## Seismic Isolation Requirements for LIGO II

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- Introduction to the two talks that follow
- Want to give a notion of the constraints and demands the LIGO II design puts on any solution for seismic isolation
- Point of departure:
  - The suspension is the hard part
  - allow its design to establish interface to isolation system

## Basic top-level requirements

- strict vacuum compatibility
- system to fit in present vacuum envelope
- present seismic isolation support piers to be used
- suspension as proposed by GEO SUSpension group
  - suspension and isolation considered separate subsystems
- systems required for both HAM and BSC chambers
  - isolation requirements similar (30 more lax for HAM, less isolation from forcibly shorter suspension)
- multiple optics, telescopes, etc. per chamber
  - two multiple suspensions plus aux. Optics in BSC
  - many optics on the HAM optics table
- some flexibility in placement of optics required
  - 'Schnupp' asymmetry, finding sweet spots

# (draft) Motion requirements (at test mass)

- 10<sup>-19</sup> m/rHz at 10 Hz for Test Masses
- 1e-9 m/sec
- 3x10<sup>-18</sup> m/rHz at 10 Hz for Mode Cleaner
- 1 cm motion to compensate tides
- all internal modes damped
- limited drift etc.
- NB: suspension provides >10<sup>-6</sup> transsmission at 10 Hz



#### Ground noise



## Decisions, decisions

- Appears that both designs are basically capable of meeting requirements
  (one last time:) An embarrassment of riches
- Neither approach quite in a state to best judge which will be 'best'
- Set of criteria established, group of advisors to help evaluate
- Plan to select an approach in April
- Meanwhile, very nice research at breakneck pace.