

### **Pre-Isolation**

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### Issue

- Ground motion at LLO with the initial LIGO seismic isolation system makes it impossible to hold the interferometers locked reliably during the day
  - » Steady-state ambient noise is higher due to anthropogenic sources
  - » Transients, particularly from logging
- Wind induced seismic noise at LHO:
  - » exceeds locking threshold at ~25 mph, or 10% of the time
  - » Expect that up-conversion is a problem at significantly lower wind speeds & a large fraction of the time
- Upgrade is required to allow both reliable locking and to allow better noise performance while locked
  - » Need 90% duty cycle & lock durations > 40 hours
  - » Need to reduce noise in the control band (< 40 Hz) to permit a smaller suspension actuator authority & lower noise
  - » Suppression in the 1-3 Hz band is most important due to excitation of the lower stack modes (Q ~ 30)



## Scope

#### Retrofit design

- » Original design included expansion capability for active control
- » No commercially available systems with acceptable performance
- » Accelerate the existing advanced LIGO R&D effort for an active pre-isolator
  - An LSC effort scientifically led by Joe Giaime (LSU) and Brian Lantz (Stanford)
  - Digital servo controls (flexibility & graceful degradation under failure)
  - Two alternative actuator designs
- » Install without disturbing in-situ optics alignment
- » BSC & HAM chambers
- » Prove performance with full scale prototypes at LASTI

#### Active Isolation with the Fine Actuation Systems (FAS) on Test Mass chambers

- » Use of the FAS actuator for active control is known as PEPI: Piezo-electric External Pre-Isolation
- » PEPI is an interim solution for LLO; installed for S1
- » PEPI is the planned solution for LHO

#### Number of retrofit systems:

- » All chambers with suspended optics at LLO (8 systems)
- » Addition of PEPI systems to Test Mass and Mode Cleaner chambers at LHO (6 systems per interferometer)



# External Pre-Isolator performance requirements

#### Basic tenets:

The pre-isolator must not increase the present noise in the GW-band, and

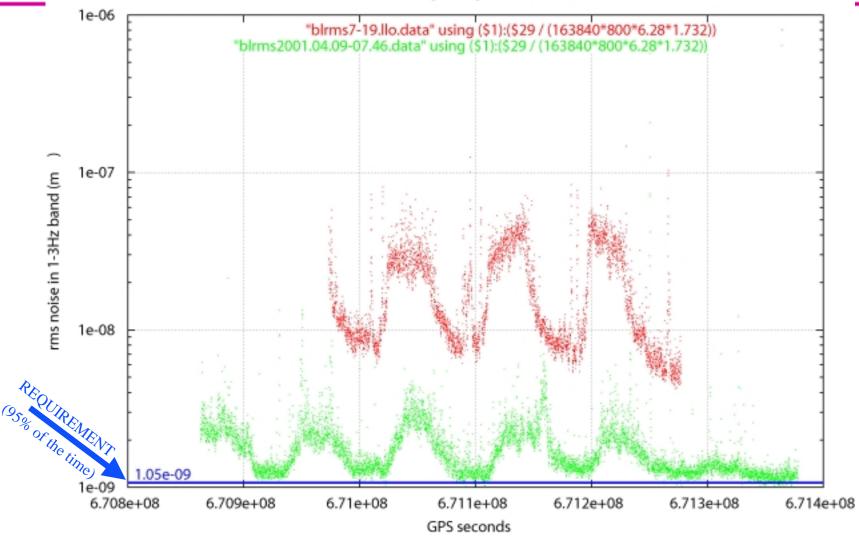
Must bring the day LLO environment to the level of the LHO night environment.

1 month 100 seconds	10 microns pk-pk 1 micron pk-pk	Presently observed stability of system
0.16 Hz	4e-7 m/√Hz	To original seismic model
1 Hz	1e-9 m/√Hz	Hanford night-time
10 Hz	4e-10 m/√Hz	spectrum in 1-3 Hz band 95% of the time
15 Hz 30 Hz 50 Hz and higher	2e-10 m/ $\sqrt{\text{Hz}}$ 6e-11 m/ $\sqrt{\text{Hz}}$ 2e-11 m/ $\sqrt{\text{Hz}}$	Not to exceed presently observed spectrum



## Daily variability – and requirement

red=livingston, green=hanford





## Initial Vibration Isolation Systems

- » Reduce in-band seismic motion by 4 6 orders of magnitude
- » Little or no attenuation below 10Hz; amplification at stack mode resonances
- » Large range actuation for initial alignment and drift compensation
- » Quiet actuation to correct for Earth tides and microseism at 0.15 Hz during observation

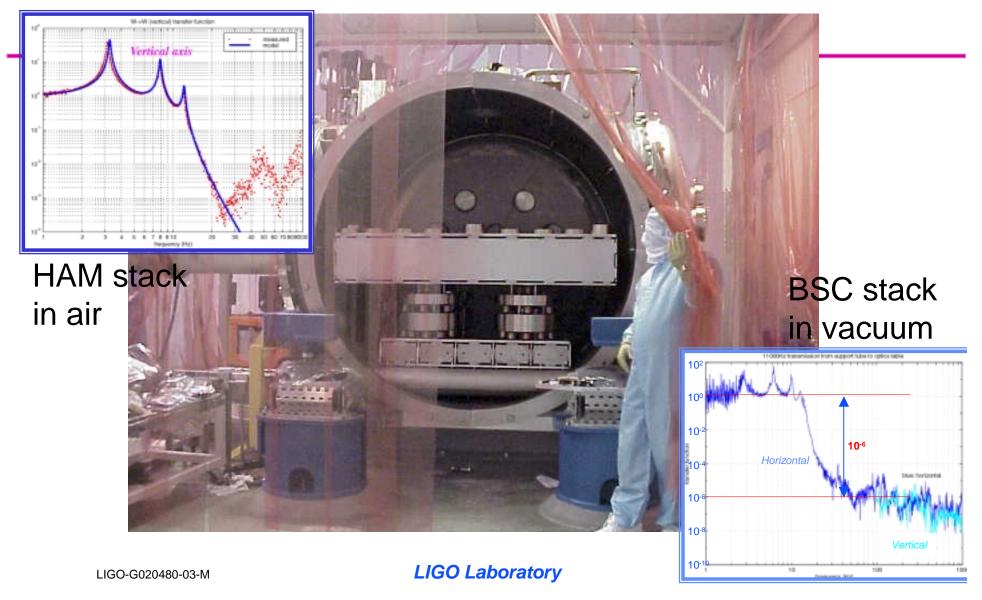






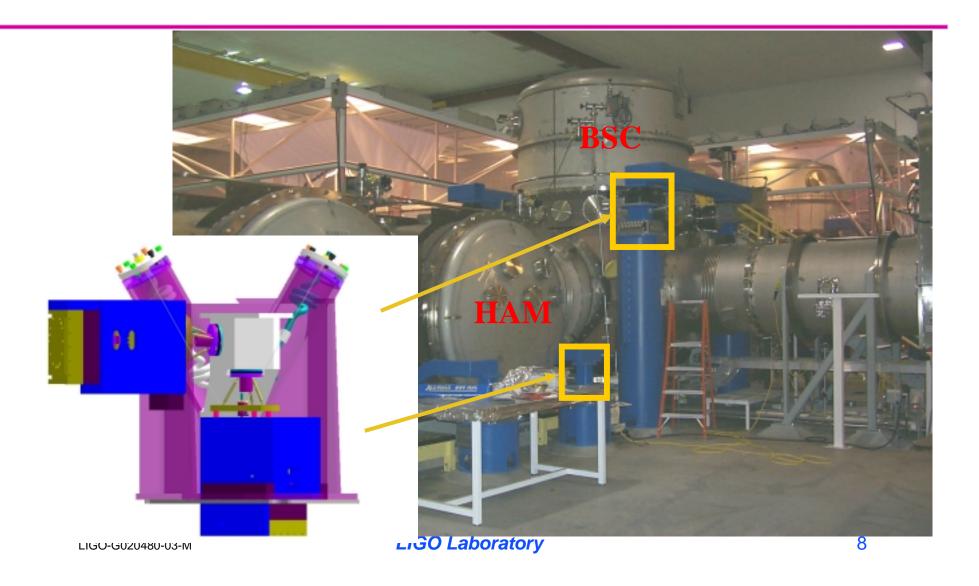


## Seismic System Performance



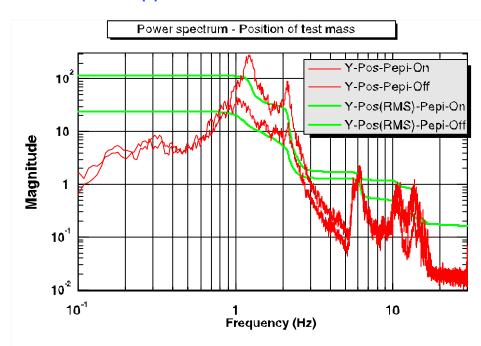
## Planned Initial Detector Modifications

LIGO



# LIGO Piezo-electric External Pre-Isolator (PEPI)

- Single (longitudinal) degree of freedom isolation
- Employs Fine Actuation System on End Test Mass Chambers (also used for Tidal and Microseismic control)
- Added to Input Test Masses at LLO for the S1 Run
- Baseline approach for LHO for S3





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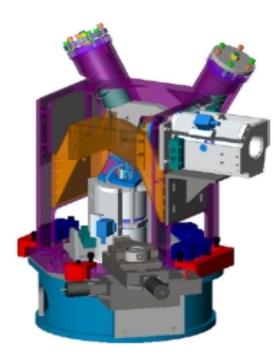
## Actuators for an External Pre-Isolation (EPI) System

- Large range (300 microns p-p) required for tidal & microseismic correction
  - » Goal of 1 mm for coarse positioning/alignment
  - » Piezo-electric actuation may not be suitable
- Quiet Hydraulic Actuators
  - » Hydraulic Wheatstone bridge
  - » Used for precision diamond turning vibration isolation
  - » High range, high stiffness, high bandwidth, high velocity
  - » Developmental system (not commercially available)
- Electro-magnetic actuators
  - » Different actuator (force instead of displacement)
  - » Increased robustness of EPI solution a second path
  - » Familiar technology (in contrast to quiet hydraulics)
  - » Reduced risk of contamination
  - » Less complexity in power supply
  - » Same performance requirements, mechanical superstructure, sensors
  - » Concern regarding EM coupling to the magnets on the suspended optics



# Hydraulic External Pre-Isolators (HEPI)

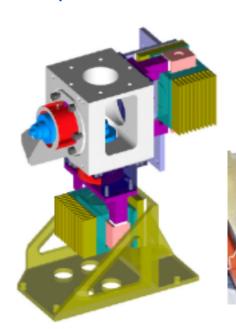




- Working fluid is high-viscosity fluid (glycerin/ethanol or mineral oil)
- Bellows hydraulic pistons apply force without sliding friction, moving seals
- Laminar-flow differential valves control forces
- Stabilized "power supply" is remote hydraulic pump with fluid-equivalent "RC" pressure filtering
- Technology adapted from precision machine tool applications

## LIGO MEPI Installed on HAM at LASTI

- Alternative to the developmental hydraulic actuator
- Uses commercially available voice-coil actuator
- 'Pin-compatible' mechanically
- Simpler electronics
- 'Soft' mechanical back impedance







### Status & Decision Points

#### MEPI

- » Installed at LASTI & under Test
- » Interaction of HAM structural support modes with the control system may limit performance, add control complexity or cause us to consider structural modifications/additions
- » Initial MEPI/HAM Chamber results are promising; Hope to demonstrate control to required performance in the next 2 months

#### HEPI

- » 3<sup>rd</sup> generation hydraulic actuator in test on the Stanford test stand
- » Pump station tests at CIT have demonstrated pressure noise performance requirements
- » Installation at LASTI to start in 2 weeks
- » Compliance of the BSC pier may likewise limit gain-bandwidth and performance
- » Initial test results are expected by early December

#### Design Review & Long-Lead Procurement Review, Jan, 2003

- » After prototype installation & some preliminary experience will decide whether to go forward with the hydraulic actuator or the electro-magnetic actuator
- » Commissioning will continue to improve performance and transition from dSpace to VME based controllers
- Installation start at LLO, April, 2003
  - » Following the S2 run