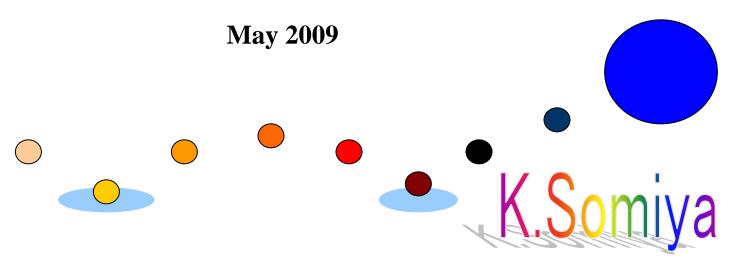
Displacement-Noise-Free Interferometry

~True Potential~

S.Kawamura, Y.Chen, K.Kokeyama, S.Sato, A.Nishizawa, and <u>K.Somiya</u>

GWADW at Ft Lauderdale



Contents

1. Invention of DFI [2004-2006]

~ Kawamura, Chen, et al

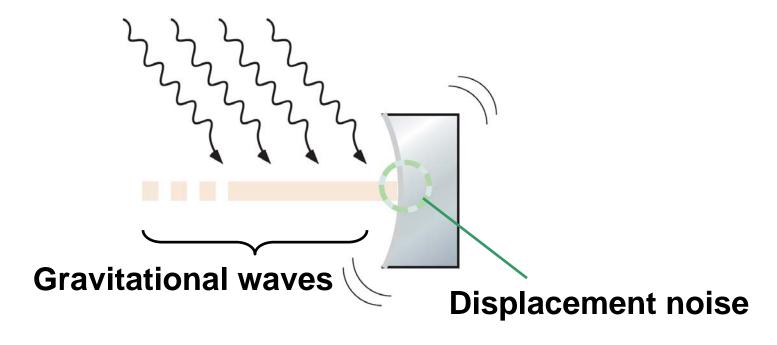
2. R&D experiment [2006-2008]

~ Kokeyama, Sato, et al

3. Recent study [2008-2009]

~ Nishizawa, Vyatchanin, et al





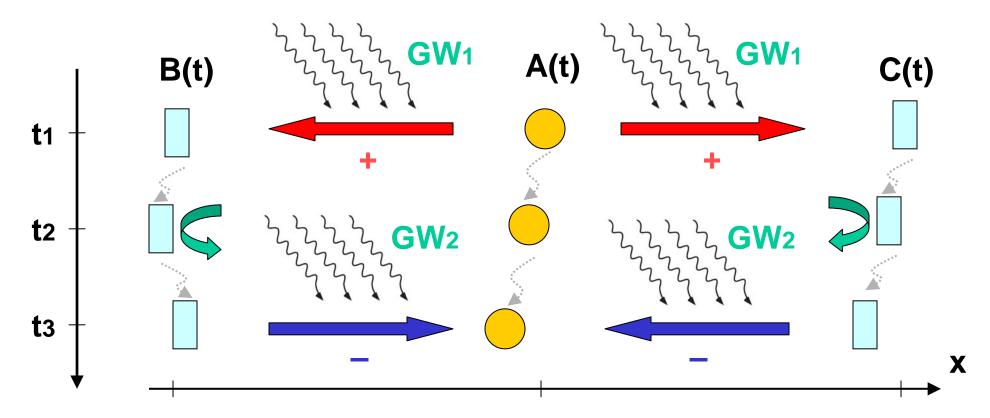
There is a difference between GW and disp. noise.

GW: probed during the light travel

disp: probed at the reflection



How can we distinguish GW and disp-noise?

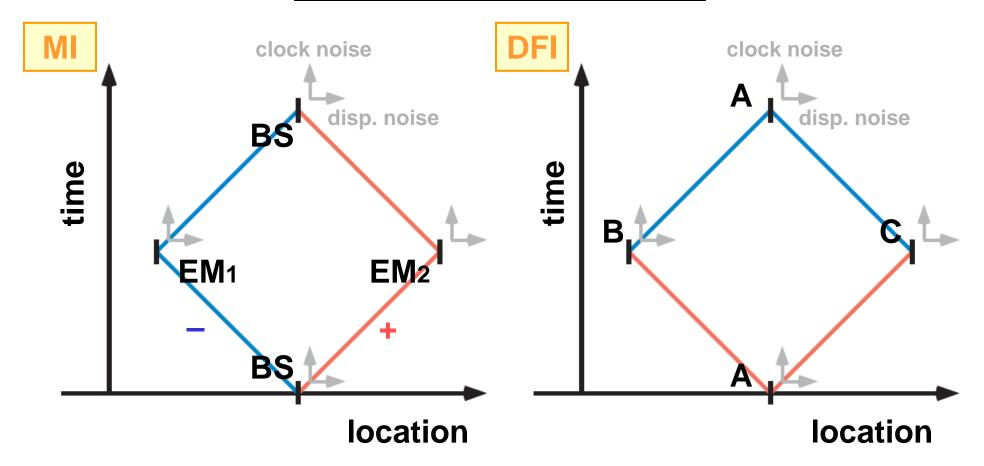


$$[A(t_1)-B(t_2)] + [C(t_2)-A(t_1)] - [A(t_3)-B(t_2)] - [C(t_2)-A(t_3)] + GW_1 - GW_2$$



displacement free + GW + clock noise

DFI vs Michelson ifo

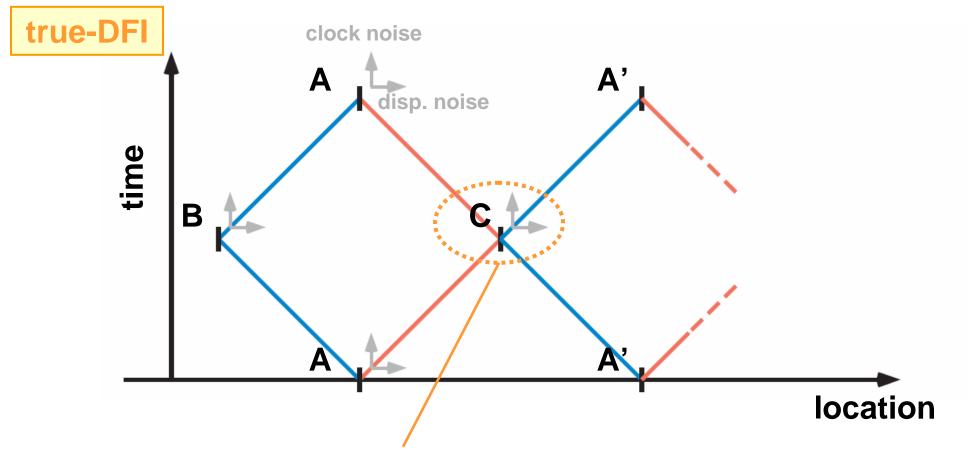


MI: no clock noise (freq noise), but disp. noise

DFI: no disp. noise, but clock noise

Actually we want to cancel both.

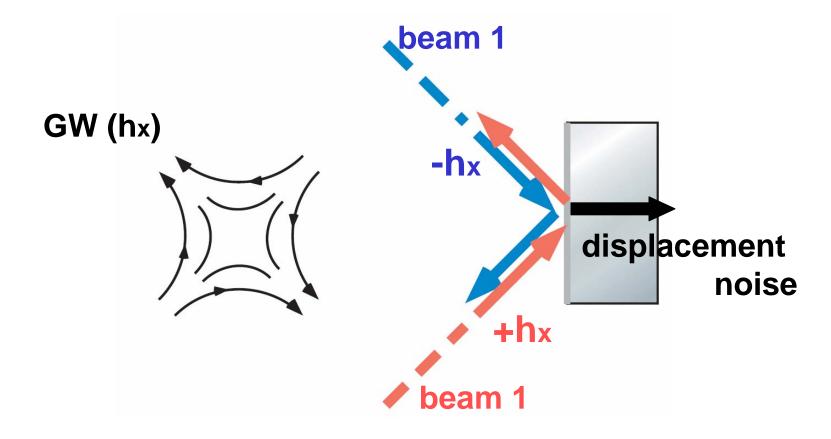
Displacement-noise and Frequency-noise free



Both disp noise and freq noise are cancelled (for C).

- We should do the same for all the components
- GW signal should not be cancelled

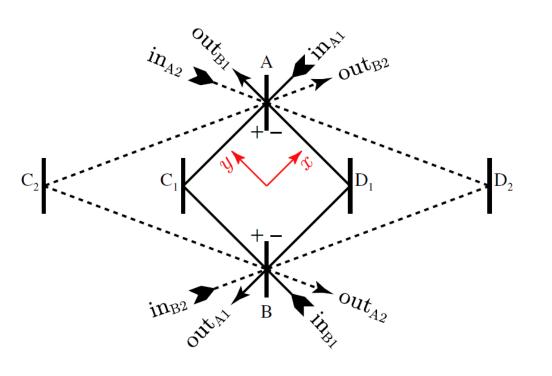
Separation of GW and disp noise



- Sign of GW is opposite between beam 1 and beam 2
- After subtraction, GW signal is prop to frequency

Bi-directional Mach-Zehnder ifo

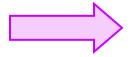
[Chen PRL 2006]



1st subtraction (mirror disp) bi-directional measurement (ex. AC₁B – BC₁A)

2nd subtraction (freq noise) MZ interferometer (ex. AC₁B – AD₁B)

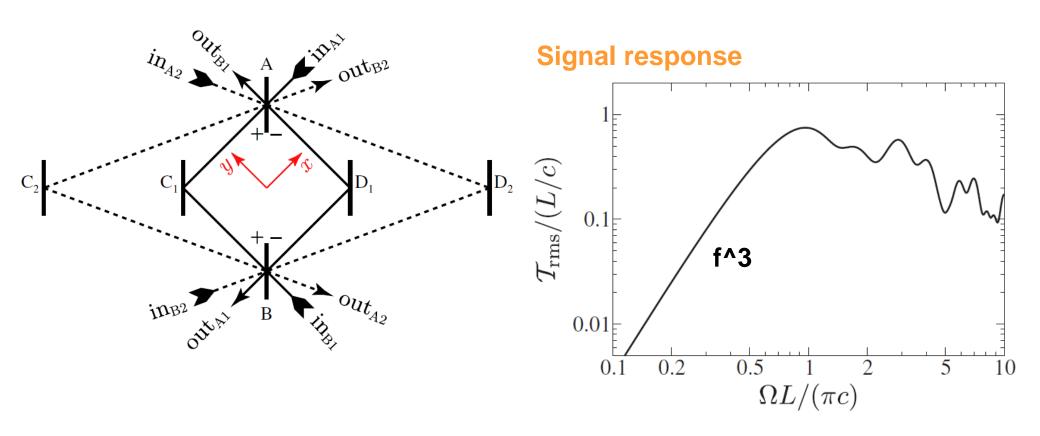
3rd subtraction (BS disp)
2 MZ combination
(ex. AC₁B - AC₂B)

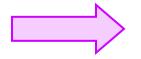


Signal is proportional to f^3 (bad)

Bi-directional Mach-Zehnder ifo

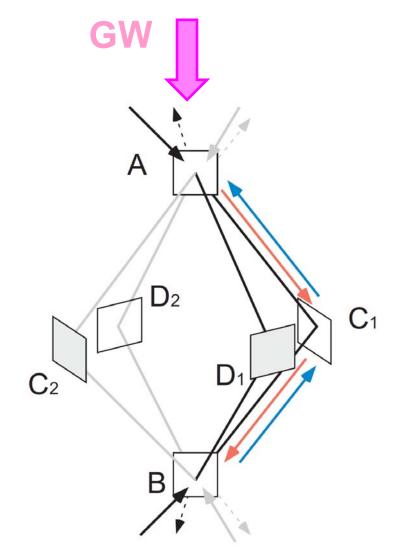
[Chen PRL 2006]

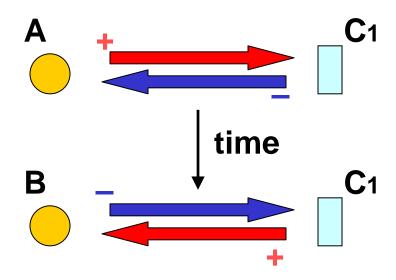




Signal is proportional to f^3 (bad)

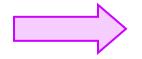






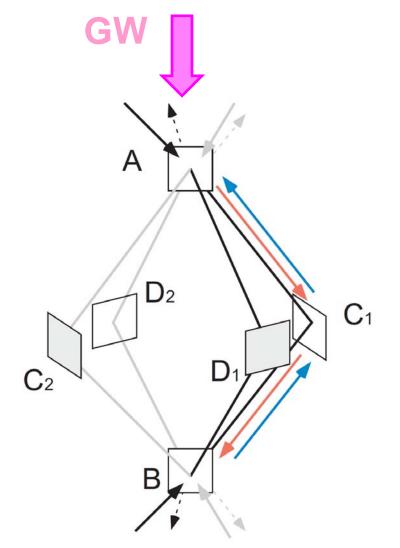
1st subtraction
DFI between A-C₁

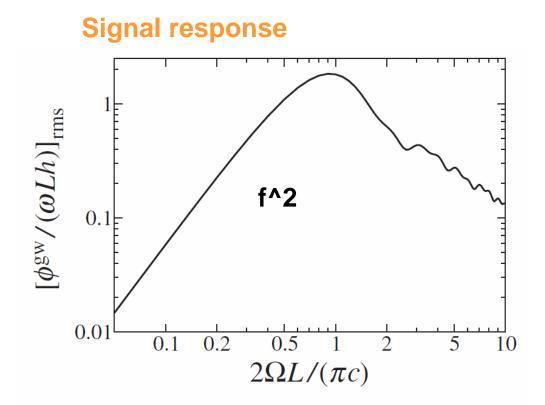
2nd subtraction
A-C₁ and B-C₁ at different time



Signal is proportional to f^2 (bad)

3D DFI



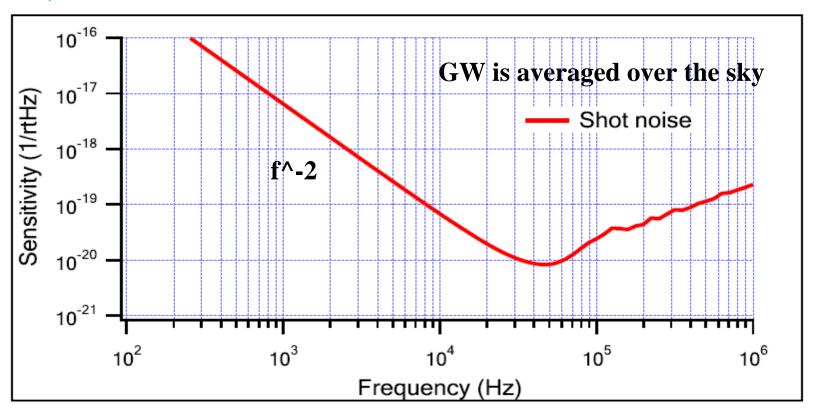




Signal is proportional to f^2 (better)

3D-DFI noise spectrum

L=4km, I=10W each



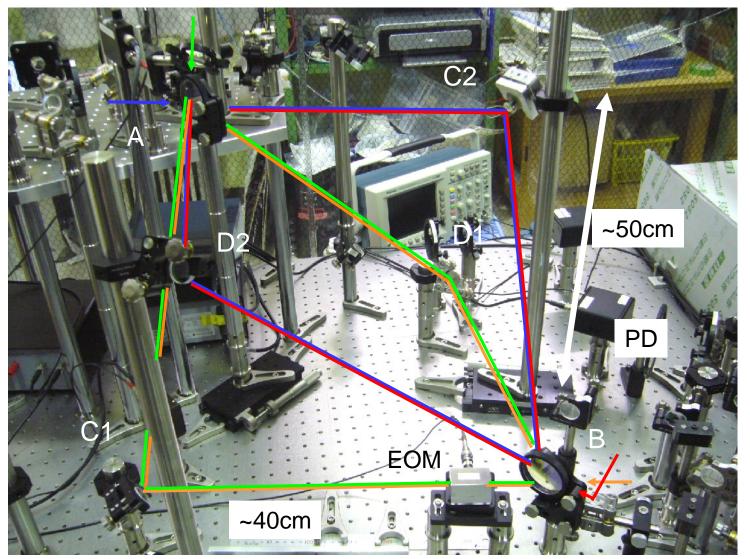
- No displacement noise at any frequency
- Shot noise is bad at low frequency

After the discovery...

- R&D experiments in Japan
 - 2D DFI
 - 3D DFI
- Effort to find a better configuration
 - time-delay device
 - narrow-band DFI
 - DFI with rigid platform
 - resonant-type DFI

3D DFI table-top experiment

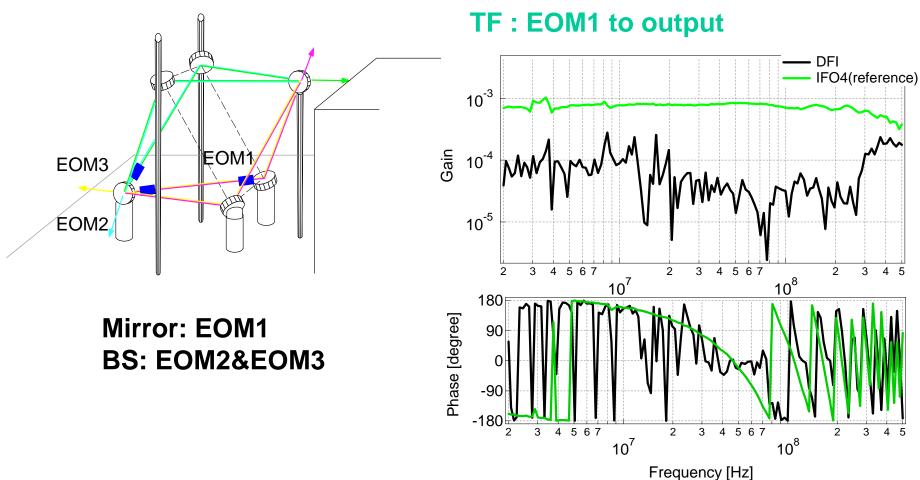
[Kokeyama 2009]



- Alignment in 3D
- High frequency (~MHz)
- TF measurement using EOMs

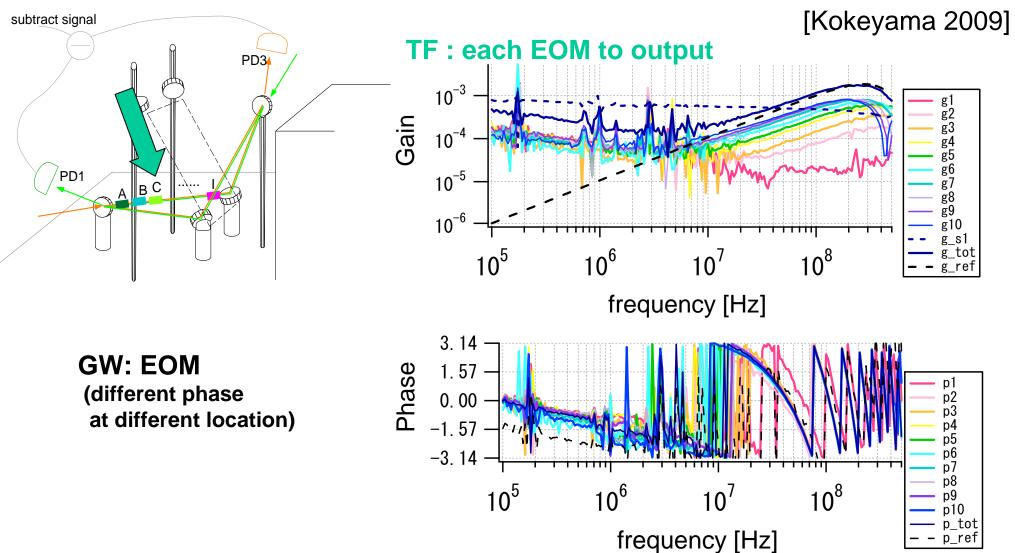
Verification of noise cancellation

[Kokeyama 2009]



Displacement components is reduced by 20~30dB

Verification of GW non-cancellation



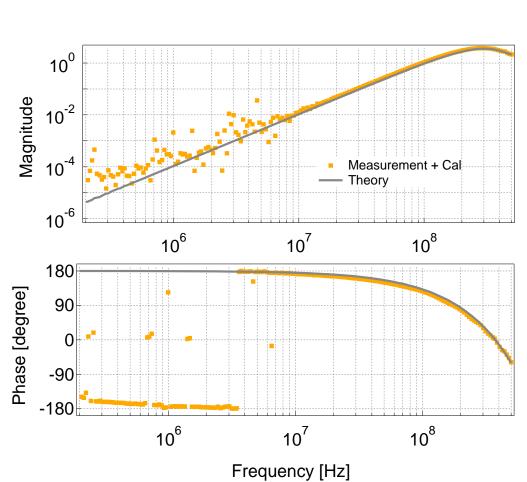
GW signal can be imitated by combining these outputs

Verification of GW non-cancellation

PD1
A
B
C
.....

GW: EOM (different phase at different location)

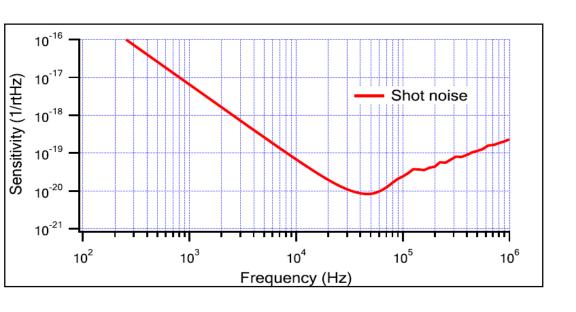
TF: combination of 10 EOMs



[Kokeyama 2009]

After the calibration, DFI signal response is obtained.

Weak-response problem



Even with the 3D DFI, shot noise \propto f^-2

Moreover...

Yanbei's theorem

1D: impossible

2D: f^-3

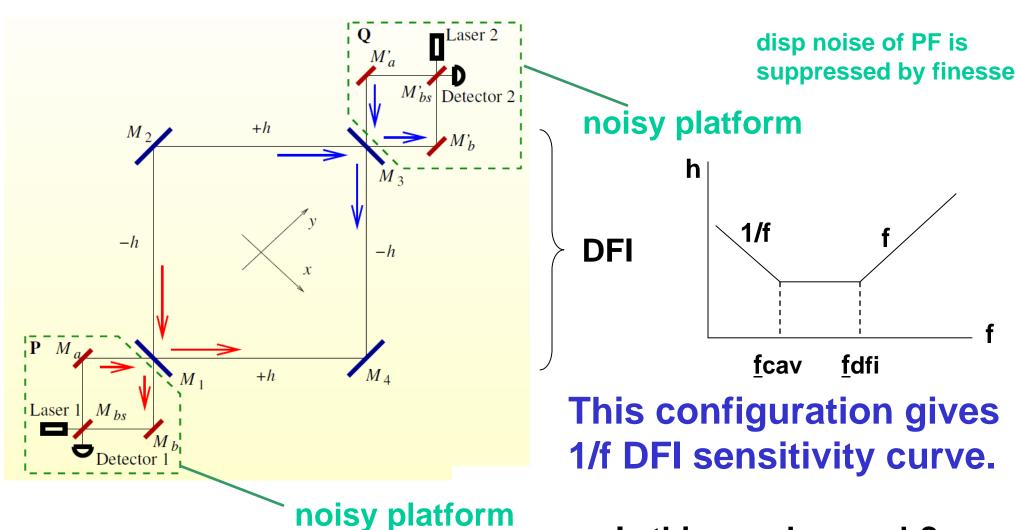
3D: f^-2

one possible ways to circumvent the problem

Make a cavity and ignore disp noise outside the cavity

Resonant speed-meter DFI

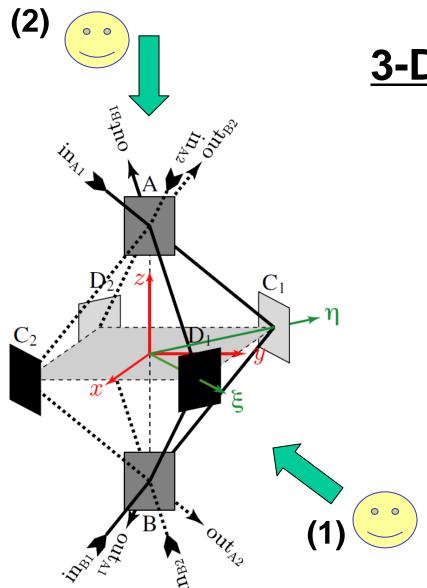
[Vyatchanin 2009]



Is this good enough?

Summary and discussions

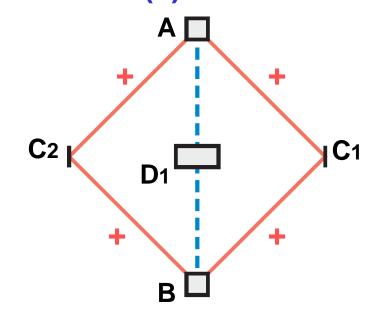
- 3D DFI with f^-2 slope is so far the best for complete DFI
- R&D experiments done
- Resonant partial-DFI seems nice with 1/f slope
- Is it possible to make the slope better?
- DFI might be useful for a space-based GW detector
- Maybe with atomic interferometer...



3-D DFI

[Chen PRL 2006]

GW from (1)'s direction



ACjB is no MZ (no GW).
ADjB's GW cancels with BDjA's.

No response to GW from (1)'s direction.

