CfA Gravitational Wave Astrophysics Group

CGWAG

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LIGO-G000285-00-D

Astrophysical Sources of Gravitational Waves

CGWAG Current Members:

Vicky Kalogera — CfA Clay Fellow

Chris Belczynski — Graduate student

Collaborators: H. Apostolatos (U. Athens), D. Lorimer (Arecibo), (non LSC members) R. Narayan (Harvard U.), and others.

WORK PLAN

Binary Inspiral

- (1) Rates
 - * formation
 - * detection

(LIGO I and II)

- (2) Binary Properties
 - * spin-orbit misalignment
 - * precession-modified waveforms

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(1) Coalescence Rates

- * Empirical (NS/NS) : w/R. Narayan, D. Spergel, J. Taylor
- Sources of uncertainty
- Rate estimates with systematic and random errors
- Large scale population syntheses
 with focus on:

* Theoretical

- extensive parameter study
- calibration of models with observational constraintswith goals:

(NS/NS, BH/NS, BH/BH)

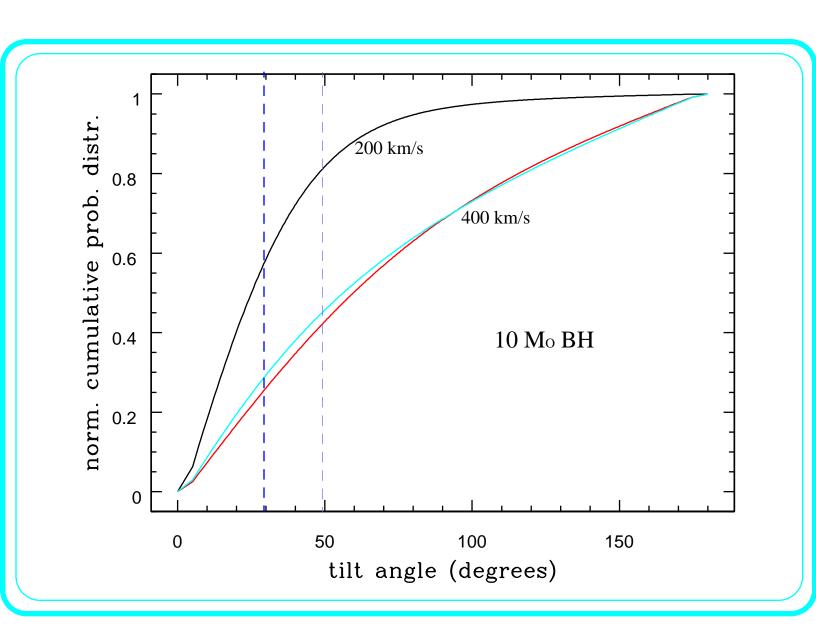
- obtain conservative ranges
- identify most important sources of uncertainty

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(2) Spin-Orbit Misalignment and Inspiral Waveforms

• Important for BH/NS binaries only

- Detection rate decreased by a factor up to ~4 (Kalogera 2000)
- Collaborative effort (with H. Apostolatos, B. Owen, A. Vecchio) to:
 - 1) identify the regions of the parameter space over which the template database must be expanded.
 - 2) estimate the number of templates needed and the associated computational costs.
 - 3) for a given choice of templates, calculate the expected decrease in the detection rate as a function of the detection threshold.
- Include a realistic spectrum of BH masses
- Examine BH spin evolution



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(3) CW Source Catalog

(ongoing effort with A. Vecchio)

with source classification, position, distance, etc.

Goal: Compile a complete catalog of known radio, X-ray, and gamma-ray sources