

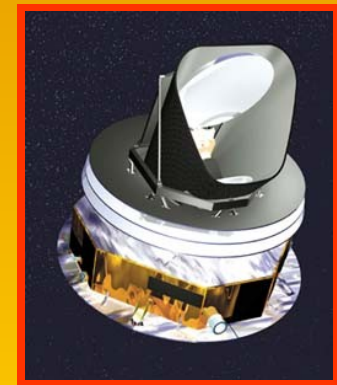
# GW-Detection



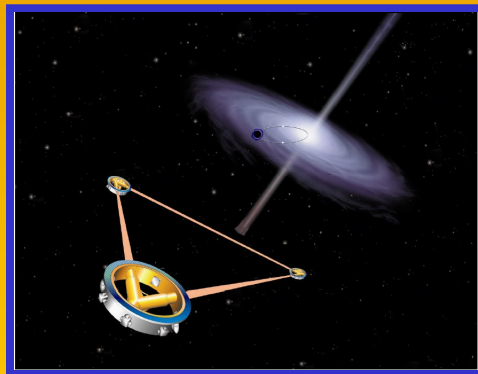
pulsar timing



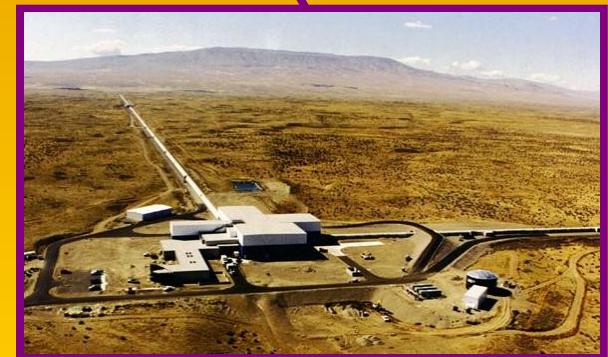
bars



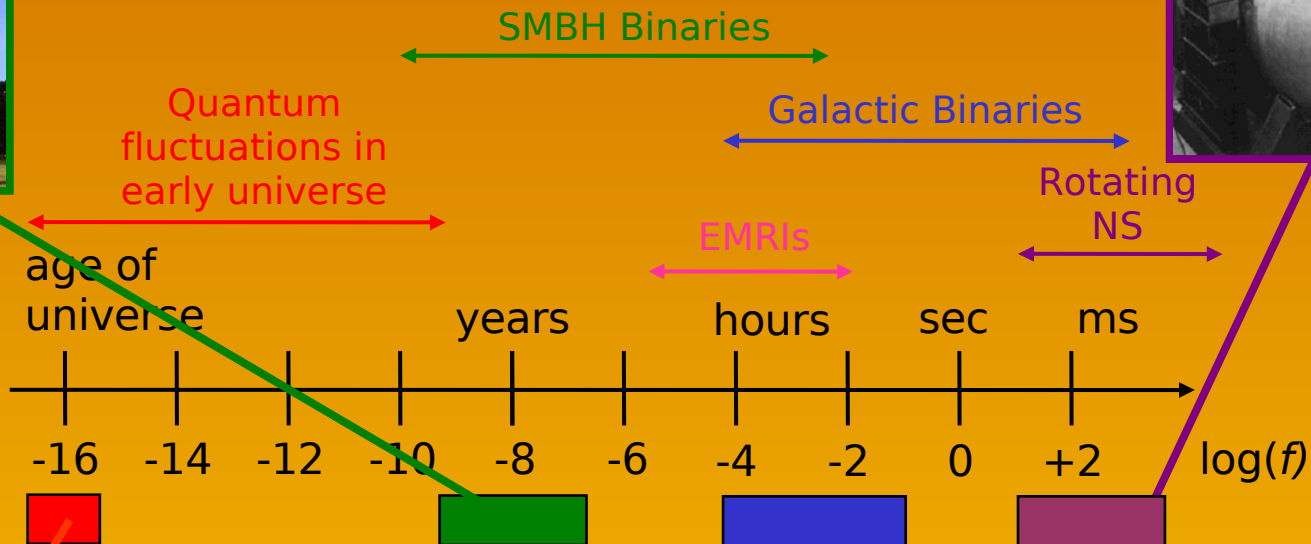
CMB polarization  
G0900657-v1



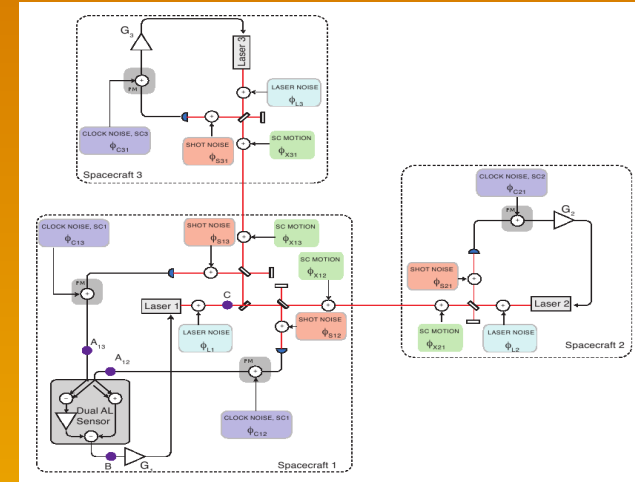
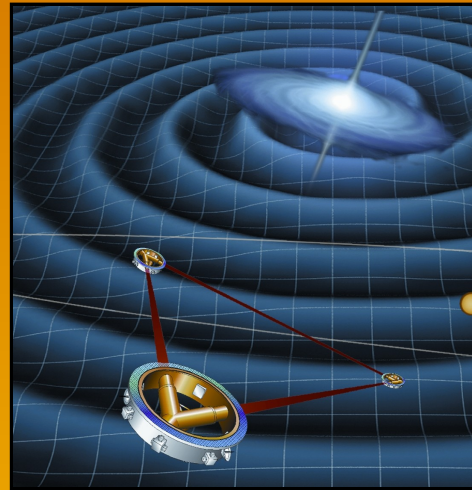
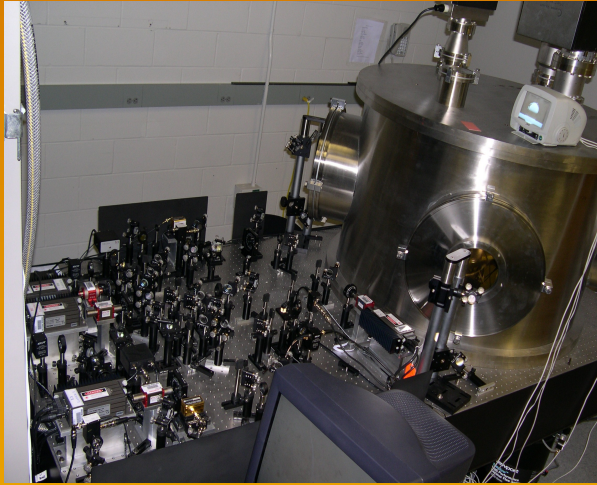
space-based  
interferometers



ground-based  
interferometers



# The long baseline Interferometry of LISA



**Guido Mueller**

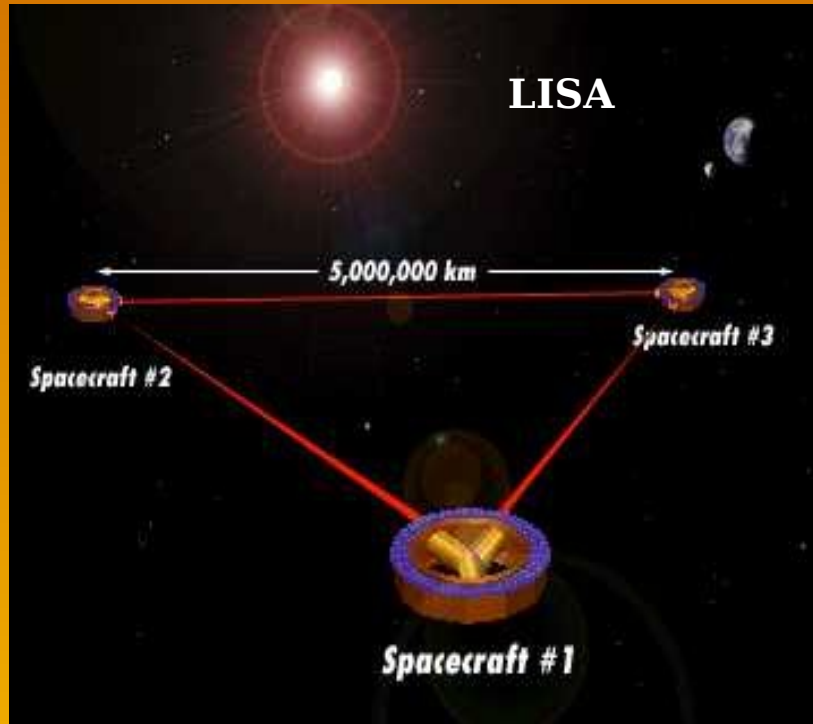
**Dept. of Physics - University of Florida**

**GWADW**

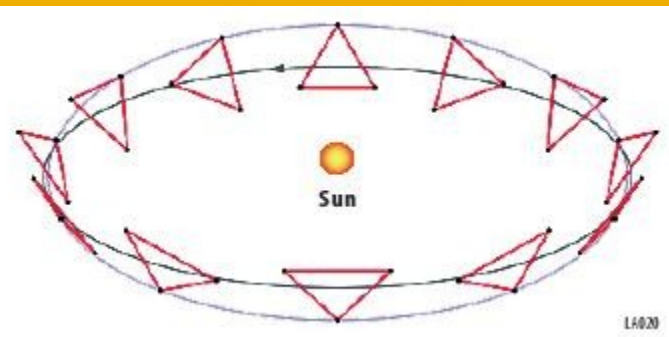
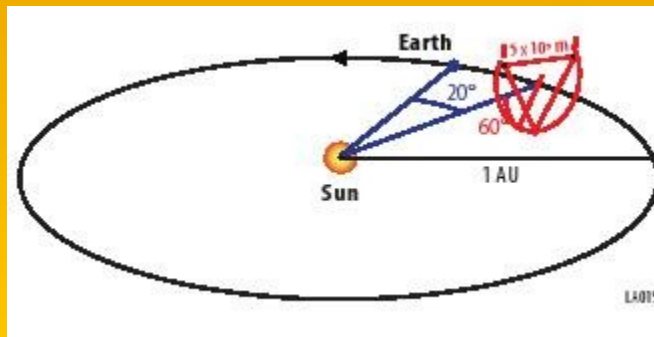
**Ft Lauderdale**

**May 2009**

# LISA Technology

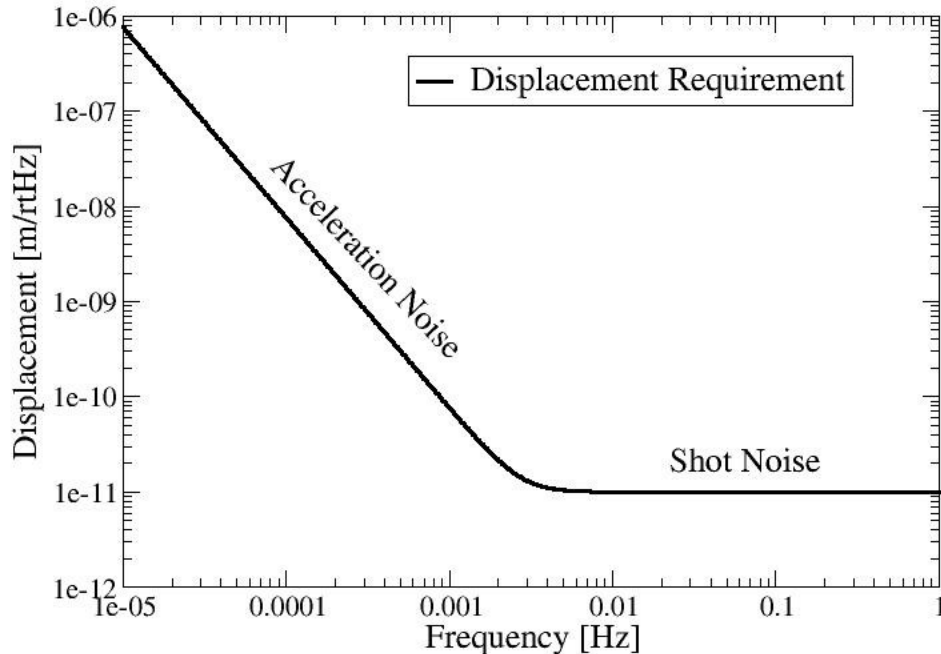


- 3 spacecraft constellation
- S/C separated by  $5 \times 10^6$  km
- Two Drag-free proof masses inside each S/C (Drag-free in sensitive direction)
- Earth-trailing solar orbit
- 5 year operational lifetime



# LISA

## Limiting Noise sources:



### Below 3mHz:

- acceleration noise  
 $3 \times 10^{-15} \text{m/s}^2/\text{rtHz}$ 
  - $1/f^2$  displacement noise  
of 12pm/rtHz @3mHz

### Above 3mHz:

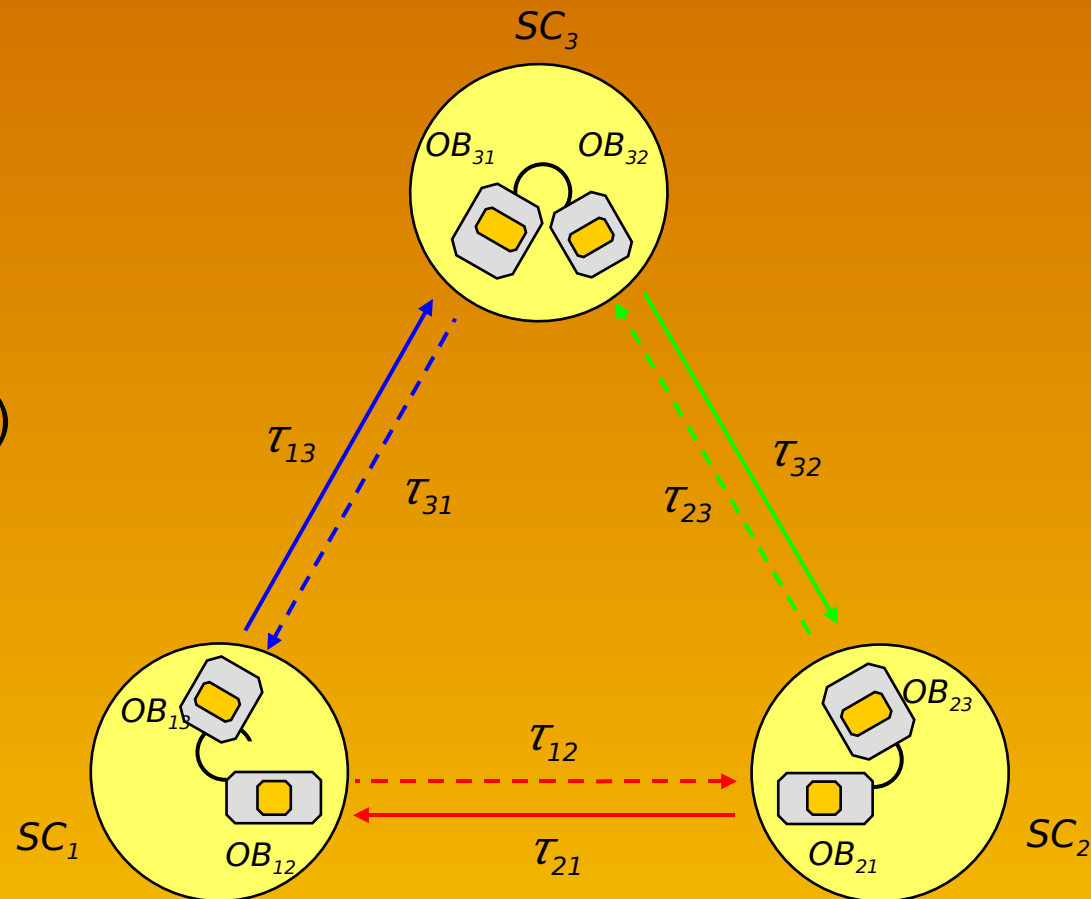
- Limited by shot noise  
based on 40pW  
received light
  - 40cm telescope

# LISA Interferometry

Goal: Measure changes in the distance between proof masses

Staged approach:

- Interferometry between PM and optical bench (OB)
- Interferometry between OBs on different S/C
- Measure differential laser phase noise on each S/C (create a beam splitter)



Note:

- $10\text{pm}/\text{rtHz} \sim 10^{-5}\text{cycl.}/\text{rtHz}$  for  $\lambda = 1\mu\text{m}$
- relative S/C motion  $\pm 10\text{m/s}$   $\Rightarrow$  Doppler shifts of  $\pm 10\text{MHz}$

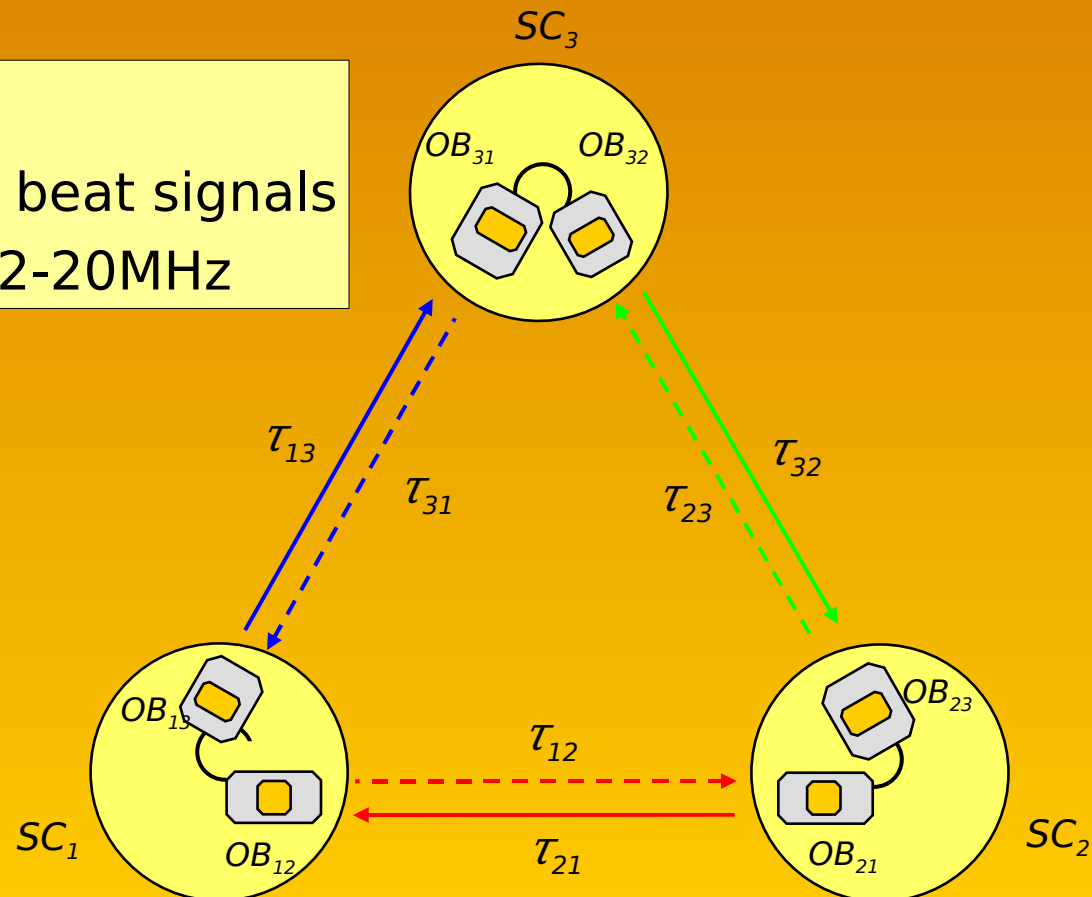
# LISA Interferometry

Note:

- $10\text{pm} \sim 10^{-5}\text{cycl. for } \lambda = 1\mu\text{m}$
- relative S/C motion  $\pm 10\text{m/s} \Rightarrow$  Doppler shifts of  $\pm 10\text{MHz}$

LISA Signals:

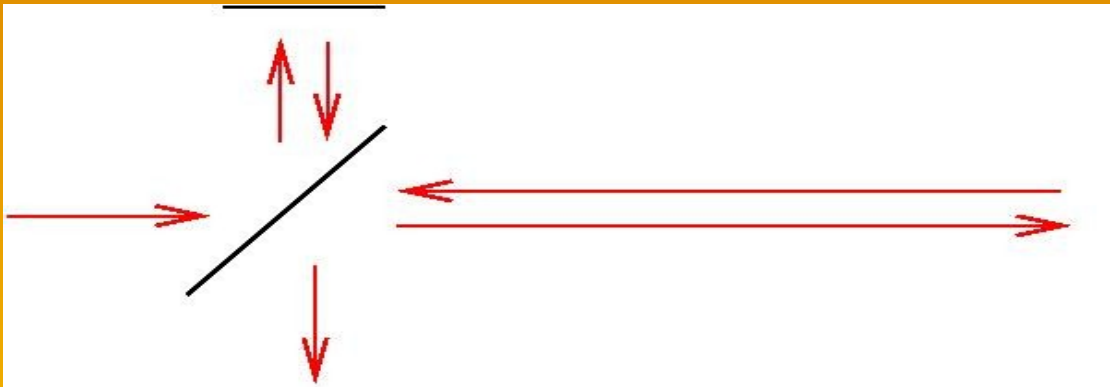
- Phase evolution in laser beat signals
- Beat frequencies betw. 2-20MHz



# LISA Interferometry

Interferometer measure a phase difference:

- Length changes  $k\delta L$
- Frequency changes  $2\pi\delta\nu\Delta L/c \Rightarrow \phi(\omega) (e^{-i\omega L_1/c} - e^{-i\omega L_2/c})$



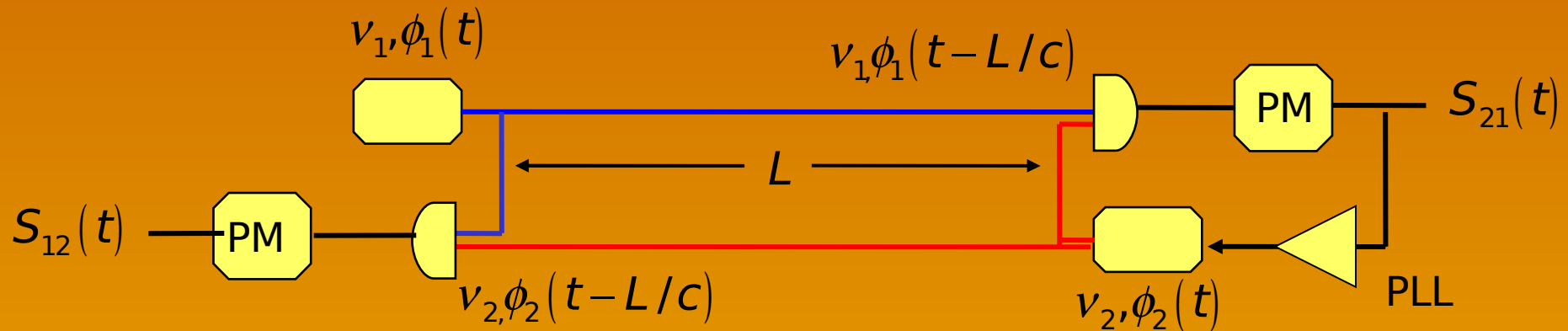
Time Domain:  
 $\phi(t-L_1/c) - \phi(t-L_2/c)$

LIGO:  $\Delta L \sim 0$  Intrinsic Common mode rejection of  $\delta\nu$

LISA:  $\Delta L \sim 50\,000$  km would require  $\delta\nu < 60\mu\text{Hz}/\text{rtHz}$

- Need to subtract phase noise in postprocessing

# Measuring a LISA arm



- Beat Signal on  $SC_2$ :  $A \sin[2\pi\nu_{12}t + \phi_1(t-L/c) - \phi_2(t)]$
- PM Signal on  $SC_2$ :  $S_{21}(t) = \phi_1(t-L/c) - \phi_2(t)$
- PLL condition:  $\phi_2(t) = \phi_1(t-L/c)$
- Beat Signal on  $SC_1$ :  $A \sin[2\pi\nu_{12}t + \phi_1(t) - \phi_2(t-L/c)]$
- PM Signal on  $SC_1$ :  $S_{12}(t) = \phi_1(t) - \phi_2(t-L/c)$   
 $= \phi_1(t) - \phi_1(t-2L/c)$



# LISA Interferometry

- Identical measurement in 2<sup>nd</sup> arm
- $SC_2$  and  $SC_3$  act as transponders

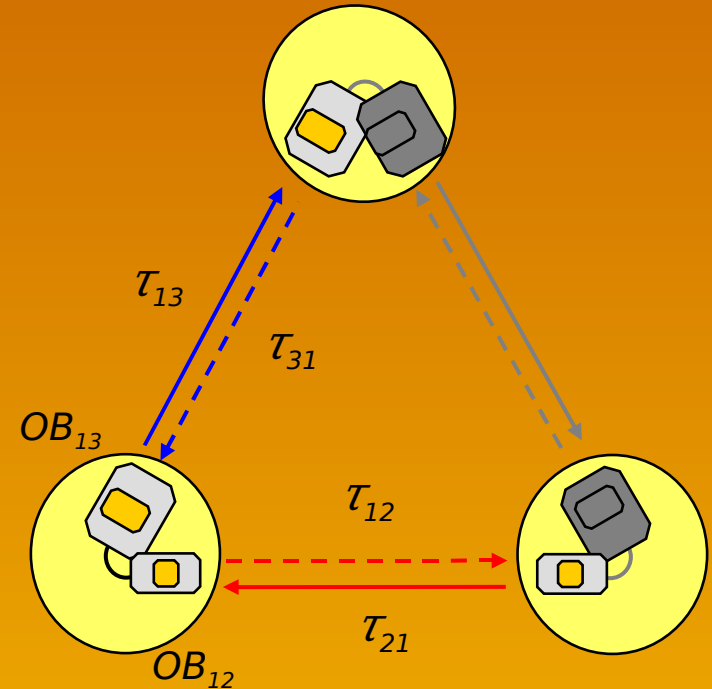
$$S_{12}(t) = \phi_1(t) - \phi_1(t - \tau_{12} - \tau_{21})$$

$$S_{13}(t) = \phi_1(t) - \phi_1(t - \tau_{13} - \tau_{31})$$

- Time-Delay Interferometry

$$X(t) = S_{12}(t) - S_{13}(t) - S_{12}(t - \tau_{13} - \tau_{31}) + S_{13}(t - \tau_{12} - \tau_{21})$$

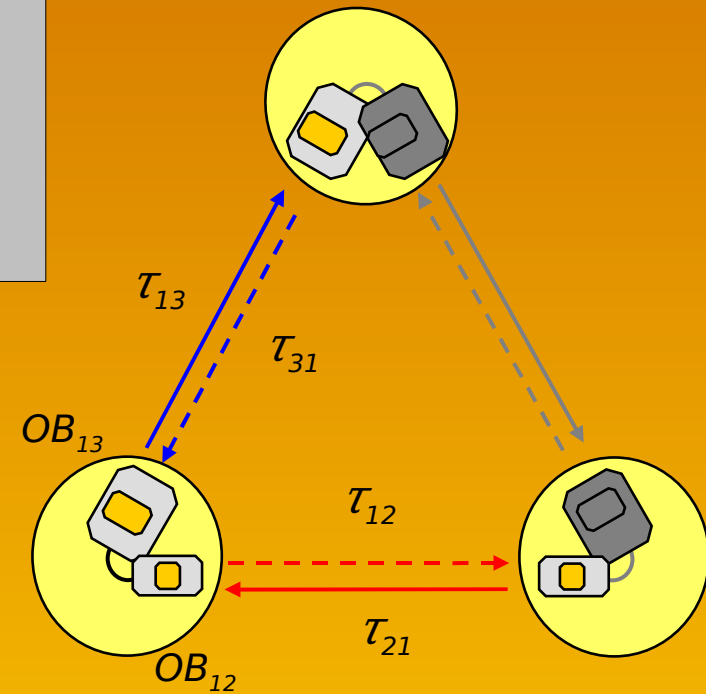
- All  $\phi(t)$ ,  $\phi(t - \tau)$ , ... - terms show up twice with opposite sign and cancel
- First-order insensitive to laser phase noise



# LISA Interferometry

What do we need to make this work?

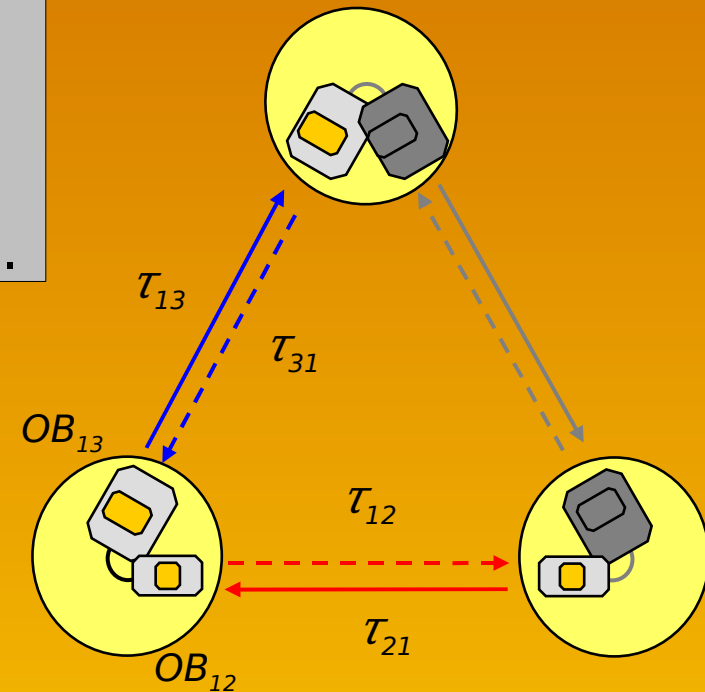
- Do we have to have low noise transponders (mirrors) at the far ends?



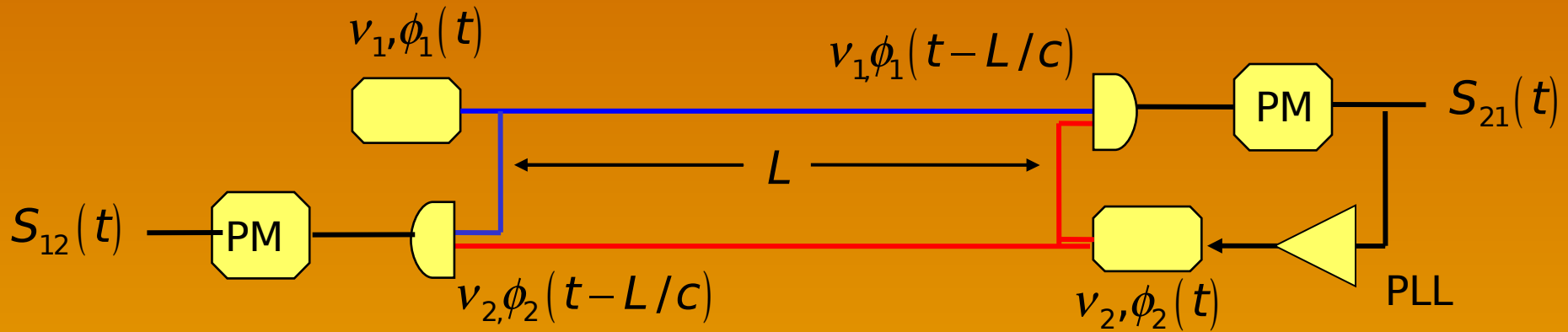
# LISA Interferometry

What do we need to make this work?

- Do we have to have low noise transponders (mirrors) at the far ends?
  - No, we only have to measure  $\phi(t)$  at the far ends and then subtract them.



# Measuring a LISA arm



- Beat Signal on SC<sub>2</sub>:  $A \sin[2\pi\nu_{12}t + \phi_1(t-L/c) - \phi_2(t)]$

- PM Signal on SC<sub>2</sub>:  $S_{21}(t) = \phi_1(t-L/c) - \phi_2(t)$

- PLL condition:  $\phi_2(t) = \phi_1(t-L/c) + \delta\phi(t)$

- Beat Signal on SC<sub>1</sub>:  $A \sin[2\pi\nu_{12}t + \phi_1(t) - \phi_2(t-L/c)]$

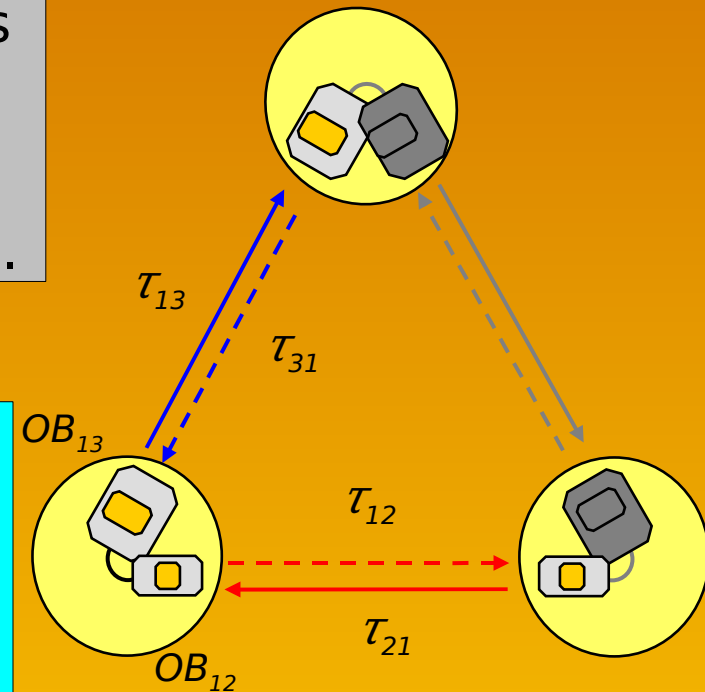
- PM Signal on SC<sub>1</sub>:  $S_{12}(t) = \phi_1(t) - \phi_2(t-L/c) - \delta\phi(t-L/c)$   
 $= \phi_1(t) - \phi_1(t-2L/c)$

# LISA Interferometry

What do we need to make this work?

- Do we have to have perfect transponders (mirrors) at the far ends?
  - No, we only have to measure  $\phi(t)$  at the far ends and then subtract them.

- Do we have to have the same phase noise enter in both arms? (Beamsplitter)

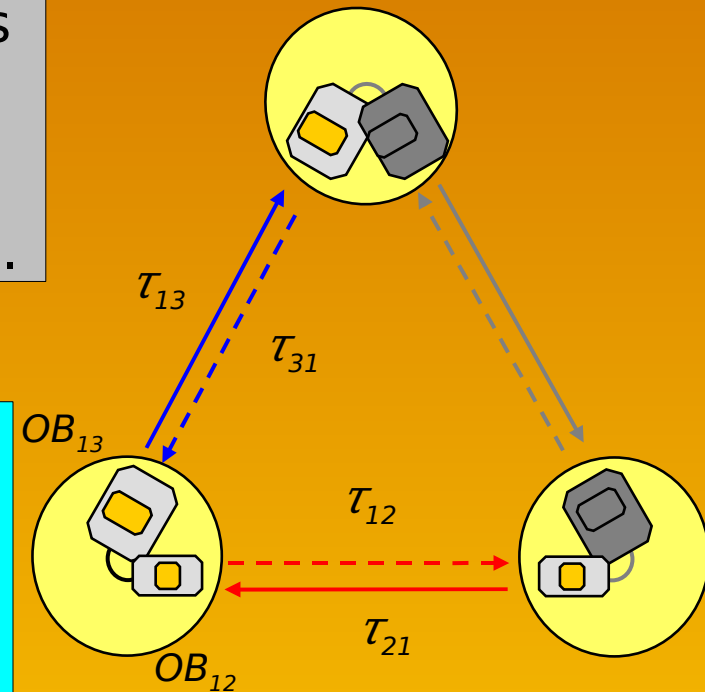


# LISA Interferometry

What do we need to make this work?

- Do we have to have perfect transponders (mirrors) at the far ends?
  - No, we only have to measure  $\phi(t)$  at the far ends and then subtract them.

- Do we have to have the same phase noise enter in both arms? (Beamsplitter)
  - No, we only have to measure the differential phase noise and subtract that as well.



# LISA Interferometry

- Repeat measurement in two arms
- $SC_2$  and  $SC_3$  act as transponders

$$S_{12}(t) = \phi_1(t) - \phi_1(t - \tau_{12} - \tau_{21})$$

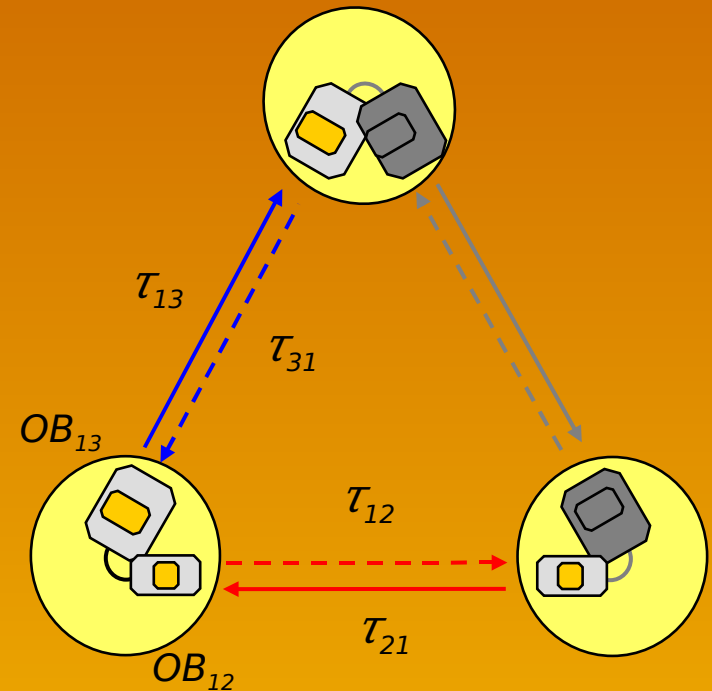
$$S_{13}(t) = \phi_1(t) - \phi_1(t - \tau_{13} - \tau_{31})$$

$$+\Delta\phi_{12}(t) - \Delta\phi_{12}(t - \tau_{13} - \tau_{31})$$

- Time-Delay Interferometry

$$X(t) = S_{12}(t) - S_{13}(t) - S_{12}(t - \tau_{13} - \tau_{31}) + S_{13}(t - \tau_{12} - \tau_{21})$$

$$-\Delta\phi_{12}(t) + \Delta\phi_{12}(t - \tau_{13} - \tau_{31})$$



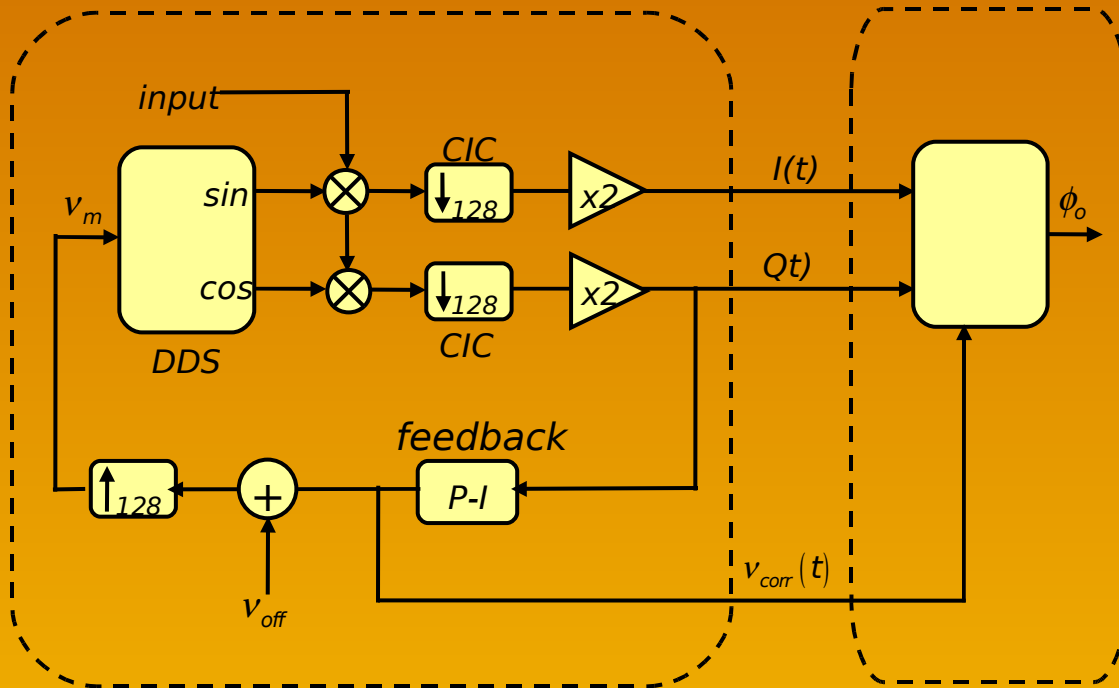
# LISA Interferometry

What do we need to make this work?

- We need to be able to measure  $\phi(t)$  in each signal with the necessary accuracy
  - Phasemeter



# Phasemeter



- Phasemeter:
- Digital
  - PLL with NCO

Numerical Controlled Oscillator    Phase Reconstruction

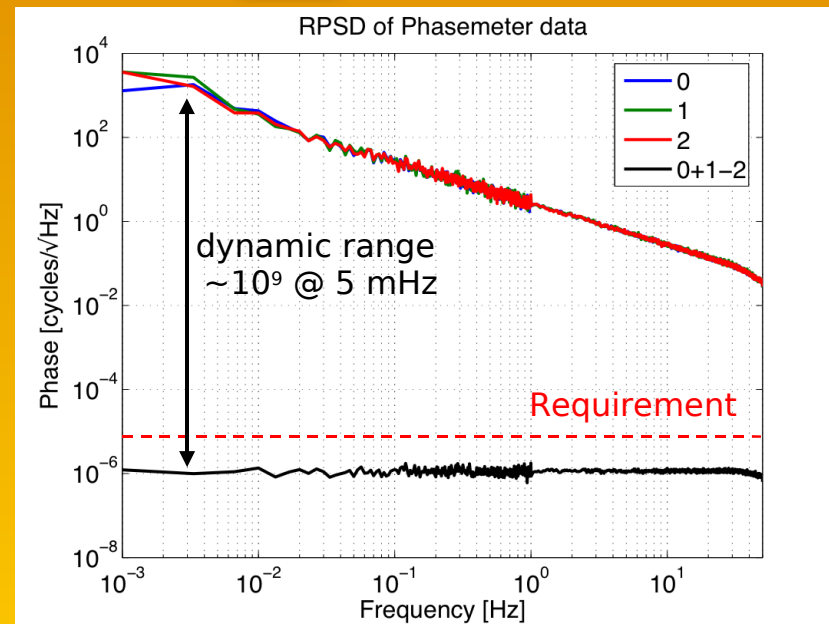
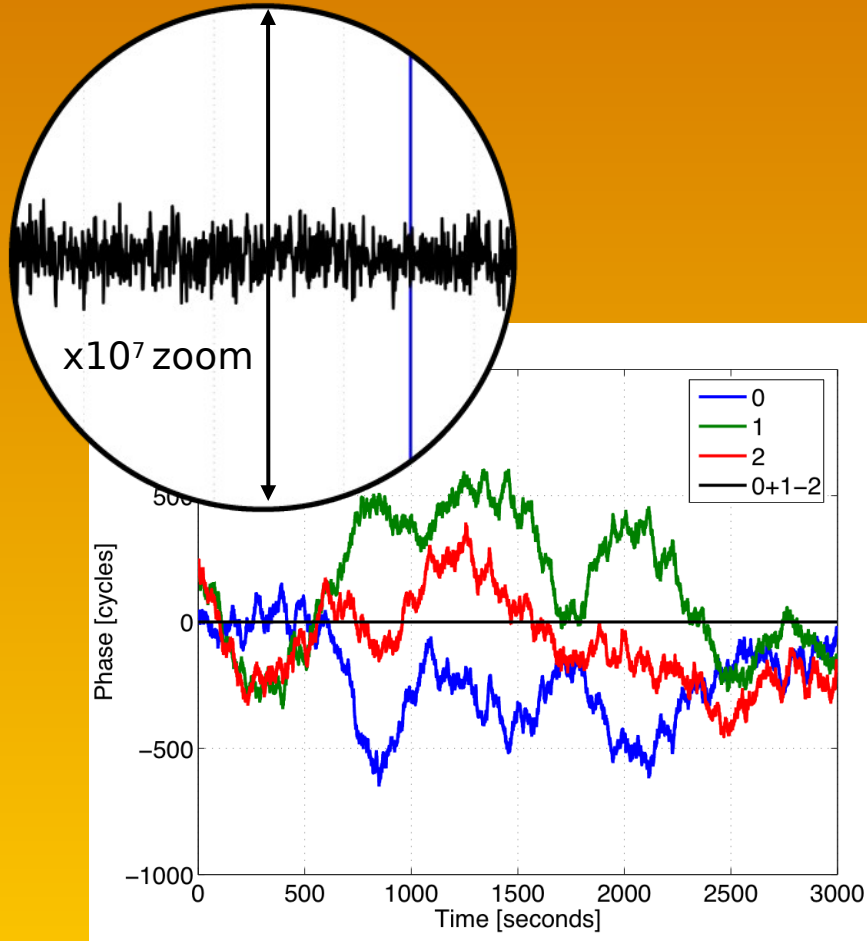
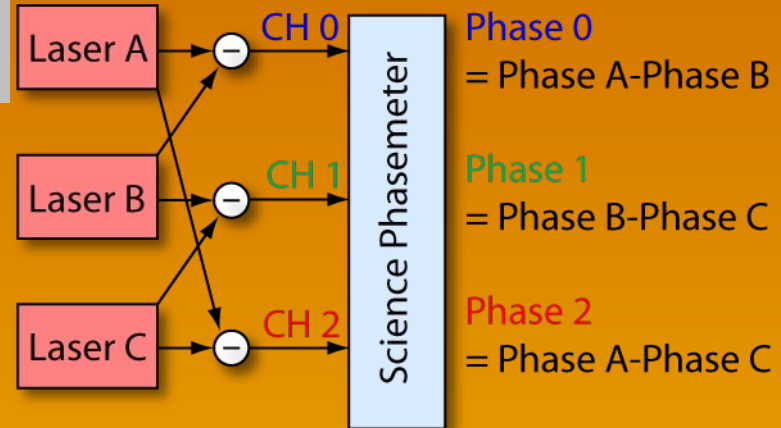
Measures the phase against the on board clock:

- Clock noise:  $\delta\tau < 5 \times 10^{-14} \text{ s}/\text{rtHz}$  for 20 MHz beat
- Requires clock stabilization between S/C

# LISA Interferometry

## Status of Phasemeter:

- It works (Digital part)
- with respect to local clock

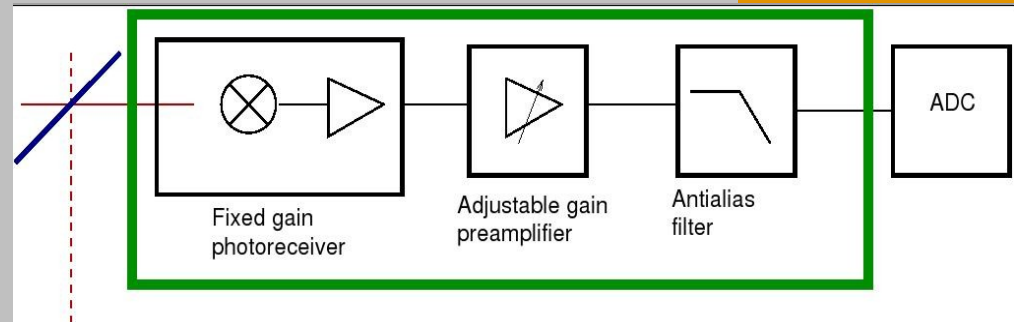


# LISA Interferometry

Status of Phasemeter:

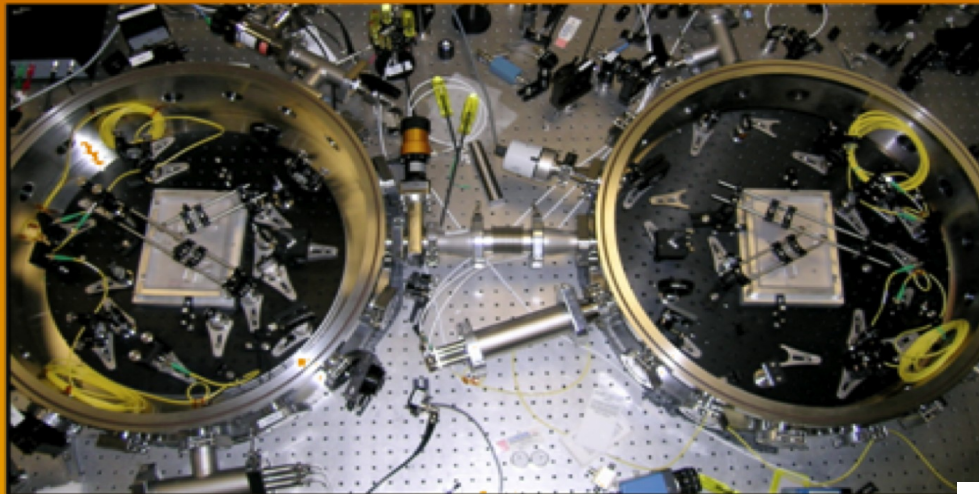
Analog Parts

- Photodiodes and Amplifiers
- Cables
- ADC timing
- ...



Ongoing R&D at JPL, AEI, GSFC, UF, ...

# LISA Interferometry

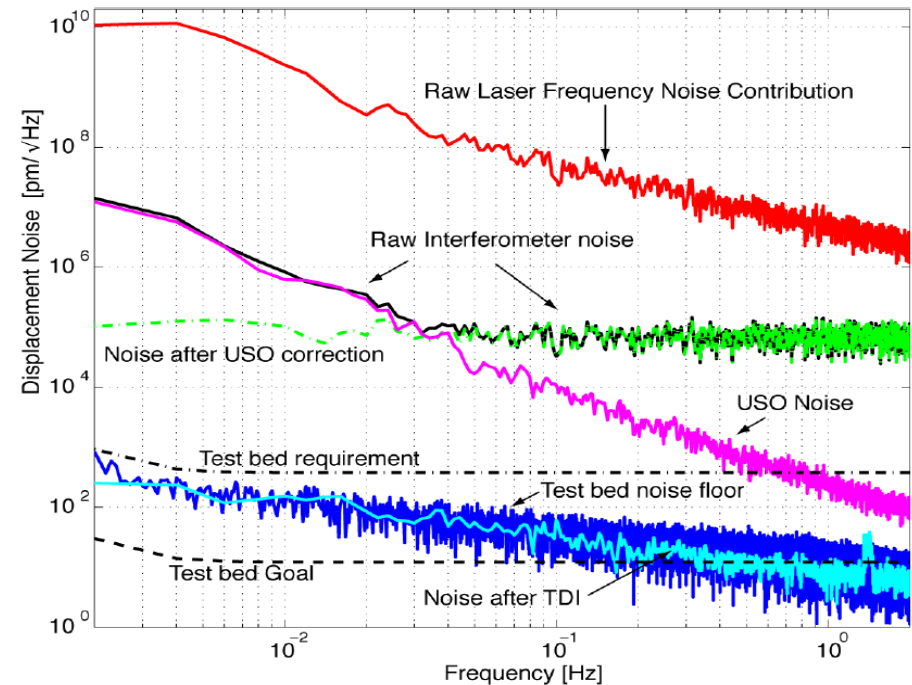


## JPL-Testbed

- Sagnac Configuration
- Laser Transponders
- Multiple Photo detectors
- ...

## Critical areas:

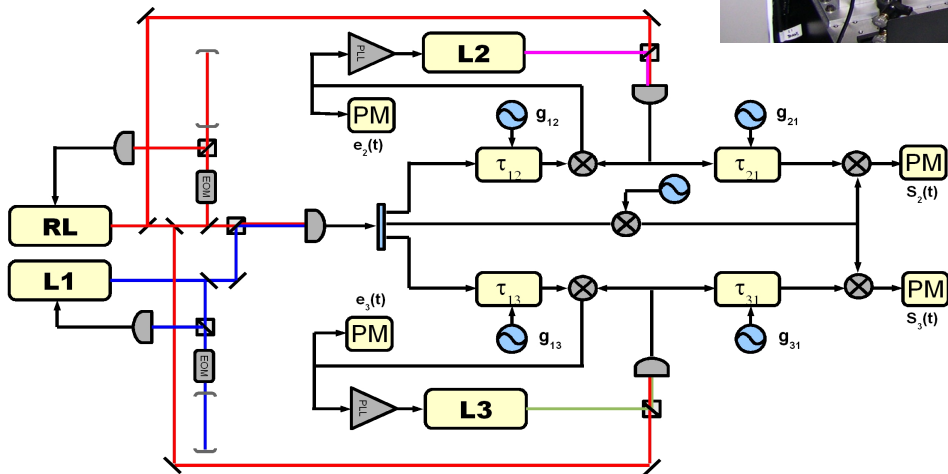
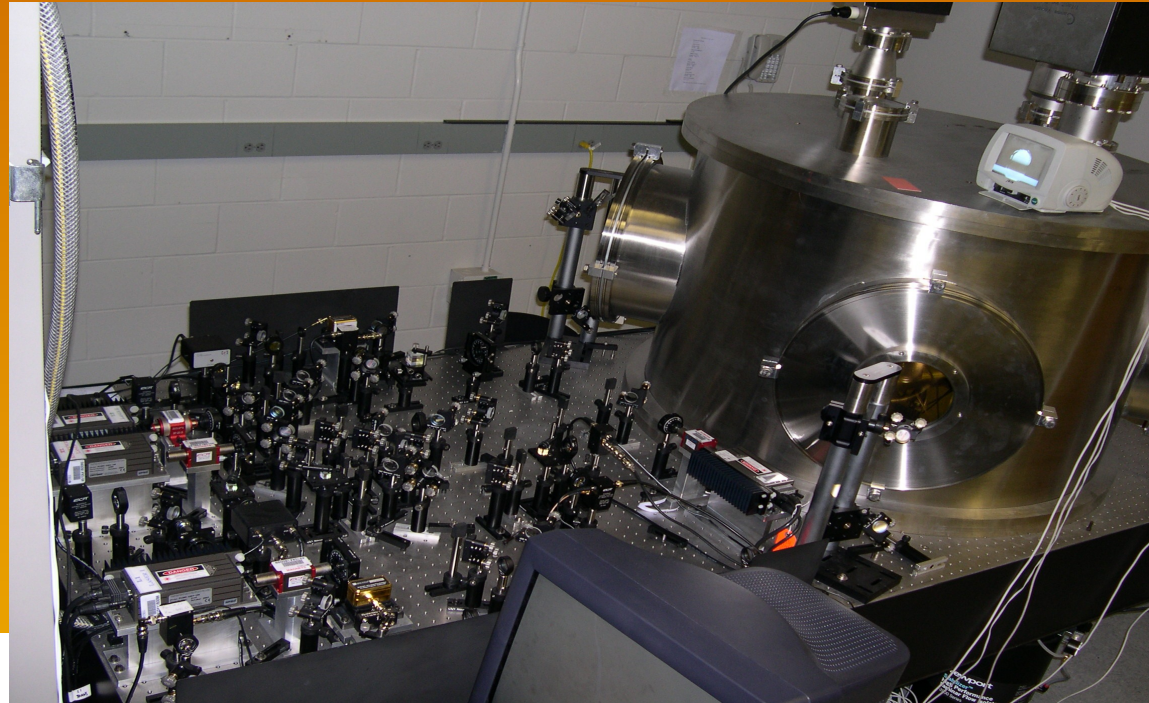
- Dispersion in PDs
- Timing noise in ADCs
  - Pilot tone (JPL)
  - Adding the pilot tone
- ...



# LISA Interferometry

UF-Testbed with:

- Laser Transponders
- Signal travel times
- Doppler shifts
- UF-Phasemeter



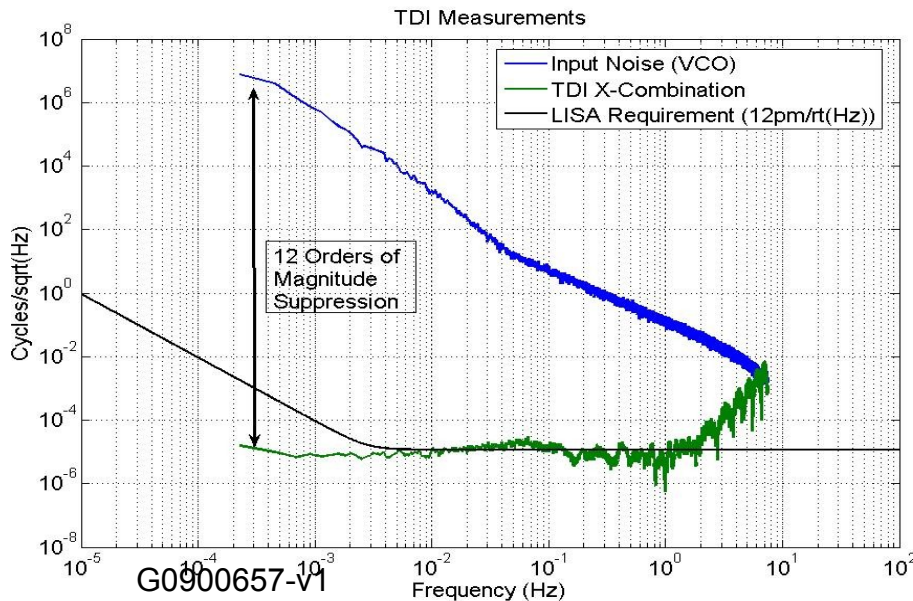
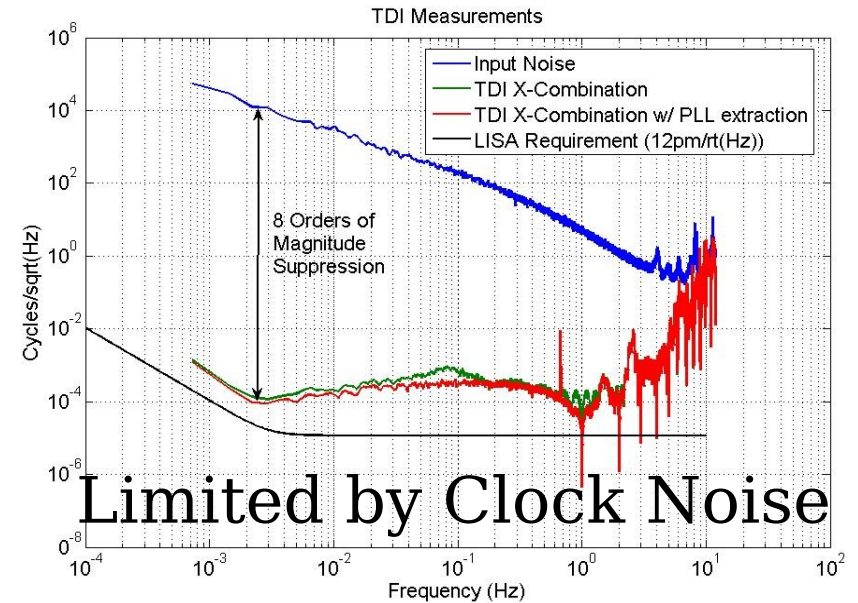
G0900657-v1



# LISA Interferometry

UF-Test with:

- Laser Transponders
- Signal travel times
- Doppler shifts
- UF-Phasemeter



Component test  
w/o clock noise  
but larger input noise

# LISA Interferometry

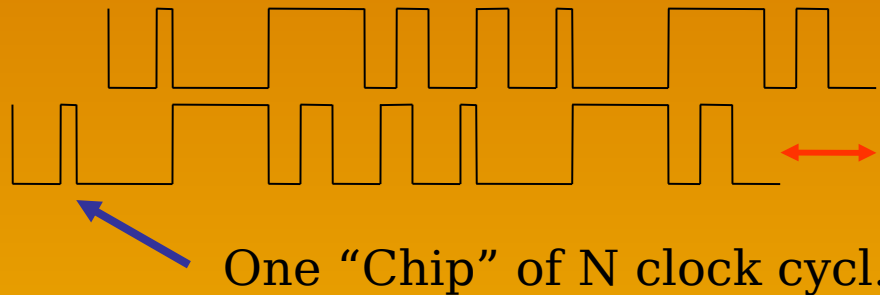
What do we need to make this work?

- We need to be able to measure  $\phi(t)$  in each signal with the necessary accuracy
  - Phasemeter
- The time stamps on each  $\phi(t)$  have to be accurate.
  - Ranging:  $\Delta L/c = \Delta\tau$
  - Laser frequency stabilization  $\delta\nu$  (to reduce the required ranging accuracy)
  - Requirement:  $\delta\nu \Delta L/c = \delta\nu \Delta\tau < 10^{-6} \text{cycl./rtHz}$

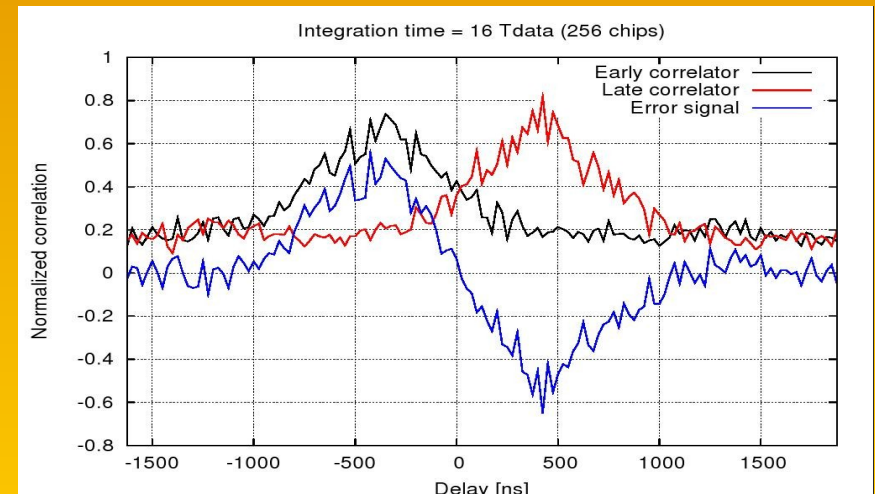
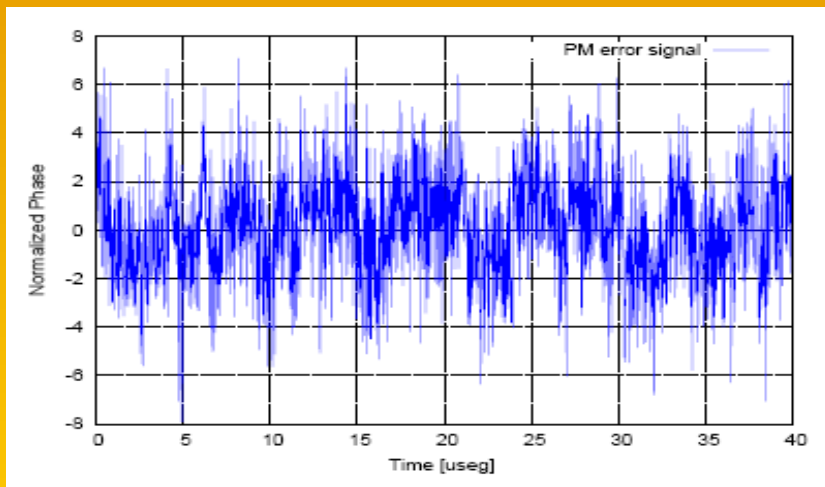
# LISA Ranging

$\Delta\tau$ : Uncertainty in Signal Travel time  $\sim$  Ranging

- Modulate Laser field with PRN code
- Current Best Estimate:  $\Delta L < 1\text{m}$



Accuracy  $<$   
Clock cycle using interpolation





# LISA Interferometry

LIGO:  $\Delta L \sim 50\,000\text{ km}$  would require  $\delta\nu < 60\mu\text{Hz}/\text{rtHz}$

- Phasemeter
- Timing System for Constellation

Requirement from  $2\pi\delta\nu\Delta L/c = 2\pi\delta\nu\Delta\tau$

$\Delta L = 1\text{ m} \longrightarrow \delta\nu = 300\text{Hz}/\text{rtHz}$  above  $3\text{mHz}$

- Goal:  $\delta\nu = 30\text{Hz}/\text{rtHz} (1 + (2.8\text{mHz}/f)^4)^{1/2}$

Multiple Options:

- Local stabilization on fixed cavity or iodine line
- Stabilization on LISA arms ( $\sim$ LIGO-like)
- Or a combination of both

# Cavity Stabilization

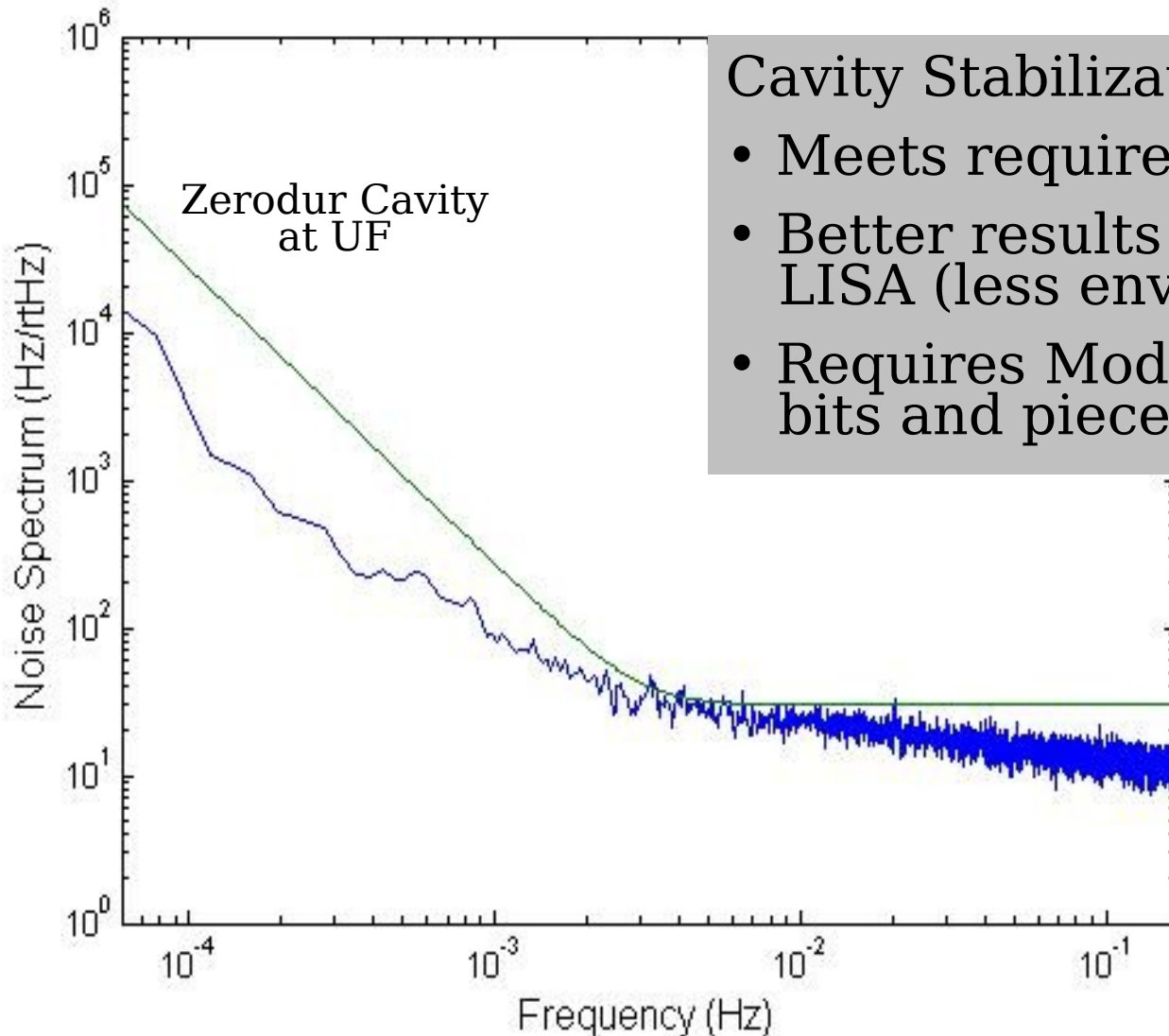
Fixed:

- Pound Drever Hall technique
- Heterodyne Interferometer (LTP-style IFO)

Tunable (to combine with Arm locking):

- PZT-cavity
- Offset phase lock
- Sideband locking

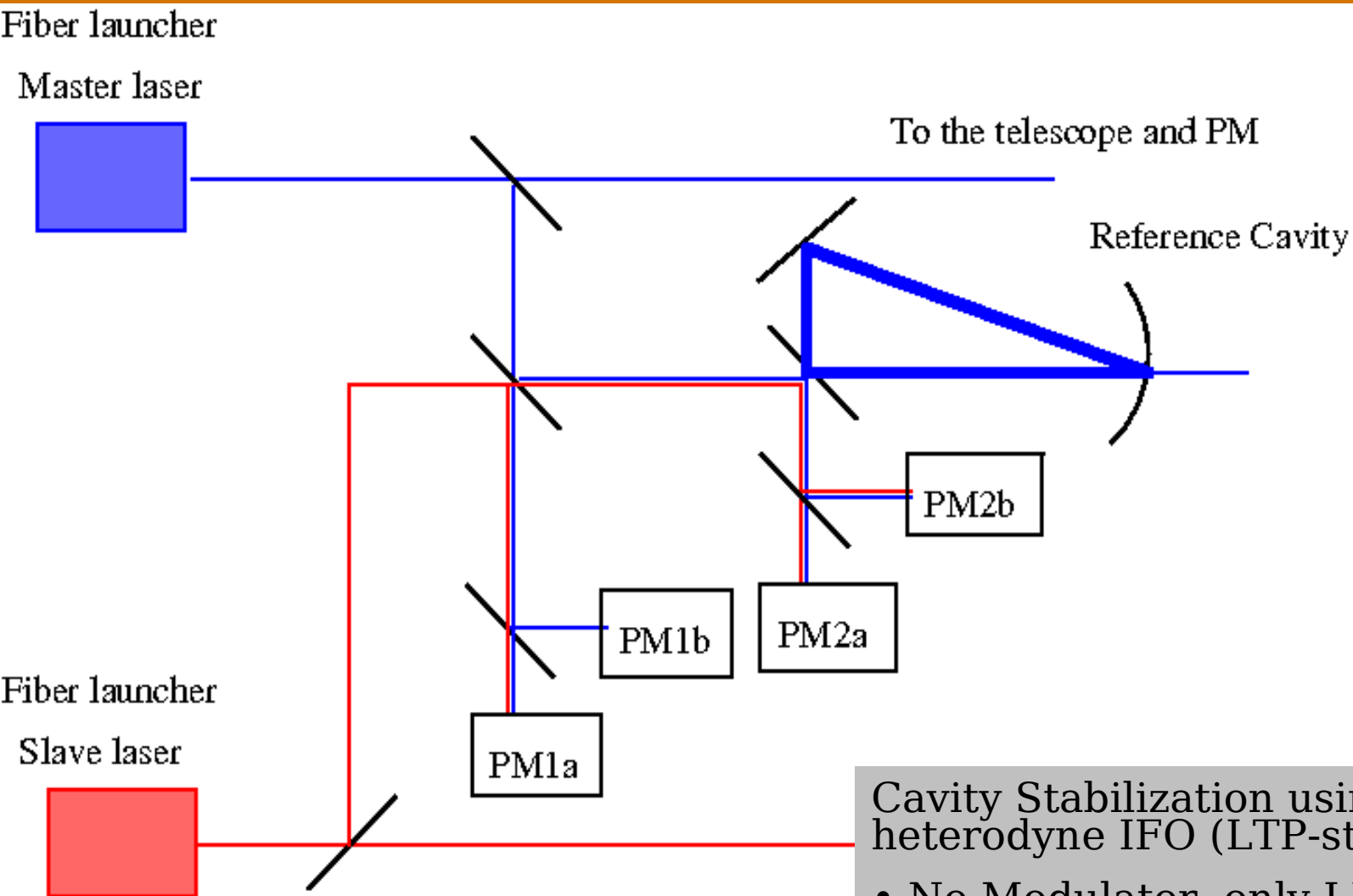
# Cavity Stabilization



## Cavity Stabilization with PDH

- Meets requirements
- Better results expected in LISA (less environmental noise)
- Requires Modulator and other bits and pieces.

# Cavity Stabilization



Cavity Stabilization using heterodyne IFO (LTP-style)

- No Modulator, only LISA PM
- Should easily meet requirements with stable cavity.

# Cavity Stabilization

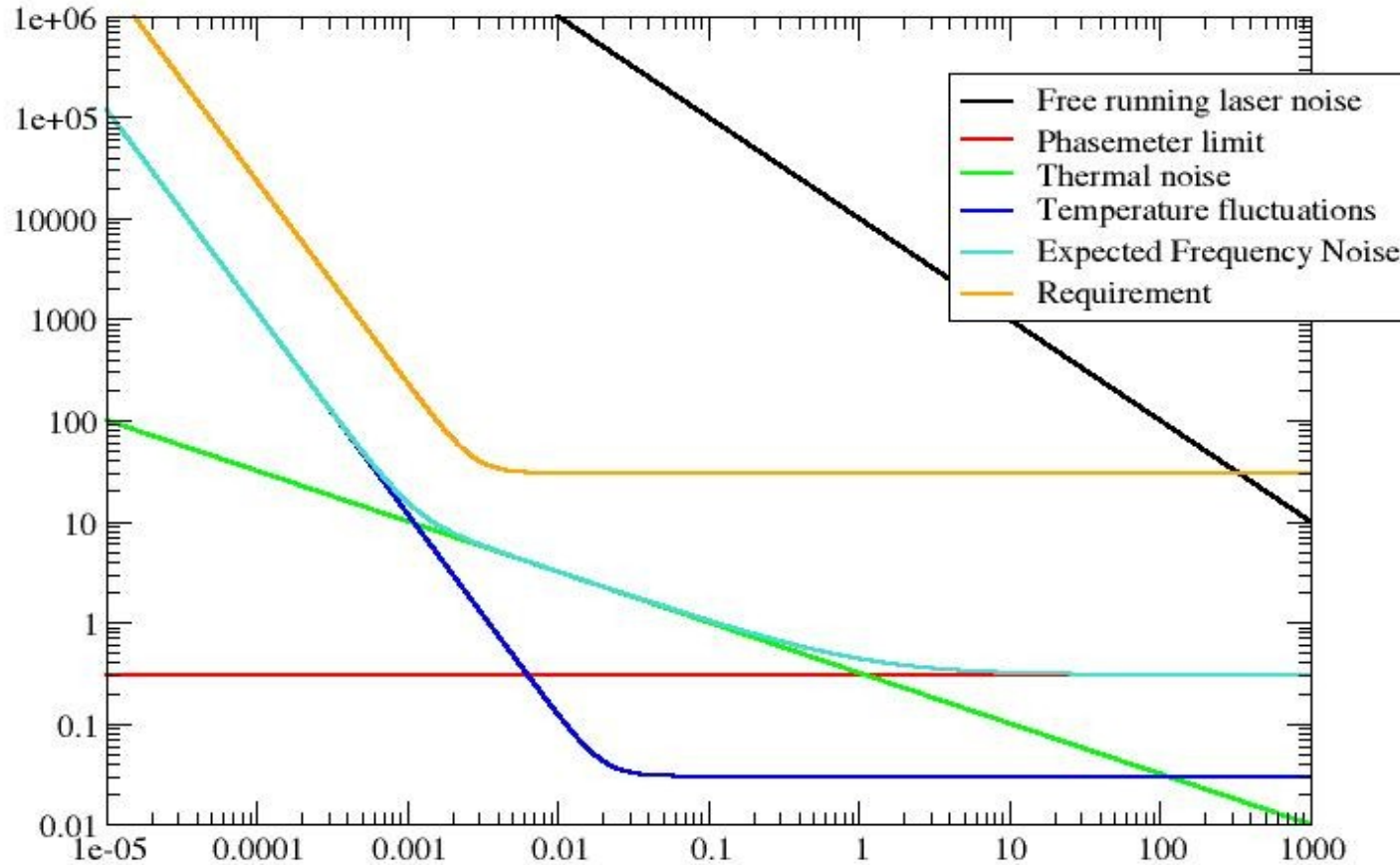
Fiber laser

Master



Fiber laser

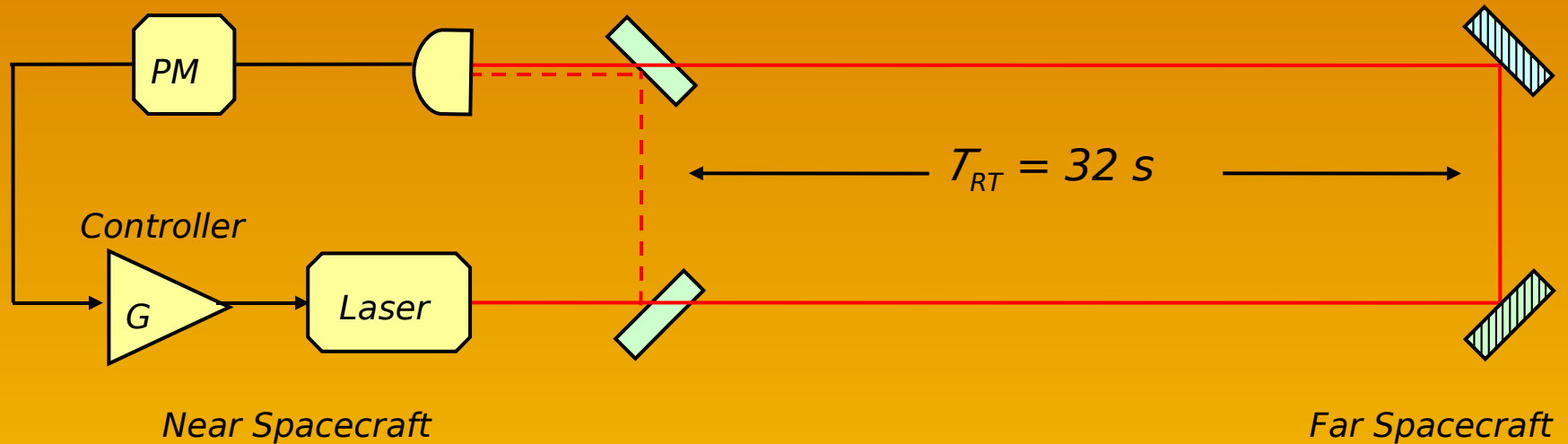
Slave laser



with stable cavity.

# Arm-Locking I

- Lasers might be pre-locked to mechanical (cavity or Mach-Zehnder) or spectroscopic reference (iodine)
- Use the arms as a stable frequency reference



- Single-arm locking error signal:  $\phi(t) - \phi(t - \tau_{RT})$
- Insensitive to noise at frequencies:  $f_n = n/\tau_{RT}$

# Arm-Locking II

- General closed-loop stability

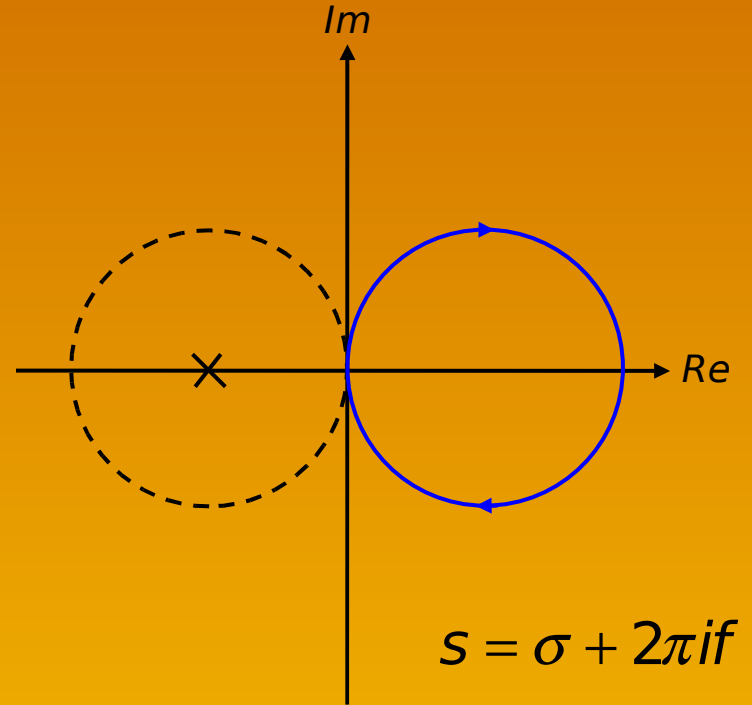
$$T_{CL}(s) = \frac{1}{1 + T_{OL}(s)}$$

- Laplace-domain arm-locking error signal

$$1 - \exp(-s\tau_{RT})$$

– nulls at  $n/\tau_{RT}$

– phase change from  $-90^\circ$  to  $+90^\circ$  at  $n/\tau_{RT}$

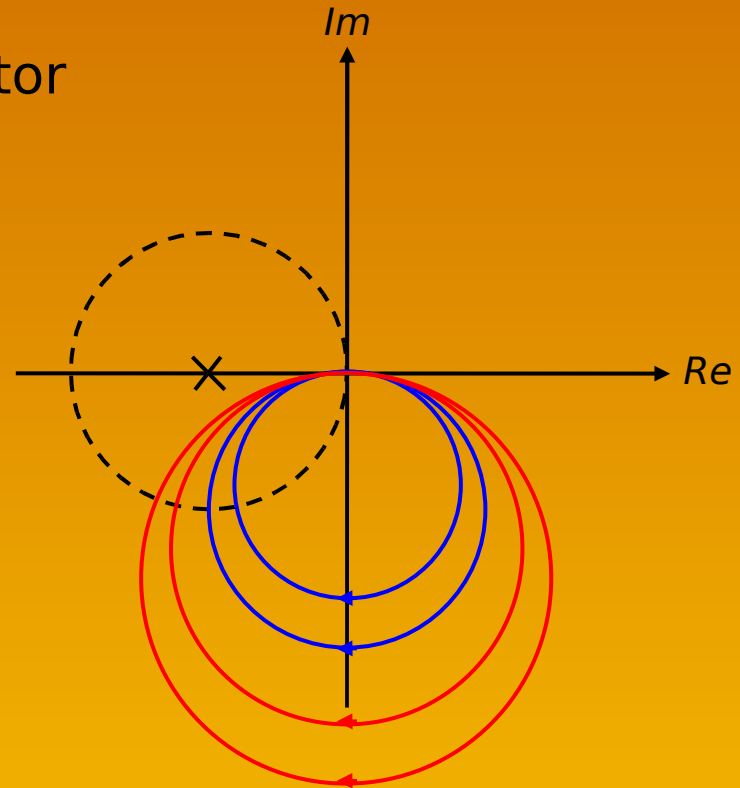


# Arm-Locking II

- Frequency actuator acts as integrator

$$\frac{1}{s} [1 - \exp(-s\tau_{RT})]$$

- System enters noise enhancement region
- Increased gain brings increased noise enhancement
- System is marginally stable, any phase loss (latency) could cause instability



- Note: Frequency actuator as integrator can be avoided when we use frequency meter instead of phasemeter (phasemeter contains one integrator as last stage)

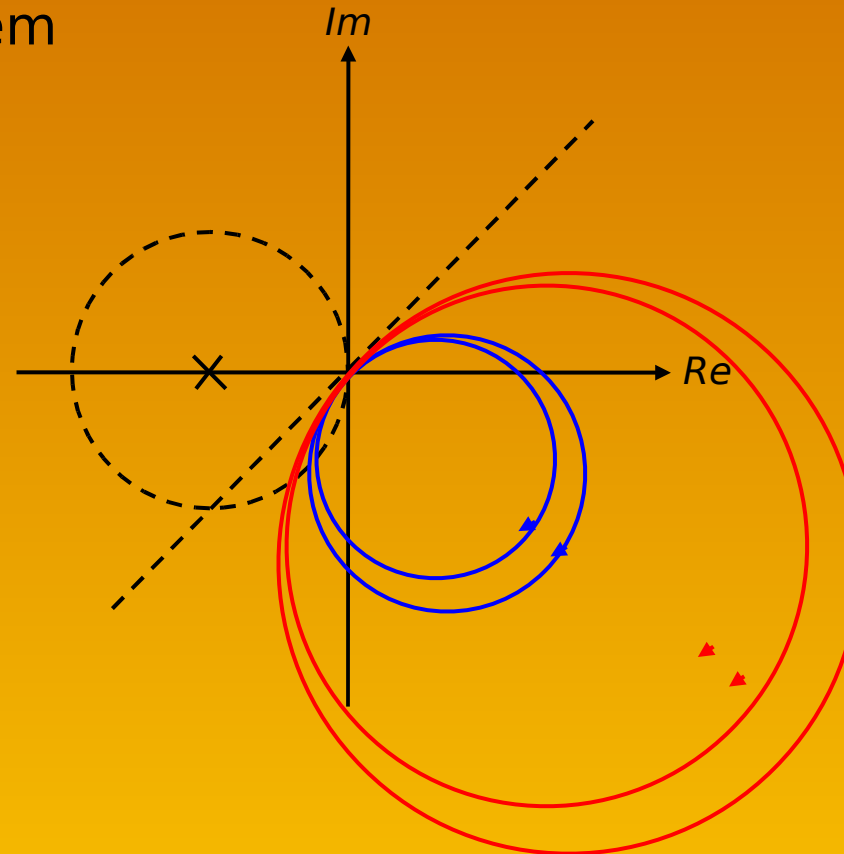


# Arm-Locking II

- Controller with  $s^p$  slope rotates system away from  $(-1,0)$

$$s^{p-1} [1 - \exp(-s\tau_{RT})]$$

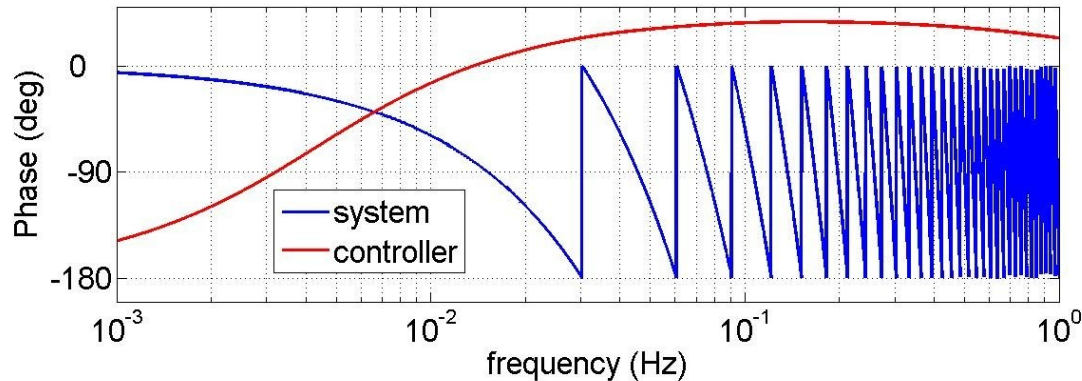
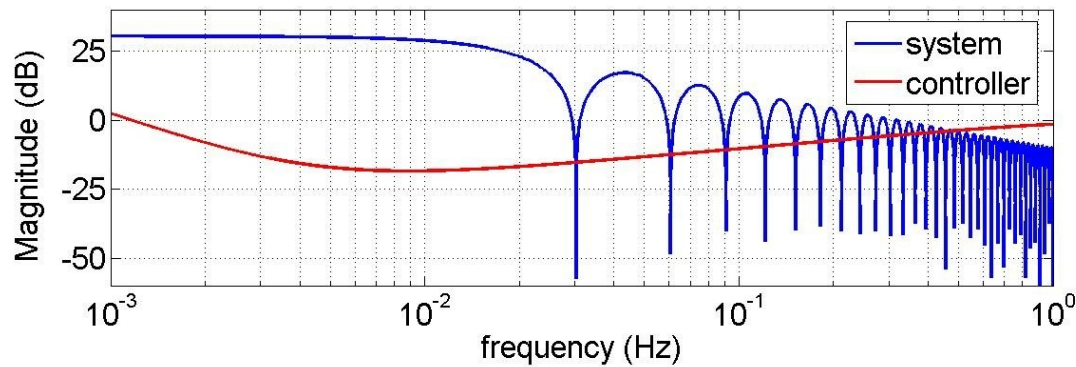
- Gain can be increased below  $1/\tau_{RT}$
- Noise enhancement at  $n/\tau_{RT}$  will remain finite



# Arm-Locking II

- Controller away from (-
- Gain can b
- Noise enha  
remain finite

Controller & Sensor Bode plot (incl. PM-integrator):



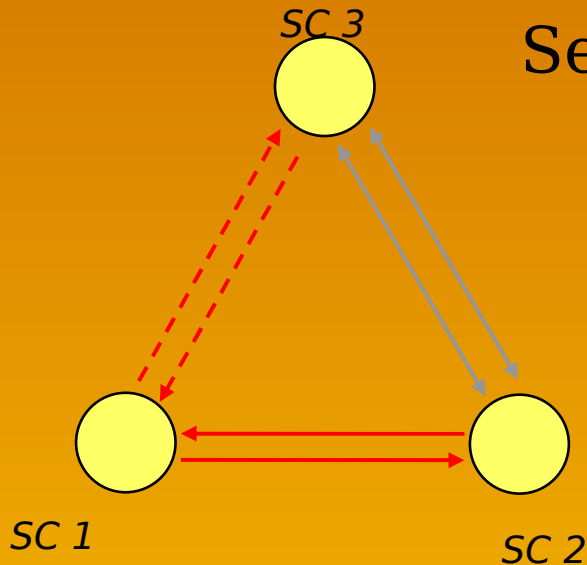
$Im$

$Re$

Very limited gain in the LISA band.

# Dual Arm-Locking Schemes

- Use other arms to get phase noise information at nulls



dual arm-locking

$$\text{Sensor: } S = T(f)_{\text{com}} S_{\text{com}} + T(f)_{\text{Diff}} S_{\text{Diff}}$$

$$S_{\text{com}} = S_1 + S_2$$

$$\Rightarrow \phi(\omega) (2 - e^{-i\omega L_1/c} - e^{-i\omega L_2/c})$$

$$S_{\text{Diff}} = S_1 - S_2$$

$$\Rightarrow \phi(\omega) (e^{-i\omega L_1/c} - e^{-i\omega L_2/c})$$

- first null between 1-3 Hz
- No peaks in LISA band
- Increased noise suppression

$S_{\text{diff}}$  has 1<sup>st</sup> zero at  $1/\Delta\tau = 1-3\text{Hz}$

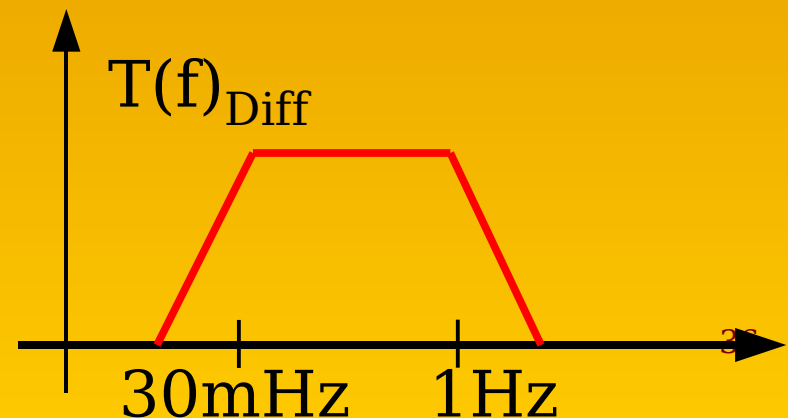
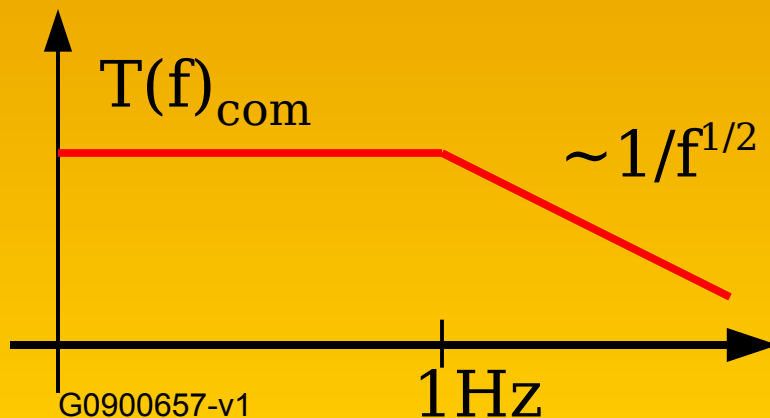
- allows to increase gain earlier
- Bad SNR at low frequencies  
(Problems with Clock noise,  
Shot Noise and S/C motion)

# Dual Arm-Locking Schemes

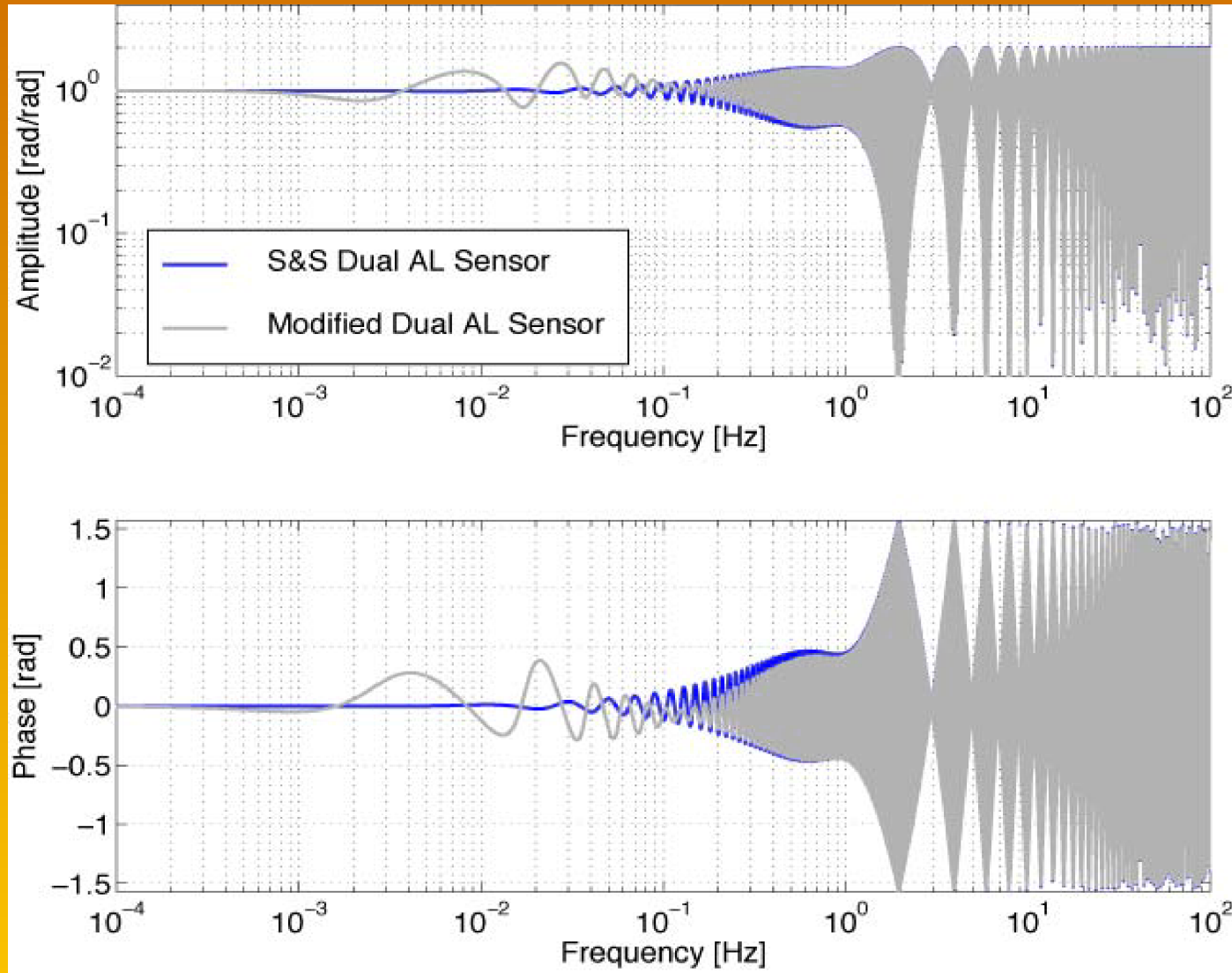
$$\text{Sensor: } S = T(f)_{\text{com}} S_{\text{com}} + T(f)_{\text{Diff}} S_{\text{Diff}}$$

Modified Dual Arm-locking sensor:

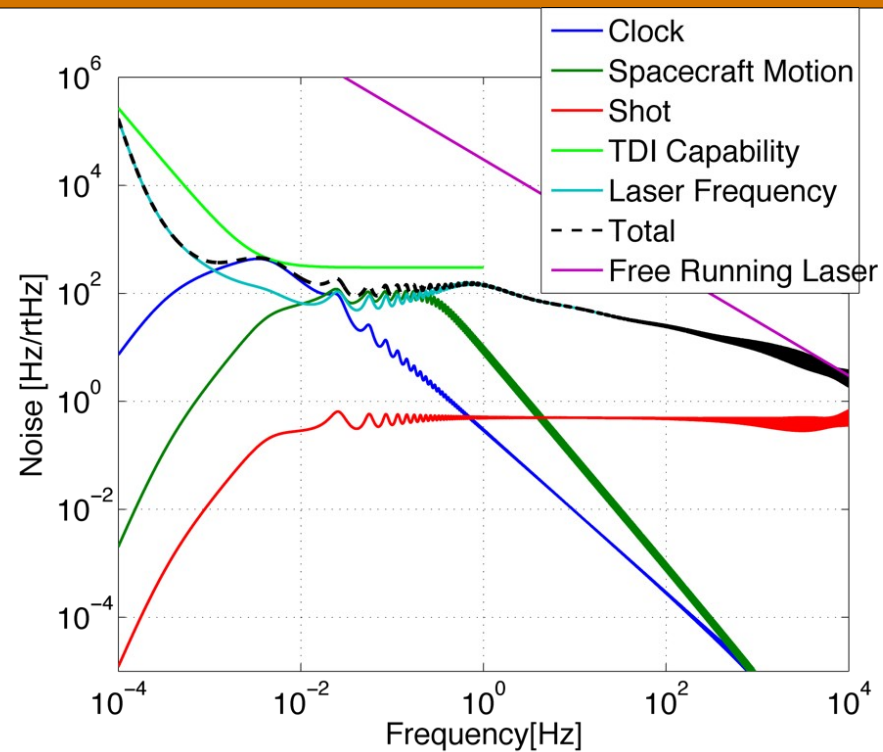
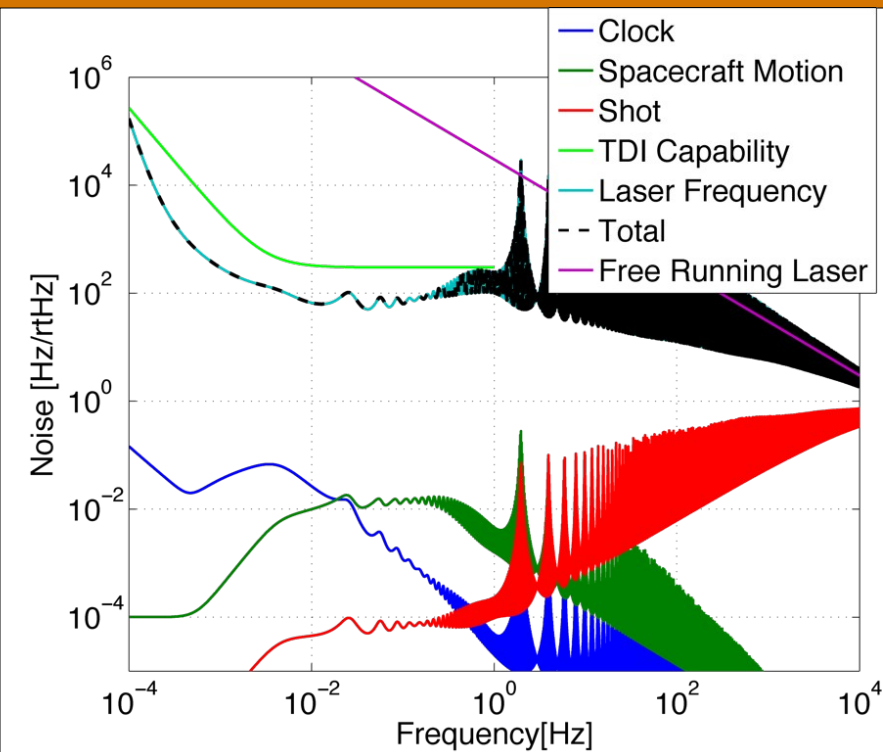
- Common arm sensor dominates below 30mHz
  - Increase gain with multiple integrators
- Differential arm sensor modifies sensor between 30mHz and 1Hz
  - Adds signal at the sensor nulls
- Common arm signal dominates above 1Hz
  - Roll down with  $1/\sqrt{f}$
  - Differential sensor signal is unstable above 1Hz



# Dual Arm-Locking Schemes



# Dual Arm-Locking Performance w/o prestabilization



Kirk McKenzie

Maximum arm length mismatch:  
 $\Delta\tau = 0.255\text{s}$ ,  $\Delta L = 75,000\text{km}$ .

Minimum arm length mismatch that  
 meets TDI capability:  $\Delta\tau = 40\mu\text{s}$ ,  $\Delta L = 12\text{km}$ .

**Gain limited, noise sources  
 negligible.**

**Noise limited for  $\Delta L < 12\text{km}$ .**

(Still meets capability)

(only see this if an interspacecraft  
 link fails)

# Dual Arm-Locking Schemes

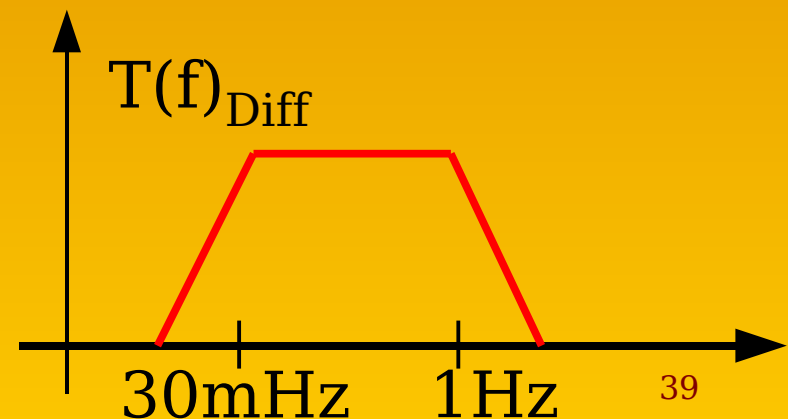
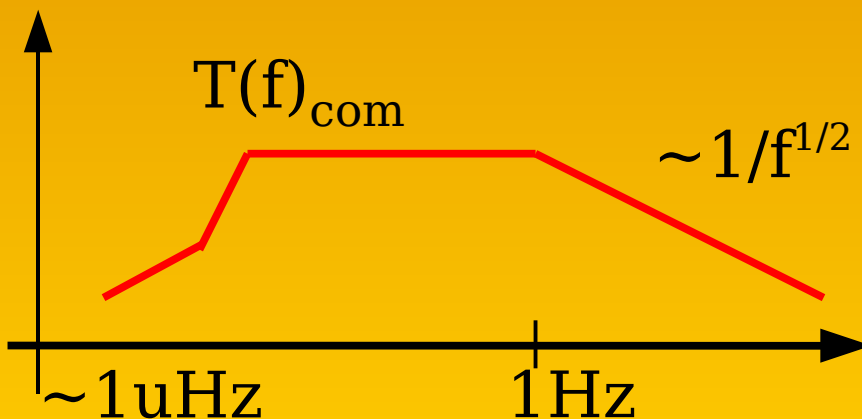
## Problem:

- Doppler shift introduces offset
  - AC-coupled sensor (corner freq.:  $\sim \mu\text{Hz}$ )
  - or estimate Doppler shift well enough
- Offset will be integrated up
  - ➔ linear ramp in laser frequency

## Time Domain:

$$\phi(t) - \phi(t - L_2/c) + \Delta\omega t$$

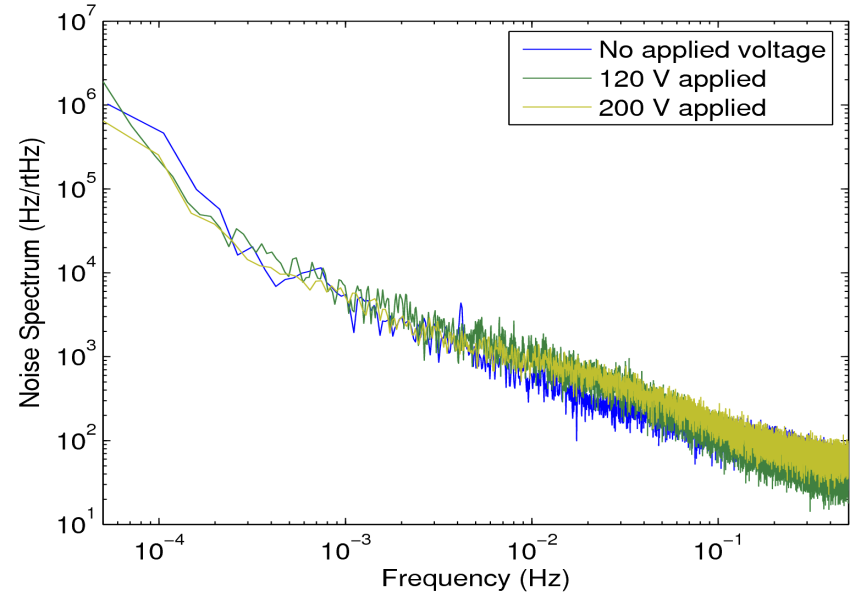
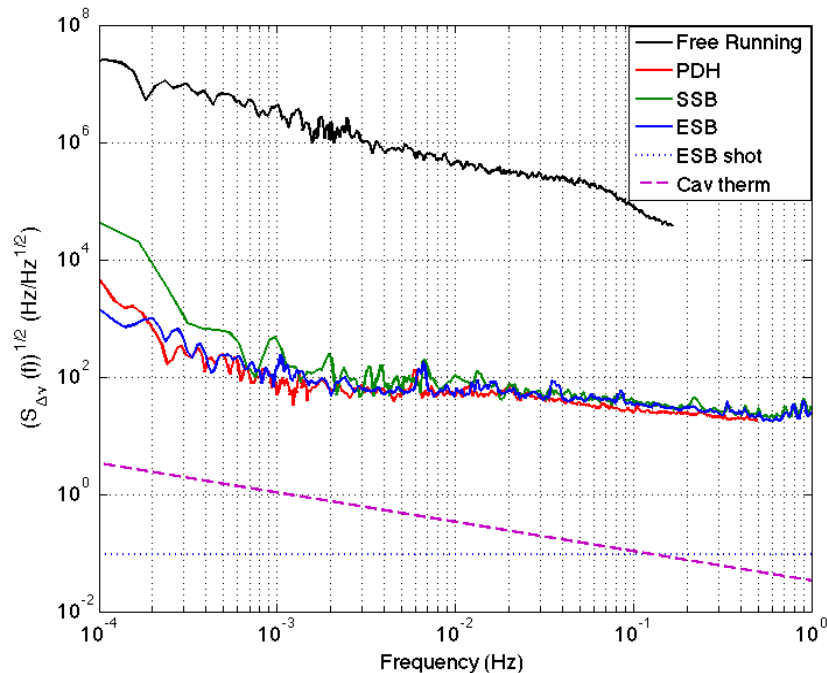
$$\Delta\omega = \omega_{\text{Doppler}} - \omega_{\text{PM}}$$



# Dual Arm-Locking Schemes

with prestabilization:

- PZT Cavity
- Sideband locking
- heterodyne IFO on MZ



PZT-Cavity:  
Alix Preston (UF)

Sideband locking:  
GSFC results



# Summary

## Long baseline IFO:

- Laser frequency noise canceled in post-processing (TDI)
- Laser ranging and clock synchronization with PRN codes with  $<1\text{m}$  accuracy
- Clock noise measurements using GHz modulation tones
- Multiple options for Laser frequency stabilization:
  - Cavity
    - with heterodyne IFO (simplest)
    - PDH (most heritage)
  - Arm Locking (most elegant)
    - modified dual arm locking (MDAL)
  - Arm Locking with pre-stabilization
    - MDAL with Sideband locking
    - MDAL with PZT
    - MDAL with LTP-style Mach Zehnder