



Bilinear noise mechanisms

e2e 15may02 dhs

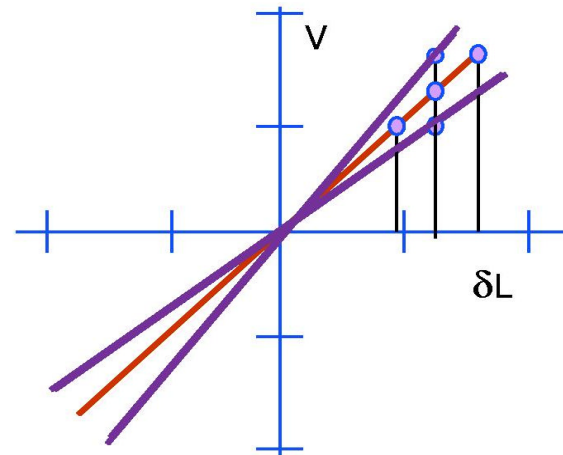
- Attempt to list the kinds of bilinear effects that could lead to noise
- Needs discussion on prioritization and ease of implementation
- Should lead to description of data needed (Detector) and models needed (e2e), ideally with people's names

Stan's Hobbyhorse: Bilinear Noise Sources

- Simplest Example:
 - » Sensing of arm length difference is proportional to input laser intensity

$$" \Delta L " (t) \approx I(t) \delta L(t)$$

$$" \Delta \tilde{L} " (\omega) \approx \tilde{I} (\omega_1) \delta \tilde{L} (\omega \pm \omega_1)$$



- Noise term linear in two variables (“bilinear”) creates output noise at sum and difference frequency



Importance of Bilinear noise mechanisms (Stan)

- Our interferometer configuration is insensitive to (most) first order noise sources
- “Traditional” noise investigation techniques (transfer functions, coherence) don’t pin-point bilinear sources
 - » Requires alternative techniques (e.g., addition of band-limited white noise)
- Understanding full nature of noise source gives experimenter two chances to reduce the output noise
- Bilinear noise sources are fairly common
- Most importantly, e^2e is a good tool for investigating bilinear noise sources



Candidates for interactions

- In general:
 - mechanical or optical offset, RMS usually (could also be DC...)
 - » Static system are susceptible to simpler solutions
 - » sensing system noise at baseband or at modulation frequency
 - A product of some two things below
 - optic axial position << Easy (...says Hiro)
 - optic angles << Easy
 - optic reflectivities/losses << Easy
 - optic ROC << Possible (reliability under investigation)
 - optic beam positions (wrt sweet spot) << Easy
 - scattered light amplitude
 - spectra at the input and the output of the Mode Cleaner
 - optic angle (dynamic) << Easy RNA FRED laser intensity fluctuations
<< Easy (data needed) RICK
 - laser frequency noise << Easy (data needed) RANA and RICK
 - laser pointing noise << Easy (data needed) DAVE O
 - modulation frequency noise << Easy (data needed)
 - modulation intensity noise << Easy (data needed)
- In the main optical system (cavities or michelson), or the mode cleaner



Some examples

- Scattered light
 - » amplitude of scatterer \times relative velocity of components
 - » Hiro: Possible (we discussed about a simple implementation of this noise)
- TM angular motion
 - » offset from 'sweet spot' \times angular noise of optics
 - » Hiro: Qualitative answer is yes, quantitiveness is not clear (reliability of time domain modal model with finite modes)



More examples

- Feedthrough of laser intensity fluctuations
 - » offset from servo null, any length servo (static or RMS dynamic) x laser intensity fluctuations at baseband
 - » Hiro: Not ready : Common mode servo and more refined PSL (and IOO) needed
- Feedthrough of laser frequency fluctuations
 - » imbalance of reflectivity of arms x laser frequency noise
 - » Hiro: Not ready : Common mode servo and more refined PSL (and IOO) needed
- Feedthrough of sideband imperfections
 - » Sideband imbalance x LO oscillator intensity noise
 - » Hiro: Not ready : Common mode servo and more refined PSL (and IOO) needed
 - » --- passage through the Mode Cleaner



More examples

- Rai: cross product of mode cleaner motion with other fluctuations - this one worries me more now as I see how the amplitude of the sidebands are modulated by the mode cleaner
 - » Hiro: Not ready : Need to make a refined mode cleaner model
- Rai: need to establish what causes the quadrature term in a given error signal to fluctuate, this is becoming a bothersome problem especially as we try to increase the gain either optically or electronically
 - » Not itself a bilinear process, but a good thing to pursue!
 - » Hiro: Possible, now that Matt has made a good LSC model



Discussion

- Most important use of e2e:
 - » May have already dreamt up 80% of the important bilinear sources
 - » The things we have not thought about are the most useful outputs from the e2e model; should not lose that focus
- Make sure that the inputs are there
- MC input data: mirror angles (HF and LF motion)
- Remote control room – 10% of someone's time
- Bi-spectrum monitor from Steve Penn
 - » <http://www.ligo.caltech.edu/docs/G/G010331-00.pdf>