

Reducing the Newtonian Noise above and below the ground



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Introduction

In future advanced interferometers Newtonian (gravity gradient) noise will be one of the fundamental limitations for the sensitivity in the low frequency region.

- Can it be estimated?
- What are the most important sources?
- Can it be reduced?



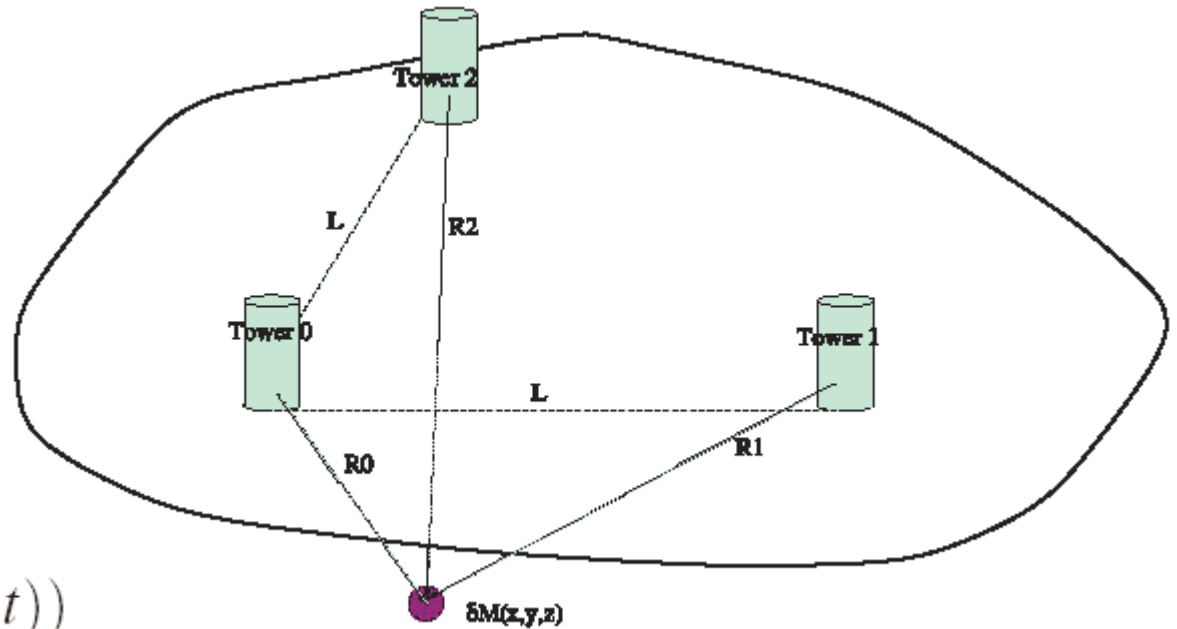
Plan of the talk

- Generalities: estimation of NN
- Models for seismic NN
- Atmospheric NN
- “Software” reduction of NN: subtraction
- “Hardware” reduction of NN: going underground

What is Newtonian Noise

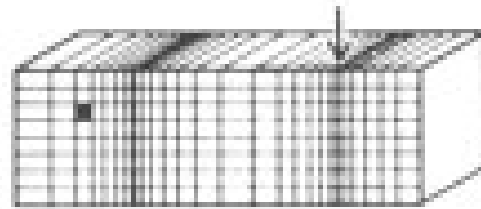
Mass density fluctuations couple directly to the test masses:

Example:
elastic
material:

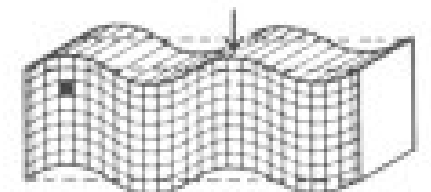


$$\delta\rho(x,t) = -\vec{\nabla} \cdot (\rho(x,t) \vec{u}(x,t))$$

$$= -\rho \vec{\nabla} \cdot \vec{u}(x,t)$$

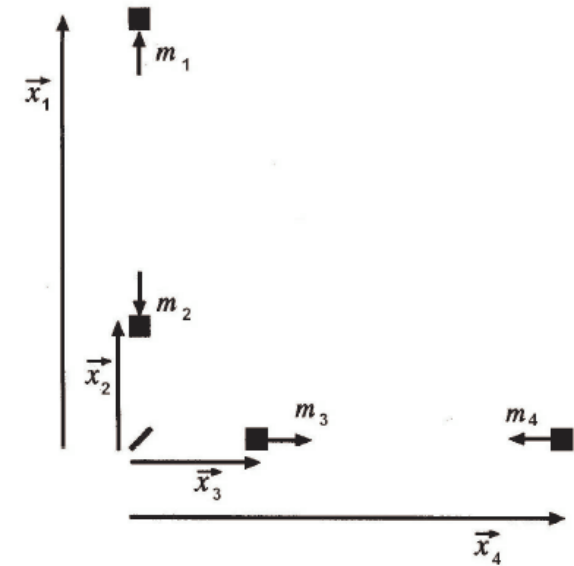
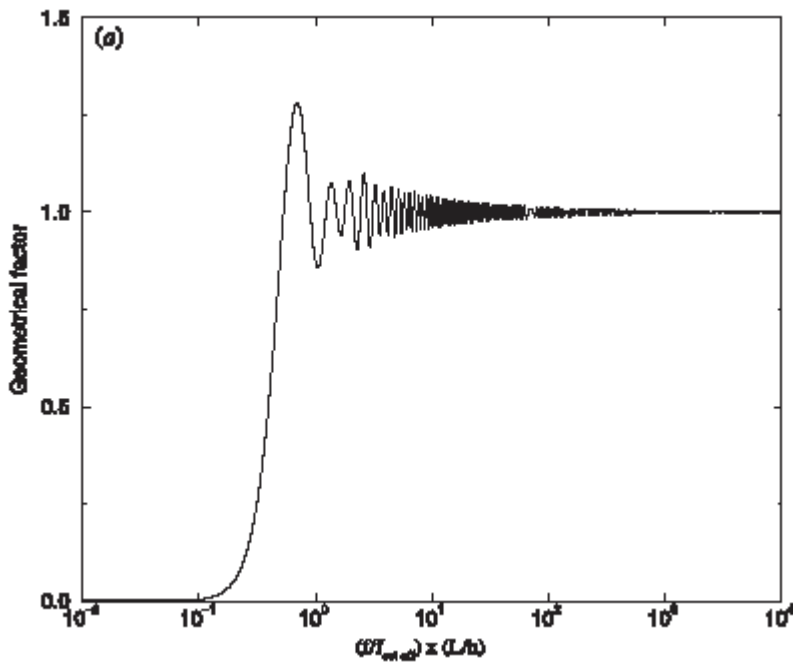


$$- \vec{u}(x,t) \cdot \vec{\nabla} \rho$$



Density fluctuations

Effect depends on geometry.



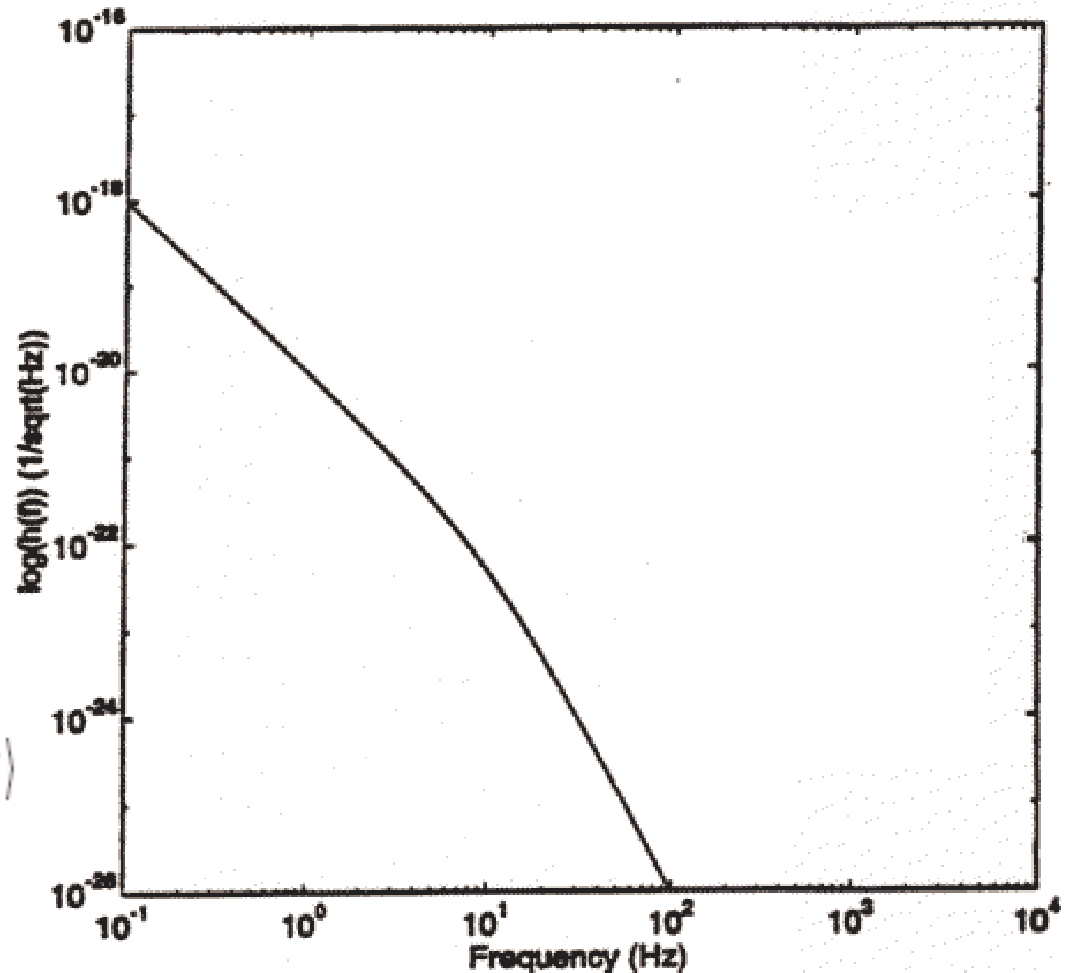
Coherently.

Seismic NN: the Saulson's model

A simple model:

- ground partitioned in square blocks (size $\lambda/2$)
- mass fluctuation uncorrelated between different blocks

$$\langle |\delta h(\omega)|^2 \rangle = \frac{16 \pi^3}{3} \frac{G^2 \rho_0^2}{L^2 |H(\omega)|^2} \langle |\delta x(\omega)|^2 \rangle$$





Seismic NN: elastic models

Elastic wave equation:

$$\partial_t^2 u_i(x, t) = c_T^2 \partial_k \partial_k u_i(x, t) + (c_L^2 - c_T^2) \partial_i \partial_k u_k(x, t)$$

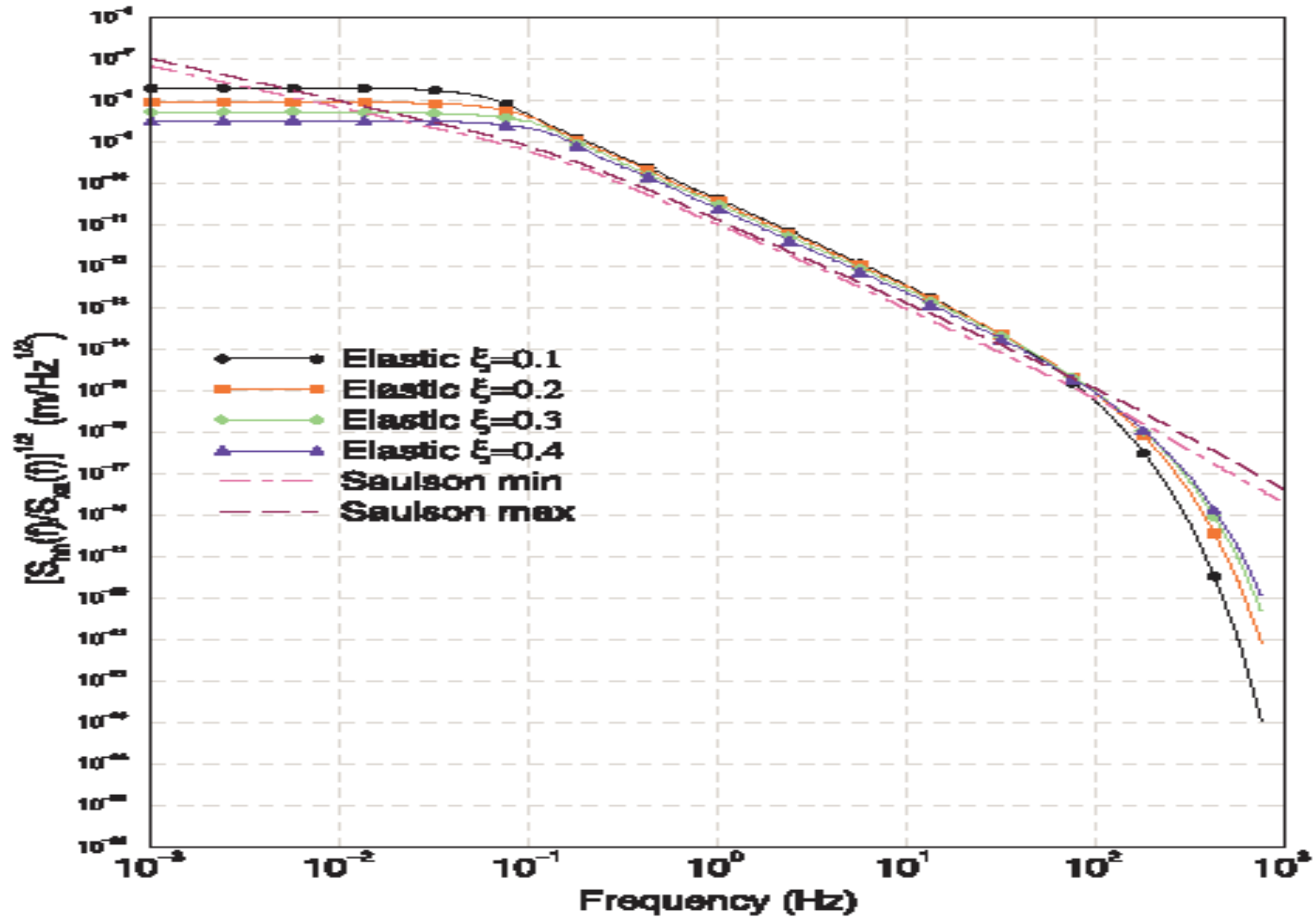
- Mode classification
- Evaluation of the contribution of each mode to NN

Transfer function between NN & seismic motion

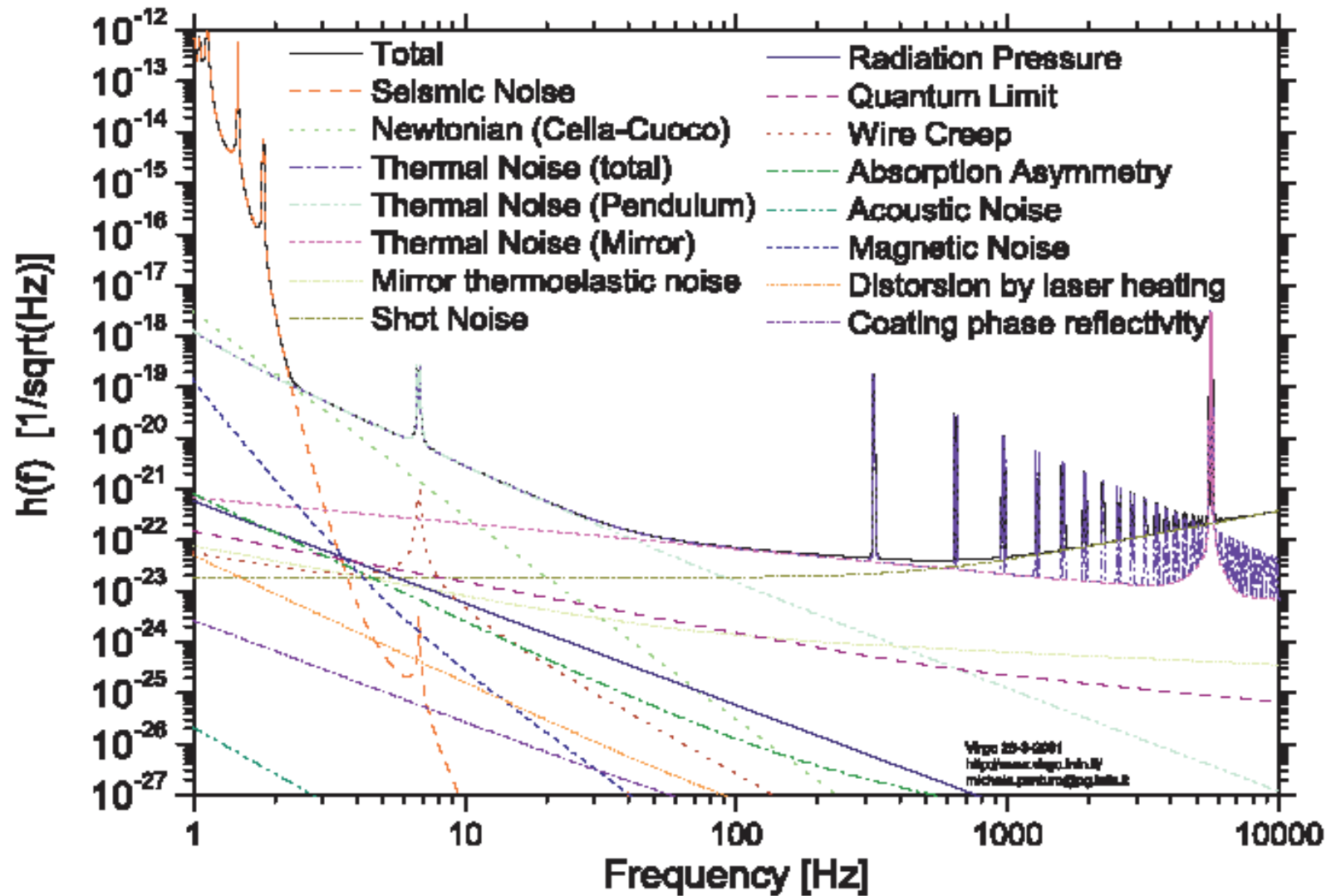
$$\begin{pmatrix} \delta \vec{a} \\ \delta a_z \end{pmatrix}(\vec{k}, z = h, \omega) = 4\pi G \rho_0 \frac{c_T^2}{\omega^2} \begin{pmatrix} \hat{k} \\ i \end{pmatrix} \{k \vec{k} \cdot \delta \vec{x} + ik^2 \delta z\} e^{-hk}$$

Using some other assumptions.....

Seismic NN: transfer function



Seismic NN: noise estimate





Atmospheric NN

- Saulson: effect of acoustic waves (negligible)
- Creighton: airborne objects, sonic booms, advection, ...

Rayleigh Bernard scenarios (G.C., E. Cuoco, P. Tomassini)

$$\begin{aligned}\partial_t \vec{u} + (\vec{u} \cdot \vec{\nabla}) \vec{u} &= -\rho^{-1} \vec{\nabla} p + \nu \nabla^2 \vec{u} + \alpha \vec{g} \theta \\ \partial_t \theta + (\vec{u} \cdot \vec{\nabla}) \theta &= \chi \nabla^2 \theta \\ \vec{\nabla} \cdot \vec{u} &= 0\end{aligned}$$

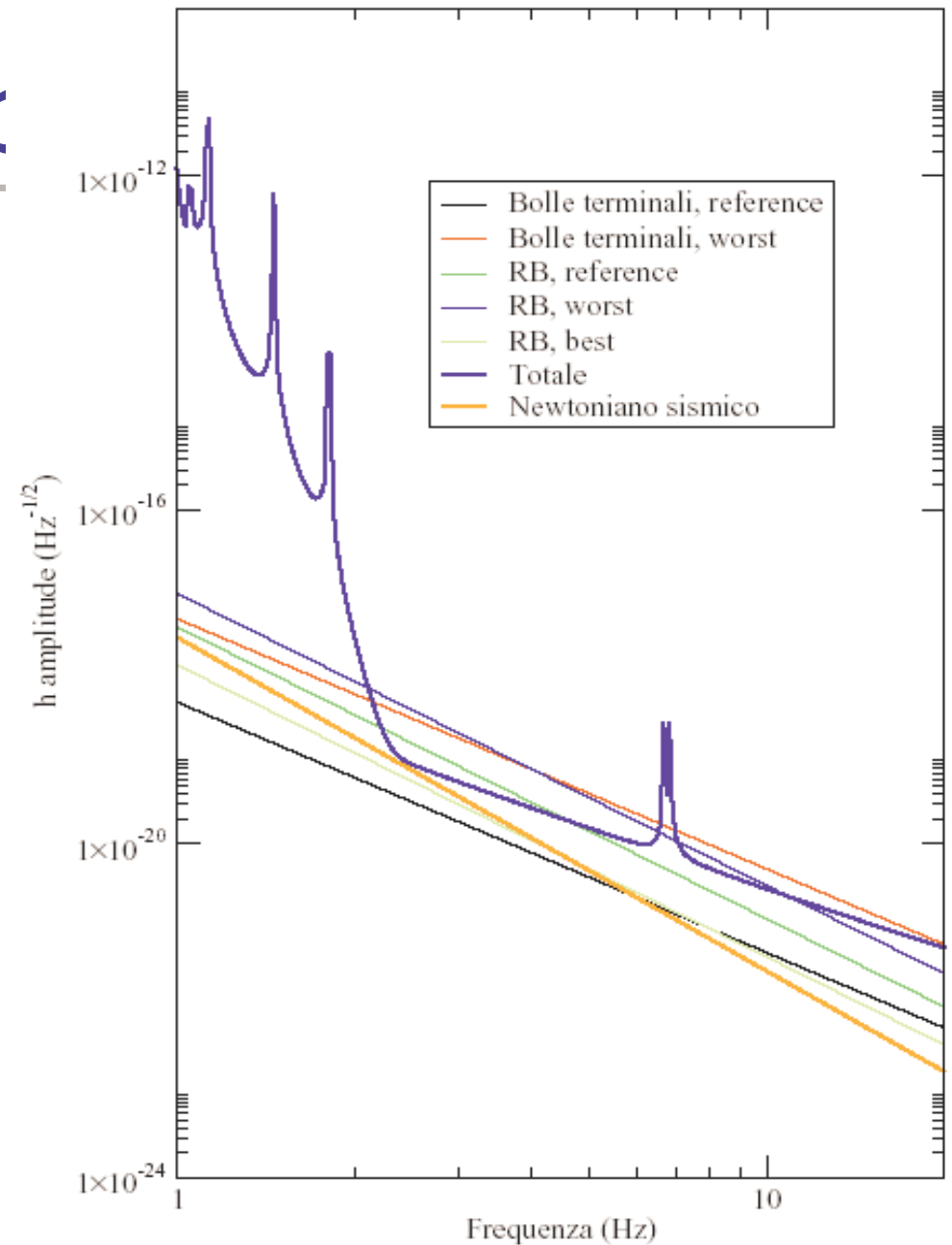
Different possibilities accordingly with the intensity of thermal gradient.

Thermal buk

- Nucleation phase (slow)
- Ascension phase

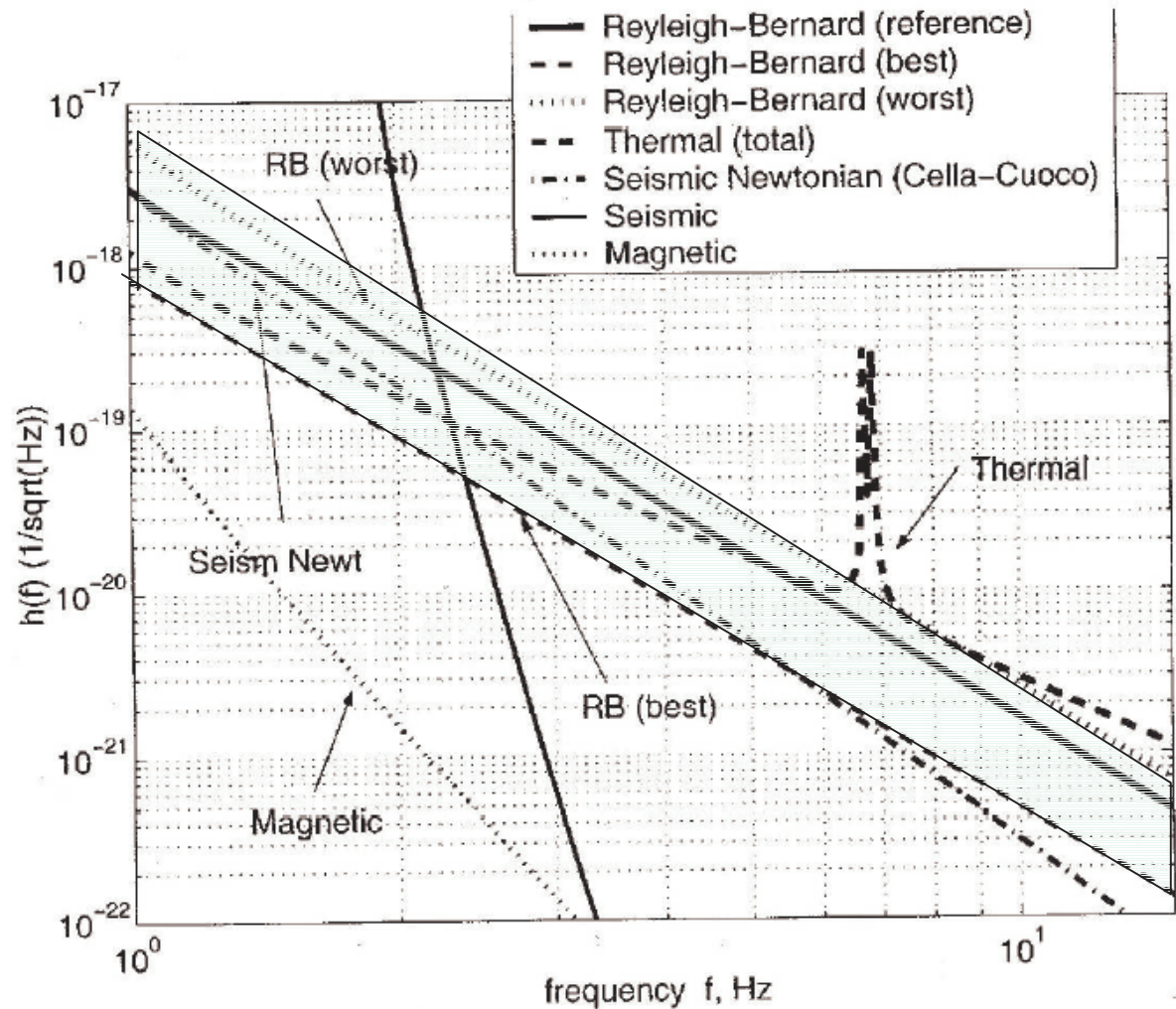
$$\partial_H z = g \left(\frac{T_b(z) - T(z)}{T(z)} \right) - 6\pi\nu r z$$

$$T_b(z) = T_b(0) + \gamma_{ad} z$$



Well developed turbulence

Method:
estimation of
the structure
functions
using simple
scaling
relations.





Other sources

- Turbulent generation of acoustic waves (Lighthill process):
Negligible (C. Cafaro, G. C.)

It seems that RB phenomenology is at least competitive with seismic NN

Models need validation:

- Atmospheric fluctuations should be well studied by astronomers (work in progress: reducing our ignorance in this field)
- Some data are available (balloons, airplanes)

Subtraction of NN: strategy



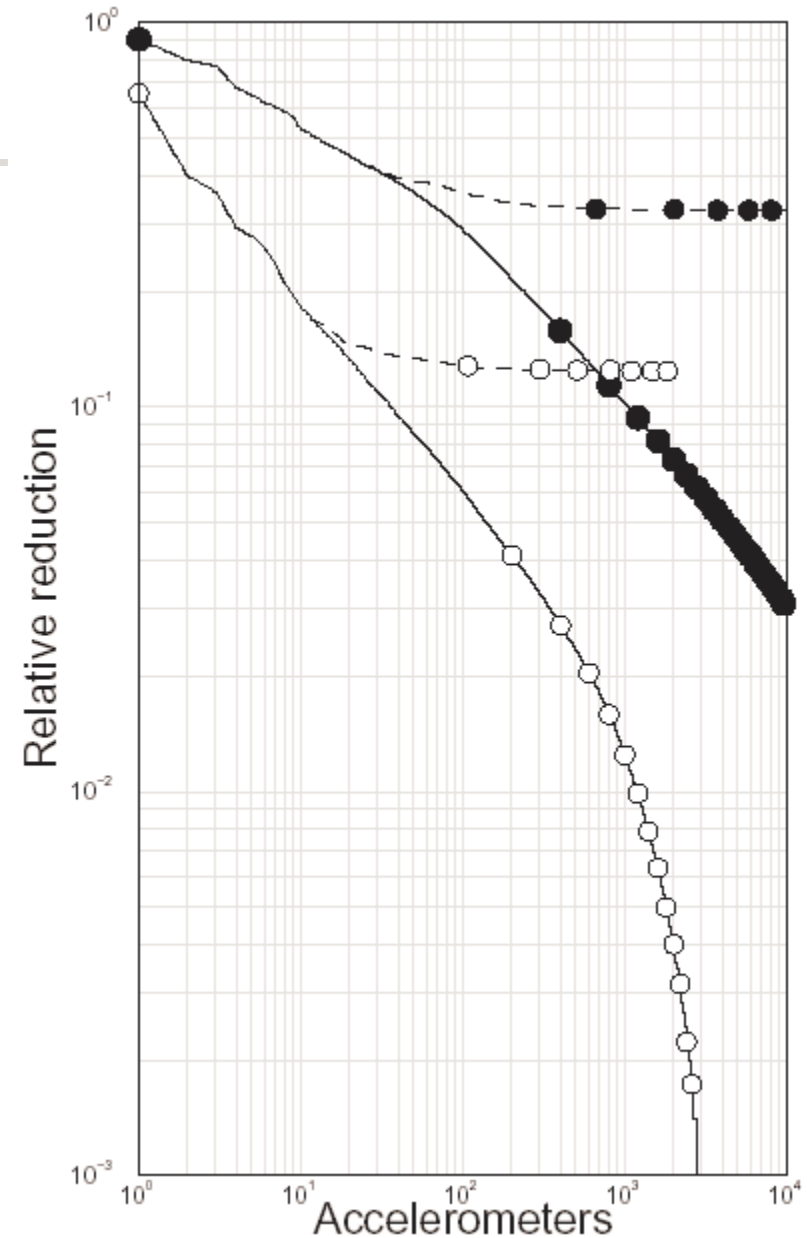
- Auxiliary channels should be correlated with the noise (easy!)
- Auxiliary channels should not be correlated with the useful signal (very easy!)

What we must put inside the box?

1. model dependent strategy  Physical modelization
2. Model independent strategy  Adaptive strategy

Qualitative discussion

- A seismic sensor give us complete information on a block
- Complete ignorance on blocks without sensors
- Blocks nearest to the test masses should be controlled first (for a given sensor budget)
- If the sensors are noisy, we can put two of them on the same block





The linear case

- In the case of additive, gaussian noise (on all the channels) we can find an explicit solution:

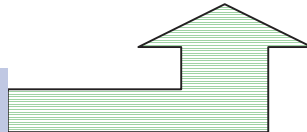
Subtracted signal:

$$X_{sub}(\omega) = X(\omega) - \sum_{i,j} \langle X(\omega) \hat{n}_i \cdot \vec{u}(x_i, \omega)^* \rangle [C^{-1}(\omega)]_{ij} \hat{n}_j \cdot \vec{u}(x_j, \omega)$$

Relative noise power spectrum reduction:

$$1 - \sum_{i,j} \frac{\langle X(\omega)^* \hat{n}_i \cdot \vec{u}(x_i, \omega) \rangle [C^{-1}(\omega)]_{ij} \langle \hat{n}_j \cdot \vec{u}(x_j, \omega)^* X(\omega) \rangle}{\langle X(\omega)^* X(\omega) \rangle}$$

Projector: noise can't increase
(if we use the correct C...)



Optimization of sensor placement

- **If we have a physical model**, we can calculate all:

Seismic correlations

$$\begin{aligned} \langle \delta x_i(x_1) \delta x_j(x_2) \rangle &= \pi \left(\frac{\omega}{c_T} \right)^2 |u_X|^2 \langle |A_4(\omega)|^2 \rangle G_{ij}^T(\vec{x}_1, \vec{x}_2) & \langle \delta x_i(x_1) \delta z(x_2) \rangle &= \pi \left(\frac{\omega}{c_T} \right)^2 \text{Re}(u_X^* u_Z) \langle |A_4(\omega)|^2 \rangle G_i^V(\vec{x}_1, \vec{x}_2) \\ \langle \delta z(x_1) \delta z(x_2) \rangle &= \pi \left(\frac{\omega}{c_T} \right)^2 |u_Z|^2 \langle |A_4(\omega)|^2 \rangle G^S(\vec{x}_1, \vec{x}_2) & \langle \delta x_i(x_1) \delta \theta_j(x_2) \rangle &= \pi \left(\frac{\omega}{c_T} \right)^3 \text{Re}(u_X^* u_\theta) \langle |A_4(\omega)|^2 \rangle G_{ij}^T(\vec{x}_1, \vec{x}_2) \\ \langle \delta \theta_i(x_1) \delta \theta_j(x_2) \rangle &= \pi \left(\frac{\omega}{c_T} \right)^4 |u_\theta|^2 \langle |A_4(\omega)|^2 \rangle G_{ij}^T(\vec{x}_1, \vec{x}_2) & \langle \delta \theta_i(x_1) \delta z(x_2) \rangle &= \pi \left(\frac{\omega}{c_T} \right)^3 \text{Re}(u_Z^* u_\theta) \langle |A_4(\omega)|^2 \rangle G_i^V(\vec{x}_1, \vec{x}_2). \end{aligned}$$

Newtonian correlations

...mixed ones...

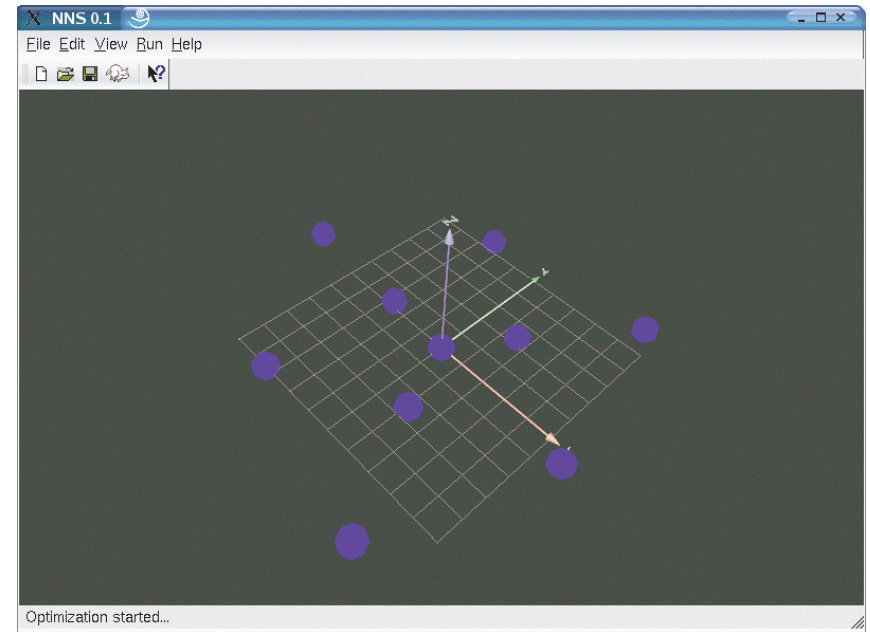
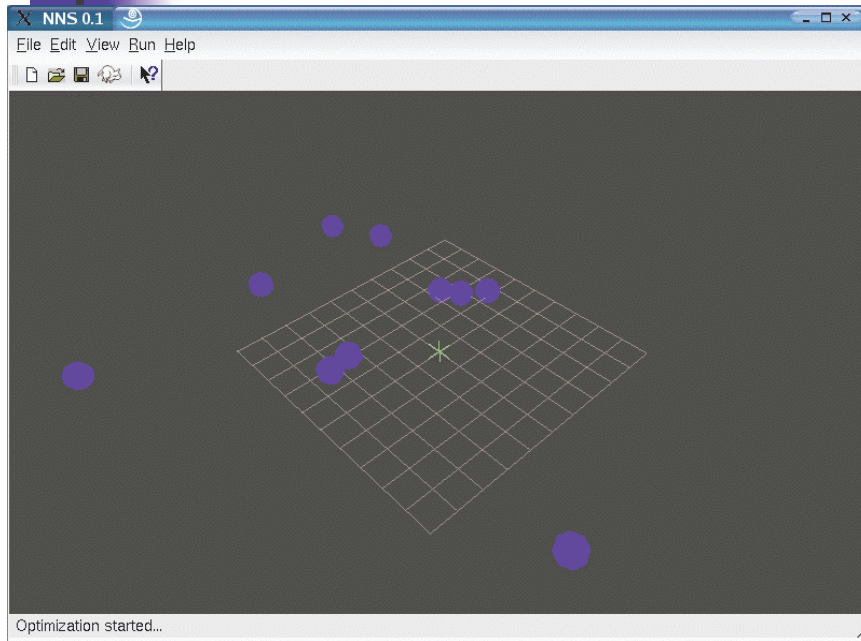
$$\begin{aligned} \langle \delta a_i(x_1) \delta a_j(x_2) \rangle &= \frac{1}{2} G^2 \rho_0^2 \left(\frac{\omega}{c_T} \right)^2 W^2 \langle |A_4(\omega)|^2 \rangle G_{ij}^T(\vec{x}_1, \vec{x}_2) \exp \left(-2h \frac{\omega}{c_T \sqrt{x}} \right) \\ \langle \delta a_i(x_1) \delta a_z(x_2) \rangle &= \frac{1}{2} G^2 \rho_0^2 \left(\frac{\omega}{c_T} \right)^2 W^2 \langle |A_4(\omega)|^2 \rangle G_i^V(\vec{x}_1, \vec{x}_2) \exp \left(-2h \frac{\omega}{c_T \sqrt{x}} \right) \\ \langle \delta a_z(x_1) \delta a_z(x_2) \rangle &= \frac{1}{2} G^2 \rho_0^2 \left(\frac{\omega}{c_T} \right)^2 W^2 \langle |A_4(\omega)|^2 \rangle G^S(\vec{x}_1, \vec{x}_2) \exp \left(-2h \frac{\omega}{c_T \sqrt{x}} \right) \end{aligned}$$

Numerically this is quite problematic...

Then we can find:

- the optimal set of sensors
- their optimal configuration

Example of optimization



Features:

- In the frequency range of interest each test mass can be studied separately
- Optimal solution is quite robust (good to know, if we ignore the physical model)
- Sensors "crystallize" in quite regular configurations

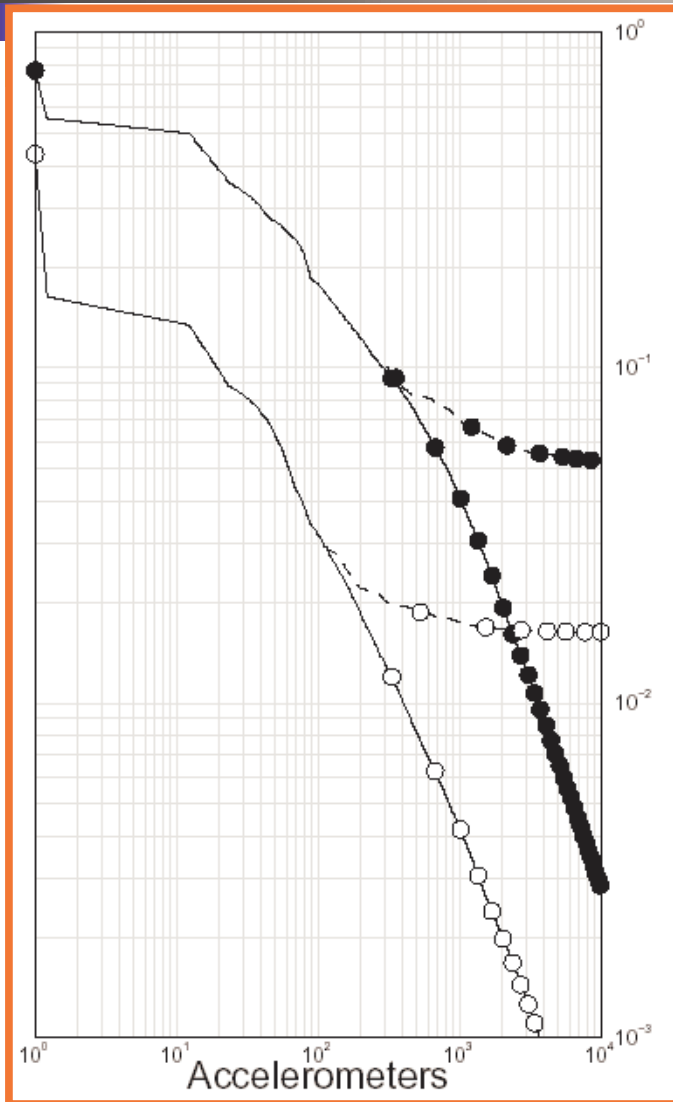


Efficiency results

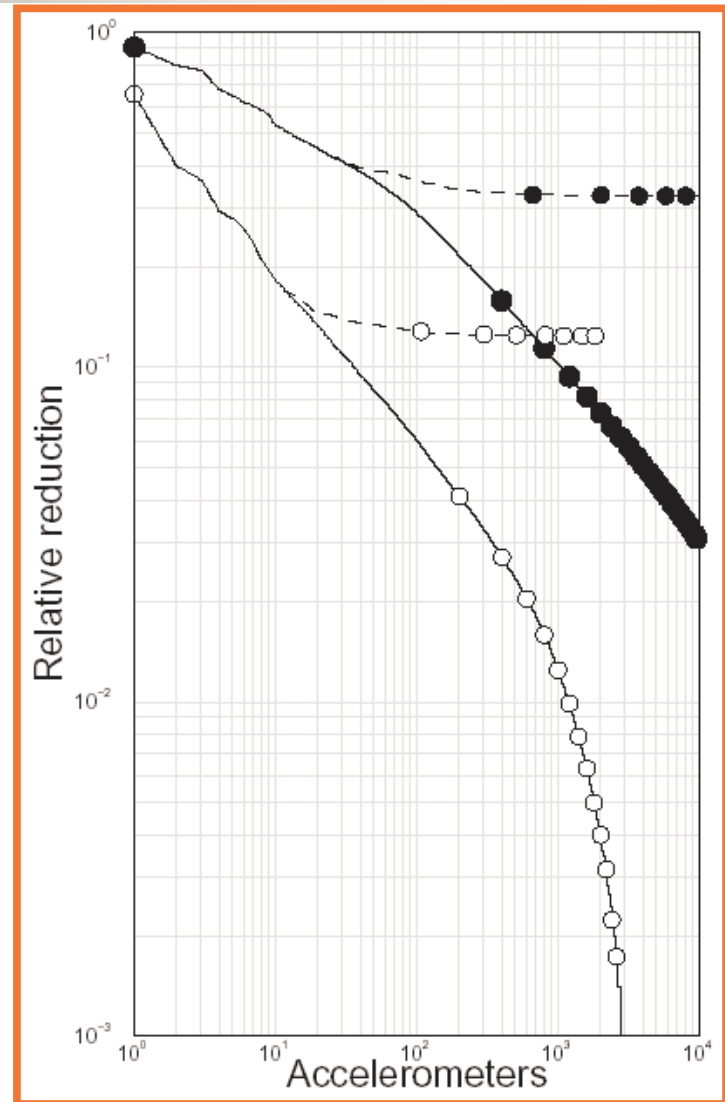
- Seismic correlations act as a repulsive potential (statistical independence)
- Mixed correlations “attract” the sensors near the test masses (maximize IF-aux correlations)
- Instrumental noise shift the eigenvalues of C , reducing the “repulsive potential”

Results: correlations are important

Efficiency results



14/04/2003

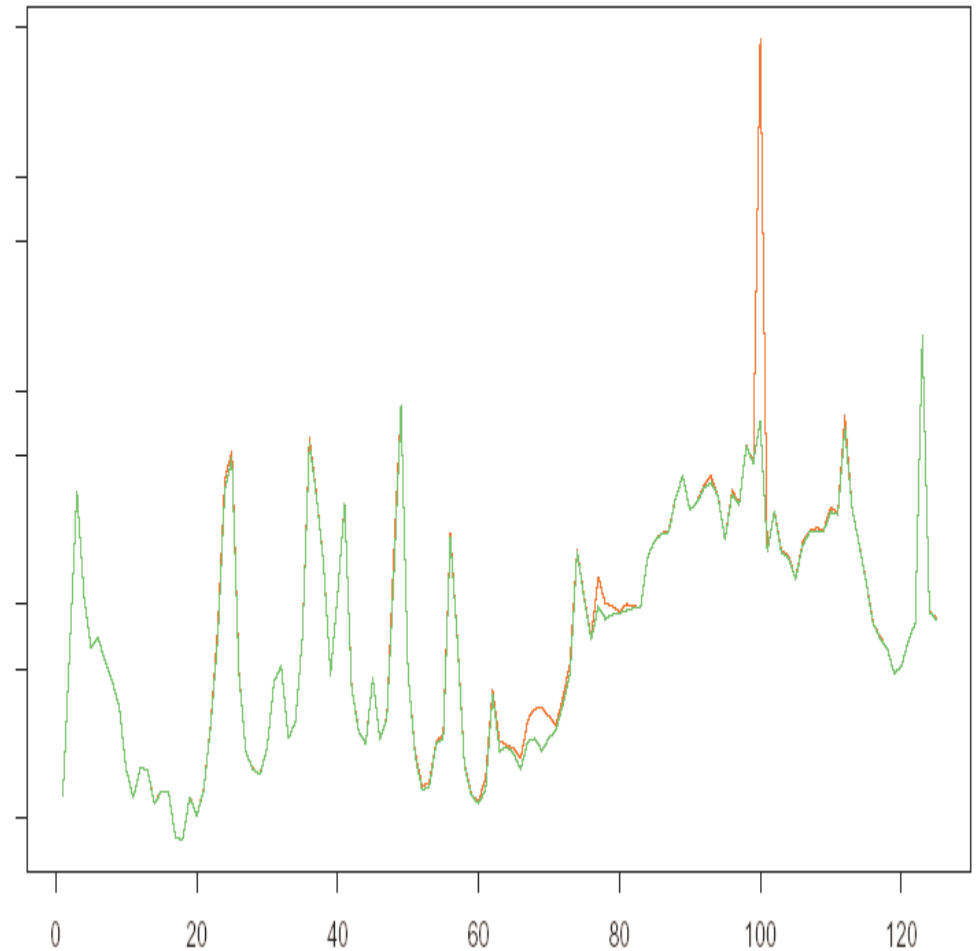


Aspen 2003

Model independent subtraction

- Can be seen as a "generalized" whitening:
- can be implemented "on line"
 - can be made adaptive

Example: subtraction of acoustic channels from IF channel (VIRGO E2 data)





Non linear case

What if the auxiliary channels contain non linear correlations with the IF channel?

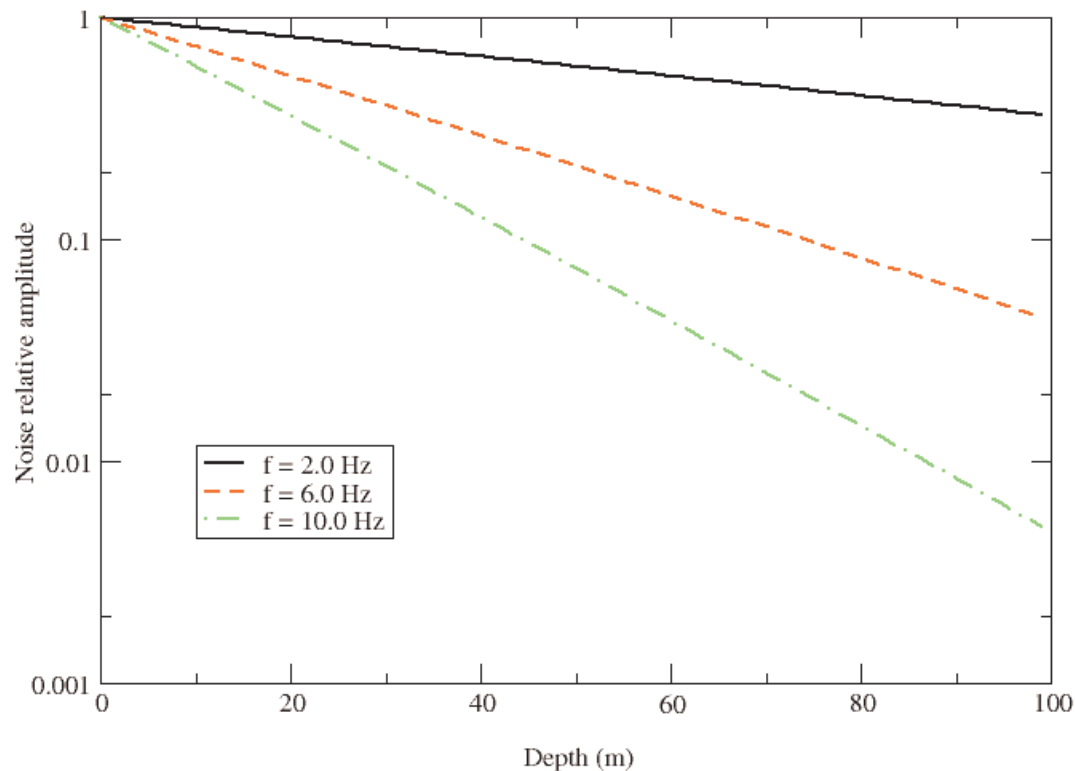
- not completely academic: a sensor for the modulus of the seismic motion

The problem can be formulated (and in some cases solved) as the design of the optimal Neyman-Pearson in presence of auxiliary channels

- can a nonlinear sensor do better?
- is the “subtraction” procedure independent by the detection problem? (There could be a well known answer (entropy?), I do not know).

Going underground: seismic NN reduction

A simple fact: surface waves die exponentially with the depth



- Surface waves are probably the most important excitations for NN
- Surface movement dominate the bulk compression effect



Other advantages:

- Horizontal seismic NN is self-regularized: weak dependence on empty space dimensions
- Atmospheric NN reduction: obvious
- Collective atmospheric effects should be damped exponentially
- Subtraction procedure should work well (but must be studied in detail)



Going underground: some possible problems

- Localized seismic waves on the gallery
 - Small masses involved
 - Easily (???) monitorable
- Acoustic (pressure waves) resonances
 - Could be avoided (close the doors!)
 - Easily monitorable too



Other hardware options

- Mechanical screen (es. cylindrical screen)
 - Delicate
 - Require hardware modifications
 - Not easily tunable
- Barriers for seismic noise
 - Should be carefully designed

Conclusions

- NNS is promising: somebody tried it yet....
- NN will be relevant for advanced IF
- Going underground is a promising option

