# LASER INTER

#### LASER INTERFEROMETER GRAVITATIONAL WAVE OBSERVATORY

## **SPECIFICATION**

E010102 -B- D

Drawing No Rev. Group

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# Substrate, Input Test Mass, 40M RSE Experiment

			APPROVALS		
AUTHOR:	CHECKED:	DATE	DCN NO.	REV	DATE
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## **Applicable Documents**

LIGO-D010104-A-D Substrate, Input Test Mass, 40M RSE Experiment LIGO-E000408-A-D Mirror Blank Material, Input Test Mass, 40M

LIGO-D000266-A-D Input Test Mass Blank, 40M

## Requirements

## **Physical Configuration**

According to

LIGO-D010104 Substrate, Input Test Mass, 40M RSE Experiment

Fabricate from

LIGO-E000408 Mirror Blank Material, Input Test Mass, 40M

## **Part and Serial Number**

The Serial number shall be per D010104 and of the format:

ITM YY Where

YY is incremental for each optic starting at 01

## **Registration Mark**

Registration mark shall be etched, ground or sandblasted coincident with the registration mark drawn on the Blank within 5 mm. The arrow orientation used on the Blank will be preserved if possible or changes reported in detail. Reference LIGO-D000266, Input Test Mass Blank.

#### Side and Bevel Polish

All sides and Bevels shall be polished from a five micrometer grit finish. These surfaces shall appear transparent with no gray, scuffs or scratches visible to the naked eye when viewed in normal room light against a black background.

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# Substrate, Input Test Mass, 40M RSE Experiment

### **Scratches and Point defects**

There shall be no scratches, sleeks or point defects within the central 10 mm The total area of scratches, sleeks and point defects within the central 30 mm diameter shall not exceed 500 square micrometers (width times length.)

The total area of scratches outside the central 30 mm diameter shall not exceed 5,000 square micrometers.

### **Minimum Point Defect**

Point defects which have a maximum dimension of 5 micrometers are disregarded.

## **Inspection Method**

- 1. The surface is examined visually by two observers independently. The examination is done against a dark background using a three-bundle fiber-optic illumination system of at least 200 W total power. A 100% inspection of the surface is carried out. Pits and scratches down to 2 micrometers in width can be detected using this method of inspection. Any scratches that are detected will be measured using a calibrated eyepiece.
- 2. Further inspection will be done with a 6X eyeglass using the same illumination conditions, again with two observers. Sleeks down to 0.5 micrometers wide can be detected using this method. The surface will be scanned along one or two chords from center to edge, then at ten positions around the edge, and ten to fifteen positions near the
- 3. An inspection is then carried out with a dark field microscope with a similar sampling frequency as described in section 2.

### Surface Figure, measured over the central 30 mm diameter

All specified quantities refer to the physical surface of the optic.

### **Surface 1: Nominally flat.**

Radius of curvature  $> \pm 5625$  meters

Astigmatism: < 20 nanometers (surface peak to valley)

## Surface 2: Nominally flat.

Radius of curvature  $> \pm 5625$  meters

Astigmatism < 20 nanometers (surface peak to valley)

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# Substrate, Input Test Mass, 40M RSE Experiment

### **Surface Errors**

The following root mean square standard deviation ( $\sigma_{rms}$ ) values are calculated from the phase maps which are to be provided with each optic.  $\sigma_{rms}$  is defined as the square root of the mean of the square of each pixel value. Known bad pixels are excluded from this calculation.

#### Surface 1

Low Spatial Frequency Band < 32 cm<sup>-1</sup>: The surface is measured using a commercial phase measurement interferometer and calibrated reference flat. With piston, tip, tilt, power (best fit spherical surface) and astigmatism removed over the central 30 mm diameter aperture:

 $\sigma_{\rm rms}$  <2 nanometers

**High Spatial Frequency Band:** Micro-roughness is measured with a commercial microscopic interferometer or surface profiler.

 $\sigma_{rms} < 0.1$  nanometers

Measured at the following locations:

- 1. The center of the mirror substrate.
- 2. Four positions equally spaced along the circumference of a centered, 30 mm diameter circle.

#### Surface 2

Low Spatial Frequency Band < 32 cm<sup>-1</sup>: The surface is measured using a commercial phase measurement interferometer and calibrated reference flat. With piston, tip, tilt, power (best fit spherical surface) and astigmatism removed over the central 30 mm diameter aperture:

 $\sigma_{rms}$  <5 nanometers

# LASER INTERFEROMETER GRAVITATIONAL WAVE OBSERVATORY LIGO

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# Substrate, Input Test Mass, 40M RSE Experiment

Table 1 Certification Data Requirements					
Specification	<b>Test Method</b>	Frequency of Inspection	Data Delivered		
Physical Dimensions	Visual Inspection	100%	Certification		
Side and Bevel Polish	Visual Inspection	100%	Certification		
Scratches and Point defects	Visual Inspection	100%	Certification		
Surface Figure	Interferometry	100%	Surface Map		
Side 1, Surface Errors – Low Spatial Frequency	Interferometry	100%	Surface Map		
Side 1, Surface Errors – High Spatial Frequency	High resolution Surface Map	100%	Certification		
Side 2, Surface Errors – Low Spatial Frequency	Interferometry	100%	Certification		

Orientation: For the purpose of full surface phase maps the substrate shall be oriented such that the point of minimum thickness shall be at the top center of the data.

Format: All Data shall be delivered according to Table 1. In addition to the hard copy an electronic data set of the phase maps shall be delivered in either ASCII or Vision.OPD format.