara Simoni)

