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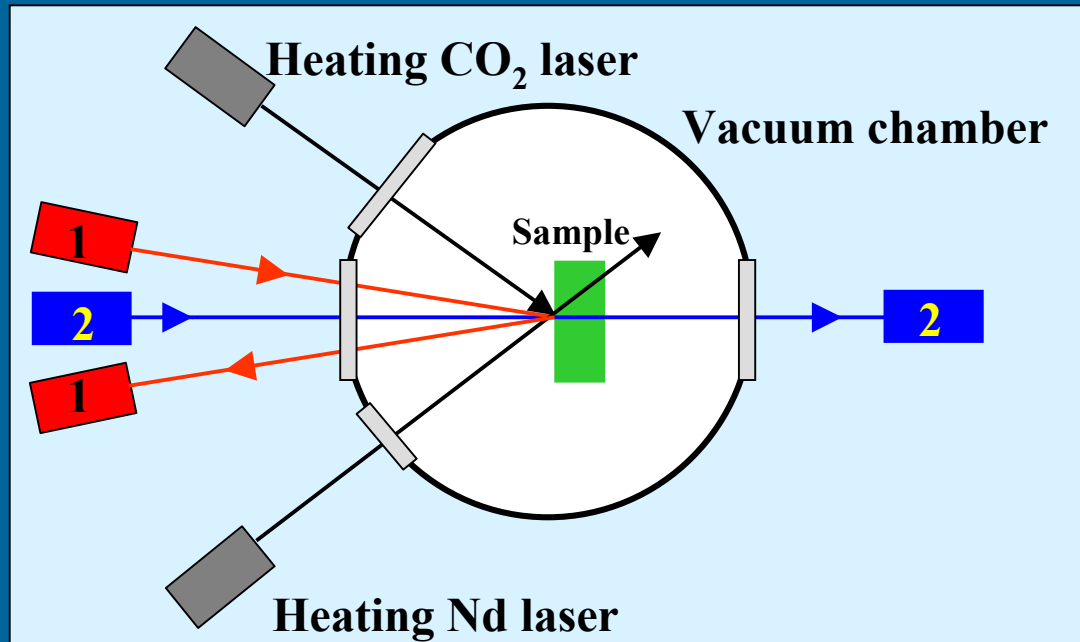
## **Remote *in situ* monitoring of weak distortions**

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# Current Research

- ◆ Remote *in situ* monitoring of weak distortions emerging under auxiliary laser heating similarly to what is expected in advanced LIGO core optics
- ◆ Proposal for remote *in situ* monitoring of weak distortions of ETM.

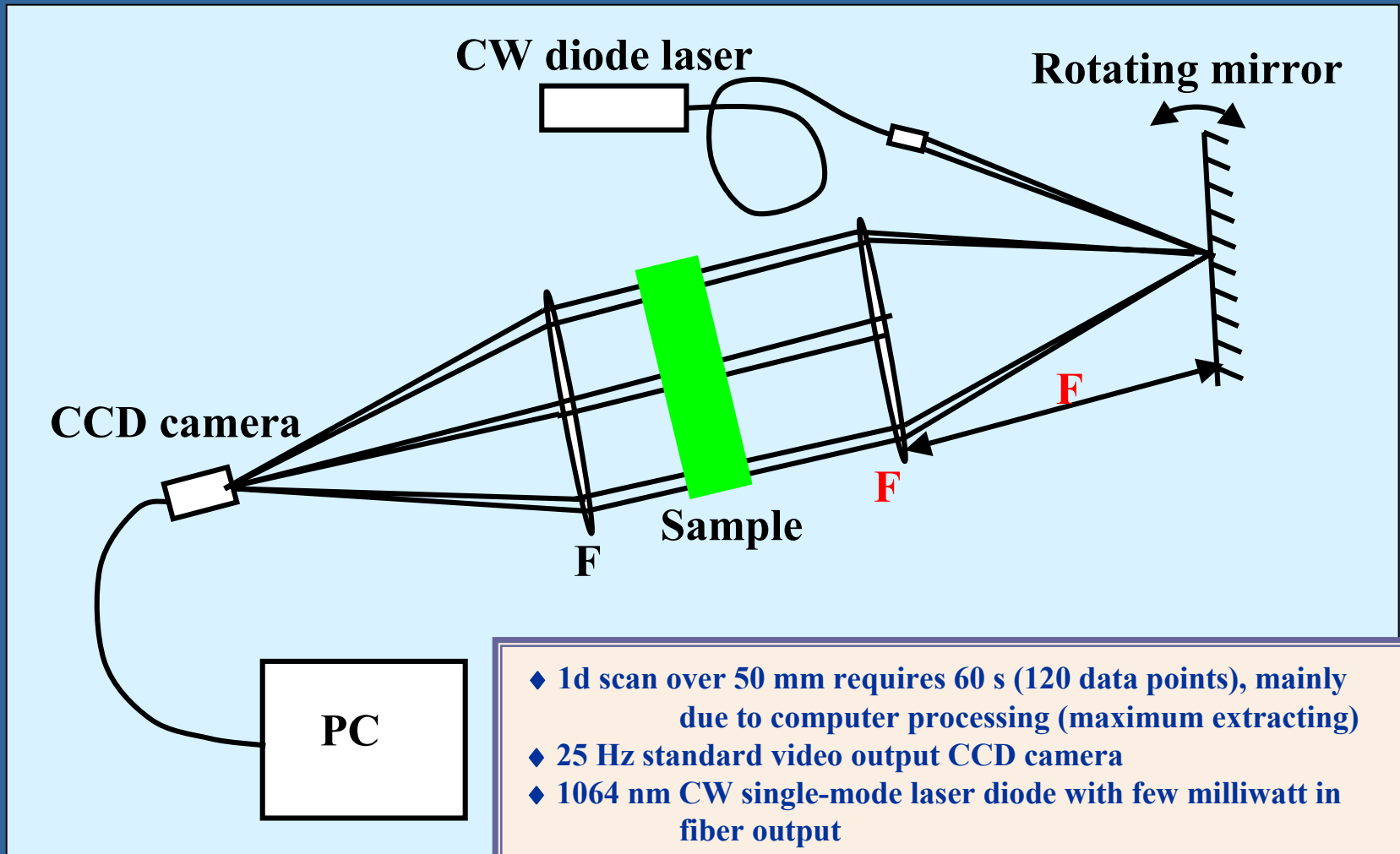
# Remote *in situ* monitoring of weak distortions emerging under auxiliary laser heating



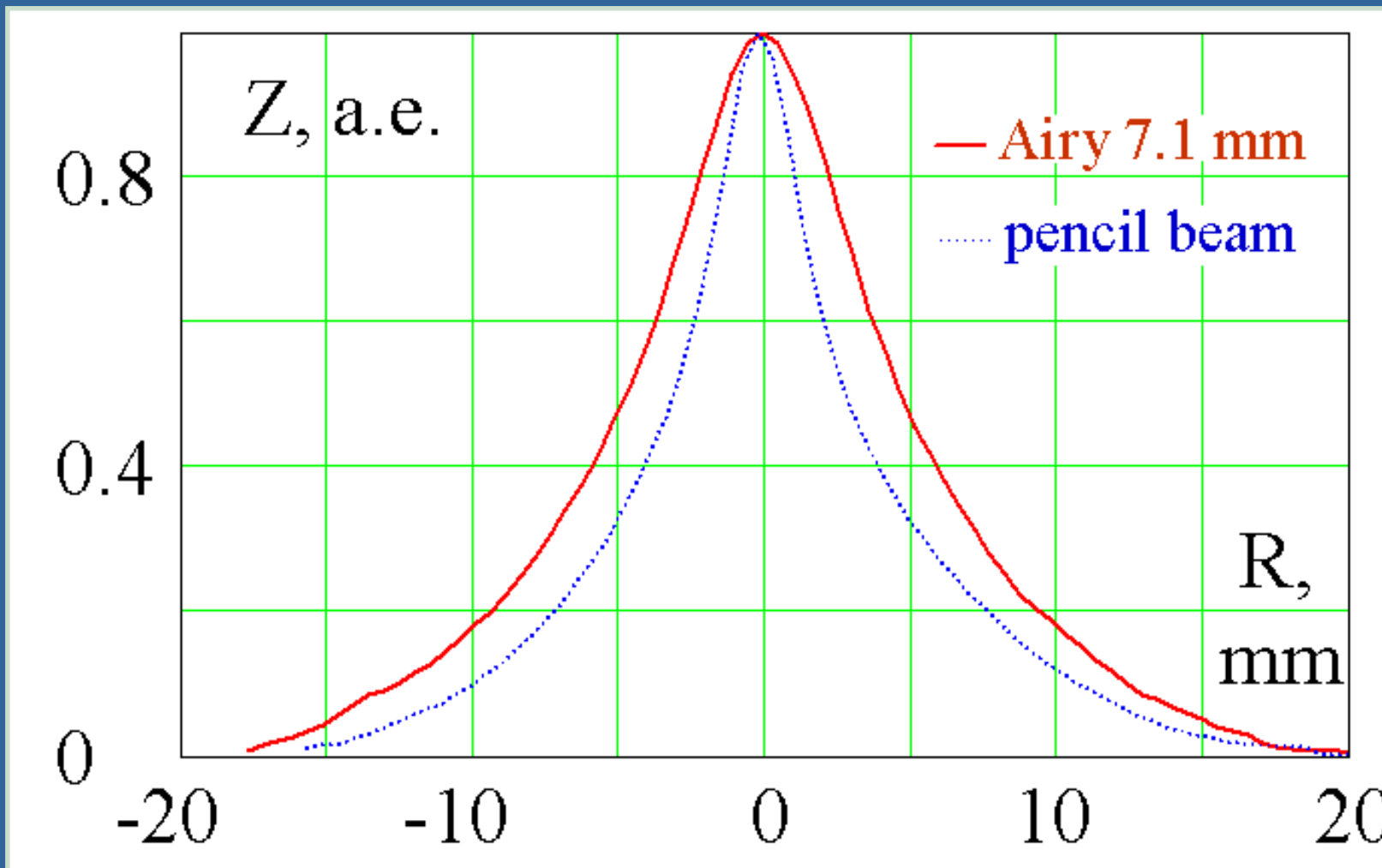
1 - **WLPMI**  
2 - **NHS and PIT**

- Optical sample bulk heating by the fundamental or second harmonic of Nd:YAG laser at a power of 10-20 W
- Surface heating with the use of a CO<sub>2</sub> laser at power of several Watts
- Inducing contamination of a small region (characteristic size of 20-100 micron) on the optical element's surface and focusing of low-power laser radiation (<100 mW) on it

# Scanning Linear Hartmann Technique



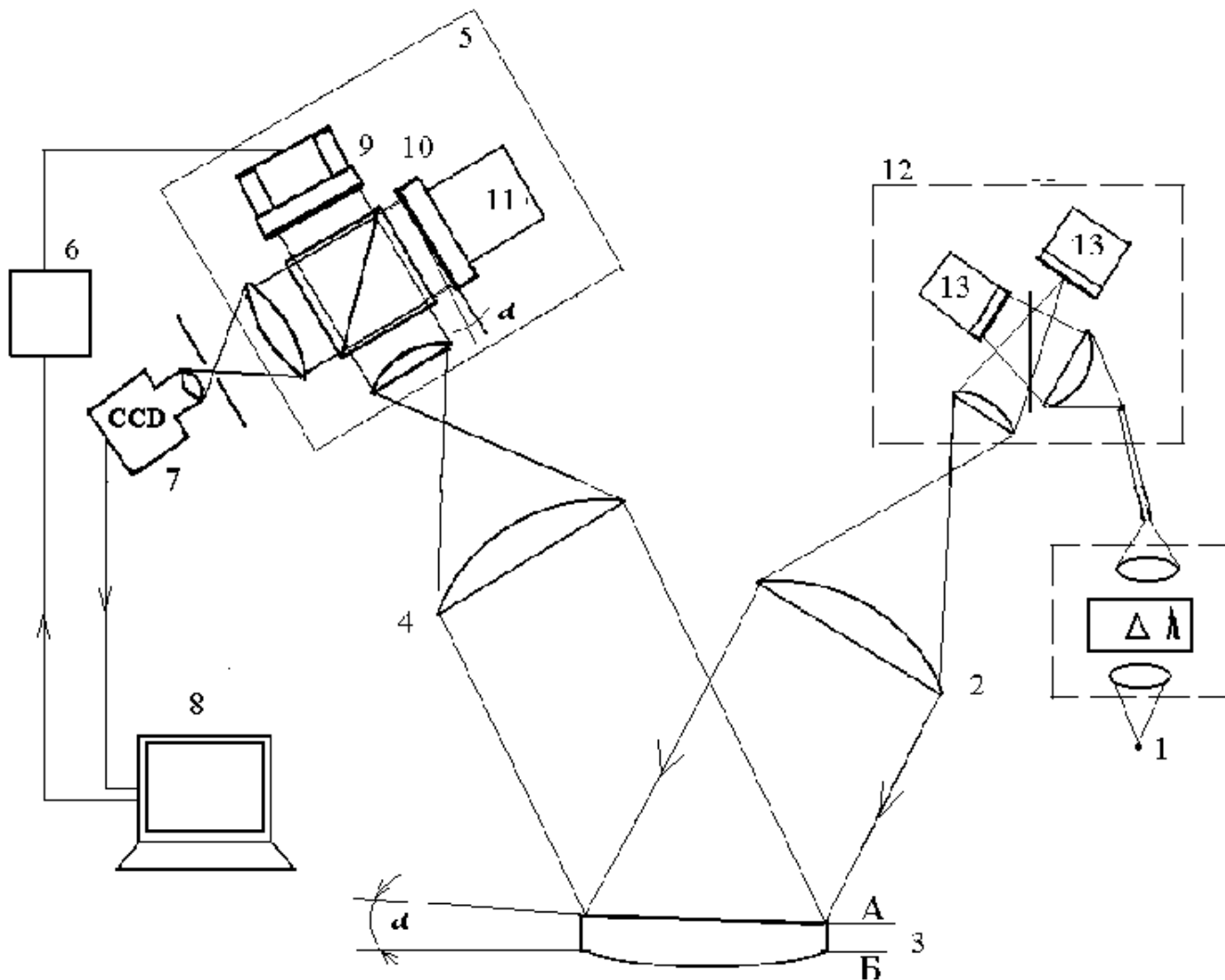
**Optical depth profile measured with scanning linear  
Hartmann sensor for two heating beam :  
7mm Airy and 1 mm pencil structures**



# White Light *In Situ* Measurement Interferometer (WLISMI)

| <b>Standard interferometers</b>   | <b>Proposed interferometers</b>  |
|---|--|
| <p>Measurement of optical length of air spacing between two surfaces.</p> <p>In profilometers one of them is a sample surface, and the other is a reference surface.</p> <p>The problem of precise measurement of phase in the interferogram is solved by phase modulation according to a known time law.</p> | <p>The proposed method relies on measurements of the phase of interferogram of radiation reflected <b>from two surfaces of one sample</b> under study.</p> <p>The precise phase measurements are ensured by the <b>modulation</b> of the probing radiation <b>spectrum</b>.</p> <p>The method provides a two-dimensional pattern of a sample's <b>optical thickness distribution</b> simultaneously over the whole aperture.</p> <p>The method is applicable to <b>remote testing</b> of optical elements with flat, spherical and cylindrical surfaces, and also with a wedge between them.</p> |

# White Light *In Situ* Measurement Interferometer. Experimental setup

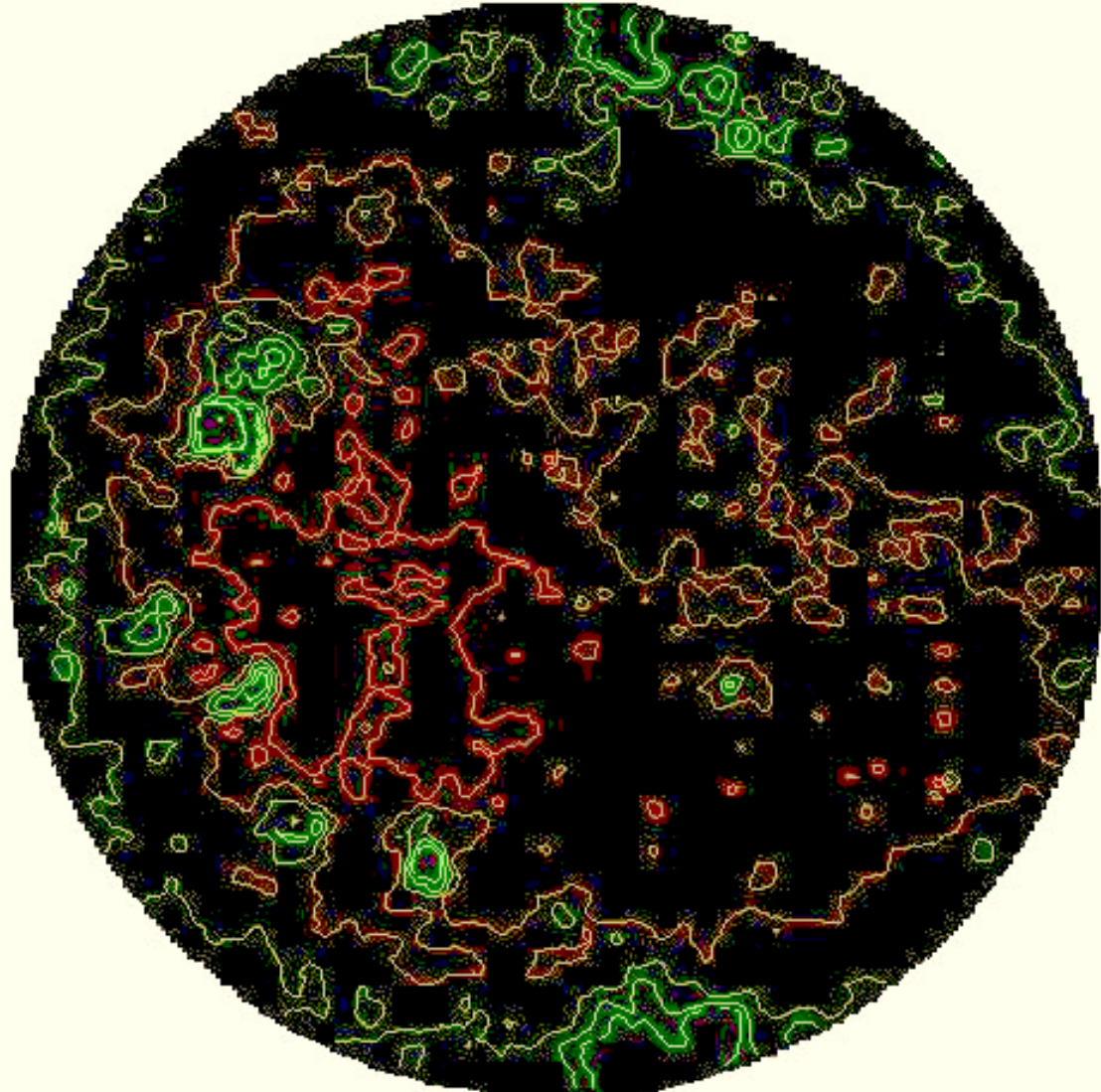


- 1 - light source;
- 2 - objective;
- 3 - sample;
- 4 - ocular;
- 5 - measurement interferometer;
- 6 - unit for synchronization and control;
- 7 - CCD camera;
- 8 - PC computer;
- 9 - modulating mirror;
- 10 - adjusting mirror;
- 11, 13 - motors;
- 12 - wave front shaper

# White Light *In Situ* Measurement Interferometer Phase Map

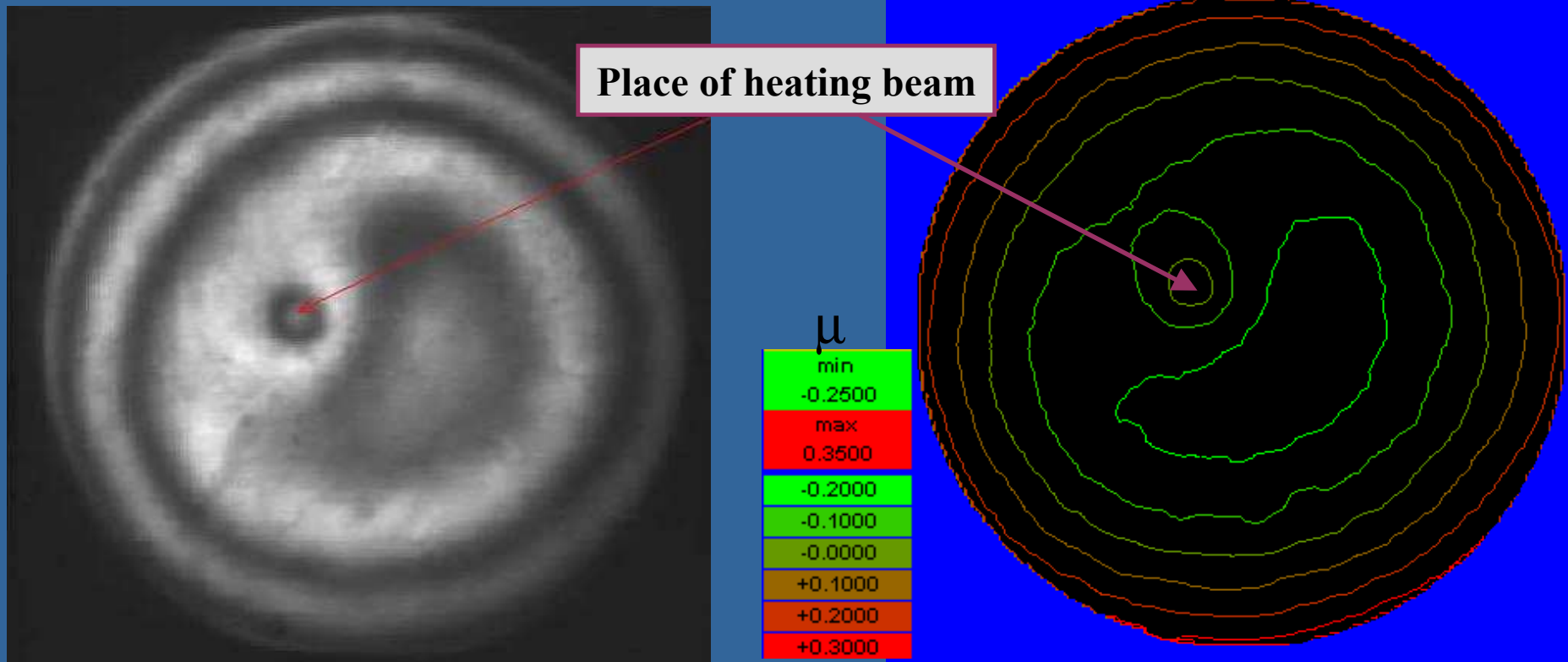
Diameter 40 mm

Thickness 10 mm





# Phase map of optical sample heated by CO<sub>2</sub> laser

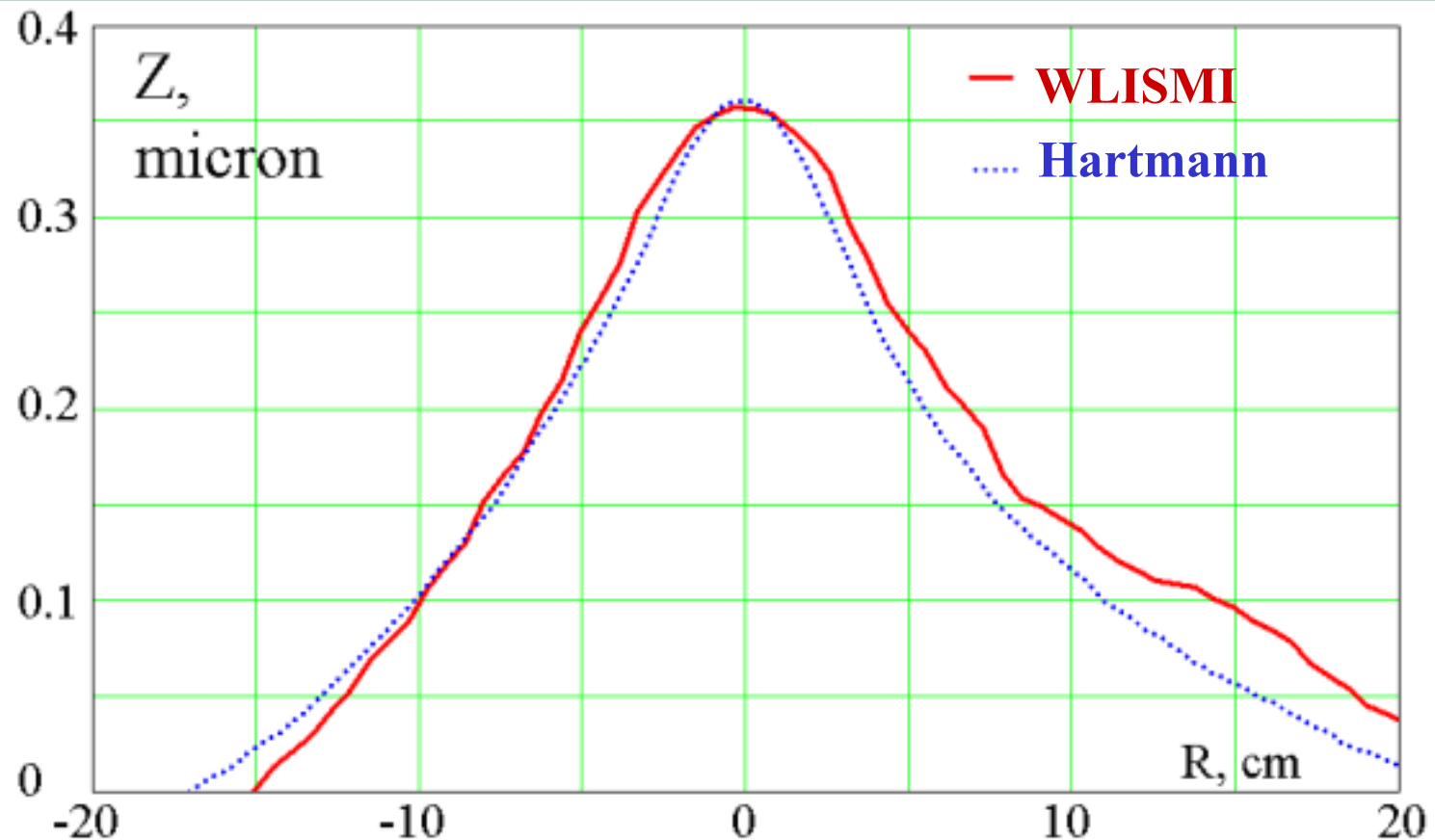


Thickness - 15 mm  
Diameter - 85 mm

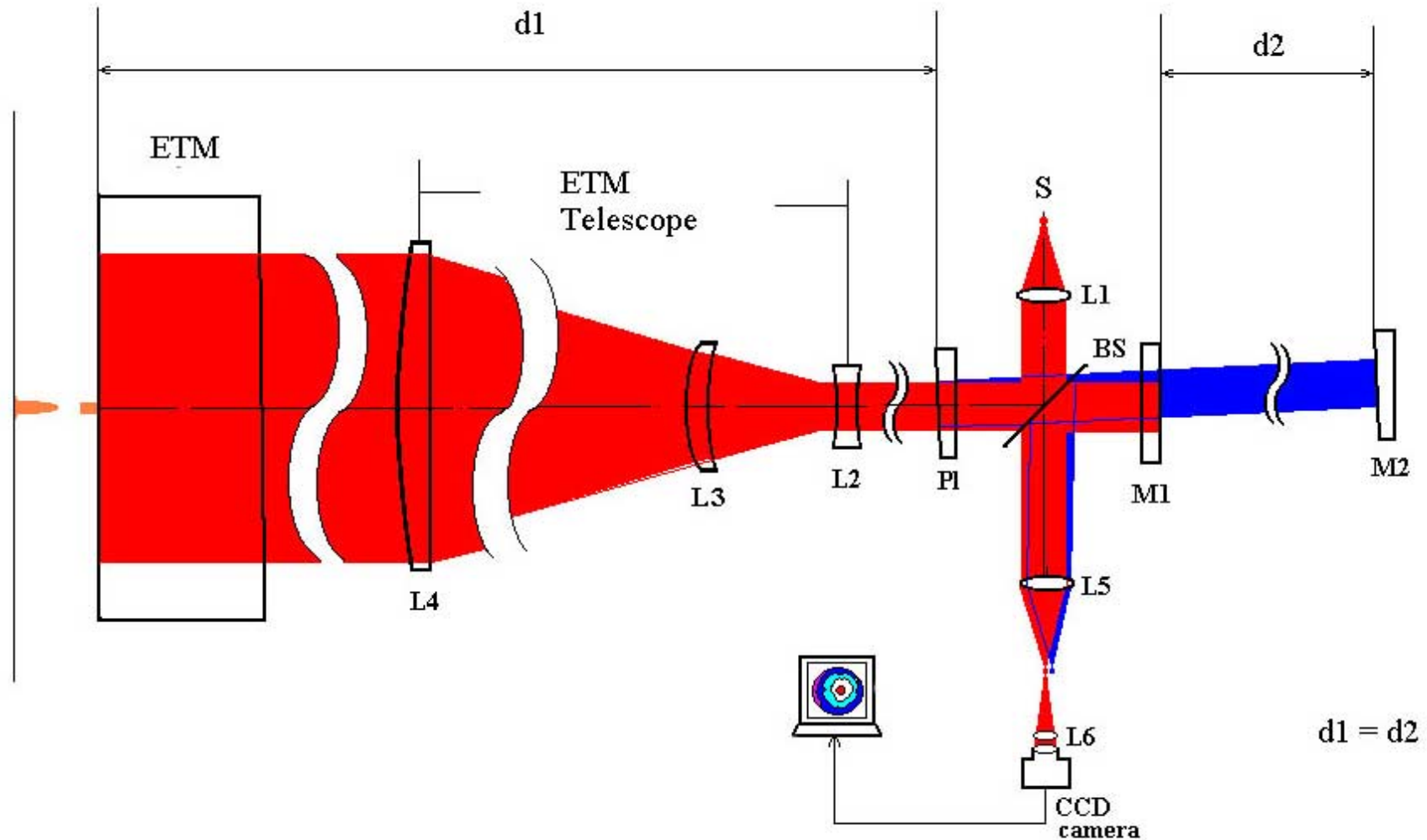
# Simultaneous measurements of optical depth profiles under heating using two different techniques

CO<sub>2</sub> laser - 120 mW

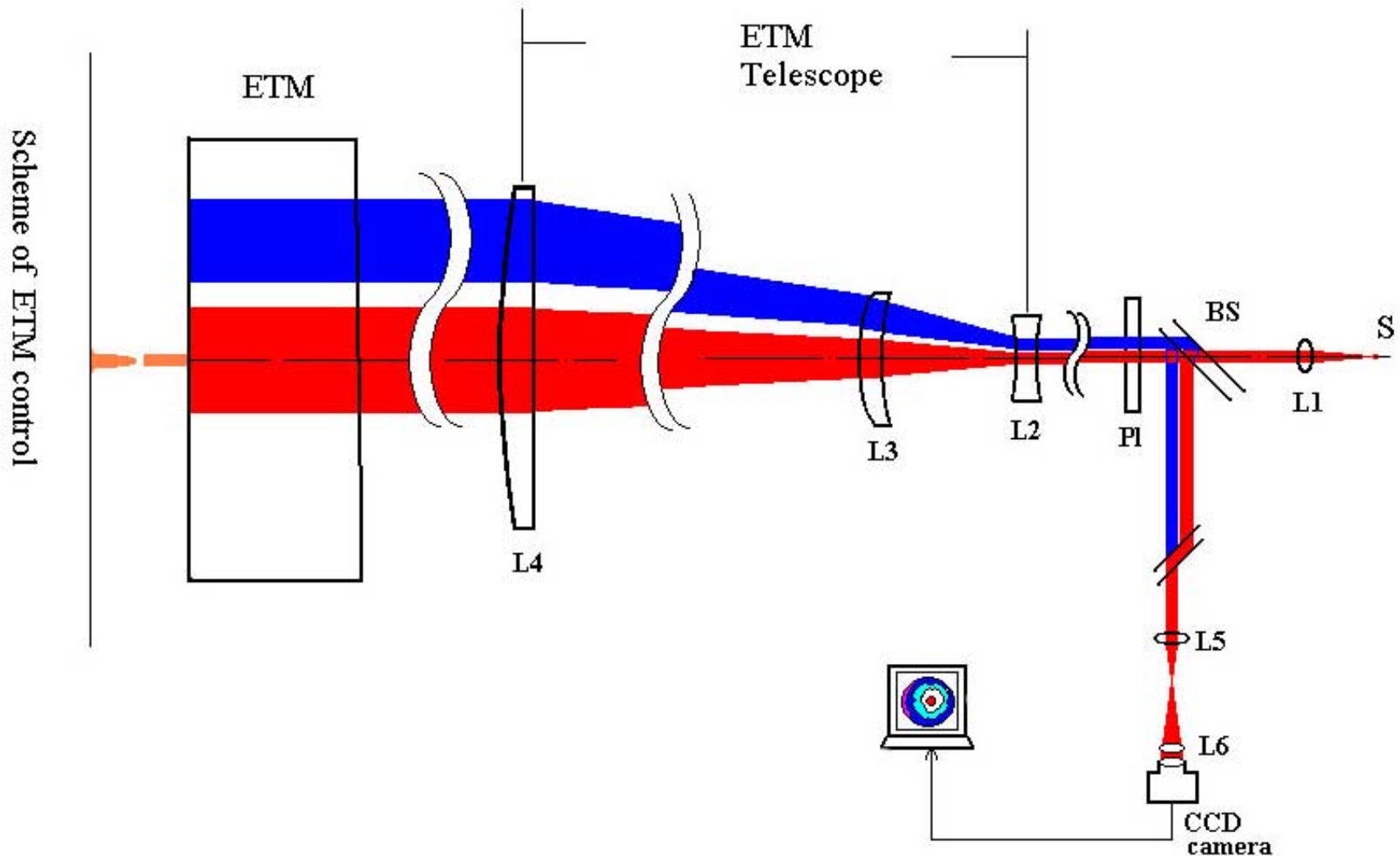
Beam size - 7 mm



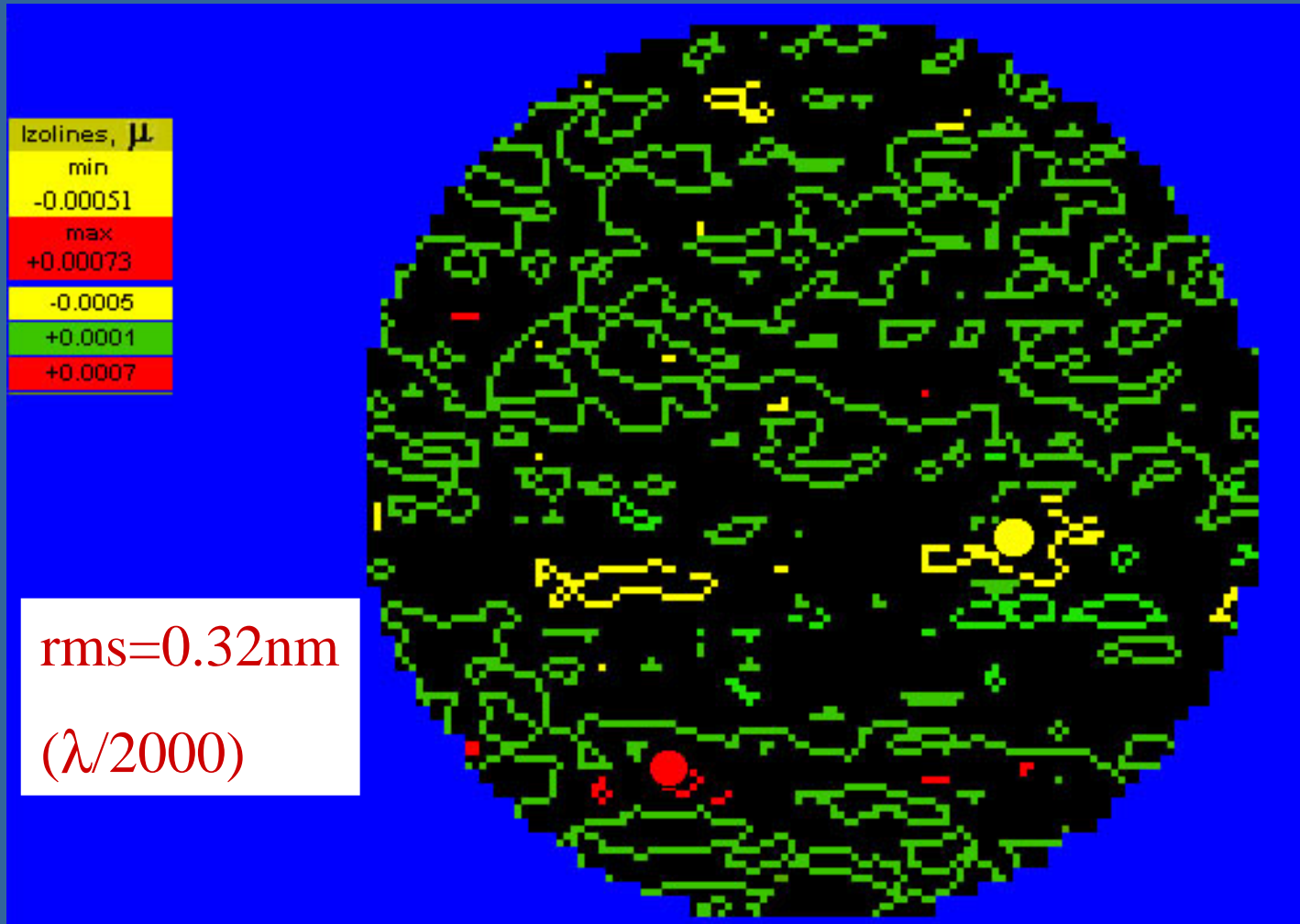
# White Light *In Situ* Measurement Interferometer. How to install in LIGO interferometer?



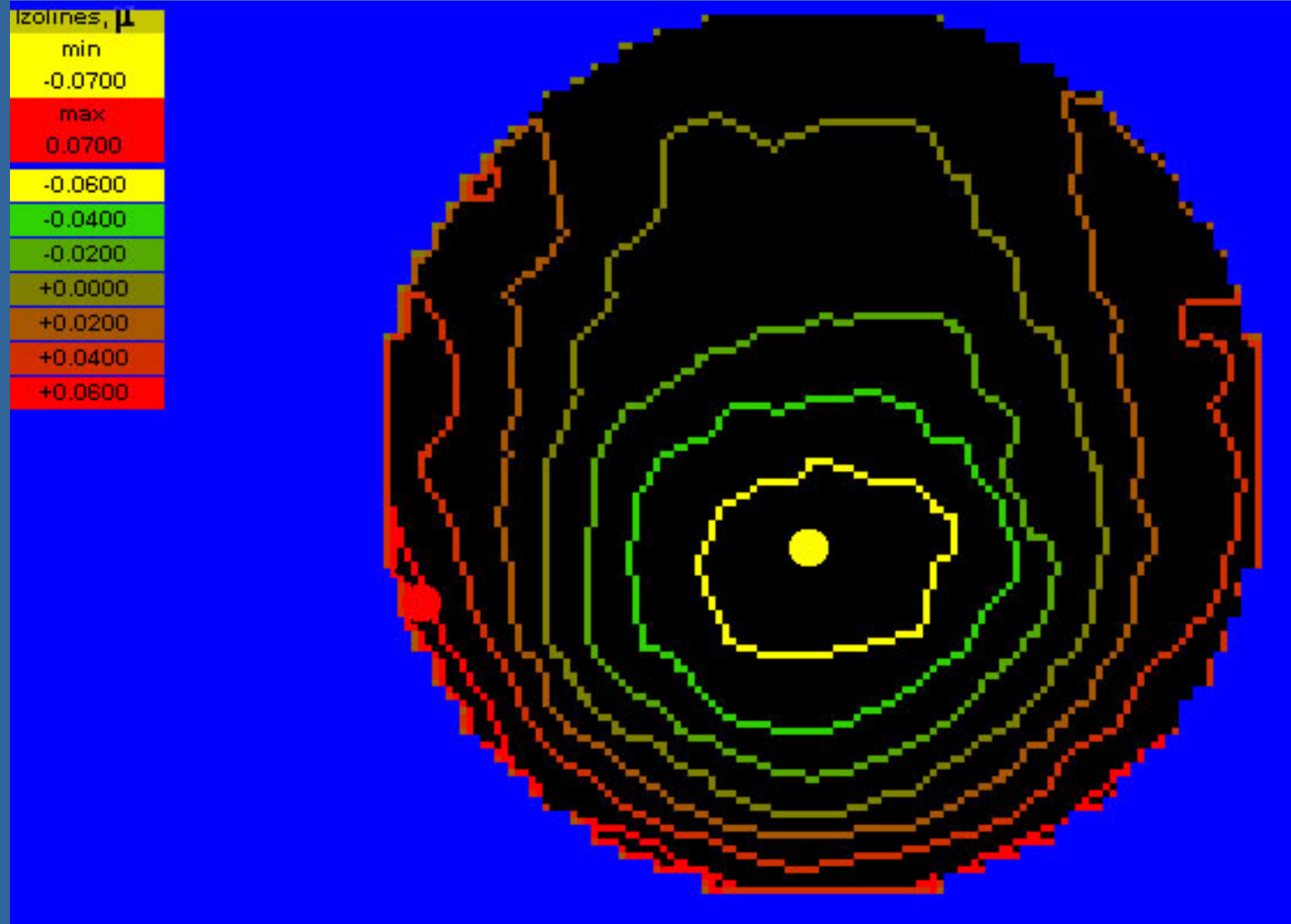
# White Light *In Situ* Measurement Interferometer. How to install in LIGO interferometer?



# Experimental results. No heating.



# Experimental results. Heated by CO<sub>2</sub> laser.



# Conclusion

- ◆ **LIGO-IAP Lab has been equipped with several instruments developed at IAP for High-Precision Characterization of LIGO Optical Components**
- ◆ **25 cm aperture white-light phase-modulated interferometer (WLPMI) for preliminary control of LIGO Core Optics has been implemented**
- ◆ **Simultaneous measurements of optical depth profiles under heating using two different techniques have been performed**
- ◆ **Version of WLPMI for installation on end station is tested experimentally.**

