

Einstein@Home

search for periodic gravitational waves
in early S5 LIGO data

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The first Einstein@Home search in S5 LIGO data



- **All-sky broad-band search for periodic gravitational waves, such as from rotating non-axisymmetric neutron stars:**
 - » Searched frequency range: $50 \text{ Hz} < f < 1500 \text{ Hz}$
 - » Searched range of frequency derivatives: $-f/\tau < \dot{f} < 0.1f/\tau$, with $\tau = 1000 \text{ yrs}$ for $f < 400\text{Hz}$ and $\tau = 8000 \text{ yrs}$ for $f > 400\text{Hz}$.
 - » Goal: high-confidence detection (not upper limits).
- **Input data: 840 hours selected from 66 days of early S5 LIGO run:**
 - » H1 4-km detector: 22 x 30-hour segments,
 - » L1 4-km detector: 6 x 30-hour, segments. } 28 30-hour segments
 - » Removed understood instrumental lines from the data.
- **Huge parameter-space search:** Distributed over more than 100 000 computers volunteered by the general public via BOINC (Berkeley Open Infrastructure for Network Computing).
- **Details:** <http://arxiv.org/abs/0905.1705>

Coherent matched-filtering over each 30-hour segment (\mathcal{F} -statistic)

» Data is convolved with a set of signal **template** waveforms corresponding to all possible sources.

- **Templates in 4D parameter space ($f, \dot{f}, \alpha, \delta$)**

» placed at approximately equal distance,
» distance measure (metric) defined from fractional loss in expected \mathcal{F} -statistic.

- **Template grid of maximum possible loss in \mathcal{F} (mismatch m):**

» $m = 15\%$ for $f < 400\text{Hz}$, $m = 40\%$ for $f > 400\text{Hz}$.

- **Template grid is a Cartesian product of**

a uniformly spaced grid in frequency f , uniformly spaced grid in \dot{f} and metric-based grid on the sky (different for each data segment).

» In total: $\approx 10^{16}$ templates!

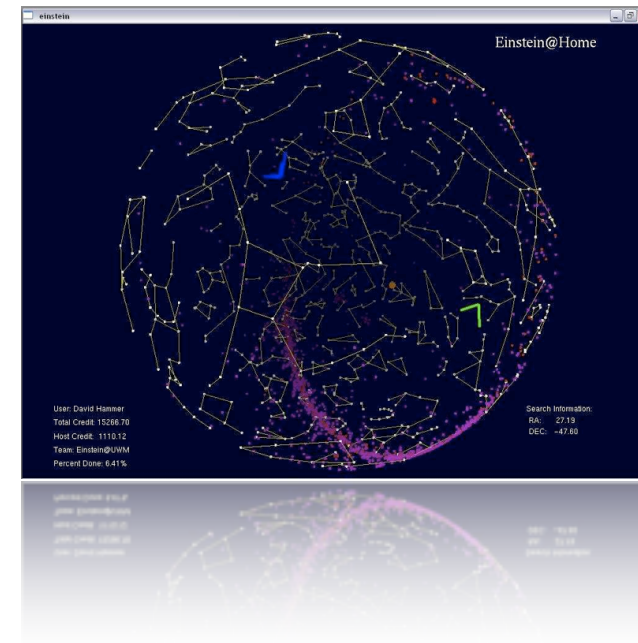
Same methods as previous
E@H S4 search:
Phys. Rev. D 79,
022001 (2009).

Einstein@Home:

Uses idle computing cycles volunteered by the general public, based upon BOINC.

- **Divide huge parameter space into small workunits:**
 - » Each workunit has computing time of order 1 day.
 - » Workunits sent out to participating hosts.
 - » Hosts return finished work in a top-list of most significant events (in \mathcal{F} -statistic).
 - » Validation: same work done on hosts owned by two different users, then results automatically compared.
 - » In total **16 446 454** workunits.
- **Total processing time on E@H project:**
 - » 6 months.
- **Total result data volume to post-process:**
 - » 0.6 TB compressed, 2.5 TB uncompressed.

Einstein@Home screensaver



- **About 100 000 active participants:**
 - » About 100 Tflops.

Goal: Find candidates from the 28 different data segments which cluster closely together in the 4D parameter space (f , \dot{f} , α , δ).

- Steps:

- » Shift candidate event frequencies to a fixed fiducial time so they can be compared
- » 'Bin' candidate events into 4D cells
- » Search for cells which have candidate events from many of the 28 data segments

- Cells are chosen to be as small as possible consistent with:

- » Cell size in sky > largest sky-grid separations (use metric-based model in declination)
- » Cell size in frequency > frequency-grid spacing + (spin-down grid spacing) x 30 hours
- » Cell size in spin-down > spin-down grid spacing

- Veto regions of parameter space

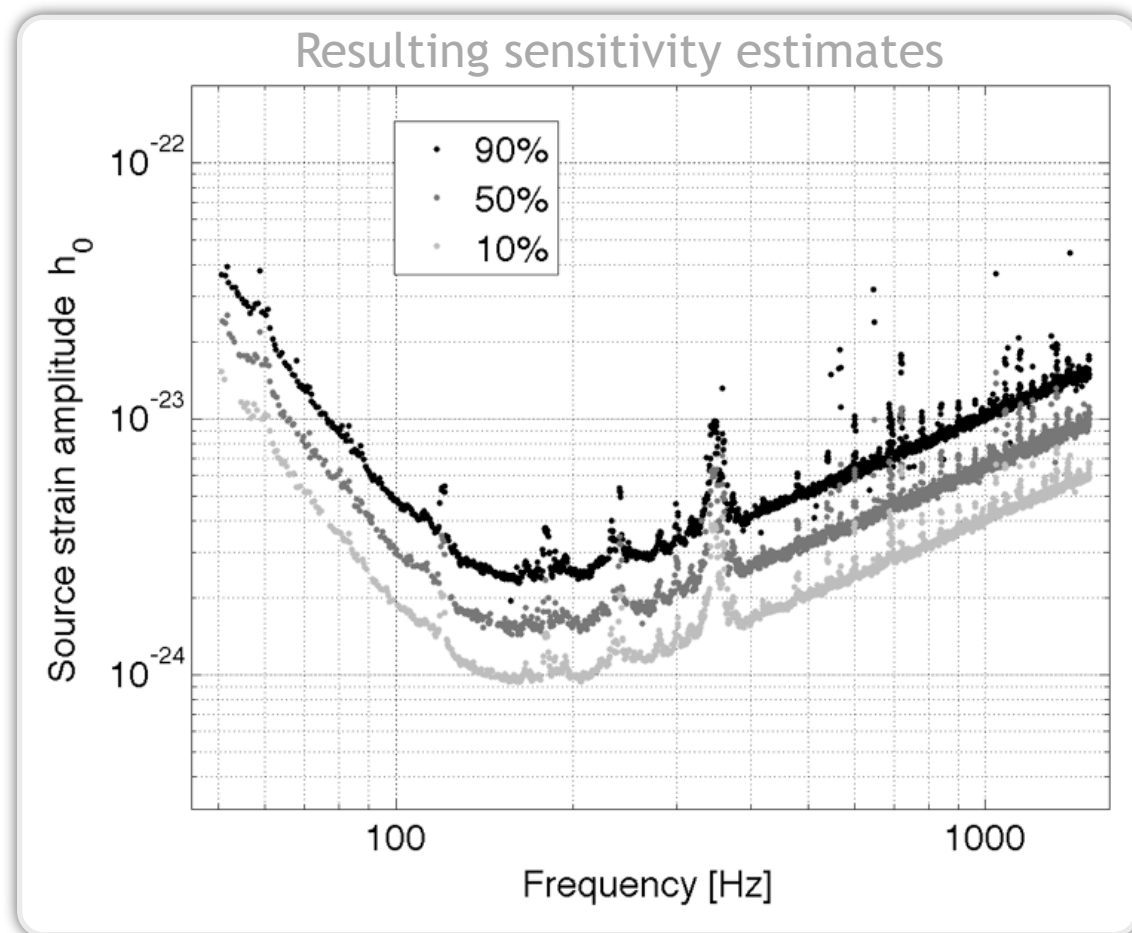
- » Exclude regions where stationary instrumental lines tend to appear due to global correlations. Physically: regions of minimum Doppler modulation.
- » In this search, about 13% of total search parameter-space is vetoed.

Pletsch, PRD 78,
102005 (2008).

- Average false alarm probability per 0.5Hz band in Gaussian noise of
 - » obtaining 10 (of 28) coincidences: 10^{-3} → expect a few in 1.5 kHz of Gaussian noise.
 - » obtaining 20 (of 28) coincidences: 10^{-21} → confident detection.

At what strain h_0 would 90% of sources be **confidently** detected?

- » Determined by MC simulations in each 0.5Hz band.
- Simulated source population probed for detection
 - » Uniformly distributed over the sky.
 - » Uniformly distributed in the nuisance parameters.
- In most sensitive band 125 - 225Hz, > 90% of sources with $h_0 \geq 3 \times 10^{-24}$ were confidently detected.
- Comparison to previous Einstein@Home S4 analysis:
 - » Improved sensitivity by a factor ≈ 3 .



Coincidence analysis results:

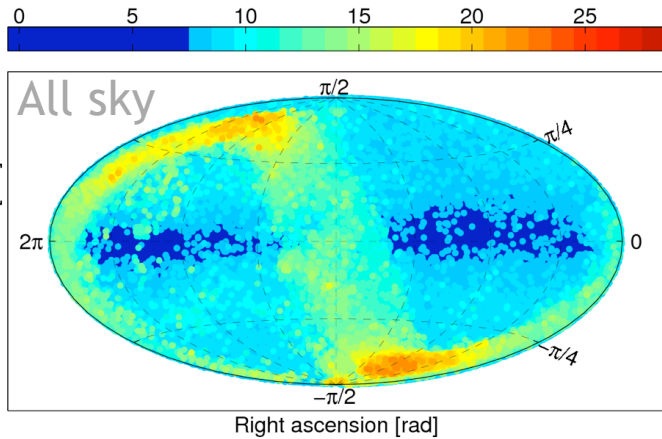
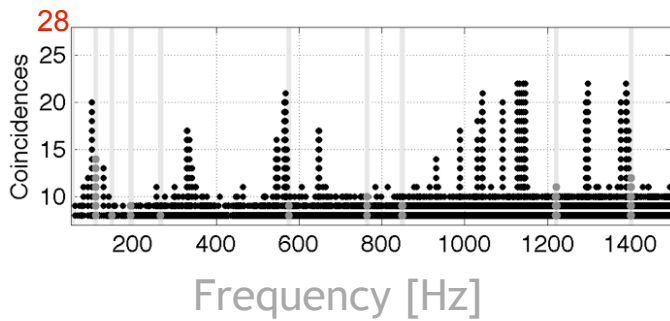


Further post-processing steps:



Final results:

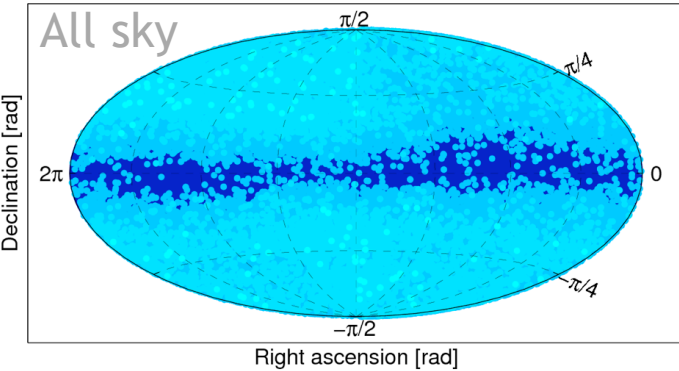
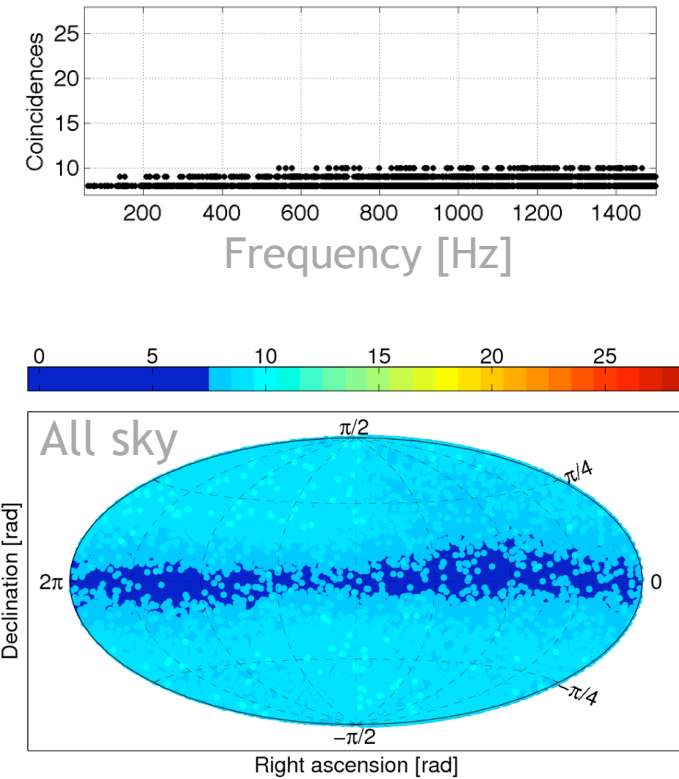
- » Left: 10 coincidences of 28 possible (but 10 also expected by chance).
- » No credible signal found.



1. A posteriori cleaning
 - » Removing instrumental lines of *known* origin.

2. Parameter-space veto
 - » Discriminating parameter-space regions of Doppler stationarity.

3. Detector coincidence veto
 - » Discriminating candidates whose coincidences come from a single detector only.



- **Einstein@Home** now running a hierarchical search in later S5 LIGO data using the **Hough transform** method.
 - » Incoherent combination step done on hosts,
 - » 121 data segments of 25 hours each (from H1 and L1),
 - » Increase in sensitivity by a factor ≈ 4 .
- **A more sensitive method** currently under development:
 - » “Global-Correlation Transform“ technique,
 - » Substantial increase in sensitivity by a factor ≈ 6 ,
 - » Planned for search of future LIGO S6 data.
- Those of you who run Einstein@Home: THANK YOU!
- Everyone else, sign-up to join at:

Krishnan et al.,
PRD 70, 082001,
(2005).

Pletsch and Allen,
ArXiv:0906.0023,
(2009).

<http://einstein.phys.uwm.edu>