

# Gravity gradient studies

Firenze, Pisa, Roma and Urbino

Francesco Fidecaro, Pisa

# Physics near the low frequency limit

Inspiral when massive objects are involved

Signal goes like

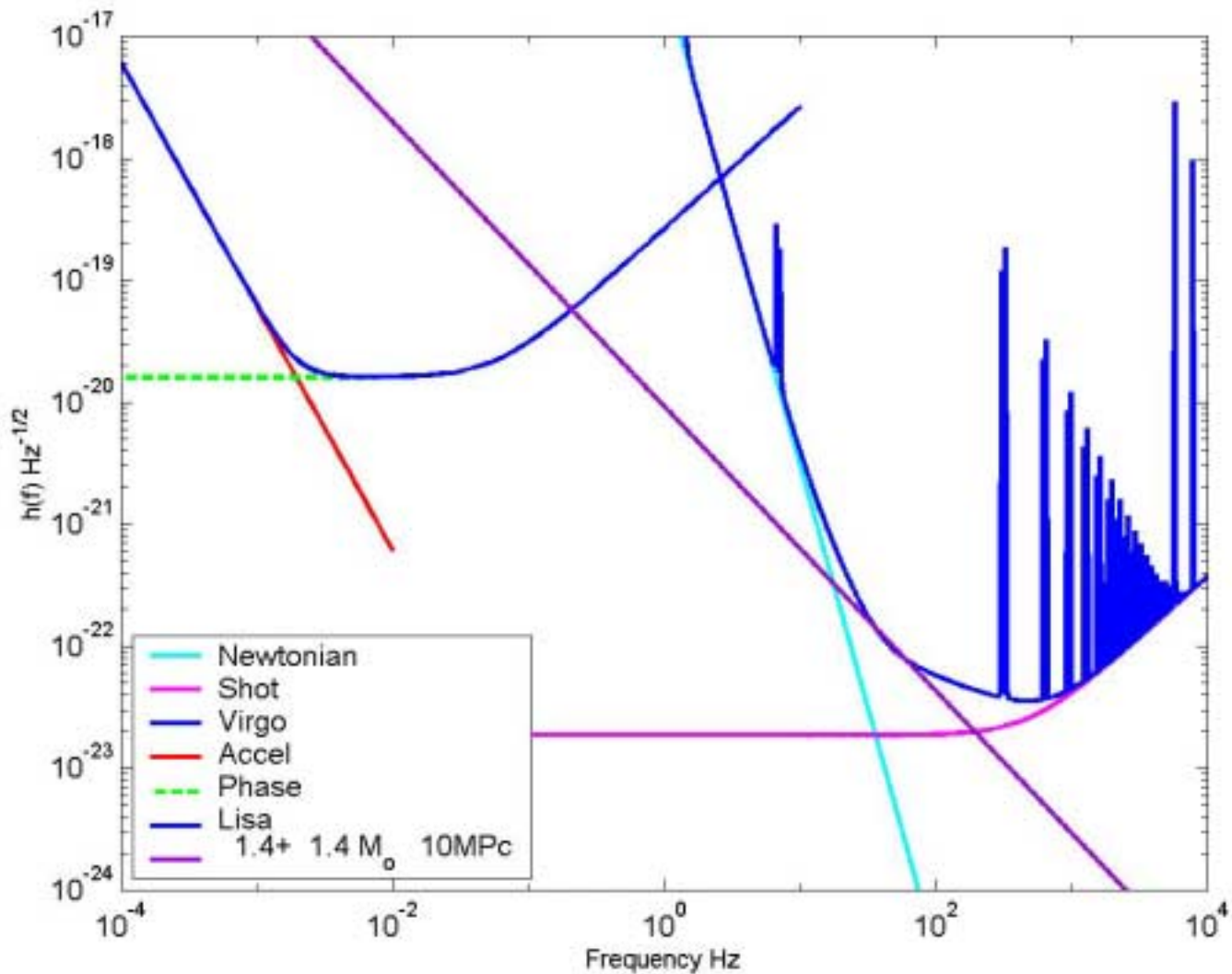
➤  $1/R, M^{5/6}, f^{-7/6}$

There is a maximum frequency

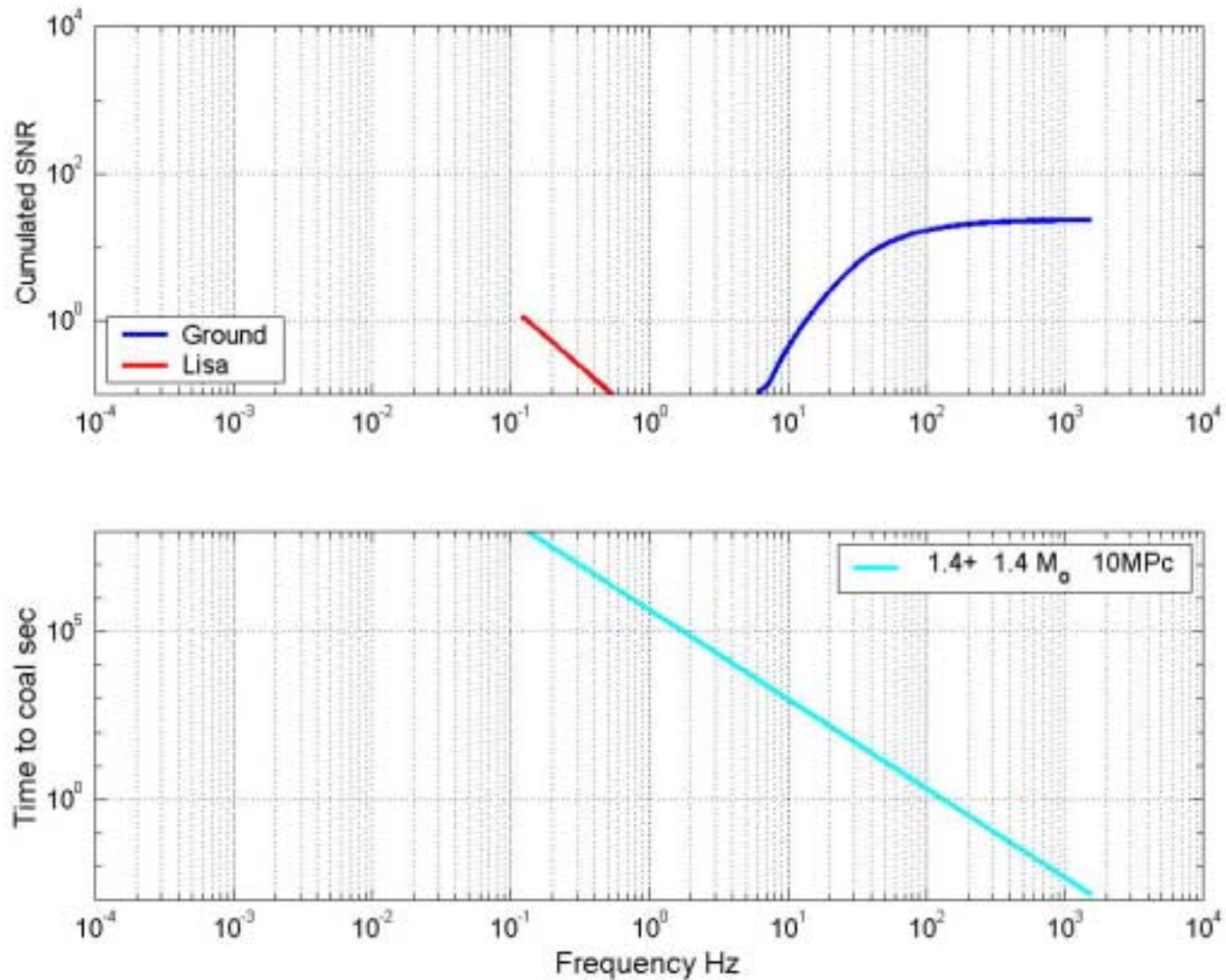
- Ringdown (Ratio:  $\sim 700$  Hz)
- Last stable orbit (Grishchuk, Lipunov, Postnov, Prokhorov and Sathyaprakash, astro-ph0008481)
- Red shift (not considered here)

GW from known pulsar

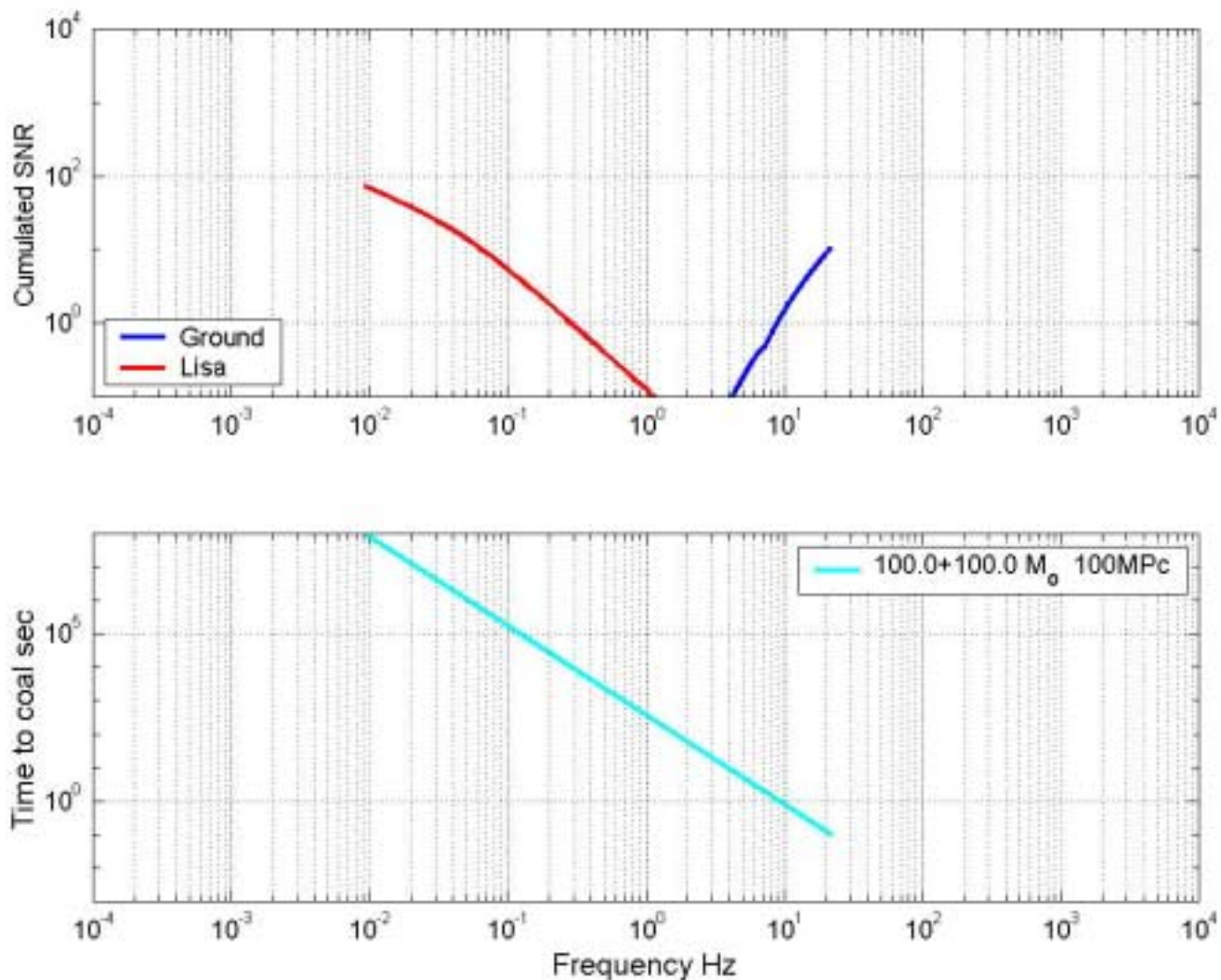
# Detectors capability

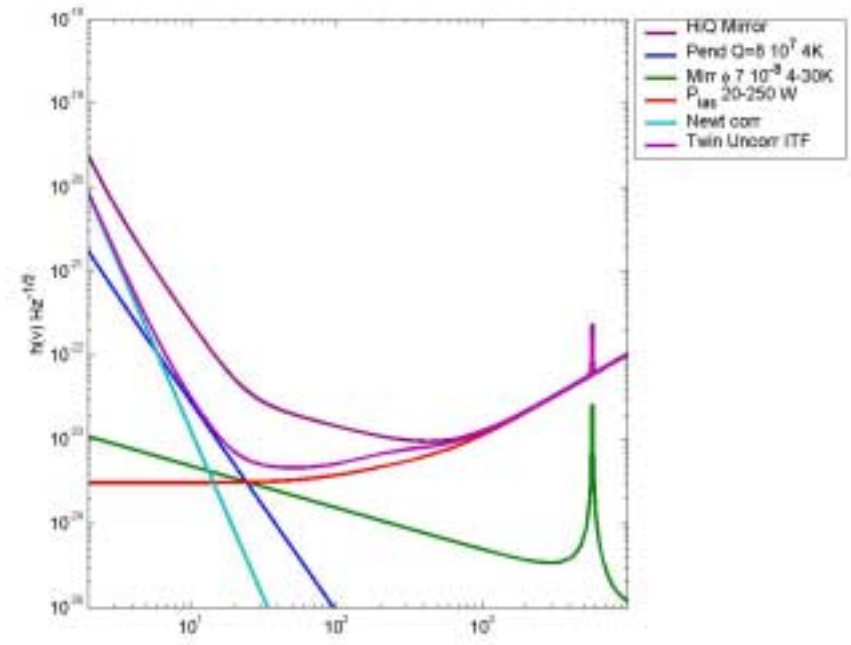
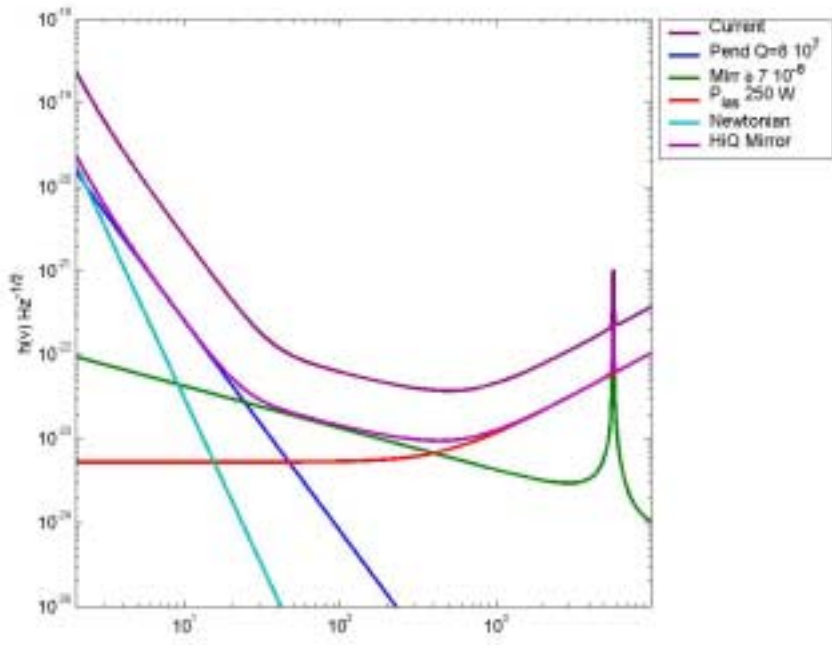


# Integrated SNR NSNS



# 100+100 $M_{\odot}$ 100 Mpc

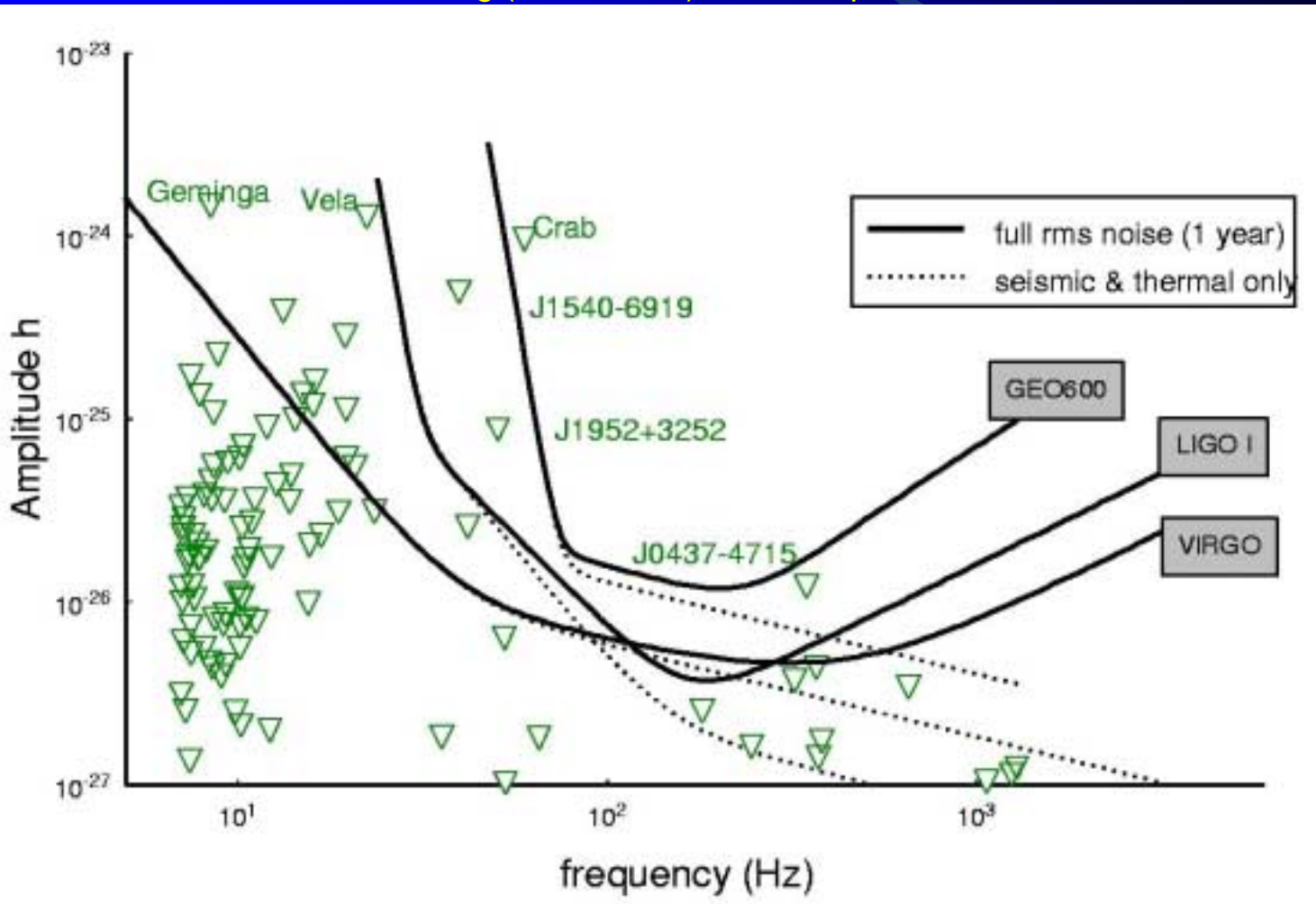




# Slowdown from pulsar

Upper limits on amplitudes from known pulsars, set by assuming spindown due to the emission of gw energy. The points represent all pulsars with gravitational wave frequencies above 7 Hz and amplitudes above  $10^{-27}$ .

Expected sensitivities of three first-generation interferometers in a one-year observation, and the thermal noise limits on narrow-banding (dotted lines). K.A.Compton and B.F.Schutz, Cascina, 1996



# Local gravity fluctuations

Can't be separated from true GW signal

Related to local mass movement or density fluctuations

“Stick out” of the seismic wall

Work by Weiss, Saulson, Hughes and Thorne, Cella



# Seism induced gravity fluctuations

Seismic waves crossing interferometer corner and end points

Horizontal shear waves have no effect

Longitudinal waves have density variation

Rayleigh waves give the main contribution:, combination of horizontal+vertical motion at free boundary.

These are surface waves, with specific propagation characteristics

# Study of gravity fluctuation

Suggested by Thorne: ground geophysical information, layer position and composition

Wave speed

Local diffusion

To construct a model that predicts the composition of a seismic wave (frequency, modes, direction) using a finite set of measurements

# A dedicated study

Groups from Pisa, Firenze, Urbino, Roma

Original Proposal by

- G. Calamai, E. Cuoco, P. Dominici, M. Mazzoni, M. Ripepe, R. Stanga, G. Losurdo, A. Bertolini, N. Beverini, F. Caratori, C. Carmisciano, G. Cella, O. Faggioni, I. Ferrante, F. Fidecaro, F. Strumia, R. Tripiccione, A. Viceré, E. Majorana, P. Puppo, P. Rapagnani, G.M. Guidi, F. Martelli, F. Vetranò

Funded independently of Virgo

Geophysical data collection and measurements

Analytical and numerical model

Development of cheap vertical accelerometers

# Goals

Establish the geology and the geophysics of site

Establish analytic and numeric model for wave propagation, compute effects on test mass

Simulate response of an array of accelerometers

Develop code to test methods for gravity gradient subtraction

Prototype vertical accelerometer

Prototype field data acquisition system

# Geophysics

## Geophysical model in traditional way

- Three homogenous plane parallel layers
- One cylindrical layer on top (Arno river sediments)

## Mathematical model in progress

- Boundary conditions between layers
- Eigenmodes classification
- Dispersion relations

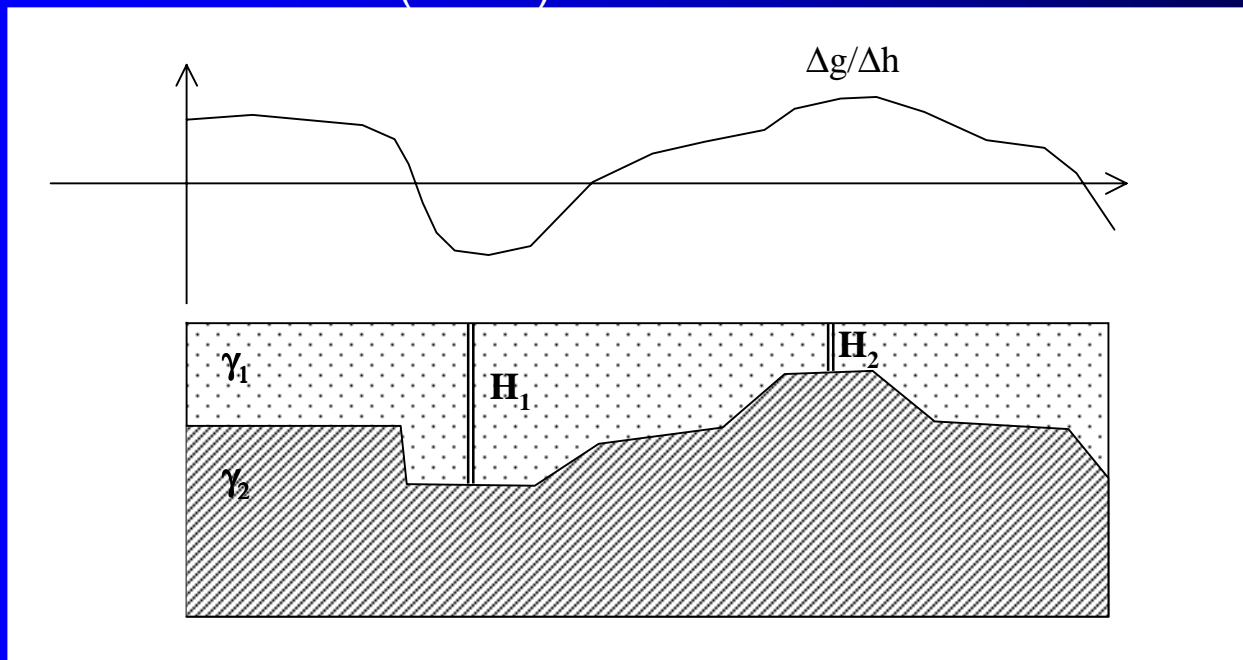
## Comparison with FEM numerical approach

Comparison with LIGO sites would be useful, already benefited from discussions with Sz. Márka, would like to make plans for joint efforts

# Seismic gravity gradient: geology

Geological layer model of Cascina site being completed

- Based on available maps, sampling and seismology
- Differential gravity measurements
  - To obtain a profile of the depth of the effective boundary between the first two layers.
  - About 150 (+/- 30) measurements needed



# Thorne Hughes paper

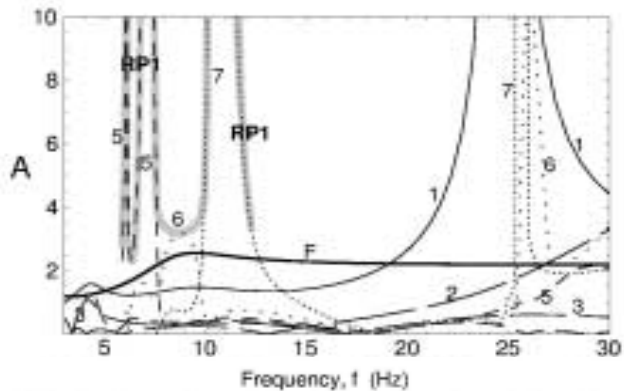


FIG. 7. Properties of the lowest 8 modes of the 4-layer Hanford model, including coupling between P- and SV-waves produced at boundaries between layers and at the earth's surface.

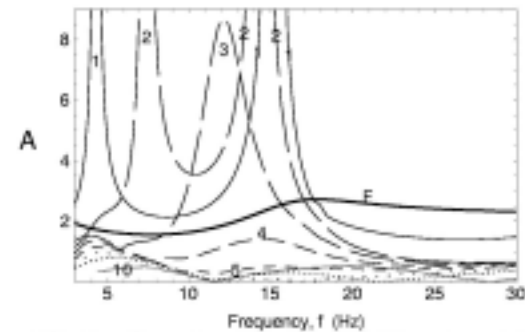
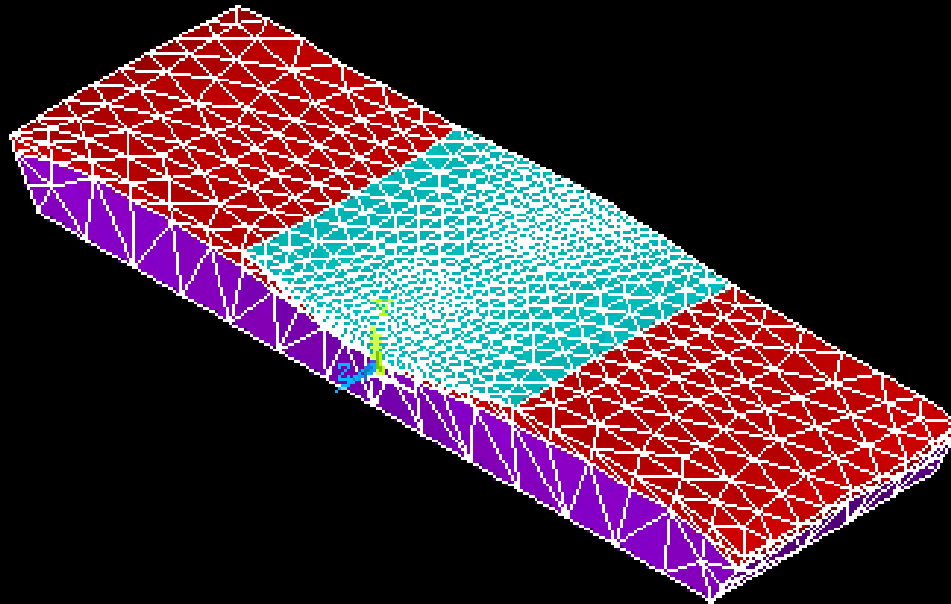


FIG. 10. Properties of the lowest 10 RS modes and the RF mode of the 4-layer Livingston model. (Modes RS8 and RS9 are not shown; their curves are sandwiched between 7 and 10.)

# ANSYS Simulation

1



```
ANSYS 5.5.1  
JUN 26 2001  
14:43:03  
DISPLACEMENT  
STEP=1  
SUB =7  
FREQ=.723777  
PowerGraphics  
EPACET=1  
AURES=Mat  
DMX =.164E-06
```

```
*DSCA=.458E+10  
XU =.57735  
YU =.57735  
ZU =.57735  
*DIST=8450  
*XF =13.72  
*YF =25.08  
*ZF =-2500  
Z-BUFFER
```



# Seismic gravity gradient correction

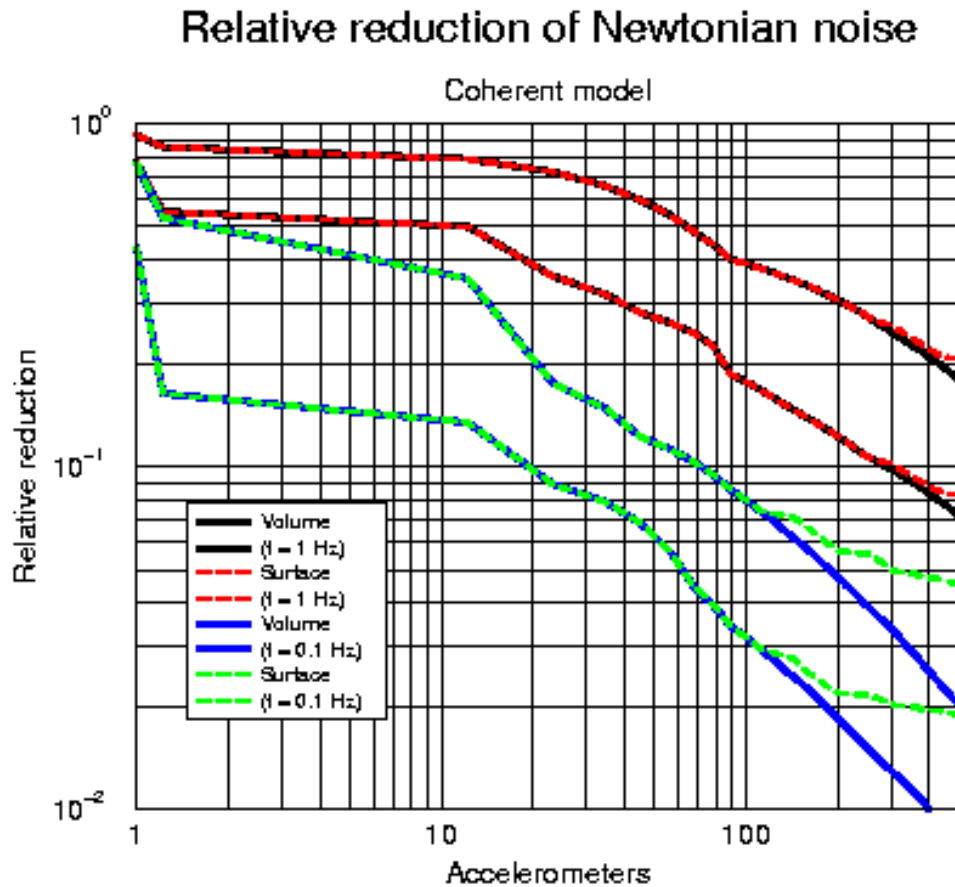
Based on ground motion measurement around mirrors

Establish expected correlation between motion and local gravity field

Linear system: take projection of seismic modes on accelerometer array

Invert accelerometer array readout to obtain expected local gravity field

# Noise subtraction (Cella)



# Vertical accelerometer

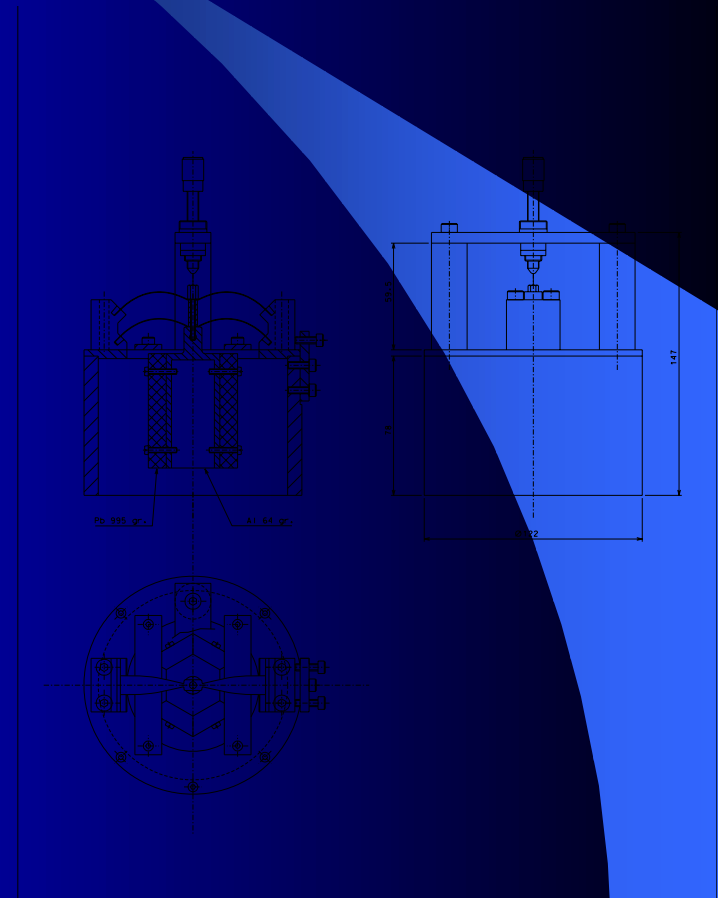
Development of a dedicated vertical accelerometer

Based on GAS concept

Capacitive readout

To be produced in hundreds

- Ease of assembly
- Reproducible
- Stable



# Atmospheric gravity gradient (not yet proposed)

Work by R Weiss, P Saulson, T Creighton

Building induced eddies density variation  
(Creighton)

Ground induced density variations

Wind induced turbulence: building, trees

Sound wave: 74 dB correspond to  $0.1 \text{ Pa} = 10^{-6}$   
bar

$10 \times 10 \times 30 \text{ m}^3$  : mass change of 3 g

Acceleration at 10 m  $2 \times 10^{-15} \text{ ms}^{-2}$

At 10 Hz  $\delta h = 1.6 \times 10^{-22}$  down to  $5 \times 10^{-23}$  with  
more appropriate treatment

# Actual conditions

Not much is known of relevant parameters and of their statistical properties

Need measurements, usual knowledge doesn't really apply

Need to know this on a scale of  $\sim 100$  m

Meteo conditions

Pressure and pressure spatial correlation

T and T spatial correlation

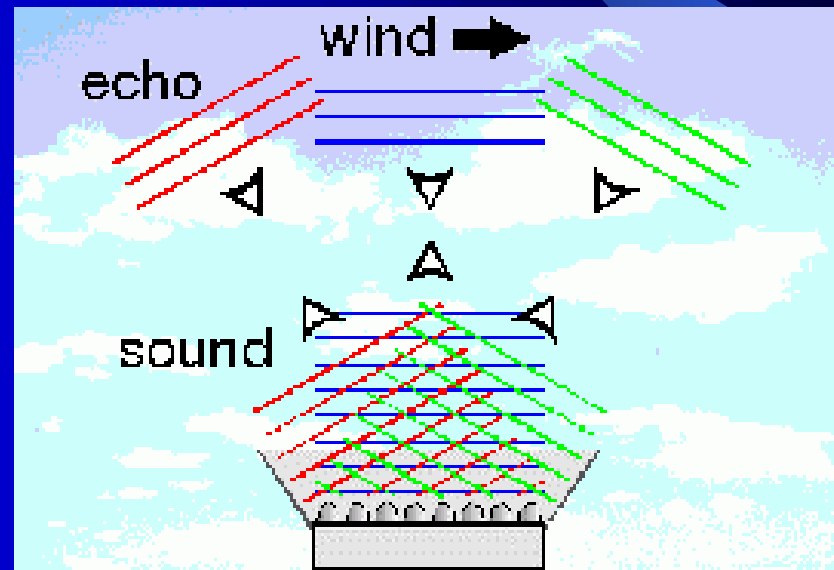
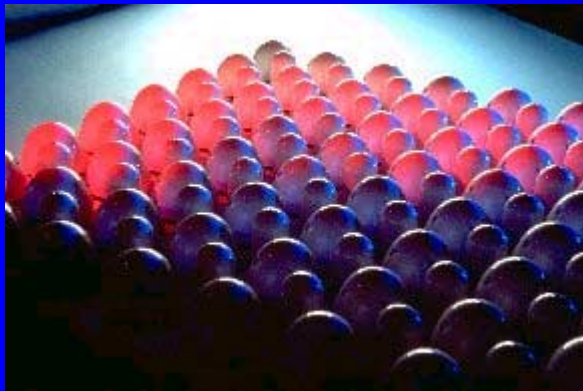
# Measurement tools

Temperature correlation on ground

Infrasound microphone

Instruments for atmospheric research

Sodar



Also measurements for active optics correction,  
scintillation on short distances

# Perspectives on atmospheric effects

Will propose to Virgo some long due measurements on site to assess effect

Infrasound

Wind

Temperature

Then evaluate whether to propose a dedicated effort