

## LIGO Laboratory California Institute of Technology *LIGO Laboratory*, MC 18-34 Pasadena CA 91125 USA

## TEL: 626.395.2129 FAX: 626.304.9834

		-	
Date:	<u>15</u> January 2008		Deleted: 4
LIGO Number	T080003- <u>02</u>		Deleted: 01
Subject:	Thermal Noise Increase due to a Gold Coated Barrel		
Authors:	Dennis Coyne		
Distribution:	aligo_coc, aligo_aos, aligo_sys		

Phil Willems has proposed the addition of a gold coating on the barrel of the test masses and the compensation plates in order to reduce the emissivity and reduce the radial temperature gradients due to radiation from the barrel<sup>1.2</sup>. Since gold is a lossy material it may also help to passively damp higher order acoustic modes which might otherwise be excited into oscillation by higher order optical cavity modes. However, there is concern that the lossy gold coating might lead to an unacceptably large increase in Brownian thermal noise in the optical readout. The purpose of this technical note is to report on a finite element analysis used to calculate the increase in thermal noise (using Levin's approach<sup>3</sup>) due to the addition of a gold coating.

An axisymmetric finite element model of the Advanced LIGO test mass was developed, as depicted in Figure 1. For simplicity (and to enable an axisymmetric model) a perfect right circular cylinder was assumed; The following geometric features were not included: wedge angle, bevels, flats and bonded "ears" (used to weld the fused silica suspension ribbons). In addition, the gold layer will need a protective SiO<sub>2</sub> overcoat layer and may need an interstitial layer to bond it to the substrate (e.g. Ni). Neither of these layers are included. The Si O<sub>2</sub> protective layer should be low loss and similar to the substrate. The interstitial layer may be exceedingly thin and inconsequential.

The elements approximating the fused silica material of the test mass substrate are comprised of parabolic quadrilateral axisymmetric solid elements. The High Reflectance (HR) and barrel coatings are comprised of axisymmetric shell elements of appropriate thickness. The mesh was refined to insure convergence. The parameters and property values used in the analysis are listed in Table 1. <u>Similar models and analysis of thermal noise were</u> performed and reported previously<sup>4</sup>. The agreement with equations from Liu & Thorne's analysis<sup>5</sup> (to 3 decimal places) plus the reasonableness of the stress contours gives one confidence in the results. However, it seems that the I-Deas finite element code<sup>6</sup> does not

masses", Physical Review D, vol. 62, 20 Nov 2000.

LIGO LABORATORY

Page 1 of 6

<sup>&</sup>lt;sup>1</sup> P. Willems, "Heating of the ITM by the Compensation Plate in Advanced LIGO", LIGO-T070123-02, 3Dec2007. <sup>2</sup> Note that the gold coating is not part of the baseline AOS or COC design as yet. This is just an memo on the feasibility of this proposed coating from the thermal noise perspective.

<sup>&</sup>lt;sup>3</sup>Y. Levin, "Internal thermal noise in the LIGO test masses: A direct approach", Phy Rev D, v57, n2, 15Jan98, p659-663. <sup>4</sup>D. Coyne, Thermal Noise Calculation with Inhomogeneous Loss using the Finite Element Method: Application to Test <u>Mass Optics with Coating Loss, Attachments and Composite Assemblies, LIGO-T020070-01, 11 Jul 2002.</u>

 $<sup>\</sup>frac{5}{2}$  Y. Liu and K. Thorne, "Thermoelastic noise and homogeneous thermal noise in finite sized gravitational-wave test

<sup>&</sup>lt;sup>6</sup> I-Deas version 9, Structural Dynamics Research Corp.

use the proper calculation when displaying strain energy<sup>7</sup>. As a consequence I have removed all of the strain energy plots from this version. I also think that a further check of the strain energy integrals (sums) is warranted.<sup>8</sup>

In lieu of strain energy plots, the stress contours are depicted in Figures 2 through 5. The non-zero stresses are  $\sigma_{r_x}$ ,  $\sigma_{\theta_x}$ ,  $\sigma_z$  and  $\sigma_{rz}$  where r is radial, z is along the symmetry axis (cylinder axis) and  $\theta$  is the angle about the z axis. The strain energy density is  $\zeta = \{\sigma\}^T \{\varepsilon\}$ 

since the stress is linearly related to the strain by the Young's modulus,  $\{\sigma\} = [E] \{\epsilon\}$ 

 $\frac{\text{The strain energy is proportional to the sums of squares of the stresses}}{\underline{\zeta} \sim (\underline{\sigma_r})^2 + (\underline{\sigma_\theta})^2 + (\underline{\sigma_z})^2 + (\underline{\sigma_{rz}})^2}$ 

The strain integrals (summations of total element strain) for the substrate, HR coating and barrel coating domains are listed in Table 2. The displacement noise (and strain), at 100 Hz, due to the Brownian motion are also compared in Table 2. The total strain noise (at 100 Hz) compares well with the Bench 6.2 calculation (3e-24  $1/\sqrt{Hz}$ ). The gold barrel coating only increases the strain noise ASD by 1%. This is due to the fact that the Brownian thermal noise is dominated by the HR coating (3e-24 HR, 0.3e-24 substrate, 0.5e-24 gold). Even if new HR coating formulations result in significantly lower loss factor, the HR coating is still likely to dominate over the gold barrel coating contribution.

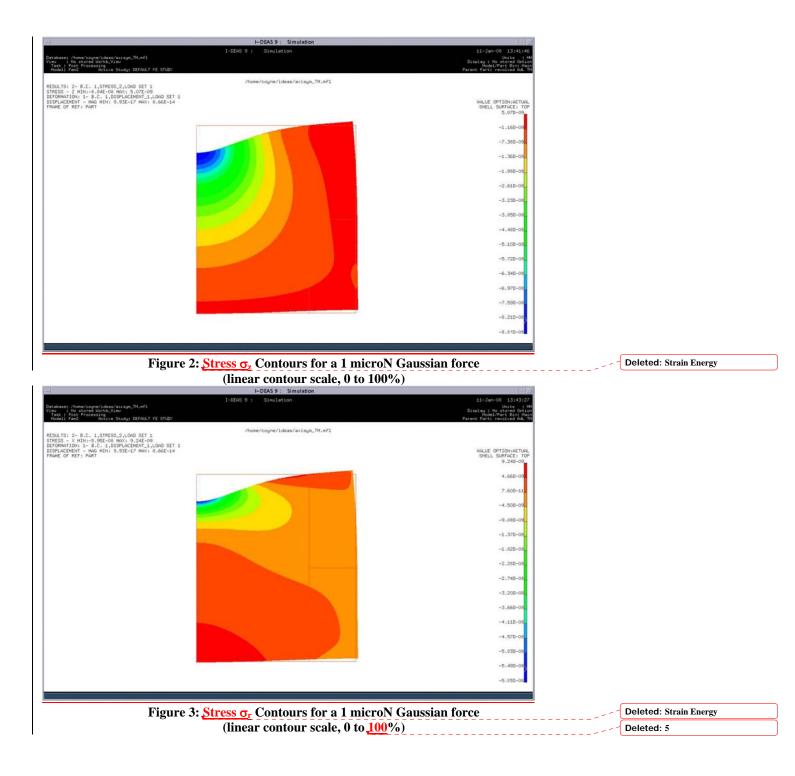
oyne/ideas/axisym\_TM.nf1 ed Workb\_View I High Reflectance (HR) Coating Gaussian pressure (axisymmetric shell distribution elements) Centerline (axis of rotation)-Barrel (axisymmetric shell elements) Substrate (axisymmetric solid elements - finer mesh near center of Kinematic degree-ofpressure) freedom used to eliminate rigid body translation) Figure 1: finite element model

<sup>7</sup> I think it was Andri Gretarsson who first questioned why the strain energy was so low at the center of the beam in LIGO-T020070. Riccardo DeSalvo noticed the same oddity in the -01 version of this memo. <sup>8</sup> A check can be performed with Liu & Thorne's formulation (I've started this, but don't have time now to complete it), or with another finite element code.

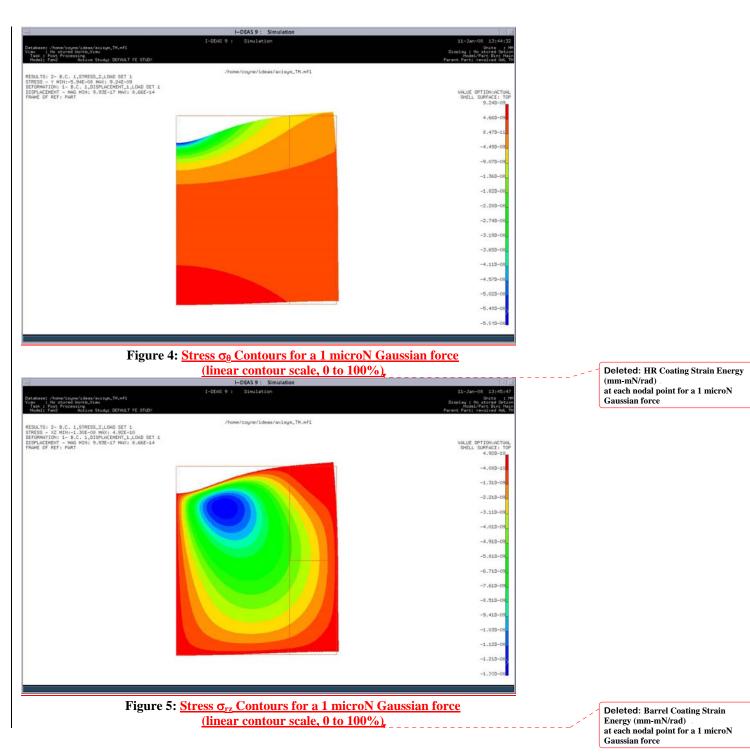
LIGO LABORATORY

Page 2 of 6

Deleted: The variation in strain energy is depicted in the contours of Figures 2 and 3. The variations of strain energy across the HR and barrel coatings are depicted in Figures 4 and 5.



Page 3 of 6



Page 4 of 6

Table 1: Parameters								
Symbol	Parameter	Value	Units	Source				
а	TM radius	0.170	m	M050397-02				
Н	TM thickness	0.200	m	Ibid				
r <sub>0</sub>	Gaussian beam radius	0.060	m	Bench 6.2				
E <sub>FS</sub>	TM elastic modulus	71.8	GPa	Fused Silica; Bench 6.2 uses 72.7GPa				
VFS	TM Poisson's ratio	0.16	-	Bench 6.2 uses 0.167				
φfs	TM bulk loss factor	2.6 x 10 <sup>-10</sup> at 100 Hz	-	Fused Silica: from Bench 6.2: $\phi = c_2 f^n, c_2 = 7.6e-12$ n = 0.77, f = frequency (Hz)				
t <sub>HR</sub>	High Reflectance dielectric coating thickness	10	μm	~40 layers @ λ/4				
E <sub>HR</sub>	HR coating elastic modulus	106.5	GPa	Ta <sub>2</sub> O <sub>5</sub> /SiO <sub>2</sub> : T970176-00 Bench 6.2 uses ~94 GPa				
$\nu_{HR}$	HR coating Poission's ratio	0.21	-	Ta <sub>2</sub> O <sub>5</sub> /SiO <sub>2</sub> : T970176-00 Bench 6.2 uses 0.2038				
Фнr	HR coating loss factor	3.38 x 10 <sup>-4</sup>	-	Eqn[23], G. Harry et. al., CQG, 19(2002), with effective coating elastic modulus, Y'=94 GPa, substrate modulus, Y=72.7 GPa, $\phi_{\perp}$ = 1.36e-4 and $\phi_{\parallel}$ = 1.79e-4				
t <sub>Au</sub>	gold barrel coating thickness	100	nm	Phil Willems estimate 11-dec-2007				
E <sub>Au</sub>	gold barrel coating elastic modulus	80	GPa	H. Anderson (ed.), AIP Physics Vade Mecum, 1981				
$\nu_{Au}$	gold barrel coating Poission's ratio	0.42	-	Ibid				
φ <sub>Au</sub>	gold barrel coating loss factor	9 x 10 <sup>-3</sup>	-	recent unpublished measurement by Andri Gretarsson				

Page 5 of 6

Symbol	Parameter	Value	Units
U <sub>FS</sub>	Total substrate strain energy normalized by the force amplitude squared	2.48e-11	m²/J
U <sub>HR</sub>	Total substrate strain energy normalized by the force amplitude squared	2.27e-15	m²/J
U <sub>Au</sub>	Total substrate strain energy normalized by the force amplitude squared	1.80e-18	m²/J
S <sub>x,FS</sub> (f=100Hz)	Displacement noise PSD contribution from the substrate	3.33e-43	m²/Hz
S <sub>x,HR</sub> (f=100Hz)	Displacement noise PSD contribution from the HR coating	3.91e-41	m²/Hz
S <sub>x,Au</sub> (f=100Hz)	Displacement noise PSD contribution from the barrel coating	8.27e-43	m²/Hz
√ S <sub>x,</sub> (f=100Hz)	Displacement noise ASD: Total	6.34e-21	m/√Hz
	Displacement noise ASD: without barrel	6.28e-21	m/√Hz
	Displacement noise ASD: substrate only	5.77e-22	m/√Hz
	Displacement noise ASD: without HR	1.08e-21	m/√Hz
h(f=100Hz)	Strain noise ASD: Total	3.17e-24	1/√Hz
	Strain noise ASD: without barrel	3.14e-24	1/√Hz
	Strain noise ASD: substrate only	2.89e-25	1/√Hz
	Strain noise ASD: without HR	5.39e-25	1/√Hz
	Strain noise ASD: HR coating only1/√Hz (compare to 2.93e-24 from Bench 6.2)	3.12e-24	1/√Hz

Page 6 of 6